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# ON THE STATIONARY AND UNIFORMLY-ROTATING SOLUTIONS OF ACTIVE SCALAR EQUATIONS 

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To my mother Cho Myeongji

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## SUMMARY

In this thesis, we study qualitative and quantitative properties of stationary/uniformly-rotating solutions of the 2D incompressible Euler equation and the generalized Surface Quasi-Geostrophic (SQG) equations. The main goal is to establish sufficient and necessary conditions for the stationary/uniformly rotating solutions to be radially symmetric. In addition, we also derive quantitative estimates for non-radial, uniformly-rotating patch solutions for the 2D Euler equation.

In chapter 1 , we briefly review basic properties of the 2D incompressible Euler equation and the generalized SQG equations. We also rigorously define stationary and uniformly-rotating solutions to those equations by means of the stream function.

Chapters 2 to 4 describe the joint work with Javier Gómez-Serrano, Jia Shi and Yao Yao [50, 52]. We establish sufficient conditions for stationary/uniformly-rotating solutions for some active scalar equations to be radially symmetric. In short, we prove that for the 2D Euler equation,
i (Patch setting) If a vortex patch $\omega=1_{D}$ is uniformly rotating with angular velocity $\Omega \in\left(0, \frac{1}{2}\right)^{c}$, then $D$ must be radially symmetric up to a translation.
ii (Smooth setting) If a smooth, non-negative compactly supported vorticity $\omega$ is uniformly rotating with angular velocity $\Omega \leq 0$, then $\omega$ is radially symmetric.

The proof is based on a variational argument that a uniformly-rotating solution can be formally thought of as a critical point of an energy functional. We apply this idea to more general active scalar equations (gSQG) and vortex sheet equation.

In chapter 5, we construct a non-radial vortex sheet with non-constant vortex strength, which is rotating with angular velocity $\Omega>0$. We obtain a curve of such non-radial solutions, bifurcating from trivial ones. This result comes from the joint work with Javier Gómez-Serrano, Jia Shi and Yao Yao [51].

In chapter 6, we describe the result in [97]. We adapt the variational argument that was used in
chapter 2 to study non-radial rotating vortex patches. It is well known that for $\Omega \in\left(0, \frac{1}{2}\right)$, there are $m$-fold symmetric rotating patches. We derive some quantitative estimates for those patches about their angular velocities and the difference with the unit disk.

## CHAPTER 1

## INTRODUCTION AND BACKGROUND

### 1.1 Two dimensional active scalar equations: Euler and gSQG

The two-dimensional incompressible Euler equation in vorticity form reads

$$
\begin{cases}\partial_{t} \omega+u \cdot \nabla \omega=0 & \text { in } \mathbb{R}^{2} \times \mathbb{R}_{+}  \tag{1.1.1}\\ u(\cdot, t)=-\nabla^{\perp}(-\Delta)^{-1} \omega(\cdot, t) & \text { in } \mathbb{R}^{2} \\ \omega(\cdot, 0)=\omega_{0} & \text { in } \mathbb{R}^{2}\end{cases}
$$

where $\nabla^{\perp}:=\left(-\partial_{x_{2}}, \partial_{x_{1}}\right)$. Note that we can express $u$ as $u(\cdot, t)=\nabla^{\perp}(\omega(\cdot, t) * \mathcal{N})$, where $\mathcal{N}(x):=$ $\frac{1}{2 \pi} \ln |x|$ is the Newtonian potential in two dimensions. This equation describes the motion of incompressible ideal fluid in two dimensional space. Mathematically, the 2D Euler equation can be seen as an example of active scalar equation, in the sense that the scalar-valued function $\omega$ is transported by the velocity $u$, meanwhile the velocity also depends on $\omega$. Another interesting active scalar equation is the inviscid Surface Quasi-Geostropic equation, where we replace $(-\Delta)^{-1}$ by $(-\Delta)^{-\frac{1}{2}}$ in (Equation 1.1.1). Besides its importance in the context of geophysics and atmosphere science, the inviscid SQG also serves as a toy model for the 3D incompressible Euler equation [32]. More generally, both the 2D Euler equation and the inviscid SQG equation belong to the following family of active scalar equations indexed by a parameter $\alpha,(0 \leq \alpha<2)$, known as the
generalized SQG equations:

$$
\begin{cases}\partial_{t} \omega+u \cdot \nabla \omega=0 & \text { in } \mathbb{R}^{2} \times \mathbb{R}_{+}  \tag{1.1.2}\\ u(\cdot, t)=-\nabla^{\perp}(-\Delta)^{-1+\frac{\alpha}{2}} \omega(\cdot, t) & \text { in } \mathbb{R}^{2} \\ \omega(\cdot, 0)=\omega_{0} & \text { in } \mathbb{R}^{2}\end{cases}
$$

Here we can also express the Biot-Savart law as

$$
\begin{equation*}
u(\cdot, t)=\nabla^{\perp}\left(\omega(\cdot, t) * K_{\alpha}\right), \tag{1.1.3}
\end{equation*}
$$

where $K_{\alpha}$ is the fundamental solution for $-(-\Delta)^{-1+\frac{\alpha}{2}}$, that is,

$$
K_{\alpha}(x)= \begin{cases}\frac{1}{2 \pi} \ln |x| & \text { for } \alpha=0  \tag{1.1.4}\\ -C_{\alpha}|x|^{-\alpha} & \text { for } \alpha \in(0,2)\end{cases}
$$

where $C_{\alpha}=\frac{1}{2 \pi} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{2^{1-\alpha} \Gamma\left(1-\frac{\alpha}{2}\right)}$ is a positive constant only depending on $\alpha$.
We either work with the patch setting, where $\omega(\cdot, t)=1_{D(t)}$ is an indicator function of a bounded set that moves with the fluid, or the smooth setting, where $\omega(\cdot, t)$ is smooth and compactly-supported in $x$. For well-posedness results for patch solutions, see the global wellposedness results [9, 26] for (Equation 1.1.1), and local well-posedness results [25, 101, 43, 75, 33] for (Equation 1.1.2) with $\alpha \in(0,2)$.

### 1.1.1 Uniformly rotating/Stationary solutions of 2D Euler and gSQG

Let us begin with the definition of a stationary/uniformly-rotating solution in the patch setting. For a bounded domain $D \subset \mathbb{R}^{2}$ with $C^{1}$ boundary, we say $\omega=1_{D}$ is a stationary patch solution to (Equation 1.1.2) for some $\alpha \in[0,2)$ if $u(x) \cdot \vec{n}(x)=0$ on $\partial D$, with $u$ given by (Equation 1.1.3).

This leads to the integral equation

$$
\begin{equation*}
1_{D} * \mathcal{K}_{\alpha} \equiv C_{i} \quad \text { on } \partial D, \tag{1.1.5}
\end{equation*}
$$

where the constant $C_{i}$ can differ on different connected components of $\partial D$. And if $\omega(x, t)=$ $1_{D}\left(R_{\Omega t} x\right)$ is a uniformly-rotating patch solution with angular velocity $\Omega$ (where $R_{\Omega t} x$ rotates a vector $x \in \mathbb{R}^{2}$ counter-clockwise by angle $\Omega t$ about the origin), then $1_{D}$ becomes stationary in the rotating frame with angular velocity $\Omega$, that is, $\left(\nabla^{\perp}\left(\omega(\cdot, t) * K_{\alpha}\right)-\Omega x^{\perp}\right) \cdot \vec{n}(x)=0$ on $\partial D$. As a result we have

$$
\begin{equation*}
1_{D} * \mathcal{K}_{\alpha}-\frac{\Omega}{2}|x|^{2} \equiv C_{i} \quad \text { on } \partial D \tag{1.1.6}
\end{equation*}
$$

where $C_{i}$ again can take different values along different connected components of $\partial D$. Note that a stationary patch $D$ also satisfies (Equation 1.1.6) with $\Omega=0$, and it can be considered as a special case of uniformly-rotating patch with zero angular velocity.

Likewise, in the smooth setting, if $\omega(x, t)=\omega_{0}\left(R_{\Omega t} x\right)$ is a uniformly-rotating solution of (Equation 1.1.2) with angular velocity $\Omega$ (which becomes a stationary solution in the $\Omega=0$ case), then we have $\left(\nabla^{\perp}\left(\omega_{0} * K_{\alpha}\right)-\Omega x^{\perp}\right) \cdot \nabla \omega_{0}=0$. As a result, $\omega_{0}$ satisfies

$$
\begin{equation*}
\omega_{0} * \mathcal{K}_{\alpha}-\frac{\Omega}{2}|x|^{2} \equiv C_{i} \quad \text { on each connected component of a regular level set of } \omega_{0}, \tag{1.1.7}
\end{equation*}
$$

where $C_{i}$ can be different if a regular level set $\left\{\omega_{0}=c\right\}$ has multiple connected components.

### 1.1.2 Uniformly rotating/Stationary vortex sheets

A vortex sheet is a weak solution of the 2D Euler equations:

$$
\begin{equation*}
v_{t}+v \cdot \nabla v=-\nabla p, \quad \nabla \cdot v=0 \tag{1.1.8}
\end{equation*}
$$

whose vorticity $\omega=\operatorname{curl}(v)$ is a delta function supported on a curve or a finite number of curves $\Gamma_{i}=z_{i}(\alpha, t)$, i.e.

$$
\begin{equation*}
\omega(x, t)=\sum_{i} \varpi_{i}(\alpha, t) \delta\left(x-z_{i}(\alpha, t)\right) . \tag{1.1.9}
\end{equation*}
$$

Here $\varpi_{i}(\alpha, t)$ is the vorticity strength on $\Gamma_{i}$ with respect to the parametrization $z_{i}$, and the above equation is understood in the sense that

$$
\int_{\mathbb{R}^{2}} \varphi(x) d \omega(x, t)=\sum_{i} \int \varphi\left(z_{i}(\alpha, t)\right) \varpi_{i}(\alpha, t) d \alpha
$$

for all test functions $\varphi(x) \in C_{0}^{\infty}\left(\mathbb{R}^{2}\right)$.
The motivation of the study of the equation (Equation 1.1.8) with vortex sheet initial data comes from the fact that in fluids with small viscosity, flows separate from rigid walls and corners [87, 102]. To model it, one may think of a solution to (Equation 1.1.8) with one incompressible fluid where the velocity changes sign in a discontinuous (tangential) way across a streamline $z$. This discontinuity induces vorticity in $z$.

The equations of motion of $\varpi_{i}$ and $z_{i}$ can be derived by means of the Birkhoff-Rott operator ([20, 81, 87, 108]), namely:

$$
\begin{equation*}
B R(z, \varpi)(x, t)=\frac{1}{2 \pi} P V \int \frac{(x-z(\beta, t))^{\perp}}{|x-z(\beta, t)|^{2}} \varpi(\beta, t) d \beta \tag{1.1.10}
\end{equation*}
$$

yielding

$$
\begin{align*}
\partial_{t} z_{i}(\alpha, t) & =\sum_{j} B R\left(z_{j}, \varpi_{i}\right)\left(z_{i}(\alpha, t)\right)+c_{i}(\alpha, t) \partial_{\alpha} z_{i}(\alpha, t)  \tag{1.1.11}\\
\partial_{t} \varpi_{i}(\alpha, t) & =\partial_{\alpha}\left(c_{i}(\alpha, t) \varpi_{i}(\alpha, t)\right), \tag{1.1.12}
\end{align*}
$$

where the term $c_{i}(\alpha, t)$ accounts for the reparametrization freedom of the curves. See the paper [70] by Izosimov-Khesin where they propose geodesic, group-theoretic, and Hamiltonian frameworks
for their description.
As in the patch/smooth setting, we first define what we mean by a stationary vortex sheet. Assume the initial data $\omega_{0}$ of (Equation 1.1.9) is supported on a finite number of curves parametrized by $z_{i}(\alpha)$, with strength $\varpi_{i}(\alpha)$ (with respect to the parametrization $z_{i}$ ) respectively. If there exists some reparametrization choice $c_{i}(\alpha)$ such that the right hand sides of (Equation 1.1.11)(Equation 1.1.12) are both identically zero for every $i$, it gives that $\omega(\cdot, t)$ is invariant in time, and we say $\omega(\cdot, t)=\omega_{0}$ is a stationary vortex sheet.

For any $x \in \mathbb{R}^{2}$ and $\Omega \in \mathbb{R}$, let $R_{\Omega t} x$ denote the rotation of $x$ counter-clockwise by angle $\Omega t$ about the origin. We say $\omega(x, t)=\omega_{0}\left(R_{\Omega t} x\right)$ is a uniformly-rotating vortex sheet with angular velocity $\Omega$ if $\omega_{0}$ is stationary in the rotating frame with angular velocity $\Omega$. (Note that in the special case $\Omega=0$, the uniformly-rotating sheet is in fact stationary.) In chapter 4, we will derive the equations satisfied by a stationary/rotating vortex sheet.

### 1.2 Main results and idea of proofs

In subsection 1.2.1 and subsection 1.2.2, we describe the results of joint work with Javier GómezSerrano, Jia Shi and Yao Yao [50, 52, 51], which concern rigidity/flexibility of uniformlyrotating/stationary solutions. The proofs for these results will be contained in chapter 2-chapter 5. In subsection 1.2.3, we describe the results obtained in [97]. The proofs for quantitative estimates will be presented in chapter 6 .

### 1.2.1 Rigidity results for 2D Euler and gSQG [50]

Chapter 2 and Chapter 3 will be devoted to establish rigidity results for 2D Euler and gSQG equations. Clearly, every radially symmetric patch/smooth function automatically satisfies (Equation 1.1.6) or (Equation 1.1.7) for all $\Omega \in \mathbb{R}$. The goal of chapter 2-chapter 3 is to address the complementary question, which can be roughly stated as following:

Question 1. In the patch or smooth setting, under what condition must a stationary/uniformlyrotating solution be radially symmetric?

Below we summarize the previous literature related to this question, and state our main results. We will first discuss the 2D Euler equation in the patch and smooth setting respectively, then discuss the gSQG equation with $\alpha \in(0,2)$.

## 2D Euler in the patch setting

Let us deal with the patch setting first. So far affirmative answers to Question 1 have only been only obtained for simply-connected patches, for angular velocities $\Omega=0, \Omega<0$ (under some additional convexity assumptions), and $\Omega=\frac{1}{2}$. For stationary patches ( $\Omega=0$ ), Fraenkel [41, Chapter 4] proved that if $D$ satisfies (Equation 1.1.6) (where $K_{\alpha}=\mathcal{N}$ ) with the same constant $C$ on the whole $\partial D$, then $D$ must be a disk. The idea is that in this case the stream function $\psi=1_{D} * \mathcal{N}$ solves a semilinear elliptic equation $\Delta \psi=g(\psi)$ in $\mathbb{R}^{2}$ with $g(\psi)=1_{\{\psi<C\}}$, where the monotonicity of the discontinuous function $g$ allows one to apply the moving plane method developed in $[105,48]$ to obtain the symmetry of $\psi$. As a direct consequence, every simplyconnected stationary patch must be a disk. But if $D$ is not simply-connected, (Equation 1.1.6) gives that $\psi=C_{i}$ on different connected components of $\partial D$, thus $\psi$ might not solve a single semilinear elliptic equation in $\mathbb{R}^{2}$. Even if $\psi$ satisfies $\Delta \psi=g(\psi), g$ might not have the right monotonicity. For these reasons, whether a non-simply-connected stationary patch must be radial still remained an open question.

For $\Omega<0$, Hmidi [59] used the moving plane method to show that a simply-connected uniformly-rotating patch $D$ satisfies some additional convexity assumption (which is stronger than star-shapedness but weaker than convexity), then $D$ must be a disk. In the special case $\Omega=\frac{1}{2}$, Hmidi [59] also showed that a simply-connected uniformly-rotating patch $D$ must be a disk, using the fact that $1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}$ becomes a harmonic function in $D$ when $\Omega=\frac{1}{2}$.

On the other hand, it is known that there can be non-radial uniformly-rotating patches for
$\Omega \in\left(0, \frac{1}{2}\right)$. The first example dates back to the Kirchhoff ellipse [74], where it was shown that any ellipse $D$ with semiaxes $a, b$ is a uniformly-rotating patch with angular velocity $\frac{a b}{(a+b)^{2}}$. DeemZabusky [36] numerically found families of patch solutions of (Equation 1.1.1) with $m$-fold symmetry by bifurcating from a disk at explicit angular velocities $\Omega_{m}^{0}=\frac{m-1}{2 m}$ and coined the term V-states. Further numerics were done in [116, 37, 86, 103]. Burbea gave the first rigorous proof of their existence by using (local) bifurcation theory arguments close to the disk [12]. There have been many recent developments in a series of works by Hmidi-Mateu-Verdera and de la Hoz-Hmidi-Mateu-Verdera [63, 67, 64] in different settings and directions (regularity of the boundary, different topologies, etc.). In particular, [67] showed the existence of $m$-fold doubly-connected non-radial patches bifurcating at any angular velocity $\Omega \in\left(0, \frac{1}{2}\right)$ from some annulus of radii $b \in(0,1)$ and 1.

There are many other interesting perspectives of the V-states, which we briefly review below, although they are not directly related to Question 1. Hassainia-Masmoudi-Wheeler [58] were able to perform global bifurcation arguments and study the whole branch of V-states. Other scenarios such as the bifurcation from ellipses instead of disks have also been studied: first numerically by Kamm [71] and later theoretically by Castro-Córdoba-Gómez-Serrano [22] and Hmidi-Mateu [60]. See also the work of Carrillo-Mateu-Mora-Rondi-Scardia-Verdera [19] for variational techniques applied to other anisotropic problems related to vortex patches. Love [82] established linear stability for ellipses of aspect ratio bigger than $\frac{1}{3}$ and linear instability for ellipses of aspect ratio smaller than $\frac{1}{3}$. Most of the efforts have been devoted to establish nonlinear stability and instability in the range predicted by the linear part. Wan [114], and Tang [110] proved the nonlinear stable case, whereas Guo-Hallstrom-Spirn [53] settled the nonlinear unstable one. See also [30]. In [112], Turkington considered $N$ vortex patches rotating around the origin in the variational setting, yielding solutions of the problem which are close to point vortices.

Our first main result is summarized in the following Theorem A, which gives a complete answer to Question 1 for 2D Euler in the patch setting. Note that $D$ is allowed to be disconnected, and each
connected component can be non-simply-connected. Figure Figure 1.1 illustrates a comparison of our result (in red color) with the previous results (in black color).

Theorem A (= Corollary 2.1.8, Theorems 2.1.10 and 2.1.12). Let $D \subset \mathbb{R}^{2}$ be a bounded domain with $C^{1}$ boundary. Assume $D$ is a stationary/uniformly-rotating patch of (Equation 1.1.1), in the sense that $D$ satisfies (Equation 1.1.6) (with $K_{\alpha}=\mathcal{N}$ ) for some $\Omega \in \mathbb{R}$. Then $D$ must be radially symmetric if $\Omega \in(-\infty, 0) \cup\left[\frac{1}{2}, \infty\right)$, and radially symmetric up to a translation if $\Omega=0$.


Figure 1.1: For 2D Euler in the patch setting, previous results on Question 1 are summarized in black color. Our results in Theorem A are colored in red.

## 2D Euler in the smooth setting

One of the main motivations of this work is to find sufficient rigidity conditions in terms of the vorticity, such that the only stationary/uniformly-rotating solutions are radial ones. Heuristically speaking, this belongs to the broader class of "Liouville Theorem" type of results, which show that solutions satisfying certain conditions must have a simpler geometric structure, such as being constant (in one direction, or all directions) or being radial. In the literature we could not find any conditions on 2D Euler that leads to radial symmetry, although several other Liouville-type results have been established for 2D fluid equations: For 2D Euler, Hamel-Nadirashvili [55, 54] proved that any stationary solution without a stagnation point must be a shear flow. (But note that this result does not apply to our setting (Equation 1.1.7), since the velocity field $u$ associated with any compactly-supported $\omega_{0}$ must have a stagnation point). See also the Liouville theorem by Koch-Nadirashvili-Seregin-Šverák for the 2D Navier-Stokes equations [76].

Let us briefly review some results on the characterization of stationary solutions to 2D Euler, although they are not directly related to Question 1. Nadirashvili [92] studied the geometry and the stability of stationary solutions, following the works of Arnold [2, 3, 4]. Izosimov-Khesin [69] characterized stationary solutions of 2D Euler on surfaces. Choffrut-Šverák [28] showed that locally near each stationary smooth solution there exists a manifold of stationary smooth solutions transversal to the foliation, and Choffrut-Székelyhidi [27] showed that there is an abundant set of stationary weak $\left(L^{\infty}\right)$ solutions near a smooth stationary one. Shvydkoy-Luo [84, 85] classified the set of stationary smooth solutions of the form $v=\nabla^{\perp}\left(r^{\gamma} f(\omega)\right)$, where $(r, \omega)$ are polar coordinates. In a different direction, Turkington [111] used variational methods to construct stationary vortex patches of a prescribed area in a bounded domain, imposing that the patch is a characteristic function of the set $\{\Psi>0\}$, and also studied the asymptotic limit of the patches tending to point vortices. Long-Wang-Zeng [80] studied their stability, as well as the regularity in the smooth setting (see also [16]). For other variational constructions close to point vortices, we refer to the work of Cao-Liu-Wei [14], Cao-Peng-Yan [15] and Smets-van Schaftingen [107]. We remark that these results do not rule out that those solutions may be radial. Musso-Pacard-Wei [91] constructed nonradial smooth stationary solutions without compact support in $\omega$. The (nonlinear $L^{1}$ ) stability of circular patches was proved by Wan-Pulvirenti [113] and later Sideris-Vega gave a shorter proof [106]. See also Beichman-Denisov [7] for similar results on the strip.

Lately, Gavrilov [45, 46] provided a remarkable construction of nontrivial stationary solutions of 3D Euler with compactly supported velocity. See also Constantin-La-Vicol for a simplified proof with extensions to other fluid equations [31].

Regarding uniformly-rotating smooth solutions $(\Omega \neq 0)$ for 2D Euler, Castro-Córdoba-Gómez-Serrano [24] were able to desingularize a vortex patch to produce a smooth $m$-fold V-state with $\Omega \sim \frac{m-1}{2 m}>0$ for $m \geq 2$. Recently García-Hmidi-Soler [44] studied the construction of V-states bifurcating from other radial profiles (Gaussians and piecewise quadratic functions).

Our second main result is the following theorem, which gives radial symmetry in the smooth
setting for $\Omega \leq 0$, under the additional assumption $\omega_{0} \geq 0$ :

Theorem B (= Theorem 2.2.5 and Corollary 2.2.6). Let $\omega_{0} \geq 0$ be smooth and compactly-supported. Assume $\omega(x, t)=\omega_{0}\left(R_{\Omega t} x\right)$ is a stationary/uniformly-rotating solution of (Equation 1.1.1) with $\Omega \leq 0$, in the sense that it satisfies (Equation 1.1.7). Then $\omega_{0}$ must be radially symmetric if $\Omega<0$, and radially symmetric up to a translation if $\Omega=0$.

The gSQG case $(0<\alpha<2)$

Recall that in the patch setting, a stationary/uniformly-rotating patch satisfies (Equation 1.1.6) with $K_{\alpha}$ given in (Equation 1.1.4). Even though the kernels $K_{\alpha}$ are qualitatively similar for all $\alpha \in[0,2)$, there is a key difference on the symmetry v.s. non-symmetry results between the cases $\alpha=0$ and $\alpha>0$ : For the 2D Euler equation ( $\alpha=0$ ), we proved in Theorem A that any rotating patch $D$ with $\Omega \leq 0$ must be radial, even if $D$ is not simply-connected. However, this result is not true for any $\alpha \in(0,2)$ : de la Hoz-Hassainia-Hmidi-Mateu [68] showed that there exist nonradial patches bifurcating from annuli at $\Omega<0$ and Gómez-Serrano [49] constructed non-radial, doubly connected stationary patches $(\Omega=0)$. Therefore we cannot expect a non-simply-connected rotating patch $D$ with $\Omega \leq 0$ to be radial for $\alpha \in(0,2)$.

However, if $D$ is a simply-connected stationary patch, then radial symmetry results were obtained in a series of works for $\alpha \in\left[0, \frac{5}{3}\right.$ ), which we review below. These works consider (Equation 1.1.6) in a more general context not limited to dimension 2: Let $K_{\alpha, d}$ be the fundamental solution of $(-\Delta)^{-1+\frac{\alpha}{2}}$ in $\mathbb{R}^{d}$ for $d \geq 2$, given by

$$
\begin{equation*}
K_{\alpha, d}:=-C_{\alpha, d}|x|^{-d+2-\alpha} \tag{1.2.1}
\end{equation*}
$$

for some $C_{\alpha, d}>0$; except that in the special case $-d+2-\alpha=0$ it becomes $K_{\alpha, d}=C_{d} \ln |x|$ for some $C_{d}>0$. Note that $K_{\alpha, d} \in L_{l o c}^{1}\left(\mathbb{R}^{d}\right)$ for all $\alpha<2$. Consider the following question:

Question 2. Let $\alpha \in[0,2)$. Assume $D \subset \mathbb{R}^{d}$ is a bounded domain such that

$$
\begin{equation*}
K_{\alpha, d} * 1_{D}-\frac{\Omega}{2}|x|^{2}=\text { const } \quad \text { on } \partial D \tag{1.2.2}
\end{equation*}
$$

for some $\Omega \leq 0$, where the constant is the same along all connected components of $\partial D$. Must $D$ be a ball in $\mathbb{R}^{d}$ ?

Positive answers to Question 2 were obtained in the $\Omega=0$ case for $\alpha<\frac{5}{3}$ in the following works. As we discussed before, Fraenkel [41] proved that $D$ must be a ball for $\alpha=0$. Also using the moving plane method, Reichel [99, Theorem 2], Lu-Zhu [83] and Han-Lu-Zhu [56] generalized this result to $\alpha \in[0,1)$. Here [83] also covered generic radially increasing potentials not too singular at the origin (which include all Riesz potentials $K_{\alpha, d}$ with $\alpha \in[0,1)$ ). Recently, Choksi-Neumayer-Topaloglu [29, Theorem 1.3] further pushed the range to $\alpha \in\left[0, \frac{5}{3}\right)$, leaving the range $\alpha \in\left[\frac{5}{3}, 2\right)$ an open problem. We point out that in all these results for $\alpha>0, \partial D$ was assumed to be at least $C^{1}$. All above results were obtained using the moving plane method.

In our third main result, we use a completely different approach to give an affirmative answer to Question 2 for all $\Omega \leq 0$ and $\alpha \in[0,2)$, under a weaker assumption on the regularity of $\partial D$.

Theorem C (= Theorem 3.1.2). Let $D$ be a bounded domain in $\mathbb{R}^{d}$ with Lipschitz boundary (and if $d=2$ we only require $\partial D$ to be rectifiable). If $D$ satisfies (Equation 1.2.2) for some $\Omega \leq 0$ and $\alpha \in[0,2)$, then it must be a ball in $\mathbb{R}^{d}$.

As a directly consequence, Theorem C implies that for the gSQG equation with $\alpha \in[0,2)$, any simply-connected rotating patch with $\Omega \leq 0$ must be a disk (see Theorem 3.1.4). In addition, in the smooth setting (Equation 1.1.7), we prove a similar result in Corollary 3.1.7 for uniformly-rotating solutions with $\Omega \leq 0$ for all $\alpha \in[0,2)$ : if the super level-sets $\left\{\omega_{0}>h\right\}$ are all simply-connected for all $h>0$, then $\omega_{0}$ must be radially decreasing.

Next we review the previous literature on uniformly-rotating solutions for the gSQG equation. Note that the case of $\alpha \in(0,2)$ is more challenging than the 2D Euler case, since
the velocity is more singular and this produces obstructions to the bifurcation theory when it comes to the choice of spaces and the regularity of the functionals involved in the construction. Hassainia-Hmidi [57] showed the existence of V-states with $C^{k}$ boundary regularity in the case $0<\alpha<1$, and in [21], Castro-Córdoba-Gómez-Serrano upgraded the result to show existence and $C^{\infty}$ boundary regularity in the remaining open cases: $\alpha \in[1,2)$ for the existence, $\alpha \in(0,2)$ for the regularity. In that case, the solutions bifurcate at angular velocities given by $\Omega_{m}^{\alpha}:=2^{\alpha-1} \frac{\Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}}\left(\frac{\Gamma\left(1+\frac{\alpha}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)}-\frac{\Gamma\left(m+\frac{\alpha}{2}\right)}{\Gamma\left(m+1-\frac{\alpha}{2}\right)}\right)$. This boundary regularity was subsequently improved to analytic in [22]. See also [62] for another family of rotating solutions, [68, 100] for the doubly connected case and [23] for a construction in the smooth setting.

One can check that $\Omega_{m}^{\alpha}$ are increasing functions of $m$ for any $\alpha$, whose limit is a finite number $\Omega^{\alpha}:=2^{\alpha-1} \frac{\Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}} \frac{\Gamma\left(1+\frac{\alpha}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)}$ for $\alpha \in[0,1)$, and $+\infty$ if $\alpha \geq 1$. It is then a natural question to ask whether there exist V-states (with area $\pi$ ) that rotate with angular velocity faster than $\Omega_{\alpha}$ for $\alpha \in(0,1)$. Our fourth main theorem answers this question among all simply-connected patches:

Theorem D (= Theorem 3.2.1). For $\alpha \in(0,1)$, let $1_{D}$ be a simply connected $V$-state of area $\pi$ and let its angular velocity be $\Omega \geq \Omega^{\alpha}$. Then $D$ must be the unit disk.

Finally, we illustrate a comparison of our results in Theorem C and D (in red color) with the previous results (in black color) in Figure Figure 1.2.


Figure 1.2: For gSQG in the patch setting, previous results on Question 1 are summarized in black color, with our results in Theorem C and D colored in red.

## Structure of the proofs: Theorem A-Theorem D

While all the previous symmetry results on Question 1 and Equation 2 [41, 83, 56, 59, 99, 29] are done by moving plane methods, our approaches are completely different, which have a variational flavor.

Theorem A is based on computing the first variation of the energy functional

$$
\mathcal{E}\left[1_{D}\right]=-\frac{1}{2} \int_{\mathbb{R}^{2}} 1_{D}(x)\left(1_{D} * \mathcal{N}\right)(x)-\frac{\Omega}{2}|x|^{2} 1_{D}(x) d x
$$

as we deform $D$ along a carefully chosen vector field. On the one hand, we show the first variation should be 0 if $D$ is a stationary/rotating patch with angular velocity $\Omega$; on the other hand, we show that the first variation must be non-zero if $\Omega \leq 0$ or $\Omega \geq \frac{1}{2}$, leading to a contradiction. In the proof, a rearrangement inequality due to Talenti [109] is crucial to get a sign condition. In the case when the patch is non-simply-connected, the choice of the right vector field is more involved since one has to deal with all the connected components at once, and even though the stream function $\psi$ is constant on each of the boundaries, it is not guaranteed that this constant will be the same for each of them, potentially yielding extra boundary terms which are not sign definite, and having to reprove a new version of the aforementioned inequality.

The smooth setting in Theorem B is based on a similar idea, but technically more difficult. The point of view is to approximate a smooth function by step functions and consider the above perturbation in each set where the step function is constant. To do this we need to obtain some quantitative (stability) estimates on our version of Talenti's rearrangement inequality, in particular in terms of the Fraenkel asymmetry of the domain in the spirit of Fusco-Maggi-Pratelli [42].

Theorem C is also based on a variational approach, but we need a different perturbation from the vector field in Theorem A, which heavily relies on the Newtonian potential, and fails for general Riesz potential $K_{\alpha}$. The key ingredient to prove Theorem C is to perturb $D$ using the continuous Steiner symmetrization [11], which has been successfully applied in other contexts by Carrillo-

Hittmeir-Volzone-Yao [18] (nonlinear aggregation models) or Morgan [90] (minimizers of the gravitational energy). This method is much more flexible and allows to treat more singular kernels than in the existing papers using moving plane methods. Due to the low regularity of the kernels, instead of computing the derivative of the energy under the perturbation, we work with finite differences instead.

Theorem D uses maximum principles and monotonicity formulas for nonlocal equations. The idea is to find the smallest disk $B(0, R)$ containing $D$ (which intersects $\partial D$ at some $x_{0}$ ), then use two different ways to compute $\nabla\left(1_{B(0, R) \backslash D} * K_{\alpha}\right)$ at $x_{0}$, and obtain a contradiction if $\Omega \geq \Omega_{\alpha}$ and $D$ is not a disk. The proof works for the full range of $\alpha \in[0,2)$, thus closing the problem raised by Hmidi [59] and de la Hoz-Hassainia-Hmidi-Mateu [66] among all simply-connected patches.

### 1.2.2 Rigidity/Flexibility results for vortex sheets [51, 52]

In chapter 4 and chapter 5, we will focus on the vortex sheet equation. There are very few known examples of nontrivial steady solutions, and in fact, other than the circle or the line, the list only comprises the segment of length $2 a$ and density

$$
\begin{equation*}
\gamma(x)=\Omega \sqrt{a^{2}-x^{2}}, \quad x \in[-a, a], \tag{1.2.3}
\end{equation*}
$$

which is a rotating solution with angular velocity $\Omega$ [6] and the family found by Protas-Sakajo [98], made out of segments rotating about a common center of rotation with endpoints at the vertices of a regular polygon. We remark that none of these are supported on a closed curve.

Numerically, O'Neil $[93,94]$ used point vortices to approximate the vortex sheet and compute uniformly rotating solutions and Elling [40] constructed numerically self-similar vortex sheets forming cusps. O'Neil [95, 96] also found numerically steady solutions which are combinations of point vortices and vortex sheets.

As in the patch/smooth setting, it is easy to see that if the $z_{i}$ 's are concentric circles with con-
stant $\varpi_{i}$ (with respect to the constant-speed parametrization) for every $i$, the solution is stationary, and it is also uniformly-rotating with any $\Omega \in \mathbb{R}$. We would like to understand the reverse implication, namely:

Question 3. Under what conditions must a stationary/uniformly-rotating vortex sheet be radially symmetric?

The next two theorems are the main results in chapter 4 and chapter 5 regarding the vortex sheet equations:

Theorem E. Let $\omega(x, t)=\omega_{0}\left(R_{\Omega t} x\right)$ be a stationary/uniformly-rotating vortex sheet with angular velocity $\Omega$. Assume that $\omega_{0}$ is concentrated on $\Gamma$, which is a finite union of smooth curves, and $\omega_{0}$ has positive vorticity strength on $\Gamma$. (See (H1)-(H3) in section 4.1 for the precise regularity and positivity assumptions.)

If $\Omega \leq 0, \Gamma$ must be a union of concentric circles, and $\omega_{0}$ must have constant strength along each circle (with respect to the constant-speed parametrization). In addition, if $\Omega<0$, all circles must be centered at the origin.

Theorem F (= Theorem 5.1.2). There exists nontrivial rotating vortex sheets with positive angular velocity $\Omega>0$, whose vortex strength is strictly positive but not concentrated on a non-radial curve.

Structure of the proofs: Theorem E-Theorem F
The proof of Theorem E is inspired by our recent rigidity result in the paper [50] on stationary and rotating solutions of the 2D Euler equations both in the smooth and vortex patch settings. To prove it, we constructed an appropriate functional and showed, on one hand, that any stationary solution had to be a critical point, and on the other, for any curve which is not a circle there existed a vector field along which the first variation was non-zero. This vector field is defined in terms of an elliptic equation in the interior of the patch. In the case of the vortex sheet, this is not possible
anymore. Instead, we desingularize the problem by considering patches of thickness $\sim \varepsilon$ which are tubular neighborhoods of the sheet. The drawback is that we lose the property that any stationary solution has to be a critical point if $\varepsilon>0$ and very careful, quantitative estimates need to be done to show that indeed the first variation of a stationary solution tends to 0 as $\varepsilon \rightarrow 0$. This setup is also reminiscent of the numerical work by Baker-Shelley [5], where they approximate the motion of a vortex sheet by a vortex patch of very small width. In [8], Benedetto-Pulvirenti proved the stability (for short time) of vortex sheet solutions with respect to solutions to 2D Euler with a thin strip of vorticity around a curve. See also the work by Caflisch-Lombardo-Sammartino [13] for more stability results with a different desingularization.

The main strategy to prove Theorem 5.1.2, is to employ bifurcation theory and try to bifurcate from the simple eigenvalue $b:=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \gamma(x, t) d x=2$. However, the standard methods (Crandall-Rabinowitz [35]) fail since the linearized operator around the circle does not satisfy the transversality condition: in other words, the nontrivial zero set is not transversal to the trivial one (disks with constant vorticity amplitude). This phenomenon is usually known in the literature as a degenerate bifurcation [73, 72]. Graphically, this can be seen in Figure Figure 5.1. The problem is that we no longer have a single branch emanating from the disk, but two, and therefore the linearized operator fails to describe the local behaviour at the bifurcation point. To overcome this issue, we first reduce the nonlinear problem to a suitable finite dimensional space by means of a Lyapunov-Schmidt reduction since the restriction of $D \mathcal{F}$ is an isomorphism between $\operatorname{Ker}(D \mathcal{F})^{\perp}$ and $\operatorname{Im}(D \mathcal{F})$. After having done so, we are left with a finite dimensional system and it is there where we perform a higher order expansion around the bifurcation point, since, as expected by the failure of the transversality condition, the first order approximation is identically zero. We obtain that in suitable coordinates, the zero sets of $\mathcal{F}$ behave as $x^{2}-y^{2}=0$ and thus two bifurcation branches emanate from the bifurcation point. The last part of the proof is devoted to handle the higher order terms, which can be controlled if we restrict the bifurcation domain to a suitable small enough neighbourhood. We mention here that this technique had been successfully employed by

Hmidi-Mateu [61] (in the hyperbolic case) and Hmidi-Renault [65] (in the elliptic case).

### 1.2.3 Quantitative estimates for rotating vortex patches [97]

The goal of chapter 6 is to establish some quantitative estimates for non-radial simply-connected rotating patches, which are known to exist. From now on, we assume that a bounded domain $D$ is simply-connected and has $C^{2}$ boundary. If $\omega(x, t):=1_{D}\left(R_{\Omega t} x\right)$ is a uniformly-rotating patch, then the net velocity in the rotating frame has no contribution to the deformation of the boundary $\partial D$, namely, $\nabla^{\perp}\left(\left(1_{D} * \mathcal{N}\right)-\frac{\Omega}{2}|x|^{2}\right) \cdot \vec{n}=0$, where $\vec{n}$ denotes the outer normal vector on $\partial D$. By integrating this along the boundary, one can derive the following equation for the relative stream function $\Psi$ :

$$
\begin{equation*}
\Psi(x):=1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}=\text { constant } \quad \text { for all } x \in \partial D \tag{1.2.4}
\end{equation*}
$$

In the rest of the paper, we say a pair $(D, \Omega)$ is a solution to (Equation 1.2.4) if $1_{D}$ and $\Omega$ satisfy the equation (Equation 1.2.4).

## Small angular velocity $\Omega$

Our first main result is about the outmost point on $\partial D$ when the angular velocity $\Omega$ is small. As mentioned earlier, ellipses are uniformly-rotating solutions. More precisely, an ellipse with semiaxes $a, b$ is rotating with angular velocity $\Omega=\frac{a b}{(a+b)^{2}}$. By imposing $b=\frac{1}{a}$ to keep the area of the patch equal to $\pi$, one can easily see that for any $0<\Omega \leq \frac{1}{4}$, there exists an ellipse that is rotating with the given $\Omega$. Moreover, the boundary is stretching as $\Omega$ tends to 0 in the sense that the length of the major axis is comparable with $\Omega^{-\frac{1}{2}}$. Note that ellipses are not the only uniformly-rotating solutions for small angular velocities. For example, the existence of secondary bifurcations from ellipses was numerically observed by Kamm in his thesis [71] and theoretically proved in [22, 60].

Thus it is a natural question whether every non-radial simply-connected rotating patch with a fixed area and $0<\Omega \ll 1$ must have its outmost point very far from the origin (center of rotation). In the next theorem, we prove this is indeed true.

Theorem G. . Let $D \subset \mathbb{R}^{2}$ be a simply connected domain such that $|D|=|B|=\pi$, where $B$ is the unit disk centered at the origin. Then there exist positive constants $\Omega_{0}$ and $\kappa_{0}$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega \in\left(0, \Omega_{0}\right)$, then either $D=B$, or

$$
\begin{equation*}
\sup _{x \in \partial D}|x|>\kappa_{0} \Omega^{-\frac{1}{2}} \tag{1.2.5}
\end{equation*}
$$

Remark 1.2.1. Note that the power $-\frac{1}{2}$ is sharp since it is achieved by ellipses. Furthermore, one can easily show that (Equation 1.2.4) is scaling invariant in the sense that if $(D, \Omega)$ is a solution, then $\left(D_{a}, \Omega\right)$ is also a solution for any $a>0$, where $D_{a}:=\left\{a x \in \mathbb{R}^{2}: x \in D\right\}$. Therefore without the restriction on the size of the patch, (Equation 1.2.5) reads as

$$
\frac{1}{\sqrt{|D|}} \sup _{x \in \partial D}|x|>\frac{\kappa_{0}}{\sqrt{\pi}} \Omega^{-\frac{1}{2}}
$$

m-fold symmetric patches
It has been known since the work of Burbea [12] that there are $m$-fold symmetric rotating patches for every integer $m \geq 2$. From the numerical results [36, 58], it appears that for $m \gg 1$, the angular velocity $\Omega$ along the bifurcation curve is very close to $\frac{1}{2}$ (i.e. $0<\frac{1}{2}-\Omega \ll 1$ for $m \gg 1$ ). But there are no such type of quantitative estimates so far. In the next theorem, we will derive a lower bound of the angular velocity by imposing large $m$.

Theorem H. There exist $m_{0} \geq 2$ and $c>0$ such that if $(\Omega, D)$ is a solution to (Equation 1.2.4)
and $D$ is simply-connected, non-radial, and $m$-fold symmetric for some $m \geq m_{0}$ then

$$
\frac{1}{2}-\Omega \leq \frac{c}{m}
$$

We emphasize that this theorem holds for a general simply-connected patch, which does not need to lie on the bifurcation curve.

For $m$-fold symmetric solutions on the global bifurcation curves constructed in [58], we will also estimate the difference between a rotating patch and the unit disk. To be precise, we will focus on the curves,

$$
\mathscr{C}_{m}:=\left\{\left(\tilde{u}_{m}(s), \tilde{\Omega}_{m}(s)\right) \in C^{2}(\mathbb{T}) \times\left(0, \frac{1}{2}\right): s \in[0, \infty)\right\} \quad \text { for } m \geq 2
$$

that satisfy the following properties (see [58, Theorem 1.1] for the details):
(A1) $\tilde{u}_{m}(s) \in\left\{u \in C^{2}(\mathbb{T}): u(\theta)=\sum_{n=1}^{\infty} a_{n} \cos (n m \theta)\right.$ for some $\left(a_{n}\right)_{n=1}^{\infty}$ and $\left.u>-1\right\}$.
(A2) $\left(D^{\tilde{u}_{m}(s)}, \tilde{\Omega}_{m}(s)\right)$ is a solution for (Equation 1.2.4),
where $D^{u}:=\left\{r(\cos \theta, \sin \theta) \in \mathbb{R}^{2}: 0 \leq r<(1+u(\theta)), \theta \in \mathbb{T}\right\}$.
(A3) $\partial_{\theta} u(\theta)<0$ for all $\theta \in\left(0, \frac{\pi}{m}\right)$, where $u=\tilde{u}_{m}(s)$.
(A4) $\left(\tilde{u}_{m}(0), \tilde{\Omega}_{m}(0)\right)=\left(0, \frac{m-1}{2 m}\right)$.
For such curves, we have the following theorem:
Theorem I. Let $\mathscr{C}_{m}:=\left\{\left(\tilde{u}_{m}(s), \tilde{\Omega}_{m}(s)\right) \in C^{2}(\mathbb{T}) \times\left(0, \frac{1}{2}\right): 0 \leq s<\infty\right\}$ be a continuous curve that satisfies the properties (A1)-(A4). Then there exist constants $c>0$ and $m_{0} \geq 3$ such that if $m \geq m_{0}$, then

$$
\left\|\tilde{u}_{m}(s)\right\|_{L^{\infty}(\mathbb{T})} \leq \frac{c}{m} \text { for all } s \geq 0
$$

Although each curve emanates from the unit disk, the possibility that $\min _{\theta \in \mathbb{T}}\left(1+\tilde{u}_{m}(s)\right)$ tends to 0 along the curve has not been completely eliminated ([58, Theorem 4.6, Lemma 6.6]), while it does certainly happen for ellipses $(m=2)$. The significant difference between $m=2$ and $m \geq 3$ is that if $m \geq 3$, then the stream function $\left(1_{D} * \mathcal{N}\right)$ behaves quite nicely, namely, $\frac{\left|\nabla\left(1_{D} * \mathcal{N}\right)\right|}{|x|}$ is globally bounded (especially near the origin) independently of $D$ (Lemma 6.2.1. See also [38, 39], and the references therein, where global boundedness of gradient of $m$-fold symmetric stream functions was proved). This will play a crucial role to eliminate the scenario that $\partial D$ almost touches the origin when $\frac{1}{2}-\Omega$ is sufficiently large compared to $\frac{1}{m}$ in Lemma 6.2.2.

We summarize the main results in Figure Figure 1.3


Figure 1.3: Illustration of the main results. a) Any patch with $\Omega \ll 1$ must have its outmost point very far from the origin. b) A bifurcation curve of $m$-fold symmetric patches cannot be continued beyond the blue dashed lines for large $m$.

## Structure of the proofs: Theorem $G$ - Theorem I

The starting point for Theorem $G$ and Theorem $H$ is the variational formulation of (Equation 1.2.4), used by Gómez-Serrano, Shi, Yao and the author in [50]. Namely, if $(D, \Omega)$ is a solution to (Equation 1.2.4) and $\partial D$ is $C^{2}$, then formally, $1_{D}$ can be thought of as a critical point of the
functional,

$$
\mathcal{I}(\rho):=\frac{1}{2} \int_{\mathbb{R}^{2}} \rho * \mathcal{N}(x) \rho(x) d x-\Omega \int_{\mathbb{R}^{2}} \frac{|x|^{2}}{2} \rho(x) d x=: \mathcal{I}_{1}(\rho)-\Omega \mathcal{I}_{2}(\rho),
$$

under measure-preserving perturbations. More precisely, it holds that

$$
\begin{equation*}
\int_{D} \vec{v} \cdot \nabla \Psi(x) d x=0, \quad \text { for any } v \in C^{2}(\bar{D}) \text { such that } \nabla \cdot \vec{v}=0 \tag{1.2.6}
\end{equation*}
$$

Indeed, (Equation 1.2.6) follows directly from (Equation 1.2.4) and the integration by parts. By choosing a specific vector field $\vec{v}:=x+\nabla p$, where $p$ is defined as the solution to the Poisson equation,

$$
\begin{cases}\Delta p=-2 & \text { in } D  \tag{1.2.7}\\ p=0 & \text { on } \partial D\end{cases}
$$

Gómez-Serrano et al. derived the following identity for uniformly-rotating patches:

$$
\begin{equation*}
2 \Omega\left(\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}\right)=(1-2 \Omega)\left(\frac{|D|^{2}}{4 \pi}-\int_{D} p d x\right) . \tag{1.2.8}
\end{equation*}
$$

Note that both parentheses are strictly positive if $D \neq B$, where $B$ is the unit disk centered at the origin. Thanks to the result by Brasco-De Philippis-Velichkov in [10, Proposition 2.1], one can find a lower bound of the right-hand side of (Equation 1.2.8) in terms of $|D \triangle B|$, namely, $\frac{|D|^{2}}{4 \pi}-\int_{D} p d x \gtrsim|D \triangle B|^{2}$. Hence (Equation 1.2.8) yields that

$$
|D \triangle B|^{2} \lesssim \Omega\left(\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}\right)<\Omega \sup _{x \in \partial D}|x|^{2}
$$

for $\Omega \ll 1$. Therefore we only need to rule out the case where $|D \triangle B|$ is small. Assuming $|D \triangle B|$ and $\Omega$ are sufficiently small, we will prove (Lemma 6.1.6) that $D$ is necessarily star-shaped and
the boundary can be parametrized by $(1+u(x)) x$, for $x \in \partial B_{1}$ and some $u \in C^{2}(\partial B)$. However, the difficulty is that we have $|D \triangle B| \sim\|u\|_{L^{1}(\partial B)}$ and $\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi} \sim\|u\|_{L^{2}(\partial B)}^{2}$, while $L^{1}$ and $L^{2}$ are not comparable. The key idea is to use a different vector $\vec{v}:=x-2 \nabla\left(1_{D} * \mathcal{N}\right)$ in (Equation 1.2.6), which gives another identity for any simply-connected rotating patches,

$$
\begin{equation*}
\left(\frac{1}{2}-\Omega\right)\left(\int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}\right)=\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \tag{1.2.9}
\end{equation*}
$$

Thanks to the result of Loeper [79, Proposition 3.1], the right-hand side in (Equation 1.2.9) can be estimated in terms of 2 -Wasserstein distance between $1_{D} d x$ and $1_{B} d x$ (see Proposition 6.1.9). In the proof of Proposition 6.1.8, we will construct an explicit transport map and obtain the bound for the right-hand side: If $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$,

$$
\begin{equation*}
\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \lesssim\left(a \int_{\mathbb{T}}|u|^{2} d \theta+\frac{1}{a} \int_{\mathbb{T}} f(\theta)^{2} d \theta\right) \tag{1.2.10}
\end{equation*}
$$

where $f(\theta):=\int_{0}^{\theta} u(s)^{2}+2 u(s) d s$ and $a \in\left(2\|u\|_{L^{\infty}(\mathbb{T})}, 1\right)$. Since $\|f\|_{L^{\infty}(\mathbb{T})} \lesssim\|u\|_{L^{1}(\mathbb{T})}$, (Equation 1.2.9) and (Equation 1.2.10) will give us $\|u\|_{L^{1}(\mathbb{T})} \sim\|u\|_{L^{2}(\mathbb{T})}$ for $0<\Omega \ll 1$, if we can choose $a$ sufficiently small.

The proof of Theorem H also relies on the identity (Equation 1.2.9). By imposing $m$-fold symmetry on the patch, we can lower the total cost of the transportation, from which we can obtain a suitable upper bound of $\frac{1}{2}-\Omega$ when $m$ is sufficiently large. Indeed, if $u$ is $\frac{2 \pi}{m}$ periodic, then $f$ is also $\frac{2 \pi}{m}$ periodic as well. Thus by choosing large $m$, we can lower $\|f\|_{L^{\infty}(\mathbb{T})}$ on the right-hand side in (Equation 1.2.10) by using Jensen's inequality.

Theorem I will be proved by showing that if $\left\|\tilde{u}_{m}(s)\right\|_{L^{\infty}}$ is too large, then $\frac{1}{2}-\Omega$ must be large enough to contradict Theorem H. The main difficulty is that $\frac{1}{2}-\Omega$ can be estimated in terms of $\left\|\tilde{u}_{m}(s)\right\|_{L^{2}(\mathbb{T})}$ by using the identity (Equation 1.2.8) (Lemma 6.2.4), while $L^{2}$ and $L^{\infty}$ are not comparable. We resolve this issue by estimating the gradient of the stream function in a very
delicate way (Lemma 6.2.5 and 6.2.6).

## CHAPTER 2 <br> RIGIDITY RESULTS FOR 2D EULER

In chapter 2-chapter 3, we study radial symmetry in 2D Euler equation and generalized SQG equations.

Notations Through chapter 2-chapter 3, we use the following notations.
For a simple closed curve $\Gamma$, denote $\operatorname{int}(\Gamma)$ by its interior, which is the bounded connected component of $\mathbb{R}^{2}$ separated by the curve $\Gamma$. Note that the Jordan-Schoenflies theorem guarantees that $\operatorname{int}(\Gamma)$ is open and simply connected.

We say that two disjoint simple closed curves $\Gamma_{1}$ and $\Gamma_{2}$ are nested if $\Gamma_{1} \subset \operatorname{int}\left(\Gamma_{2}\right)$ or vice versa. We say that two connected domains $D_{1}, D_{2}$ are nested if one is contained in a hole of the other one.

For a bounded connected domain $D \subset \mathbb{R}^{2}$, we denote by $\partial_{\text {out }} D$ its outer boundary. And if $D$ is doubly-connected, we denote by $\partial_{\text {in }} D$ its inner boundary,

For a set $D$, we use $1_{D}(x)$ to denote its indicator function. And for a statement $S$, we let $\mathbb{1}_{S}=\left\{\begin{array}{ll}1 & \text { if } S \text { is true } \\ 0 & \text { if } S \text { is false }\end{array} .\left(\right.\right.$ e.g. $\left.\mathbb{1}_{\pi<3}=0\right)$.

For a domain $U \subset \mathbb{R}^{2}$, in the boundary integral $\int_{\partial U} \vec{n} \cdot \vec{f} d \sigma$, the vector $\vec{n}$ is taken as the outer normal of the domain $U$ in that integral.

### 2.1 Radial symmetry of steady/rotating patches for 2D Euler equation

Throughout this section we work with the 2D Euler equation (Equation 1.1.1) in the patch setting. For a stationary or uniformly-rotating patch $D$ with angular velocity $\Omega \in \mathbb{R}$, let

$$
f_{\Omega}(x):=\left(1_{D} * \mathcal{N}\right)(x)-\frac{\Omega}{2}|x|^{2}
$$

Recall that in (Equation 1.1.6) we have shown that $f_{\Omega} \equiv C_{i}$ on each connected component of $\partial D$, where the constants can be different on different connected components.

Our goal in this section is to prove Theorem A, which completely answers Question 1 for 2D Euler patches. As we described in the introduction, our proof has a variational flavor, which is done by perturbing $D$ by a carefully chosen vector field, and compute the first variation of an associated energy functional in two different ways. In subsection 2.1.1, we will define the energy functional and the perturbation vector field, and give a one-page proof in Theorem 2.1.2 that answers Question 1 among simply-connected patches. (Note that even among simply-connected patches, it is an open question whether every rotating patch with $\Omega>\frac{1}{2}$ or $\Omega<0$ must be a disk.) In the following subsections, we further develop this method, and modify our perturbation vector field to cover non-simply-connected patches.

### 2.1.1 Warm-up: radial symmetry of simply-connected rotating patches

We begin by providing a sketch and some motivations of our approach, and then give a rigorous proof afterwards in Theorem 2.1.2. Suppose that $D$ is a $C^{1}$ simply-connected rotating patch with angular velocity $\Omega$ that is not a disk. We perturb $D$ in "time" (here the "time" $t$ is just a name for our perturbation parameter, and is irrelevant with the actual time in the Euler equation) with a velocity field $\vec{v}(x) \in C^{1}(D) \cap C(\bar{D})$ that is divergence-free in $D$, which we will fix later. That is, consider the transport equation

$$
\rho_{t}+\nabla \cdot(\rho \vec{v})=0
$$

with $\rho(\cdot, 0)=1_{D}$. We then investigate how the "energy functional"

$$
\mathcal{E}[\rho]:=-\int_{\mathbb{R}^{2}} \frac{1}{2} \rho(x)(\rho * \mathcal{N})(x)-\frac{\Omega}{2}|x|^{2} \rho(x) d x
$$

changes in time under the perturbation. Formally, we have

$$
\begin{align*}
\left.\frac{d}{d t} \mathcal{E}[\rho]\right|_{t=0} & =-\int_{\mathbb{R}^{2}} \rho_{t}(x, 0)\left((\rho(\cdot, 0) * \mathcal{N})(x)-\frac{\Omega}{2}|x|^{2}\right) d x  \tag{2.1.1}\\
& =-\int_{D} \vec{v}(x) \cdot \nabla\left(\left(1_{D} * \mathcal{N}\right)(x)-\frac{\Omega}{2}|x|^{2}\right) d x
\end{align*}
$$

The above transport equation and the energy functional only serve as our motivation, and will not appear in the proof. In the actual proof we only focus on the right hand side of (Equation 2.1.1), which is an integral that is well-defined by itself:

$$
\begin{equation*}
\mathcal{I}:=-\int_{D} \vec{v}(x) \cdot \nabla\left(\left(1_{D} * \mathcal{N}\right)(x)-\frac{\Omega}{2}|x|^{2}\right) d x=-\int_{D} \vec{v} \cdot \nabla f_{\Omega} d x . \tag{2.1.2}
\end{equation*}
$$

We will use two different ways to compute $\mathcal{I}$, and show that if $D$ is not a disk, the two ways lead to a contradiction for $\Omega \leq 0$ or $\Omega \geq \frac{1}{2}$.

On the one hand, since $f_{\Omega}$ is a constant on $\partial D$ (denote it by $c$ ), the divergence theorem yields the following for every $\vec{v} \in C^{1}(D) \cap C(\bar{D})$ that is divergence-free in $D$ :

$$
\begin{equation*}
\mathcal{I}=-c \int_{\partial D} \vec{n} \cdot \vec{v} d \sigma+\int_{D}(\nabla \cdot \vec{v}) f_{\Omega} d x=-c \int_{D} \nabla \cdot \vec{v} d x+\int_{D}(\nabla \cdot \vec{v}) f_{\Omega} d x=0 \tag{2.1.3}
\end{equation*}
$$

On the other hand, we fix $\vec{v}$ as follows, which is at the heart of our proof. Let $\vec{v}(x):=-\nabla \varphi(x)$ in $D$, where

$$
\begin{equation*}
\varphi(x):=\frac{|x|^{2}}{2}+p(x) \quad \text { in } D \tag{2.1.4}
\end{equation*}
$$

with $p(x)$ being the solution to Poisson's equation

$$
\begin{cases}\Delta p(x)=-2 & \text { in } D  \tag{2.1.5}\\ p(x)=0 & \text { on } \partial D\end{cases}
$$

Note that $\varphi$ is harmonic in $D$, thus $\vec{v}$ is indeed divergence-free in $D$. This definition of $\vec{v}$ is motivated by the fact that among all divergence-free vector fields in $D$, such $\vec{v}$ is the closest one to $-\vec{x}$ in the $L^{2}(D)$ distance. (In fact, such $\vec{v}$ is connected to the gradient flow of $\int_{D} \frac{|x|^{2}}{2} d x$ in the metric space endowed by 2-Wasserstein distance, under the constraint that $|D(t)|$ must remain constant $[88,89,1]$.) Formally, one expects that $D$ becomes "more symmetric" as we perturb it by $\vec{v}$, which inspires us to consider the first variation of $\mathcal{E}$ under such perturbation.

In the proof we will show that with such choice of $\vec{v}$, we can compute $\mathcal{I}$ in another way and obtain that $\mathcal{I}>0$ for $\Omega \leq 0$ and $\mathcal{I}<0$ for $\Omega \geq \frac{1}{2}$. Therefore in both cases, we obtain a contradiction with $\mathcal{I}=0$ in (Equation 2.1.3).

Our proof makes use of a rearrangement inequality for solutions to elliptic equations, which is due to Talenti [109]. Below is the form that we will use; the original theorem works for a more general class of elliptic equations.

Proposition 2.1.1 ([109], Theorem 1). Let $D \subset \mathbb{R}^{2}$ be a bounded domain with $C^{1}$ boundary, and let $p$ be defined as in (Equation 2.1.5). Let $B$ be a disk centered at the origin with $|B|=|D|$, and let $p_{B}$ solve (Equation 2.1.5) in B. Then we have $p^{*} \leq p_{B}$ pointwise in $B$, where $p^{*}$ is the radial decreasing rearrangement of $p^{*}$. As a consequence, we have

$$
\int_{D} p(x) d x \leq \frac{1}{4 \pi}|D|^{2}
$$

and the equality is achieved if and only if $D$ is a disk.

Now we are ready to prove the following theorem, saying that any simply-connected station-
ary/rotating patch with $\Omega \leq 0$ or $\Omega \geq \frac{1}{2}$ must be a disk. Interestingly, the same proof can treat the two disjoint intervals $\Omega \leq 0$ and $\Omega \geq \frac{1}{2}$ all at once.

Theorem 2.1.2. Let $D$ be a simply-connected domain with $C^{1}$ boundary. If $D$ is a rotating patch solution with angular velocity $\Omega$, where $\Omega \leq 0$ or $\Omega \geq \frac{1}{2}$, then $D$ must be a disk, and it must be centered at the origin unless $\Omega=0$.

Proof. Let $D$ be a rotating patch with $\Omega \in(-\infty, 0] \cup\left[\frac{1}{2}, \infty\right)$. As we described above, in this theorem we will use two different ways to compute the integral $\mathcal{I}$ defined in (Equation 2.1.2), where we fix $\vec{v}(x):=-\nabla \varphi(x)$, with $\varphi$ and $p$ defined as in (Equation 2.1.4) and (Equation 2.1.5).

On the one hand, we have that $\vec{v}$ is divergence free in $D$, and elliptic regularity theory immediately yields that $\vec{v} \in C^{1}(D) \cap C(\bar{D})$. Using the assumption that $D$ is a rotating patch, we know $f_{\Omega}$ is a constant on $\partial D$. (Note that $\partial D$ is a connected closed curve since we assume $D$ is simply-connected). Thus the computation in (Equation 2.1.3) directly gives that $\mathcal{I}=0$.

On the other hand, we compute $\mathcal{I}$ as follows:

$$
\begin{equation*}
\mathcal{I}=-\int_{D} \vec{v} \cdot \nabla f_{\Omega} d x=\int_{D} \nabla \varphi \cdot \nabla f_{\Omega} d x=\underbrace{\int_{D} x \cdot \nabla f_{\Omega} d x}_{=: \mathcal{I}_{1}}+\underbrace{\int_{D} \nabla p \cdot \nabla f_{\Omega} d x}_{=: \mathcal{I}_{2}} . \tag{2.1.6}
\end{equation*}
$$

For $\mathcal{I}_{1}$, we have

$$
\begin{align*}
\mathcal{I}_{1} & =\int_{D} x \cdot \nabla\left(1_{D} * \mathcal{N}\right) d x-\int_{D} x \cdot \Omega x d x \\
& =\frac{1}{2 \pi} \int_{D} \int_{D} \frac{x \cdot(x-y)}{|x-y|^{2}} d y d x-\Omega \int_{D}|x|^{2} d x \\
& =\frac{1}{4 \pi} \int_{D} \int_{D} \frac{x \cdot(x-y)-y \cdot(x-y)}{|x-y|^{2}} d y d x-\Omega \int_{D}|x|^{2} d x  \tag{2.1.7}\\
& =\frac{1}{4 \pi}|D|^{2}-\Omega \int_{D}|x|^{2} d x
\end{align*}
$$

where the third equality is obtained by exchanging $x$ with $y$ in the first integral, then taking average with the original integral. To compute $\mathcal{I}_{2}$, using the divergence theorem (and the fact that $p=0$ on
$\partial D)$, we have

$$
\begin{equation*}
\mathcal{I}_{2}=-\int_{D} p \Delta f_{\Omega} d x=(2 \Omega-1) \int_{D} p d x \tag{2.1.8}
\end{equation*}
$$

Plugging (Equation 2.1.7) and (Equation 2.1.8) into (Equation 2.1.6) gives

$$
\begin{equation*}
\mathcal{I}=\frac{1}{4 \pi}|D|^{2}-\Omega \int_{D}|x|^{2} d x+(2 \Omega-1) \int_{D} p d x \tag{2.1.9}
\end{equation*}
$$

When $\Omega=0$, Proposition 2.1.1 directly gives that $\mathcal{I}>0$ if $D$ is not a disk, contradicting $\mathcal{I}=0$.
When $\Omega \in(-\infty, 0) \cup\left[\frac{1}{2}, \infty\right)$, let $B$ be a disk centered at the origin with the same area as $D$. Towards a contradiction, assume $D \neq B$. Among all sets with the same area as $D$, the disk $B$ is the unique one that minimizes the second moment, thus we have

$$
\int_{D}|x|^{2} d x>\int_{B}|x|^{2} d x=\frac{1}{2 \pi}|D|^{2}
$$

where the last step follows from an elementary computation. Plugging this into (Equation 2.1.9) gives the following inequality for $\Omega \in\left[\frac{1}{2}, \infty\right)$ :

$$
\mathcal{I}<\frac{1}{4 \pi}|D|^{2}-\frac{\Omega}{2 \pi}|D|^{2}+(2 \Omega-1) \int_{D} p d x=(1-2 \Omega)\left(\frac{1}{4 \pi}|D|^{2}-\int_{D} p d x\right) \leq 0 .
$$

On the other hand, for $\Omega \in(-\infty, 0)$, we have

$$
\mathcal{I}>\frac{1}{4 \pi}|D|^{2}-\frac{\Omega}{2 \pi}|D|^{2}+(2 \Omega-1) \int_{D} p d x=(1-2 \Omega)\left(\frac{1}{4 \pi}|D|^{2}-\int_{D} p d x\right)>0
$$

and we get a contradiction to $\mathcal{I}=0$ in all the cases, thus the proof is finished.

Remark 2.1.3. In fact, one can easily check that Theorem 2.1.2 holds for a bounded disconnected patch $D=\dot{U}_{i=1}^{N} D_{i}$ with $C^{1}$ boundary, as long as each connected component $D_{i}$ is simply-
connected. Here the proof remains the same, except a small change in the $\mathcal{I}=0$ proof: since now we have $f_{\Omega}=c_{i}$ on $\partial D_{i}$, (Equation 2.1.3) should be replaced by

$$
\mathcal{I}=-\sum_{i=1}^{N}\left(c_{i} \int_{\partial D_{i}} \vec{n} \cdot \vec{v} d \sigma+\int_{D_{i}}(\nabla \cdot \vec{v}) f_{\Omega} d x\right)=0 .
$$

Even in the regime $\Omega \in\left(0, \frac{1}{2}\right)$, where non-radial rotating patches are known to exist (recall that there exist patches bifurcating from a disk at $\Omega_{m}=\frac{m-1}{2 m}$ for all $m \geq 2$ ), our approach still allows us to obtain the following quantitative estimate, saying that if a simply-connected patch $D$ rotates with angular velocity $\Omega \in\left(0, \frac{1}{2}\right)$ that is very close to $\frac{1}{2}$, then $D$ must be very close to a disk, in the sense that their symmetric difference must be small.

Corollary 2.1.4. Let $D$ be a simply-connected domain with $C^{1}$ boundary. Assume $D$ is a rotating patch solution with angular velocity $\Omega$, where $\Omega \in\left(\frac{1}{4}, \frac{1}{2}\right)$. Let $\delta:=\frac{1}{2}-\Omega$. Then we have

$$
|D \triangle B| \leq 2 \sqrt{2 \delta}|D|
$$

where $B$ is the disk centered at the origin with the same area as $D$.

Proof. In the proof of Theorem 2.1.2, combining the equation $\mathcal{I}=0$ and (Equation 2.1.9) together, we have that

$$
\frac{1}{4 \pi}|D|^{2}-\Omega \int_{D}|x|^{2} d x-(1-2 \Omega) \int_{D} p d x=0
$$

Dividing both sides by $\Omega$ and rearranging the terms, we obtain

$$
\int_{D}|x|^{2} d x-\frac{1}{2 \pi}|D|^{2}=\frac{1-2 \Omega}{\Omega}\left(\frac{1}{4 \pi}|D|^{2}-\int_{D} p d x\right) \leq \frac{2 \delta|D|^{2}}{\pi},
$$

where in the inequality we used that $2 \delta:=1-2 \Omega, \Omega>\frac{1}{4}$, and $\int_{D} p d x \geq 0$.

Since $\int_{B}|x|^{2} d x=\frac{1}{2 \pi}|D|^{2}$, the above inequality implies that

$$
\begin{equation*}
\int_{D \backslash B}|x|^{2} d x-\int_{B \backslash D}|x|^{2} d x \leq \frac{2 \delta|D|^{2}}{\pi} \tag{2.1.10}
\end{equation*}
$$

Since $D$ and $B$ has the same area, let us denote $\beta:=|D \backslash B|=|B \backslash D|$. Among all sets $U \subset B^{c}$ with area $\beta, \int_{U}|x|^{2} d x$ is minimized when $U$ is an annulus with area $\beta$ and inner circle coinciding with $\partial B$, thus an elementary computation gives

$$
\int_{D \backslash B}|x|^{2} d x \geq \inf _{U \subset B^{c},|U|=\beta} \int_{U}|x|^{2} d x=\frac{\beta(2|B|+\beta)}{2 \pi}
$$

Likewise, among all sets $V \subset B$ with area $\beta, \int_{V}|x|^{2} d x$ is maximized when $V$ is an annulus with area $\beta$ and outer circle coinciding with $\partial B$, thus

$$
\int_{B \backslash D}|x|^{2} d x \leq \sup _{V \subset B,|V|=\beta} \int_{V}|x|^{2} d x=\frac{\beta(2|B|-\beta)}{2 \pi} .
$$

Subtracting these two inequalities yields

$$
\int_{D \backslash B}|x|^{2} d x-\int_{B \backslash D}|x|^{2} d x \geq \frac{\beta^{2}}{\pi},
$$

and combining this with (Equation 2.1.10) immediately gives

$$
\beta^{2} \leq 2 \delta|D|^{2},
$$

thus $|D \triangle B|=2 \beta \leq 2 \sqrt{2 \delta}|D|$.

### 2.1.2 Radial symmetry of non-simply-connected stationary patches

In this subsection, we aim to prove radial symmetry of a connected rotating patch $D$ with $\Omega \leq 0$, where $D$ is allowed to be non-simply-connected. Let $D \subset \mathbb{R}^{2}$ be a bounded connected domain
with $C^{1}$ boundary. Assume $D$ has $n$ holes with $n \geq 0$, and then let $h_{1}, \cdots, h_{n} \subset \mathbb{R}^{2}$ denote the $n$ holes of $D$ (each $h_{i}$ is a bounded open set). Note that $\partial D$ has $n+1$ connected components: they include the outer boundary of $D$, which we denote by $\partial D_{0}$, and the inner boundaries $\partial h_{i}$ for $i=1, \ldots, n$.

To begin with, we point out that even for the steady patch case $\Omega=0$, the proof of Theorem 2.1.2 cannot be directly adapted to the non-simply-connected patch. If we define $\vec{v}$ in the same way, then the second way to compute $\mathcal{I}$ still goes through (since Proposition 2.1 .1 still holds for non-simply-connected $D$ ), and leads to $\mathcal{I}>0$ if $D$ is not a disk. But the first way to compute $\mathcal{I}$ no longer gives $\mathcal{I}=0$ : if $D$ is stationary and not simply-connected, $f(x):=\left(1_{D} * \mathcal{N}\right)(x)$ may take different constant values on different connected components of $\partial D$, thus the identity (Equation 2.1.3) no longer holds.

In order to fix this issue, we still define $\vec{v}=-\nabla \varphi=-\nabla\left(\frac{|x|^{2}}{2}+p\right)$, but modify the definition of $p$ in the following lemma. Compared to the previous definition (Equation 2.1.5), the difference is that $p$ now takes different values $0, c_{1}, \ldots, c_{n}$ on each connected component of $\partial D$. The lemma shows that there exist values of $\left\{c_{i}\right\}_{i=1}^{n}$, such that $\int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma=-2\left|h_{i}\right|$ along the boundary of each hole. As we will see later, this leads to $\int_{\partial h_{i}} \vec{v} \cdot \vec{n} d \sigma=0$ for $i=1, \ldots, n$, which ensures $\mathcal{I}=0$. (Of course, with $p$ defined in the new way, the second way of computing $\mathcal{I}$ no longer follows from Proposition 2.1.1, and we will take care of this later in Proposition 2.1.6.)

Lemma 2.1.5. Let $D, h_{i}$ and $\partial D_{0}$ be given as in the first paragraph of Section subsection 2.1.2. Then there exist positive constants $\left\{c_{i}\right\}_{i=1}^{n}$, such that the solution $p: \bar{D} \rightarrow \mathbb{R}$ to the Poisson equation

$$
\begin{cases}\Delta p=-2 & \text { in } D  \tag{2.1.11}\\ p=c_{i} & \text { on } \partial h_{i} \text { for } i=1, \ldots, n \\ p=0 & \text { on } \partial D_{0}\end{cases}
$$

satisfies

$$
\begin{equation*}
\int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma=-2\left|h_{i}\right| \quad \text { for } i=1, \ldots, n \tag{2.1.12}
\end{equation*}
$$

Here $\left|h_{i}\right|$ is the area of the domain $h_{i} \subset \mathbb{R}^{2}$.

Proof. Let $u$ satisfy that

$$
\begin{cases}\Delta u=-2 & \text { in } D \\ u=0 & \text { on } \partial D\end{cases}
$$

Furthermore let the function $v_{j}$ for $j=1, \ldots, n$ be the solution to

$$
\begin{cases}\Delta v_{j}=0 & \text { in } D \\ v_{j}=0 & \text { on } \partial D \backslash \partial h_{j} \\ v_{j}=1 & \text { on } \partial h_{j}\end{cases}
$$

Now we consider the following linear equation,

$$
\begin{equation*}
A x=b, \tag{2.1.13}
\end{equation*}
$$

where $A_{i, j}=\int_{\partial h_{i}} \nabla v_{j} \cdot \vec{n} d \sigma$ and $b_{i}=-2\left|h_{i}\right|-\int_{\partial h_{i}} \nabla u \cdot \vec{n} d \sigma$. We argue that (Equation 2.1.13) has a unique solution. Thanks to the divergence theorem, we have

$$
0=\int_{D} \Delta v_{j} d x=\int_{\partial D_{0}} \nabla v_{j} \cdot \vec{n} d \sigma-\sum_{i=1}^{n} \int_{\partial h_{i}} \nabla v_{j} \cdot \vec{n} d \sigma
$$

Therefore,

$$
\sum_{i=1}^{n} A_{i, j}=\int_{\partial D_{0}} \nabla v_{j} \cdot \vec{n} d \sigma<0
$$

where the last inequality follows from the Hopf Lemma since $v_{j}$ attains its minimum value 0 on
$\partial D_{0}$, and $v_{j} \not \equiv 0$ on $\partial D$. A similar argument gives that $A_{i, j}>0$ for $i \neq j$ and $A_{j, j}<0$. Thus $A$ is invertible by Gershgorin circle theorem [47], leading to a unique solution of (Equation 2.1.13). Let us denote the solution by $x=\left(c_{1}, \ldots, c_{n}\right)^{t}$. Then the function $p$ defined by

$$
p:=u+\sum_{i=1}^{n} c_{i} v_{i}
$$

satisfies the desired properties (Equation 2.1.12).
Now we prove that $c_{i}>0$ for $i \geq 1$. Suppose that $c_{i *}:=\min _{i} c_{i} \leq 0$. Then by the minimum principle, $p$ attains its minimum on $\partial h_{i^{*}}$. Therefore,

$$
0 \leq \int_{\partial h_{i^{*}}} \nabla p \cdot \vec{n} d \sigma=-2\left|h_{i^{*}}\right|<0
$$

which is a contradiction.

Next we prove a parallel version of Talenti's theorem for the function $p$ constructed in Lemma 2.1.5. We will use this result throughout Section section 4.1-section 4.2.

Proposition 2.1.6. Let $D \subset \mathbb{R}^{2}$ be a bounded connected domain with $C^{1}$ boundary. Assume $D$ has $n$ holes with $n \geq 0$, and denote by $h_{1}, \cdots, h_{n} \subset \mathbb{R}^{2}$ the holes of $D$ (each $h_{i}$ is a bounded open set). Let $p: \bar{D} \rightarrow \mathbb{R}$ be the function constructed in Lemma 2.1.5. Then the following two estimates hold:

$$
\begin{equation*}
\sup _{\bar{D}} p \leq \frac{|D|}{2 \pi} \tag{2.1.14}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{D} p(x) d x \leq \frac{|D|^{2}}{4 \pi} \tag{2.1.15}
\end{equation*}
$$

Furthermore, for each of the two inequalities above, the equality is achieved if and only if $D$ is either a disk or an annulus.

Proof. The proof is divided into two parts: In step 1 we prove the two inequalities
(Equation 2.1.14) and (Equation 2.1.15), and in step 2 we show that equality can be achieved if and only if $D$ is a disk or an annulus.

Step 1. When $D$ is simply-connected, (Equation 2.1.14) and (Equation 2.1.15) directly follow from Talenti's theorem Proposition 2.1.1. Next we consider a non-simply-connected domain $D$, and prove that these inequalities also hold when $p: \bar{D} \rightarrow \mathbb{R}$ is defined as in Lemma 2.1.5.

For $k \in \mathbb{R}^{+}$, let us denote $D_{k}:=\{x \in D: p(x)>k\}, g(k):=\left|D_{k}\right|$ and $\tilde{D}_{k}:=$ $D_{k} \dot{\cup}\left(\dot{U}_{\left\{i: c_{i}>k\right\}} \overline{h_{i}}\right)$. Elliptic regularity theory gives that $p \in C^{\infty}(D)$, thus by Sard's theorem, $k$ is a regular value for almost every $k \in\left(0, \sup _{D} p\right)$, that is, $|\nabla p(x)|>0$ on $\{x \in D: p(x)=k\}$. Thus $\{x \in D: p(x)=k\}$ is a union of smooth simply closed curves and equal to $\partial \tilde{D}_{k}$ for almost every $k \in\left(0, \sup _{D} p\right)$.

Since $\partial D_{k}=\partial \tilde{D}_{k} \dot{\cup}\left(\dot{U}_{\left\{i: c_{i}>k\right\}} \partial h_{i}\right)$ for $k \notin\left\{c_{1}, \ldots, c_{n}\right\}$, we compute

$$
\begin{aligned}
g(k) & =-\frac{1}{2} \int_{D_{k}} \Delta p(x) d x=-\frac{1}{2} \int_{\partial D_{k}} \nabla p \cdot \vec{n} d \sigma \\
& =-\frac{1}{2} \int_{\partial \tilde{D}_{k}} \nabla p \cdot \vec{n} d \sigma+\frac{1}{2} \sum_{\left\{i: c_{i}>k\right\}} \int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma \\
& =-\frac{1}{2} \int_{\partial \tilde{D}_{k}} \nabla p \cdot \vec{n} d \sigma-\sum_{\left\{i: c_{i}>k\right\}}\left|h_{i}\right|,
\end{aligned}
$$

where the last identity is due to (Equation 2.1.12). Therefore, it follows that

$$
\begin{equation*}
g(k)+\sum_{\left\{i: c_{i}>k\right\}}\left|h_{i}\right|=-\frac{1}{2} \int_{\partial \tilde{D}_{k}} \nabla p \cdot \vec{n} d \sigma=\frac{1}{2} \int_{\partial \tilde{D}_{k}}|\nabla p| d \sigma, \tag{2.1.16}
\end{equation*}
$$

where the last equality follows from the fact that $\nabla p$ is perpendicular to the tangent vector on the level set.

On the other hand, the coarea formula yields that

$$
g(k)=\int_{\mathbb{R}} \int_{\partial \tilde{D}_{s}} 1_{D_{k}} \frac{1}{|\nabla p|} d \sigma d s=\int_{k}^{\infty} \int_{\partial \tilde{D}_{s}} \frac{1}{|\nabla p|} d \sigma d s .
$$

Therefore, it follows that for almost every $k \in\left(0, \sup _{D} p\right)$,

$$
\begin{equation*}
g^{\prime}(k)=-\int_{\partial \tilde{D}_{k}} \frac{1}{|\nabla p|} d \sigma \tag{2.1.17}
\end{equation*}
$$

Thus it follows from (Equation 2.1.16) and (Equation 2.1.17) that

$$
\begin{equation*}
g^{\prime}(k)\left(g(k)+\sum_{\left\{i: c_{i}>k\right\}}\left|h_{i}\right|\right)=-\frac{1}{2}\left(\int_{\partial \tilde{D}_{k}}|\nabla p| d \sigma\right)\left(\int_{\partial \tilde{D}_{k}} \frac{1}{|\nabla p|} d \sigma\right) \leq-\frac{1}{2} P\left(\tilde{D}_{k}\right)^{2} \tag{2.1.18}
\end{equation*}
$$

where $P(E)$ denotes the perimeter of a rectifiable curve $\partial E$. Note that the last inequality becomes equality if and only if $|\nabla p|$ is a constant on $\partial \tilde{D}_{k}$. Also, the isoperimetric inequality gives that

$$
\begin{equation*}
P\left(\tilde{D}_{k}\right)^{2} \geq 4 \pi\left|\tilde{D}_{k}\right| \tag{2.1.19}
\end{equation*}
$$

where equality holds if and only if $\tilde{D}_{k}$ is a disk. This yields that

$$
\begin{equation*}
g^{\prime}(k)\left(g(k)+\sum_{\left\{i: c_{i}>k\right\}}\left|h_{i}\right|\right) \leq-2 \pi\left|\tilde{D}_{k}\right|=-2 \pi\left(g(k)+\sum_{\left\{i: c_{i}>k\right\}}\left|h_{i}\right|\right) . \tag{2.1.20}
\end{equation*}
$$

Therefore, $g^{\prime}(k) \leq-2 \pi$ for almost every $k \in\left(0, \sup _{D} p\right)$. Combining it with the fact that $g(0)=$ $|D|$, we have

$$
g(k) \leq(g(0)-2 \pi k)_{+}=(|D|-2 \pi k)_{+} \text {for almost every } k \geq 0
$$

This proves that $\sup _{\bar{D}} p \leq \frac{|D|}{2 \pi}$. It follows that

$$
\int_{D} p(x) d x=\int_{D} \int_{0}^{\frac{|D|}{2 \pi}} 1_{\{k<p(x)\}} d k d x=\int_{0}^{\frac{|D|}{2 \pi}} g(k) d k \leq \int_{0}^{\frac{|D|}{2 \pi}}(|D|-2 \pi k)_{+} d x=\frac{|D|^{2}}{4 \pi}
$$

Step 2. Now we show that for the two inequalities (Equation 2.1.14) and (Equation 2.1.15),
the equality is achieved if and only if $D$ is either a disk or an annulus. First, if $D$ is either a disk or annulus centered at some $x_{0} \in \mathbb{R}^{2}$, then uniqueness of solution to Poisson's equation gives that $p$ is radially symmetric about $x_{0}$. Since we have $\Delta p=-2$ in $D$ and $p=0$ on the outer boundary of $D$, this gives an explicit formula $p(x)=-\frac{\left|x-x_{0}\right|^{2}}{2}+\frac{R^{2}}{2}$ for $x \in D$, where $R$ is the outer radius of $D$. For either a disk or an annulus, one can explicitly compute $\sup _{D} p$ and $\int_{D} p d x$ to check that equalities in (Equation 2.1.14) and (Equation 2.1.15) are achieved.

To prove the converse, assume that either (Equation 2.1.14) or (Equation 2.1.15) achieves equality, and we aim to show that $D$ is either a disk or an annulus. In order for either equality to be achieved, (Equation 2.1.20) needs to achieve equality at almost every $k \in\left(0, \sup _{D} p\right)$. In addition, $g(k)$ needs to be continuous in $k$ since $g(k)$ is decreasing. Since (Equation 2.1.20) follows from a combination of the Cauchy-Schwarz inequality in (Equation 2.1.18) and the isoperimetric inequality in (Equation 2.1.19), we need to have all the three conditions below in order for either (Equation 2.1.14) or (Equation 2.1.15) to achieve equality:
(1) $|\nabla p|$ is a constant on each level set $\partial \tilde{D}_{k}$ for almost every $k \in\left(0, \sup _{D} p\right)$;
(2) $\tilde{D}_{k}$ is a disk for almost every $k \in\left(0, \sup _{D} p\right)$.
(3) $g(k)=\left|D_{k}\right|$ is continuous in $k$. As a result, $\left|\tilde{D}_{k}\right|$ is continuous in $k$ at all $k \neq c_{i}$, with $c_{i}>0$ defined in (Equation 2.1.11).

Next we will show that if all these three conditions are satisfied, then $D$ must be an annulus or disk. First, note that by sending $k \searrow 0$ in condition (2), and combining it with the continuity of $\left|\tilde{D}_{k}\right|$ as $k \searrow 0$, it already gives that the outer boundary of $D$ must be a circle. Therefore if $D$ is simply-connected, it must be a disk.

If $D$ is non-simply-connected, using condition (2) and (3), we claim that $D$ can have only one hole, which must be a disk, and $p$ must achieve its maximum value in $\bar{D}$ on the boundary of the hole. To see this, let $h_{i}$ be any hole of $D$, and recall that $\left.p\right|_{\partial h_{i}}=c_{i}$. As we consider the set limit of
$\tilde{D}_{k}$ as $k$ approaches $c_{i}$ from below and above, by definition of $\tilde{D}_{k}$ we have

$$
\lim _{k \nearrow c_{i}} \tilde{D}_{k}=\lim _{k \searrow c_{i}} \tilde{D}_{k} \dot{\cup}\left(\dot{U}_{\left\{j: c_{j}=c_{i}\right\}} \overline{h_{j}}\right)
$$

By (2) and (3), the left hand side $\lim _{k \nearrow c_{i}} \tilde{D}_{k}$ is a disk, and the set $\lim _{k \backslash c_{i}} \tilde{D}_{k}$ on the right hand side is also a disk (if the limit is non-empty). But after taking union with the holes $\left\{h_{j}: c_{j}=c_{i}\right\}$ (each is a simply-connected set), the right hand side will be a disk if and only if $\lim _{k \backslash c_{i}} \tilde{D}_{k}$ is empty, $\dot{U}_{\left\{j: c_{j}=c_{i}\right\}} \overline{h_{j}}=\overline{h_{i}}$, and $h_{i}$ is a disk. This implies $c_{i}=\sup _{D} p$ and $c_{j}<c_{i}$ for all $j \neq i$. But since $h_{i}$ is chosen to be any hole of $D$, we know $D$ can have only one hole (call it $h$ ), which is a disk, and $\sup _{D} p=\left.p\right|_{\partial h}$. Finally, note that condition (1) gives that all the disks $\left\{\tilde{D}_{k}\right\}$ are concentric, and as a result we have $D$ is an annulus, finishing the proof.

Finally, we are ready to show that every connected stationary patch $D$ with $C^{1}$ boundary must be either a disk or an annulus.

Theorem 2.1.7. Let $D \subset \mathbb{R}^{2}$ be a bounded domain with $C^{1}$ boundary. Suppose that $\omega(x):=1_{D}(x)$ is a stationary patch solution to the $2 D$ Euler equation in the sense of (Equation 1.1.5). Then $D$ is either a disk or an annulus.

Proof. If $D$ has $n$ holes (where $n \geq 0$ ), denote them by $h_{1}, \ldots, h_{n}$. By (Equation 1.1.5), the function $f:=1_{D} * \mathcal{N}$ is constant on each of connected component of $\partial D$, and let us denote

$$
f(x)= \begin{cases}a_{i} & \text { on } \partial h_{i}  \tag{2.1.21}\\ a_{0} & \text { on } \partial D_{0}\end{cases}
$$

Let $p: \bar{D} \rightarrow \mathbb{R}$ be defined as in Lemma 2.1.5, and let $\varphi:=\frac{|x|^{2}}{2}+p$. Similar to the proof of Theorem 2.1.2, we calculate $\mathcal{I}:=\int_{D} \nabla \varphi \cdot \nabla f d x$ in two different ways. Note that $\nabla f=\nabla\left(f-a_{0}\right)$ in $D$. Applying the divergence theorem to $\mathcal{I}$ and using (Equation 2.1.21) and $\Delta \varphi=0$ in $D$, it
follows that

$$
\begin{equation*}
\mathcal{I}=\int_{\partial D}(\nabla \varphi \cdot \vec{n})\left(f-a_{0}\right) d \sigma-\int_{D} \Delta \varphi\left(f-a_{0}\right) d x=-\sum_{i=1}^{n}\left(a_{i}-a_{0}\right) \int_{\partial h_{i}} \nabla \varphi \cdot \vec{n} d \sigma . \tag{2.1.22}
\end{equation*}
$$

By definition of $\varphi$, and combining it with the property of $p$ in (Equation 2.1.12), we have

$$
\begin{align*}
\int_{\partial h_{i}} \nabla \varphi \cdot \vec{n} d \sigma & =\int_{\partial h_{i}} \nabla\left(\frac{|x|^{2}}{2}\right) \cdot \vec{n} d \sigma+\int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma  \tag{2.1.23}\\
& =\int_{h_{i}} 2 d x+\int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma=0 .
\end{align*}
$$

Plugging this into (Equation 2.1.22) gives $\mathcal{I}=0$. On the other hand, we also have

$$
\mathcal{I}=\int_{D} x \cdot \nabla f d x+\int_{D} \nabla p \cdot \nabla f d x=: E_{1}+E_{2}
$$

We compute

$$
\begin{equation*}
E_{1}=\int_{D} x \cdot\left(1_{D} * \nabla \mathcal{N}\right) d x=\int_{D} \int_{D} \frac{1}{2 \pi} \frac{x \cdot(x-y)}{|x-y|^{2}} d y d x=\frac{|D|^{2}}{4 \pi} \tag{2.1.24}
\end{equation*}
$$

where the last equality is obtained by exchanging $x$ with $y$ and taking the average with the original integral. For $E_{2}$, the divergence theorem yields that

$$
E_{2}=\int_{\partial D} p \nabla f \cdot \vec{n} d \sigma-\int_{D} p \Delta f d x=\int_{\partial D} p \nabla f \cdot \vec{n} d \sigma-\int_{D} p d x .
$$

Using the property of $p$ in (Equation 2.1.11) and the fact that $\Delta f=0$ in $h_{i}$, the divergence theorem yields

$$
\begin{equation*}
\int_{\partial D} p \nabla f \cdot \vec{n} d \sigma=-\sum_{i=1}^{n} \int_{\partial h_{i}} p \nabla f \cdot \vec{n} d \sigma=-\sum_{i=1}^{n} c_{i} \int_{h_{i}} \Delta f d x=0 \tag{2.1.25}
\end{equation*}
$$

As a result, we have $E_{2}=-\int_{D} p d x$. If $D$ is neither a disk nor an annulus, Proposition 2.1.6 gives

$$
\mathcal{I}=E_{1}+E_{2}=\frac{|D|^{2}}{4 \pi}-\int_{D} p d x>0
$$

contradicting $\mathcal{I}=0$.

In the next corollary, we generalize the above result to a nonnegative stationary patch with multiple (disjoint) patches.

Corollary 2.1.8. Let $\omega(x):=\sum_{i=1}^{n} \alpha_{i} 1_{D_{i}}$, where $\alpha_{i}>0$, each $D_{i}$ is a bounded connected domain with $C^{1}$ boundary, and $D_{i} \cap D_{j}=\emptyset$ if $i \neq j$. Assume that $\omega$ is a stationary patch solution, that is, the function $f(x):=\omega * \mathcal{N}$ satisfies $\nabla^{\perp} f \cdot \vec{n}=0$ on $\partial D_{i}$ for all $i=1, \ldots, n$. Then $\omega$ is radially symmetric up to a translation.

Proof. Following similar notations as the beginning of Section subsection 2.1.2, we denote the outer boundary of $D_{i}$ by $\partial D_{i 0}$, and the holes of each $D_{i}$ (if any) by $h_{i k}$ for $k=1, \ldots, N_{i}$. Let $p_{i}: \overline{D_{i}} \rightarrow \mathbb{R}$ be defined as in Lemma 2.1.5, that is, $p_{i}$ satisfies

$$
\begin{cases}\Delta p_{i}=-2 & \text { in } D_{i} \\ p_{i}=c_{i k} & \text { on } \partial h_{i k} \\ p_{i}=0 & \text { on } \partial D_{i 0}\end{cases}
$$

where $c_{i k}$ is chosen such that $\int_{\partial h_{i k}} \nabla p_{i} \cdot \vec{n} d \sigma=-2\left|h_{i k}\right|$. We then define $\varphi: \cup_{i=1}^{n} \bar{D}_{i} \rightarrow \mathbb{R}$, such that in each $\bar{D}_{i}$ we have $\varphi=\varphi_{i}:=\frac{|x|^{2}}{2}+p_{i}$.

Similar to Theorem 2.1.7, we compute

$$
\mathcal{I}:=\int_{\mathbb{R}^{2}} \omega \nabla \varphi \cdot \nabla f d x=\sum_{i=1}^{n} \int_{D_{i}} \alpha_{i} \nabla \varphi_{i} \cdot \nabla f d x
$$

in two different ways. On the one hand, since $f=\omega * \mathcal{N}$ is a constant on each connected component
of $\partial D_{i}$, the same computation of Theorem 2.1.7 yields that $\int_{D_{i}} \nabla \varphi_{i} \cdot \nabla f d x=0$, therefore $\mathcal{I}=0$. On the other hand, since $\nabla \varphi=x+\nabla p_{i}$ in each $D_{i}$, we break $\mathcal{I}$ into

$$
\mathcal{I}=\sum_{i, j=1}^{n} \alpha_{i} \alpha_{j} \int_{D_{i}} x \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x+\sum_{i, j=1}^{n} \alpha_{i} \alpha_{j} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x=: \mathcal{I}_{1}+\mathcal{I}_{2}
$$

For $\mathcal{I}_{1}$, we compute

$$
\begin{align*}
\mathcal{I}_{1} & =\sum_{i, j=1}^{n} \frac{\alpha_{i} \alpha_{j}}{2}\left(\int_{D_{i}} x \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x+\int_{D_{j}} x \cdot \nabla\left(1_{D_{i}} * \mathcal{N}\right) d x\right) \\
& =\sum_{i, j=1}^{n} \frac{\alpha_{i} \alpha_{j}}{2}\left(\int_{D_{i}} \int_{D_{j}} \frac{x \cdot(x-y)}{2 \pi|x-y|^{2}} d y d x+\int_{D_{j}} \int_{D_{i}} \frac{x \cdot(x-y)}{2 \pi|x-y|^{2}} d y d x\right) \\
& =\sum_{i, j=1}^{n} \frac{\alpha_{i} \alpha_{j}}{4 \pi}\left|D_{i}\right|\left|D_{j}\right| \tag{2.1.26}
\end{align*}
$$

where we exchanged $i$ with $j$ to get the first equality. For $\mathcal{I}_{2}$, we have

$$
\mathcal{I}_{2}=\sum_{i=1}^{n} \alpha_{i}^{2} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{i}} * \mathcal{N}\right) d x+\sum_{i \neq j} \alpha_{i} \alpha_{j} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x=: I_{21}+I_{22}
$$

By the same computation for $E_{2}$ in the proof of Theorem 2.1.7, we have

$$
\begin{equation*}
I_{21}=-\sum_{i=1}^{n} \alpha_{i}^{2} \int_{D_{i}} p_{i} d x \tag{2.1.27}
\end{equation*}
$$

For $i \neq j$, we denote $j \prec i$ if $D_{j}$ is contained in a hole of $D_{i}$. (And if $D_{j}$ is not contained in any hole of $D_{i}$, we say $j \nprec i$.) Using this notation, the divergence theorem directly yields that

$$
\begin{equation*}
\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma=-\sum_{k=1}^{N_{i}} \int_{\partial h_{i k}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma=0 \quad \text { if } j \nprec i . \tag{2.1.28}
\end{equation*}
$$

And if $j \prec i$, then the divergence theorem and (Equation 2.1.14) in Proposition 2.1.6 yield

$$
\begin{equation*}
\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma \geq-\sup _{\partial D_{i}} p_{i}\left|D_{j}\right| \geq-\frac{1}{2 \pi}\left|D_{i}\right|\left|D_{j}\right| \quad \text { if } j \prec i . \tag{2.1.29}
\end{equation*}
$$

Hence it follows that

$$
\begin{equation*}
I_{22} \geq-\sum_{i \neq j} \mathbb{1}_{j \prec i} \frac{\alpha_{i} \alpha_{j}}{2 \pi}\left|D_{i}\right|\left|D_{j}\right|=-\sum_{i \neq j}\left(\mathbb{1}_{j \prec i}+\mathbb{1}_{i \prec j}\right) \frac{\alpha_{i} \alpha_{j}}{4 \pi}\left|D_{i}\right|\left|D_{j}\right|, \tag{2.1.30}
\end{equation*}
$$

where the last step is obtained by exchanging $i, j$ and taking average with the original sum. Note that we have $\mathbb{1}_{j \prec i}+\mathbb{1}_{i \prec j} \leq 1$ for any $i \neq j$. From (Equation 2.1.26), (Equation 2.1.27) and (Equation 2.1.30), we obtain

$$
\begin{equation*}
\mathcal{I} \geq \sum_{i=1}^{n} \alpha_{i}^{2}\left(\frac{\left|D_{i}\right|^{2}}{4 \pi}-\int_{D_{i}} p_{i} d x\right)+\sum_{\substack{\nprec i \\ \text { and } i \nless j \\ i \neq j}} \alpha_{i} \alpha_{j} \frac{\left|D_{i}\right|\left|D_{j}\right|}{4 \pi} . \tag{2.1.31}
\end{equation*}
$$

Since we already know that $\mathcal{I}=0$ and all the summands in (Equation 2.1.31) are nonnegative, it follows that

$$
\frac{\left|D_{i}\right|^{2}}{4 \pi}=\int_{D_{i}} p_{i} d x \text { for all } i=1, \ldots, n \text { and }\{(i, j): i \neq j, i \nprec j \text { and } j \nprec i\}=\emptyset .
$$

Therefore every $D_{i}$ is either a disk or an annulus by Proposition 2.1.6 and they are nested. By relabeling the indices, we can assume that $i \prec i+1$ for $i=1, \ldots, n-1$.

Next we prove that all $D_{i}$ 's are concentric by induction. For $k \geq 1$, suppose $D_{1}, \ldots, D_{k}$ are known to be concentric about some $o \in \mathbb{R}^{2}$. To show $D_{k+1}$ is also centered at o, we break $f$ into

$$
f=\sum_{i=1}^{k}\left(\alpha_{i} 1_{D_{i}}\right) * \mathcal{N}+\sum_{i=k+1}^{n}\left(\alpha_{i} 1_{D_{i}}\right) * \mathcal{N} .
$$

In the first sum, each $D_{i}$ is centered at $o$ for $i \leq k$, thus Lemma 2.1.9(a) (which we prove right after
this theorem) yields that $\sum_{i=1}^{k}\left(\alpha_{i} 1_{D_{i}}\right) * \mathcal{N}=\frac{C}{2 \pi} \ln |x-o|$ on $\partial_{\text {in }} D_{k+1}$, where $C=\sum_{i=1}^{k} \alpha_{i}\left|D_{i}\right|>0$. In the second sum, for each $i \geq k+1$, since each $D_{i}$ is an annulus with $\partial_{\text {in }} D_{k+1}$ in its hole, Lemma 2.1.9(b) gives that $1_{D_{i}} * \mathcal{N} \equiv$ const on $\partial_{\text {in }} D_{k+1}$ for all $i \geq k+1$. Thus overall we have $f=\frac{C}{2 \pi} \ln |x-o|+C_{2}$ on $\partial_{\text {in }} D_{k+1}$ for $C>0$. Combining it with the assumption that $f$ is a constant on $\partial_{\text {in }} D_{k+1}$, we know $D_{k+1}$ must also be centered at $o$, finishing the induction step.

Now we state and prove the lemma used in the proof of Corollary 2.1.8, which follows from standard properties of the Newtonian potential.

Lemma 2.1.9. Assume $g \in L^{\infty}\left(\mathbb{R}^{2}\right)$ is radially symmetric about some $o \in \mathbb{R}^{2}$, and is compactly supported in $B(o, R)$. Then $\eta:=g * \mathcal{N}$ satisfies the following:
(a) $\eta(x)=\frac{\int_{\mathbb{R}^{2}} g d x}{2 \pi} \ln |x-o|$ for all $x \in B(0, R)^{c}$.
(b) If in addition we have $g \equiv 0$ in $B(o, r)$ for some $r \in(0, R)$, then $\eta=$ const in $B(o, r)$.

Proof. To show (a), we take any $x \in B(o, R)^{c}$ and consider the circle $\Gamma \ni x$ centered at $o$. By radial symmetry of $\eta$ about $o$ and the divergence theorem, we have

$$
\nabla \eta \cdot \frac{x}{|x|}=\frac{1}{|\Gamma|} \int_{\Gamma} \nabla \eta \cdot \vec{n} d \sigma=\frac{1}{|\Gamma|} \int_{\operatorname{int}(\Gamma)} \Delta \eta d x=\frac{\int_{\mathbb{R}^{2}} g(x) d x}{2 \pi|x-o|},
$$

which implies $\eta(x)=\frac{\int g d x}{2 \pi} \ln |x-o|+C$. To show that $C=0$, for $|x|$ sufficiently large we have

$$
|C|=\left|\int_{B(o, R)} g(x)(\mathcal{N}(x-y)-\mathcal{N}(x-o)) d y\right| \leq\|g\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \sup _{y \in B(o, R)}|\mathcal{N}(x-o)-\mathcal{N}(x-y)|
$$

and by sending $|x| \rightarrow \infty$ we have $C=0$, which gives (a). To show (b), it suffices to prove that $\nabla \eta \equiv 0$ in $B(o, r)$. Take any $x \in B(o, r)$, and consider the circle $\Gamma_{2} \ni x$ centered at $o$. Again, symmetry and the divergence theorem yield that

$$
|\nabla \eta(x)|=\frac{1}{\left|\Gamma_{2}\right|} \int_{\Gamma_{2}} \nabla \eta \cdot \vec{n} d \sigma=\frac{1}{\left|\Gamma_{2}\right|} \int_{\operatorname{int}\left(\Gamma_{2}\right)} \Delta \eta d x=\frac{\int_{\operatorname{int}(\Gamma)} g(x) d x}{|\Gamma|}=0,
$$

finishing the proof of (b).

### 2.1.3 Radial symmetry of non-simply-connected rotating patches with $\Omega<0$

In this subsection, we show that a nonnegative uniformly rotating patch solution (with multiple disjoint patches) must be radially symmetric if the angular velocity $\Omega<0$.

Theorem 2.1.10. For $i=1, \ldots, n$, let $D_{i}$ be a connected domain with $C^{1}$ boundary, and assume $D_{i} \cap D_{j}=\emptyset$ for $i \neq j$. If $\omega=\sum_{i=1}^{n} \alpha_{i} 1_{D_{i}}$ is a nonnegative rotating patch solution with $\alpha_{i}>0$ and angular velocity $\Omega<0$, then $\omega$ must be radially symmetric.

Proof. In this proof, let

$$
f_{\Omega}(x):=\omega * \mathcal{N}-\frac{\Omega}{2}|x|^{2} .
$$

In each $D_{i}$, let us define $p_{i}$ as in Lemma 2.1.5. Let $\varphi_{i}:=\frac{|x|^{2}}{2}+p_{i}$ in each $D_{i}$. As in Theorem 2.1.10, we compute $\mathcal{I}:=\sum_{i=1}^{n} \alpha_{i} \int_{D_{i}} \nabla \varphi_{i} \cdot \nabla f_{\Omega} d x$ in two different ways. Since $f_{\Omega}$ is a constant on each connected component of $\partial D_{i}$ and $\nabla \varphi_{i}$ is divergence free in $D_{i}$, we still have $\mathcal{I}=0$ as in the proof of Theorem 2.1.7.

On the other hand, we have

$$
\begin{aligned}
\mathcal{I} & =\sum_{i=1}^{n} \alpha_{i} \int_{D_{i}}\left(x+\nabla p_{i}\right) \cdot \nabla(\omega * \mathcal{N}) d x+\underbrace{(-\Omega)}_{\geq 0} \sum_{i=1}^{n} \alpha_{i} \int_{D_{i}}\left(x+\nabla p_{i}\right) \cdot x d x \\
& =: \mathcal{I}_{1}+(-\Omega) \mathcal{I}_{2} .
\end{aligned}
$$

As in the proof of Corollary 2.1.8, we have

$$
\begin{equation*}
\mathcal{I}_{1}=\sum_{i=1}^{n} \alpha_{i}^{2}\left(\frac{\left|D_{i}\right|^{2}}{4 \pi}-\int_{D_{i}} p_{i} d x\right)+\sum_{\substack{j \nless i \operatorname{and} i \nless j \\ i \neq j}} \alpha_{i} \alpha_{j} \frac{\left|D_{i}\right|\left|D_{j}\right|}{4 \pi} \geq 0 . \tag{2.1.32}
\end{equation*}
$$

Note that $\mathcal{I}_{1}=0$ as long as all $D_{i}$ 's are nested annuli/disk, even if they are not concentric. For $\mathcal{I}_{2}$, using Cauchy-Schwarz inequality in the second step, and Lemma 2.1.11 in the third step (which
we will prove right after this theorem), we have

$$
\begin{align*}
\mathcal{I}_{2} & =\sum_{i=1}^{n} \alpha_{i}\left(\int_{D_{i}}|x|^{2} d x+\int_{D_{i}} \nabla p_{i} \cdot x d x\right) \\
& \geq \sum_{i=1}^{n} \alpha_{i}\left(\int_{D_{i}}|x|^{2} d x-\left(\int_{D_{i}}|\nabla p|^{2} d x\right)^{1 / 2}\left(\int_{D_{i}}|x|^{2} d x\right)^{1 / 2}\right) \geq 0 . \tag{2.1.33}
\end{align*}
$$

Combining (Equation 2.1.32) and (Equation 2.1.33) gives us $\mathcal{I} \geq 0$. If there is any $D_{i}$ that is not a disk or annulus centered at the origin, Lemma 2.1.11 would give a strict inequality in the last step of (Equation 2.1.33), which leads to $\mathcal{I}>0$ and thus contradicts with $\mathcal{I}=0$.

Now we state and prove the lemma that is used in the proof of Theorem 2.1.10.

Lemma 2.1.11. Let $D$ be a connected domain with $C^{1}$ boundary, and let p be as in Lemma 2.1.5. Then we have

$$
\begin{equation*}
-\int_{D} \nabla p \cdot x d x=\int_{D}|\nabla p|^{2} d x \leq \int_{D}|x|^{2} d x \tag{2.1.34}
\end{equation*}
$$

Furthermore, in the inequality, " $="$ is achieved if and only if $D$ is a disk or annulus centered at the origin.

Proof. We compute

$$
\begin{aligned}
\int_{D}|\nabla p|^{2} d x & =\int_{\partial D} p \nabla p \cdot \vec{n} d \sigma+\int_{D} 2 p d x \\
& =-\int_{\partial D} p x \cdot \vec{n} d \sigma+\int_{D} 2 p d x
\end{aligned}
$$

where in the last equality we use that $p$ is constant along each $\partial h_{i}$, as well as the following identity due to (Equation 2.1.12) and the divergence theorem (here $\vec{n}$ is the outer normal of $h_{i}$ ):

$$
\int_{\partial h_{i}} \nabla p \cdot \vec{n} d \sigma=-2\left|h_{i}\right|=-\int_{h_{i}} \Delta \frac{|x|^{2}}{2} d x=-\int_{\partial h_{i}} x \cdot \vec{n} d \sigma .
$$

On the other hand, the divergence theorem yields

$$
-\int_{D} \nabla p \cdot x d x=-\int_{\partial D} p x \cdot \vec{n} d \sigma+\int_{D} 2 p d x .
$$

Therefore using Young's inequality $-\nabla p \cdot x \leq \frac{1}{2}|\nabla p|^{2}+\frac{1}{2}|x|^{2}$ (where the equality is achieved if and only if $-\nabla p=x$ ), we have

$$
\int_{D}|\nabla p|^{2} d x=-\int_{D} \nabla p \cdot x d x \leq \frac{1}{2} \int_{D}|\nabla p|^{2} d x+\frac{1}{2} \int_{D}|x|^{2} d x
$$

which proves (Equation 2.1.34). Here the equality is achieved if and only if $-\nabla p=x$ in $D$, which is equivalent with $p+\frac{|x|^{2}}{2}$ being a constant in $D$, and it can be extended to $\bar{D}$ due to continuity of $p$. By our construction of $p$ in Lemma 2.1.5, $p$ is already a constant on each connected component of $\partial D$, implying $\frac{|x|^{2}}{2}$ is constant on each piece of $\partial D$, hence $\partial D$ must be a family of circles centered at the origin. By the assumption that $D$ is connected, it must be either a disk or annulus centered at the origin.

### 2.1.4 Radial symmetry of non-simply-connected rotating patches with $\Omega \geq \frac{1}{2}$

In this final subsection for patches, we consider a bounded domain $D$ with $C^{1}$ boundary. $D$ can have multiple connected components, and each connected component can be non-simplyconnected. If $1_{D}$ is a rotating patch solution to the Euler equation with angular velocity $\Omega \geq \frac{1}{2}$, we will show $D$ must be radially symmetric and centered at the origin.

To do this, one might be tempted to proceed as in Theorem 2.1.2 and replace $p: D \rightarrow \mathbb{R}$ by the function defined in Lemma 2.1.5. Here the first way of computing $\mathcal{I}=\int_{D}(x+\nabla p) \cdot \nabla f_{\Omega} d x$ still yields $\mathcal{I}=0$, but the second way gives some undesired terms caused by the holes $h_{i}$ :

$$
\mathcal{I}=\frac{1}{4 \pi}|D|^{2}-\Omega \int_{D}|x|^{2} d x+(2 \Omega-1) \int_{D} p d x+\left.2 \Omega \sum_{i=1}^{n} p\right|_{\partial h_{i}}\left|h_{i}\right|
$$

Due to the last term on the right hand side, we are unable to show $\mathcal{I} \leq 0$ when $\Omega \geq \frac{1}{2}$ as we did before in Theorem 2.1.2. For this reason, we take a different approach in the next theorem. Instead of defining $p$ as a function in $D$ and $\mathcal{I}$ as an integral in $D$, we want to define them in $D^{c}$. But since $D^{c}$ is unbounded, we define $p^{R}$ and $\mathcal{I}_{R}$ in a truncated set $B(0, R) \backslash D^{c}$, and then use two different ways to compute $\mathcal{I}_{R}$. By sending $R \rightarrow \infty$, we will show that the two ways give a contradiction unless $D$ is radially symmetric.

Theorem 2.1.12. For a bounded domain $D$ with $C^{1}$ boundary, assume that $1_{D}$ is a rotating patch solution to the Euler equation with angular velocity $\Omega \geq \frac{1}{2}$. Then $D$ is radially symmetric and centered at the origin.

Proof. Since $D$ is bounded, let us choose $R_{0}>0$ such that $B_{R_{0}} \supset D$. For any $R>R_{0}$, consider the domain $B_{R} \backslash \bar{D}$, which may have multiple connected components. We call the component touching $\partial B_{R}$ as $D_{0, R}$, and name the other connected components by $U_{1}, \ldots, U_{n}$. Throughout this proof we assume that $n \geq 1$ : if not, then each connected component of $D$ is simply connected, which has been already treated in Theorem 2.1.2 and Remark 2.1.3. We also define $V:=B_{R} \backslash D_{0, R}$, which is the union of $D$ and all its holes. Note that $V$ may have multiple connected components, but each must be simply-connected. See Figure Figure 2.1 for an illustration of $D_{0, R},\left\{U_{i}\right\}_{i=1}^{n}$ and $V$.


Figure 2.1: For a set $D \subset B\left(0, R_{0}\right)$ (the whole yellow region on the left), the middle figure illustrates the definition of $D_{0, R}$ (the blue region), $\left\{U_{i}\right\}$ (the gray regions), and right right figure illustrate $V=B_{R} \backslash D_{0, R}$ (the green region).

To prove the theorem, the key idea is to define $p^{R}$ and $\mathcal{I}_{R}$ in $B_{R} \backslash D$, instead of in $D$. Let $p_{0, R}$ and $p_{i}$ be defined as in Lemma 2.1.5 in $D_{0, R}$ and $U_{i}$ respectively, then set $\varphi_{0, R}:=p_{0, R}+\frac{|x|^{2}}{2}$ in $D_{0, R}$, and $\varphi_{i}:=p_{i}+\frac{|x|^{2}}{2}$ in $U_{i}$ for $i=1, \ldots, n$. Finally, define $p^{R}$ and $\varphi^{R}: \mathbb{R}^{2} \rightarrow \mathbb{R}$ as

$$
p^{R}:=p_{0, R} 1_{D_{0, R}}+\sum_{i=1}^{n} p_{i} 1_{U_{i}}, \quad \varphi^{R}:=\varphi_{0, R} 1_{D_{0, R}}+\sum_{i=1}^{n} \varphi_{i} 1_{U_{i}} .
$$

Since $1_{D}$ rotates with angular velocity $\Omega \geq \frac{1}{2}$, we know $f_{\Omega}:=1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}$ is constant on each connected component of $\partial D$. Next we will compute

$$
\begin{equation*}
\mathcal{I}_{R}:=\int_{B_{R} \backslash \bar{D}} \nabla f_{\Omega} \cdot \nabla \varphi^{R} d x \tag{2.1.35}
\end{equation*}
$$

in two different ways. If some connected component of $\partial D$ is not a circle, we will derive a contradiction by sending $R \rightarrow \infty$. We point out that as we increase $R$, the domain $D_{0, R}$ will change, but the domains $\left\{U_{i}\right\}_{i=1}^{n}$ and $V$ all remain unchanged.

On the one hand, we break $\mathcal{I}_{R}$ into

$$
\mathcal{I}_{R}=\int_{D_{0, R}} \nabla f_{\Omega} \cdot \nabla \varphi_{0, R} d x+\sum_{i=1}^{n} \int_{U_{i}} \nabla f_{\Omega} \cdot \nabla \varphi_{i} d x=: \mathcal{I}_{R}^{1}+\mathcal{I}_{R}^{2}
$$

Since $f_{\Omega}$ is constant on each connected component of $\partial U_{i}$, the same computation as (Equation 2.1.22)-(Equation 2.1.23) gives $\mathcal{I}_{R}^{2}=0$. For $\mathcal{I}_{R}^{1}$, note that although $f_{\Omega}$ is a constant along the boundary of each hole of $D_{0, R}$, it is not a constant along $\partial_{\text {out }} D_{0, R}=\partial B_{R}$. Thus similar computations as (Equation 2.1.22)-(Equation 2.1.23) now give

$$
\begin{align*}
\mathcal{I}_{R}^{1} & =\int_{\partial B_{R}}\left(1_{D} * \mathcal{N}-\frac{\Omega}{2} R^{2}\right) \nabla \varphi_{0, R} \cdot \vec{n} d \sigma  \tag{2.1.36}\\
& =\int_{\partial B_{R}}\left(\left(1_{D} * \mathcal{N}\right)(x)-|D| \mathcal{N}(x)\right) \nabla \varphi_{0, R} \cdot \vec{n} d \sigma(x)
\end{align*}
$$

where in the second equality we used $\int_{\partial B_{R}} \nabla \varphi_{0, R} \cdot \vec{n} d \sigma=0$ and the fact that $\mathcal{N}(x)$ is constant on
$\partial B_{R}$. For any $x \in \partial B_{R}$, since $D \subset B_{R_{0}}$ and $R>R_{0}$, we can control $\left(1_{D} * \mathcal{N}\right)(x)-|D| \mathcal{N}(x)$ as

$$
\begin{equation*}
\left|\left(1_{D} * \mathcal{N}\right)(x)-|D| \mathcal{N}(x)\right| \leq \frac{1}{2 \pi} \int_{D}|\log | x-y|-\log | x| | d y \leq \frac{|D|}{2 \pi}\left|\log \left(1-\frac{R_{0}}{R}\right)\right| \text { on } \partial B_{R} \tag{2.1.37}
\end{equation*}
$$

We introduce the following lemma to control $\left|\nabla \varphi_{0, R} \cdot \vec{n}\right|$ on $\partial B_{R}$, whose proof is postponed to the end of this subsection.

Lemma 2.1.13. Let $D \subset B_{R_{0}}$ be a domain with $C^{1}$ boundary. For any $R>R_{0}$, let $D_{0, R}, V, p_{0, R}$ and $\varphi_{0, R}$ be defined as in the proof of Theorem 2.1.12. Then we have

$$
\begin{equation*}
\left|\nabla \varphi_{0, R} \cdot \vec{n}\right| \leq \frac{N R_{0}^{2}}{2 R \log \left(R / R_{0}\right)} \quad \text { on } \partial B_{R} \tag{2.1.38}
\end{equation*}
$$

where $N>0$ is the number of connected components of $V$ (and is independent of $R$ ).
Once we have this lemma, plugging (Equation 2.1.38) and (Equation 2.1.37) into (Equation 2.1.36) yields

$$
\left|\mathcal{I}_{R}^{1}\right| \leq \frac{N|D| R_{0}^{2}}{2}\left|\log \left(1-\frac{R_{0}}{R}\right)\right|\left(\log \left(R / R_{0}\right)\right)^{-1} \leq|D| \frac{C\left(D, R_{0}\right)}{R \log R} \rightarrow 0 \quad \text { as } R \rightarrow \infty
$$

Combining this with $\mathcal{I}_{R}^{2}=0$ gives

$$
\begin{equation*}
\lim _{R \rightarrow \infty} \mathcal{I}_{R}=0 . \tag{2.1.39}
\end{equation*}
$$

Next we compute $\mathcal{I}_{R}$ in another way. Note that $1_{B_{R}} * \mathcal{N}-\frac{|x|^{2}}{4}$ is a radial harmonic function in $B_{R}$, thus is equal to some constant $C_{R}$ in $B_{R}$. Using this fact, we can rewrite $f_{\Omega}$ as

$$
f_{\Omega}=1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}=\left(1_{D}-1_{B_{R}}\right) * \mathcal{N}-\left(\frac{\Omega}{2}-\frac{1}{4}\right)|x|^{2}+C_{R}
$$

As a result, $\mathcal{I}_{R}$ can be rewritten as

$$
\begin{align*}
\mathcal{I}_{R} & =-\int_{B_{R} \backslash \bar{D}} \nabla\left(1_{B_{R} \backslash \bar{D}} * \mathcal{N}\right) \cdot \nabla \varphi^{R} d x-\frac{(2 \Omega-1)}{2} \int_{B_{R} \backslash \bar{D}} x \cdot \nabla \varphi^{R} d x  \tag{2.1.40}\\
& =:-\mathcal{J}_{R}^{1}-\frac{(2 \Omega-1)}{2} \mathcal{J}_{R}^{2}
\end{align*}
$$

Next we will show $\mathcal{J}_{R}^{1}, \mathcal{J}_{R}^{2} \geq 0$, leading to $\mathcal{I}_{R} \leq 0$. Let us start with $\mathcal{J}_{R}^{2}$. Applying Lemma 2.1.11 to each of $D_{0, R}$ and $\left\{U_{i}\right\}_{i=1}^{n}$ immediately gives

$$
\begin{equation*}
\mathcal{J}_{R}^{2} \geq \int_{D_{0, R}}|x|^{2}+\nabla p_{0, R} \cdot x d x+\sum_{i=1}^{n} \int_{U_{i}}|x|^{2}+\nabla p_{i} \cdot x d x=: T_{0, R}+\sum_{i=1}^{n} T_{i} \geq 0 \tag{2.1.41}
\end{equation*}
$$

Note that the $T_{i}$ 's are independent of $R$ for $i=1, \ldots, n$, and we know $T_{i} \geq 0$ with equality achieved if and only if $U_{i}$ is an annulus or a disk centered at the origin. This will be used later to show all $\left\{U_{i}\right\}_{i=1}^{n}$ are centered at the origin in the $\Omega>\frac{1}{2}$ case. (When $\Omega=\frac{1}{2}$, the coefficient of $\mathcal{J}_{R}^{2}$ becomes 0 in (Equation 2.1.40), thus a different argument is needed in this case.)

We now move on to $\mathcal{J}_{R}^{1}$. We first break it as

$$
\mathcal{J}_{R}^{1}=\int_{B_{R} \backslash \bar{D}} \nabla\left(1_{B_{R} \backslash \bar{D}} * \mathcal{N}\right) \cdot x d x+\int_{B_{R} \backslash \bar{D}} \nabla\left(1_{B_{R} \backslash \bar{D}} * \mathcal{N}\right) \cdot \nabla p^{R} d x=: J_{11}+J_{12}
$$

An identical computation as (Equation 2.1.24) gives $J_{11}=\frac{1}{4 \pi}\left(\left|D_{0, R}\right|+\sum_{i=1}^{n}\left|U_{i}\right|\right)^{2}$. For $J_{12}$, the same computation as (Equation 2.1.27)-(Equation 2.1.29) gives the following (where we used that each $U_{i}$ lies in a hole of $D_{0, R}$ for $\left.i=1, \ldots, n\right)$ :

$$
J_{12} \geq-\int_{D_{0, R}} p_{0, R} d x-\sum_{i=1}^{n} \int_{U_{i}} p_{i} d x-\sum_{1 \leq i \leq n}\left|U_{i}\right| \sup _{\substack{D_{0, R}}} p_{0, R}-\sum_{\substack{1 \leq i, j \leq n \\ j<i}} \frac{\left|U_{i}\right|\left|U_{j}\right|}{2 \pi}
$$

Adding up the estimates for $J_{11}$ and $J_{12}$, we get

$$
\begin{align*}
\mathcal{J}_{R}^{1} \geq & \left(\frac{1}{4 \pi}\left|D_{0, R}\right|^{2}-\int_{D_{0, R}} p_{0, R} d x\right)+\left(\sum_{1 \leq i \leq n}\left|U_{i}\right|\right)\left(\frac{1}{2 \pi}\left|D_{0, R}\right|-\sup _{D_{0, R}} p_{0, R}\right) \\
& +\sum_{i=1}^{n}\left(\frac{1}{4 \pi}\left|U_{i}\right|^{2}-\int_{U_{i}} p_{i} d x\right)+\sum_{\substack{j \nless i \text { and } i \nless j \\
i \neq j}} \frac{1}{4 \pi}\left|U_{i}\right|\left|U_{j}\right| . \tag{2.1.42}
\end{align*}
$$

By Proposition 2.1.6, all terms on the right hand side are nonnegative. But note that only the two terms in the second line are independent of $R$. Plugging (Equation 2.1.42) and (Equation 2.1.41) into (Equation 2.1.40) gives the following (where we only keep the terms independent of $R$ on the right hand side):

$$
\liminf _{R \rightarrow \infty}\left(-\mathcal{I}_{R}\right) \geq \sum_{i=1}^{n}\left(\frac{1}{4 \pi}\left|U_{i}\right|^{2}-\int_{U_{i}} p_{i} d x\right)+\sum_{\substack{j \nprec i \\ i \not a n d}}^{i \neq j}<1 ~ \frac{1}{4 \pi}\left|U_{i}\right|\left|U_{j}\right|+\frac{2 \Omega-1}{2} \sum_{i=1}^{n} T_{i} \geq 0 .
$$

Combining this with the previous limit (Equation 2.1.39), we know $U_{i}$ must be an annulus or a disk for $i=1, \ldots, n$, and they must be nested in each other. In addition, if $\Omega>\frac{1}{2}$, we have $T_{i}=0$ for $i=1, \ldots, n$, implying that each $U_{i}$ is centered at the origin.

The radial symmetry of $D_{0, R}$ is more difficult to obtain. Although the first two terms on the right hand side of (Equation 2.1.42) are both strictly positive if $D_{0, R}$ is not an annulus, we need some uniform-in- $R$ lower bound to get a contradiction in the $R \rightarrow \infty$ limit. Between these two terms, it turns out the second term is easier to control. This is done in the next lemma, whose proof we postpone to the end of this subsection.

Lemma 2.1.14. Let $D \subset B_{R_{0}}$ be a domain with $C^{1}$ boundary. For any $R>R_{0}$, let $D_{0, R}, V$ and $p_{0, R}$ be given as in the proof of Theorem 2.1.12. If $V$ is not a single disk, there exists some constant
$C(V)>0$ only depending on $V$, such that

$$
\liminf _{R \rightarrow \infty}\left(\frac{1}{2 \pi}\left|D_{0, R}\right|-\sup _{D_{0, R}} p_{0, R}\right) \geq C(V)>0
$$

If $V$ is not a disk, Lemma 2.1 .14 gives $\liminf _{R \rightarrow \infty} \mathcal{J}_{R}^{1}>\left(\sum_{1 \leq i \leq n}\left|U_{i}\right|\right) C(V)>0$. (Recall that in the beginning of this proof we assume $\sum_{1 \leq i \leq n}\left|U_{i}\right|>0$, and it is independent of $R$.) This implies $\lim \inf _{R \rightarrow \infty}\left(-\mathcal{I}_{R}\right) \geq C(V)>0$, contradicting (Equation 2.1.39).

So far we have shown that $\partial D$ is a union of nested circles, and it remains to show that they are all centered at 0 . For the $\Omega>\frac{1}{2}$ case, we already showed all $\left\{U_{i}\right\}_{i=1}^{n}$ are centered at 0 , so it suffices to show the outmost circle $\partial V$ (denote by $B(\tilde{o}, \tilde{r})$ ) is also centered at 0 . By definition of $\left\{U_{i}\right\}_{i=1}^{n}$, we have $D=B(\tilde{o}, \tilde{r}) \backslash\left(\dot{\cup}_{i=1}^{n} U_{i}\right)$. Note that $1_{B(\tilde{o}, \tilde{r})} * \mathcal{N}=\frac{|x-\tilde{o}|^{2}}{4}+C$ for some constant $C$, and $1_{\dot{ن}_{i=1}^{n} U_{i}} * \mathcal{N}$ is radially increasing. Therefore $f_{\Omega}$ can be written as

$$
f_{\Omega}=1_{B(\tilde{o}, \tilde{r})} * \mathcal{N}-1_{\dot{U}_{i=1}^{n} U_{i}} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}=\frac{|x-\tilde{o}|^{2}}{4}-g(x),
$$

where $g$ is radially symmetric, and strictly increasing in the radial variable. Since both $f_{\Omega}$ and $\frac{\mid x-\tilde{\left.\right|^{2}}}{4}$ are known to take constant values on $\partial B(\tilde{o}, \tilde{r})$, it implies $g$ must be constant on $\partial B(\tilde{o}, \tilde{r})$ too, and the fact that $g$ is a radially increasing function gives that $\tilde{o}=0$. This finishes the proof for $\Omega>\frac{1}{2}$.

For $\Omega=\frac{1}{2}$, we do not know whether $\left\{U_{i}\right\}_{i=1}^{n}$ are centered at 0 yet. Denote by $U_{1}$ be the innermost one. Then we have

$$
\begin{equation*}
f_{\Omega}(x)=\frac{|x-\tilde{o}|^{2}}{4}-1_{\dot{\cup}_{i=1}^{n} U_{i}} * \mathcal{N}-\frac{|x|^{2}}{4}=\frac{\tilde{o} \cdot x}{2}+\text { const } \quad \text { for } x \in \partial_{\mathrm{out}} U_{1}, \tag{2.1.43}
\end{equation*}
$$

where the second equality follows from Lemma $2.1 .9(\mathrm{~b})$, where we used that $1 \prec j$ for all $2 \leq$ $j \leq n$. Combining (Equation 2.1.43) with the fact that $f_{\Omega}=$ const on $\partial_{\text {out }} U_{1}$ gives $\tilde{o}=0$, that is, the outmost circle must be centered at 0 . This leads to $f_{\Omega}=-\sum_{i=1}^{n} 1_{U_{i}} * \mathcal{N}$. Since $f_{\Omega}=\mathrm{const}$
on each connected component of $\partial U_{i}$, we can apply the last part in the proof of Corollary 2.1.8 to show that all $\left\{U_{i}\right\}_{i=1}^{n}$ are all concentric. Denoting their center by $o_{1}$, we can show that $o_{1}=0$ : Lemma 2.1.9(a) gives $f_{\Omega}(x)=C \ln \left|x-o_{1}\right|$ for some $C<0$ on $\partial B(\tilde{o}, \tilde{r})$, and since we have $f=$ const on $\partial B(\tilde{o}, \tilde{r})$ and $\tilde{o}=0$, it implies that $o_{1}=0$, finishing the proof.

Proof of Lemma 2.1.13. For notational simplicity, we shorten $p_{0, R}, D_{0, R}$ and $\varphi_{0, R}$ into $p_{R}, D_{R}$ and $\varphi_{R}$ thoughout this proof. Recall that $\partial D_{R}=\partial B_{R} \cup \partial V$. Clearly we have $\varphi_{R}=\frac{R^{2}}{2}$ on $\partial B_{R}$, due to $p_{R}=0$ on $\partial_{\text {out }} D_{R}=\partial B_{R}$. We claim that

$$
\begin{equation*}
-\frac{N R_{0}^{2}}{2} \leq \varphi_{R}-\frac{R^{2}}{2} \leq \frac{R_{0}^{2}}{2} \quad \text { on } \partial V, \tag{2.1.44}
\end{equation*}
$$

where $N \geq 1$ is the number of connected components of $V$. Once it is proved, we apply the comparison principle to the functions $\varphi_{R}-\frac{R^{2}}{2}$ and $\pm g$, where

$$
g(x):=\frac{N R_{0}^{2}}{2 \log \left(R / R_{0}\right)} \log \frac{R}{|x|} .
$$

Note that $g \equiv 0$ on $\partial B_{R}$, and $g \geq \frac{N R_{0}^{2}}{2}$ on $\partial V$ since $\partial V \subset B_{R_{0}}$. If $0 \notin \overline{D_{R}}$, then the functions $\varphi_{R}-\frac{R^{2}}{2}$ and $\pm g$ are all harmonic in $D_{R}$, their values on $\partial B_{R}$ are all 0 , and their boundary values on $\partial V$ are ordered due to (Equation 2.1.44). The comparison principle in $D_{R}$ then yields

$$
\begin{equation*}
-g(x) \leq \varphi_{R}(x)-\frac{R^{2}}{2} \leq g(x) \quad \text { in } D_{R} \tag{2.1.45}
\end{equation*}
$$

Since $\varphi_{R}-\frac{R^{2}}{2} \equiv g \equiv 0$ on $\partial B_{R}$, (Equation 2.1.45) gives $\left|\nabla \varphi_{R} \cdot \vec{n}\right| \leq|\nabla g \cdot \vec{n}|=\frac{N R_{0}^{2}}{2 R \log \left(R / R_{0}\right)}$ on $\partial B_{R}$, which is the desired estimate (Equation 2.1.38). And if $0 \in \overline{D_{R}}$, then (Equation 2.1.45) still holds in $D_{R} \backslash B_{\epsilon}$ for all sufficiently small $\epsilon>0$ by applying the comparison principle in this set, and (Equation 2.1.38) again follows as a consequence.

In the rest of the proof we will show (Equation 2.1.44). Its second inequality is straightforward:

$$
\varphi_{R}-\frac{R^{2}}{2} \leq \frac{R_{0}^{2}}{2}+\sup _{\overline{D_{R}}} p_{R}-\frac{R^{2}}{2} \leq \frac{R_{0}^{2}}{2} \quad \text { on } \partial V,
$$

here the first inequality follows from the definition of $\varphi_{R}$ and the fact that $V \subset B_{R_{0}}$, and the second inequality is due to $\sup _{\overline{D_{R}}} p_{R} \leq \frac{\left|D_{R}\right|}{2 \pi} \leq \frac{R^{2}}{2}$ in Proposition 2.1.6.

It remains to prove the first inequality of (Equation 2.1.44). Let us fix any $R>R_{0}$. Denote the $N$ connected components of $\partial V$ by $\left\{\Gamma_{i}\right\}_{i=1}^{N}$, and let $\Gamma_{0}:=\partial B_{R}$. These notations lead to $\partial D_{R}=\cup_{i=0}^{N} \Gamma_{i}$. For $i=0, \ldots, N$, let $L_{i} \subset \mathbb{R}$ be the range of $\varphi_{R}-\frac{R^{2}}{2}$ on $\Gamma_{i}$. By continuity of $\varphi_{R}$, each $L_{i}$ is a closed bounded interval, which can be a single point. Clearly, $L_{0}=\{0\}$ due to $\left.\varphi_{R}\right|_{\partial B_{R}}=\frac{R^{2}}{2}$. Towards a contradiction, suppose

$$
\begin{equation*}
v_{\min }:=\min _{1 \leq i \leq N} \inf L_{i}=\inf _{\partial V}\left(\varphi_{R}-\frac{R^{2}}{2}\right)=:-\frac{N\left|R_{0}\right|^{2}}{2}-\delta \quad \text { for some } \delta>0 \tag{2.1.46}
\end{equation*}
$$

As for the maximum value, since $L_{0}=\{0\}$ we have

$$
\begin{equation*}
v_{\max }:=\max _{0 \leq i \leq N} \sup L_{i} \geq 0 \tag{2.1.47}
\end{equation*}
$$

For $i=1, \ldots, N$, using $\left.p_{R}\right|_{\Gamma_{i}}=$ const, $\varphi_{R}=p_{R}+\frac{|x|^{2}}{2}$ and $\Gamma_{i} \subset B_{R_{0}}$, the length of each interval $L_{i}$ satisfies

$$
\begin{equation*}
\left|L_{i}\right|=\operatorname{osc}_{\Gamma_{i}} \frac{|x|^{2}}{2} \leq \frac{R_{0}^{2}}{2} \quad \text { for } i=1, \ldots, N \tag{2.1.48}
\end{equation*}
$$

Comparing (Equation 2.1.48) with (Equation 2.1.46)-(Equation 2.1.47), we know the union of $\left\{L_{i}\right\}_{i=0}^{N}$ cannot fully cover the interval $\left[v_{\min }, v_{\text {max }}\right]$, thus they can be separated in the following sense: there exists a nonempty proper subset $S \subset\{0, \ldots, N\}$, such that the range of $L_{i}$ for indices in $S$ and $S^{c}:=\{0, \ldots, N\} \backslash S$ are strictly separated by at least $\delta$, i.e. $\min _{i \in S} \inf L_{i} \geq$
$\max _{i \in S^{c}} \sup L_{i}+\delta$. In terms of $\varphi_{R}$, we have

$$
\begin{equation*}
\min _{i \in S} \inf _{\Gamma_{i}} \varphi_{R} \geq \max _{i \in S^{c}} \max _{\Gamma_{i}} \varphi_{R}+\delta \tag{2.1.49}
\end{equation*}
$$

Since $\varphi_{R}$ is harmonic in $D_{R}$, whose boundary is $\cup_{i=0}^{N} \Gamma_{i}$, it is a standard comparison principle exercise to show that (Equation 2.1.49) implies

$$
\begin{equation*}
\sum_{i \in S} \int_{\Gamma_{i}} \nabla \varphi_{R} \cdot \vec{n} d \sigma>0 \tag{2.1.50}
\end{equation*}
$$

where $\vec{n}$ denotes the outer normal of $D_{R}$. But on the other hand, we have

$$
\begin{equation*}
\int_{\Gamma_{i}} \nabla \varphi_{R} \cdot \vec{n} d \sigma=0 \quad \text { for } i=0, \ldots, N \tag{2.1.51}
\end{equation*}
$$

To see this, the cases $i=1, \ldots, N$ can be done by an identical computation as (Equation 2.1.23), and the $i=0$ case follows from $\int_{\partial D_{R}} \varphi_{R} \cdot \vec{n} d \sigma=\int_{D_{R}} \Delta \varphi_{R} d x=0$ and the fact that $\partial D_{R}=$ $\cup_{i=0}^{N} \Gamma_{i}$. Thus we have obtained a contradiction between (Equation 2.1.50) and (Equation 2.1.51), completing the proof.

Proof of Lemma 2.1.14. Assume $V$ has $N$ connected components $\left\{V_{j}\right\}_{j=1}^{N}$ for $N \geq 1$. For notational simplicity, we shorten $D_{0, R}, p_{0, R}$ and $\varphi_{0, R}$ into $D_{R}, p_{R}$ and $\varphi_{R}$ in this proof. Let $\epsilon_{R}:=\frac{1}{2 \pi}\left|D_{R}\right|-\sup _{D_{R}} p_{R}$, which is nonnegative by Proposition 2.1.6. Towards a contradiction, assume there exists a diverging subsequence $\left\{R_{i}\right\}_{i=1}^{\infty}$ such that $\lim _{i \rightarrow 0} \epsilon_{R_{i}}=0$.

Define $\tilde{\varphi}_{R_{i}}:=\varphi_{R_{i}}-\frac{R_{i}^{2}}{2}$. We claim that $\left\{\tilde{\varphi}_{R_{i}}\right\}_{i=1}^{\infty}$ has a subsequence that converges locally uniformly to some bounded harmonic function $\varphi_{\infty}$ in $\mathbb{R}^{2} \backslash V$.

To show this, we will first obtain a uniform bound of $\left\{\tilde{\varphi}_{R_{i}}\right\}_{i=1}^{\infty}$. Note that (Equation 2.1.44) gives that $\sup _{\partial V}\left|\tilde{\varphi}_{R_{i}}\right| \leq \frac{N R_{0}^{2}}{2}$ for all $i \in \mathbb{N}^{+}$. Since $\tilde{\varphi}_{R_{i}} \equiv 0$ on $\partial B_{R_{i}}$ for all $i \in \mathbb{N}^{+}$, the maximum principle for harmonic function gives $\sup _{D_{R_{i}}}\left|\tilde{\varphi}_{R_{i}}\right| \leq \frac{N R_{0}^{2}}{2}$ for all $i \in \mathbb{N}^{+}$.

For any $R>2 R_{0}$, we will obtain a uniform gradient estimate for $\left\{\tilde{\varphi}_{R_{i}}\right\}$ in $D_{R}$ for all $R_{i}>2 R$.

First note that since $\partial B_{R}$ is in the interior of $D_{R_{i}}$ (due to $R_{i}>2 R$ ), interior estimate for harmonic function (together with the above uniform bound) gives that $\left\|\tilde{\varphi}_{R_{i}}\right\|_{C^{2}\left(\partial B_{R}\right)} \leq C\left(N, R_{0}\right)$. On the other boundary $\partial V$, recall that $\tilde{\varphi}_{R_{i}} \left\lvert\, \partial V_{j}=\frac{|x|^{2}}{2}+c_{i, j}\right.$, with $\left|c_{i, j}\right| \leq \frac{(N+1) R_{0}^{2}}{2}$. Thus $\left\|\tilde{\varphi}_{R_{i}}\right\|_{C^{2}\left(\partial D_{R}\right)} \leq$ $C\left(N, R_{0}\right)$ for all $R_{i}>2 R$, and the standard elliptic regularity theory gives the uniform gradient estimate $\sup _{D_{R}}\left|\nabla \tilde{\varphi}_{R_{i}}\right| \leq C(V)$. This allows us to take a further subsequence (which we still denote by $\left.\left\{\tilde{\varphi}_{R_{i}}\right\}\right)$ that converges uniformly in $\overline{D_{R}}$ to some harmonic function $\tilde{\varphi}_{\infty} \in C\left(\overline{D_{R}}\right)$. Since $R>2 R_{0}$ is arbitrary, we can repeat this procedure (for countably many times) to obtain a subsequence that converges locally uniformly to a harmonic function $\tilde{\varphi}_{\infty}$ in $\mathbb{R}^{2} \backslash V$, where $\left.\tilde{\varphi}_{\infty}\right|_{\partial V_{j}}=\frac{|x|^{2}}{2}+c_{j}$ with $\left|c_{j}\right| \leq \frac{(N+1) R_{0}^{2}}{2}$. This finishes the proof of the claim.

Now define

$$
\tilde{p}_{R_{i}}:=p_{R_{i}}-\frac{R_{i}^{2}}{2}=\tilde{\varphi}_{R_{i}}-\frac{|x|^{2}}{2}
$$

which is known to converge locally uniformly to $\tilde{p}_{\infty}:=\tilde{\varphi}_{\infty}-\frac{|x|^{2}}{2}$ in $\mathbb{R}^{2} \backslash V$. Note that $\tilde{p}_{\infty}$ is not radially symmetric up to any translation: To see this, recall that $\left.\tilde{p}_{\infty}\right|_{\partial V_{j}} \equiv c_{j}$. If $\tilde{p}_{\infty}$ is radial about some $x_{0}$, it must be of the form $-\frac{\left|x-x_{0}\right|^{2}}{2}+c$ due to $\Delta \tilde{p}_{\infty}=-2$. As a result, the level sets of $\tilde{p}_{\infty}$ are all nested circles, thus $V$ must be a single disk (where we used that each connected component of $V$ is simply-connected).

Next we will show that $\lim _{i \rightarrow 0} \epsilon_{R_{i}}=0$ implies $\tilde{p}_{\infty}$ is radial up to a translation, leading to a contradiction. For $k \in \mathbb{R}$, let $g_{i}(k):=\left|\left\{x \in D_{R_{i}}: p_{R_{i}}(x)>k\right\}\right|$. In the proof of Proposition 2.1.6, we have shown that $g_{i}(0)=\left|D_{R_{i}}\right|, g_{i}$ is decreasing in $k$, with $g_{i}^{\prime}(k) \leq-2 \pi$ for almost every $k \in\left(0, \sup _{D_{R_{i}}} p_{R_{i}}\right)$. Since $\sup _{D_{R_{i}}} p_{R_{i}}=\frac{1}{2 \pi}\left|D_{R_{i}}\right|-\epsilon_{R_{i}}$, we can control $g_{i}(k)$ from below and above as follows:

$$
\begin{equation*}
\left(\left|D_{R_{i}}\right|-2 \pi k-2 \pi \epsilon_{R_{i}}\right)_{+} \leq g_{i}(k) \leq\left(\left|D_{R_{i}}\right|-2 \pi k\right)_{+} \quad \text { for all } k \geq 0 \tag{2.1.52}
\end{equation*}
$$

Likewise, define $\tilde{g}_{i}(k):=\left|\left\{x \in D_{R_{i}}: \tilde{p}_{R_{i}}(x)>k\right\}\right|$, and $\tilde{g}_{\infty}(k):=\left|\left\{x \in D_{R_{i}}: \tilde{p}_{\infty}(x)>k\right\}\right|$. Since $\tilde{p}_{R_{i}}=p_{R_{i}}-\frac{R_{i}^{2}}{2}$, we have $\tilde{g}_{i}(k)=g_{i}\left(k+\frac{R_{i}^{2}}{2}\right)$ for all $k \geq-\frac{R_{i}^{2}}{2}$, thus (Equation 2.1.52) is
equivalent to

$$
\left(-|V|-2 \pi k-2 \pi \epsilon_{R_{i}}\right)_{+} \leq \tilde{g}_{i}(k) \leq(-|V|-2 \pi k)_{+} \quad \text { for all } k \geq-\frac{R_{i}^{2}}{2}
$$

The locally uniform convergence of $p_{R_{i}}$ gives $\lim _{i \rightarrow \infty} \tilde{g}_{i}=\tilde{g}_{0}$, and since we assume $\lim _{i \rightarrow \infty} \epsilon_{R_{i}}=$ 0 , we take the $i \rightarrow \infty$ limit in the above inequality and obtain

$$
\tilde{g}_{\infty}(k)=(-2 \pi k-|V|)_{+} \quad \text { for all } k \in \mathbb{R}
$$

which implies

$$
\begin{equation*}
\tilde{g}_{\infty}^{\prime}(k)=-2 \pi \quad \text { for all } k \in\left(-\infty, \sup _{\mathbb{R}^{2} \backslash V} \tilde{p}_{\infty}\right) \tag{2.1.53}
\end{equation*}
$$

Applying the proof of Proposition 2.1.6 to $\tilde{p}_{\infty}$ (note that the proof still goes through even though $\tilde{p}_{\infty}$ takes negative values, and is defined in an unbounded domain), we have that (Equation 2.1.53) can happen only if $\tilde{D}_{k}:=\left\{\tilde{p}_{\infty}>k\right\} \dot{\cup}\left(\dot{U}_{\left\{j: c_{j}>k\right\}} \overline{V_{j}}\right)$ is a disk for almost every $k \in\left(-\infty, \sup _{\mathbb{R}^{2} \backslash V} \tilde{p}_{\infty}\right)$, and $\left|\nabla \tilde{p}_{\infty}\right|$ is a constant on almost every $\partial \tilde{D}_{k}$. These two conditions imply that all regular sets of $\tilde{p}_{\infty}$ are concentric circles, thus $\tilde{p}_{\infty}$ is radial up to a translation, and we have obtained a contradiction.

### 2.2 Radial symmetry of nonnegative smooth stationary solutions

Let $\omega$ be a nonnegative compactly supported smooth stationary solution of the 2D Euler equation. Note that $\omega$ being stationary is equivalent with

$$
\nabla^{\perp} \omega \cdot \nabla(\omega * \mathcal{N})=0 \quad \text { in } \mathbb{R}^{2}
$$

As a result, along every regular level set of $\omega$, we have $f:=\omega * \mathcal{N}$ is a constant.
In this section, we prove that such $\omega$ must be radially symmetric up to a translation. In the
proof, the two key steps are to show that every regular level set of $\omega$ is a circle, and these circles are concentric. These are done by approximating $\omega$ by a step function $\omega_{n}=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}$ such that the sets $\left\{D_{i}\right\}$ are disjoint, and $\left\|\omega-\omega_{n}\right\|_{L^{\infty}}=O(1 / n)$. We then define $\varphi^{n}=\frac{|x|^{2}}{2}+\sum_{i=1}^{M_{n}} 1_{D_{i}} p_{i}$ corresponding to this step function $\omega_{n}$, and compute $\mathcal{I}_{n}=\int \omega_{n} \nabla \varphi^{n} \cdot \nabla f d x$ in two ways.

Due to the $O(1 / n)$ error in the approximation, the qualitative statement in Proposition 2.1.6 that "the equality is achieved if and only if $D$ is a disk or annulus" is no longer good enough for us. We need to obtain various quantitative versions of (Equation 2.1.14) for doubly-connected domains, and two such versions are stated below.

In Lemma 2.2.2, the quantitative constant $c_{0}>0$ depends on the Fraenkel asymmetry of the outer boundary defined in Definition 2.2.1.

Definition 2.2.1 (c.f. [42, Section 1.2]). For a bounded domain $E \subseteq \mathbb{R}^{2}$, we define the Fraenkel asymmetry $\mathcal{A}(E) \in[0,1)$ as

$$
\mathcal{A}(E):=\inf _{x_{0}}\left\{\frac{\left|E \Delta\left(x_{0}+r B\right)\right|}{|E|}: x_{0} \in \mathbb{R}^{2}, \pi r^{2}=|E|\right\}
$$

where $B$ is a unit disk in $\mathbb{R}^{2}$.

Lemma 2.2.2. Let $D$ be a doubly connected set. Let us denote the hole of $D$ by an open set $h$, and let $\tilde{D}:=D \cup \bar{h}$. We define $p$ in $D$ as in Lemma 2.1.5. Then if $\mathcal{A}(\tilde{D})>0$, there is a constant $c_{0}$ that only depends on $\mathcal{A}(\tilde{D})$, such that

$$
\left.p\right|_{\partial h} \leq \frac{|D|}{2 \pi}\left(1-c_{0}\right)
$$

Lemma 2.2.2 will be used in the main theorem to show that all level sets of $\omega$ are circles. To obtain radial symmetry of $\omega$, we also need to show all these level sets are concentric. To do this, we need to obtain some quantitative lemmas for a region between two non-concentric disks. In Lemma 2.2.3 we consider a thin tubular region between two non-concentric disks whose radii are
close to each other, and obtain a quantitative version of (Equation 2.1.14) for such domain.

Lemma 2.2.3. For $\epsilon>0$, consider two open disks $B_{1}:=B\left(o_{1}, 1\right)$ and $B_{2}=B\left(o_{2}, 1+\epsilon\right)$ such that $B_{1} \subset B_{2}$. Suppose $\left|o_{1}-o_{2}\right|=a \epsilon$ with $a \in(0,1)$, and let $p$ be defined as in Lemma 2.1.5 in $D:=\overline{B_{2}} \backslash B_{1}$. Then if $\epsilon$ and a satisfy that $0<\epsilon<\frac{a^{2}}{64}$, we have

$$
\begin{equation*}
\left.p\right|_{\partial B_{1}} \leq \frac{|D|}{2 \pi}\left(1-\frac{a^{2}}{16}\right) . \tag{2.2.1}
\end{equation*}
$$

In Lemma 2.2.4 we consider a region between two non-concentric disks (that is not necessarily a thin tubular region), and obtain a quantitative version of (Equation 2.1.14) for such domain.

Lemma 2.2.4. Consider two open disks $B_{r}:=B\left(o_{1}, r\right)$ and $B_{R}=B\left(o_{2}, R\right)$ such that $B_{r} \subset B_{R}$. Let $p$ be defined as in (2.1.5) in $D:=B_{R} \backslash \overline{B_{r}}$. Suppose $l:=\left|o_{1}-o_{2}\right|>0$ and there exist $\delta_{1}>0$, and $\delta_{2}>0$ such that $\delta_{1}<r<R<\delta_{2}$. Then there exist a constant $c_{0}$ that only depends on $\delta_{1}, \delta_{2}$ and $l$ such that

$$
\left.p\right|_{\partial B_{r}} \leq \frac{|D|}{2 \pi}\left(1-c_{0}\right)
$$

The proofs of the above quantitative lemmas will be postponed to Section subsection 2.2.1. Now we are ready to prove the main theorem.

Theorem 2.2.5. Let $\omega$ be a compactly supported smooth nonnegative stationary solution to the $2 D$ Euler equation. Then $\omega$ is radially symmetric up to a translation.

Proof. Note that as mentioned in step 1 of Proposition 2.1.6, we have that for almost every $k \in$ $\left(0,\|\omega\|_{L^{\infty}}\right), \omega^{-1}(\{k\})$ is a smooth 1-manifold. Furthermore, since $\omega$ is compactly supported, each such level set is a disjoint union of finite number of simply closed curves. For any such closed curve, we call it a "level set component" in this proof.

We split the proof into several steps. Throughout step 1, 2 and 3, we prove that all level set components of $\omega$ must be circles. In step 4, we will prove that any two level set components are
nested, i.e. one is contained in the other. Lastly we present the proof that all level set components are concentric in step 5 and 6.

Step 1. Towards a contradiction, suppose there is $k>0$ that is a regular value of $\omega$, and $\omega^{-1}(\{k\})$ has a connected component $\Gamma$ that differs from a circle. Recall that int $(\Gamma)$ denotes the interior of $\Gamma$, which is open and simply connected. Since $\Gamma$ is not a circle, we have $\mathcal{A}(\operatorname{int}(\Gamma))>0$, with $\mathcal{A}$ as in Definition 2.2.1.

In this step, we investigate level set components near $\Gamma$. Since $k$ is a regular value, we can find an open neighborhood $U$ of $\Gamma$ and a constant $\eta>0$ such that $|\nabla \omega|>\eta$ in $U$. For any $x \in \Gamma$, consider the flow map $\Phi_{t}(x)$ originating from $x$, given by

$$
\frac{d}{d t} \Phi_{t}(x)=\frac{\nabla \omega\left(\Phi_{t}(x)\right)}{\left|\nabla \omega\left(\Phi_{t}(x)\right)\right|^{2}}
$$

with initial condition $\Phi_{0}(x)=x$. Since $\frac{\nabla \omega}{|\nabla \omega|^{2}}$ is smooth and bounded in $U$, we can choose $\delta_{1}>0$ so that $\Phi_{t}(\Gamma):=\left\{\Phi_{t}(x): x \in \Gamma\right\}$ lies in $U$ for any $t \in\left(-\delta_{1}, \delta_{1}\right)$. Note that $\Phi_{t}$ is a 1-parameter group of diffeomorphisms, thus $\Phi_{t}(\Gamma)$ is also a smooth simply closed curve for $t \in\left(-\delta_{1}, \delta_{1}\right)$. Then we observe that

$$
\begin{equation*}
\frac{d}{d t} \omega\left(\Phi_{t}(x)\right)=\nabla \omega\left(\Phi_{t}(x)\right) \cdot \frac{\nabla \omega\left(\Phi_{t}(x)\right)}{\left|\nabla \omega\left(\Phi_{t}(x)\right)\right|^{2}}=1 \quad \text { for }(t, x) \in\left(-\delta_{1}, \delta_{1}\right) \times \Gamma . \tag{2.2.2}
\end{equation*}
$$

Hence for each $t \in\left(-\delta_{1}, \delta_{1}\right), \Phi_{t}(\Gamma)$ is a level set component and

$$
\begin{equation*}
\omega\left(\Phi_{t_{1}}(\Gamma)\right) \neq \omega\left(\Phi_{t_{2}}(\Gamma)\right) \text { if } t_{1} \neq t_{2} \tag{2.2.3}
\end{equation*}
$$

By continuity of the map $(t, x) \mapsto \Phi_{t}(x)$, we can find $\delta_{2} \in\left(0, \delta_{1}\right)$ such that

$$
\begin{equation*}
\mathcal{A}\left(\operatorname{int}\left(\Phi_{t}(\Gamma)\right)\right)>\frac{1}{2} \mathcal{A}(\operatorname{int}(\Gamma)) \text { for any } t \in\left(-\delta_{2}, \delta_{2}\right) . \tag{2.2.4}
\end{equation*}
$$

Since two different level sets cannot intersect, we can assume without loss of generality that $\operatorname{int}\left(\Phi_{-\delta_{2}}(\Gamma)\right) \subset \operatorname{int}\left(\Phi_{\delta_{2}}(\Gamma)\right)$. Then it follows from the intermediate value theorem and (Equation 2.2.2) that

$$
\begin{equation*}
\operatorname{int}\left(\Phi_{-\delta_{2}}(\Gamma)\right) \subset \Phi_{t}(\Gamma) \subset \operatorname{int}\left(\Phi_{\delta_{2}}(\Gamma)\right), \text { for any } t \in\left(-\delta_{2}, \delta_{2}\right) \tag{2.2.5}
\end{equation*}
$$

Lastly we denote $V:=\operatorname{int}\left(\Phi_{\delta_{2}}(\Gamma)\right) \backslash \overline{\operatorname{int}\left(\Phi_{-\delta_{2}}(\Gamma)\right)}$ which is a nonempty open doubly connected set, therefore $|V|>0$.

Step 2. For any integer $n>1$, we claim that we can approximate $\omega$ by a step function $\omega_{n}$ of the form $\omega_{n}(x)=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}(x)$, which satisfies all the following properties.
(a) Each $D_{i}$ is a connected open domain with smooth boundary and possibly has a finite number of holes.
(b) Each connected component of $\partial D_{i}$ is a level set component of $\omega$.
(c) $D_{i} \cap D_{j}=\emptyset$ if $i \neq j$.
(d) $\left\|\omega_{n}-\omega\right\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \leq \frac{2}{n}\|\omega\|_{L^{\infty}\left(\mathbb{R}^{n}\right)}$.

To construct such $\omega_{n}$ for a fixed $n>1$, let $r_{0}=0$ and $r_{n+1}=\|\omega\|_{L^{\infty}}$. We pick $r_{1}, \ldots, r_{n}$ to be regular values of $\omega$, such that $0<r_{1}<\cdots<r_{n}<\|\omega\|_{L^{\infty}}$, and $r_{i+1}-r_{i}<\frac{2}{n}\|\omega\|_{L^{\infty}}$ for $i=0, \ldots, n$. We denote $D_{i}:=\left\{x \in \mathbb{R}^{2}: r_{i}<\omega(x)<r_{i+1}\right\}$ for $i=1, \ldots, n-1$, and let $D_{n}:=\left\{x \in \mathbb{R}^{2}: \omega(x)>r_{n}\right\}$. Thus for each $i=1, \cdots, n, D_{i}$ is a bounded domain with smooth boundary. We can then write it as $D_{i}=\dot{\cup}_{l=1}^{m_{i}} D_{i}^{l}$ for some $m_{i} \in \mathbb{N}$ where $D_{i}^{l}$,s are connected components of $D_{i}$. Then let $\omega_{n}(x):=\sum_{i=1}^{n} r_{i} \sum_{l=1}^{m_{i}} 1_{D_{i}^{l}}$. By relabeling the indices, we rewrite $\omega_{n}(x)=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}$, where $M_{n}=\sum_{i=1}^{n} m_{i}$, and each $\alpha_{i} \in\left\{r_{1}, \ldots, r_{n}\right\}$. One can easily check that such $\omega_{n}$ satisfies properties (a)-(d).

Of course, there are many ways to choose the values $r_{1}, \cdots, r_{n}$, with each choice leading to a different $\omega_{n}$. From now on, for any $n>1$, we fix $\omega_{n}(x):=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}(x)$ as any function constructed in the above way. (Note that $\alpha_{i}$ and $D_{i}$ all depend on $n$, but we omit their $n$ dependence for notational simplicity.)

Finally, let us point out that for $\omega_{n}(x)$ constructed above, if $D_{i} \subset V$ for some $i$, then $D_{i}$ must be doubly connected, since step 1 shows that all level set components in $V$ are nested curves. We will use this in step 3 and 5.

Step 3. For any $n>1$, let $\omega_{n}(x)=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}(x)$ be constructed in Step 2. For each $D_{i}$, let we define $p_{i}^{n}$ in $D_{i}$ as in Lemma 2.1.5. We set

$$
\left\{\begin{array}{l}
p^{n}:=\sum_{i=1}^{M_{n}} p_{i}^{n} 1_{D_{i}}  \tag{2.2.6}\\
\varphi_{i}^{n}:=p_{i}+\frac{|x|^{2}}{2} \text { in } D_{i} \\
\varphi^{n}:=\sum_{i=1}^{M_{n}} \varphi_{i}^{n} 1_{D_{i}}
\end{array}\right.
$$

As in Theorem 2.1.7, let $f:=\omega * \mathcal{N}$, and we will compute

$$
\begin{equation*}
\mathcal{I}^{n}:=\int_{\mathbb{R}^{2}} \omega_{n}(x) \nabla \varphi^{n}(x) \cdot \nabla f(x) d x \tag{2.2.7}
\end{equation*}
$$

in two different ways and derive a contradiction by taking the $n \rightarrow \infty$ limit.
On the one hand, the same computation as in (Equation 2.1.22)-(Equation 2.1.23) yields that

$$
\begin{equation*}
\mathcal{I}^{n}=\sum_{i=1}^{M_{n}} \alpha_{i}\left(\int_{\partial D_{i}} f(x) \nabla \varphi_{i}^{n}(x) \cdot \vec{n} d \sigma-\int_{D_{i}} f(x) \Delta \varphi_{i}^{n}(x) d x\right)=0 . \tag{2.2.8}
\end{equation*}
$$

On the other hand,

$$
\begin{aligned}
\mathcal{I}^{n} & =\int_{\mathbb{R}^{2}} \omega_{n}(x) x \cdot \nabla f(x) d x+\int_{\mathbb{R}^{2}} \omega_{n}(x) \nabla p^{n}(x) \cdot \nabla f(x) d x \\
& =: \mathcal{I}_{1}^{n}+\mathcal{I}_{2}^{n} .
\end{aligned}
$$

Since $\omega_{n}$ satisfies property (d) in step 2, it follows that

$$
\lim _{n \rightarrow \infty} \mathcal{I}_{1}^{n}=\int_{\mathbb{R}^{2}} \omega(x) x \cdot \nabla f(x) d x
$$

A similar computation as in (Equation 2.1.24) yields that

$$
\begin{align*}
\int_{\mathbb{R}^{2}} \omega(x) x \cdot \nabla f(x) d x & =\frac{1}{2 \pi} \int_{\mathbb{R}^{2}} \int_{\mathbb{R}^{2}} \omega(x) \omega(y) \frac{x \cdot(x-y)}{|x-y|^{2}} d x d y \\
& =\frac{1}{4 \pi} \int_{\mathbb{R}^{2}} \int_{\mathbb{R}^{2}} \omega(x) \omega(y) d x d y \\
& =\frac{1}{4 \pi}\left(\int_{\mathbb{R}^{2}} \omega(x) d x\right)^{2} \tag{2.2.9}
\end{align*}
$$

where we used the symmetry of the integration domain to get the second equality.
Now we estimate the limit of $\mathcal{I}_{2}^{n}$. By Lemma 2.1.11, we have $\int_{D_{i}}\left|\nabla p_{i}^{n}\right|^{2} d x \leq \int_{D_{i}}|x|^{2} d x$, hence $\left\|\omega_{n} \nabla p\right\|_{L^{2}\left(\mathbb{R}^{2}\right)}$ is uniformly bounded. Since $\omega_{n} \rightarrow \omega$ in $L^{\infty}$, the bounded convergence theorem yields that

$$
\lim _{n \rightarrow \infty} \int_{\mathbb{R}^{2}} \omega_{n} \nabla p^{n} \cdot \nabla\left(\left(\omega_{n}-\omega\right) * \mathcal{N}\right)(x) d x=0
$$

therefore

$$
\liminf _{n \rightarrow \infty} \mathcal{I}_{2}^{n}=\liminf _{n \rightarrow \infty} \int_{\mathbb{R}^{2}} \omega_{n}(x) \nabla p^{n}(x) \cdot \nabla\left(\omega_{n} * \mathcal{N}\right) d x
$$

From now on, we will omit the $n$ dependence in $p_{i}^{n}$ for notational simplicity. Let us break the integral in the right hand side as

$$
\begin{align*}
\int_{\mathbb{R}^{2}} \omega_{n}(x) \nabla p^{n}(x) \cdot \nabla\left(\omega_{n} * \mathcal{N}\right) d x & =\sum_{i, j=1}^{M_{n}} \alpha_{i} \alpha_{j} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x \\
& =\sum_{i=1}^{M_{n}} \alpha_{i}^{2} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{i}} * \mathcal{N}\right) d x+\sum_{i \neq j} \alpha_{i} \alpha_{j} \int_{D_{i}} \nabla p_{i} \cdot \nabla\left(1_{D_{j}} * \mathcal{N}\right) d x \\
& =: F_{1}+F_{2} \tag{2.2.10}
\end{align*}
$$

For $F_{1}$, the divergence theorem yields

$$
\begin{equation*}
F_{1}=\sum_{i=1}^{M_{n}} \alpha_{i}^{2}\left(\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{i}} * \mathcal{N}\right) \cdot \vec{n} d \sigma-\int_{D_{i}} p_{i} d x\right)=-\sum_{i=1}^{M_{n}} \alpha_{i}^{2} \int_{D_{i}} p_{i} d x \tag{2.2.11}
\end{equation*}
$$

where the second equality follows from an identical computation as in (Equation 2.1.25). Then by Proposition 2.1.6, we have

$$
\begin{equation*}
F_{1} \geq-\frac{1}{4 \pi} \sum_{i=1}^{M_{n}} \alpha_{i}^{2}\left|D_{i}\right|^{2} \tag{2.2.12}
\end{equation*}
$$

For $F_{2}$, the divergence theorem yields
$F_{2}=\sum_{i \neq j} \alpha_{i} \alpha_{j}\left(\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma-\int_{D_{i}} p_{i} 1_{D_{j}} d x\right)=\sum_{i \neq j} \alpha_{i} \alpha_{j} \int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma$,
where we use property (c) in step 2 to get the last equality.
For $i \neq j$, recall that as in the proof of Corollary 2.1.8, we denote $j \prec i$ if $D_{j}$ is contained in
a hole of $D_{i}$. Then divergence theorem gives

$$
\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma \begin{cases}=0 & \text { if } j \nprec i  \tag{2.2.13}\\ \geq-\sup _{D_{i}} p_{i}\left|D_{j}\right| & \text { if } j \prec i\end{cases}
$$

Next we will improve this inequality for $j \prec i$ and $i \in L$, where $L:=\left\{i: D_{i} \subset V\right\}$. (Note that $L$ depends on $\omega_{n}$, where we omit this dependence for notational simplicity.) From the discussion at the end of step 2 , we know that $D_{i}$ has exactly one hole for all $i \in L$. Using the divergence theorem together with this observation, (Equation 2.2.13) becomes

$$
\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma \begin{cases}=0 & \text { if } j \nprec i,  \tag{2.2.14}\\ \geq-\sup _{D_{i}} p_{i}\left|D_{j}\right| & \text { if } j \prec i \text { and } i \notin L \\ =-p_{i}| |_{\partial_{\text {in }} D_{i}}\left|D_{j}\right| & \text { if } j \prec i \text { and } i \in L .\end{cases}
$$

For the second case on the right hand side, we simply use the crude bound $\sup _{D_{i}} p_{i} \leq \frac{\left|D_{i}\right|}{2 \pi}$ from Proposition 2.1.6. For the third case we can have a better bound: for any $i \in L$, by Lemma 2.2 .2 and (Equation 2.2.4), there exists an $\epsilon>0$ that only depends on $\mathcal{A}(\operatorname{int}(\Gamma))$ (and in particular is independent of $n$ ), such that $\left.p_{i}\right|_{\partial_{\text {in }} D_{i}} \leq\left(\frac{1}{2 \pi}-\epsilon\right)\left|D_{i}\right|$. Thus (Equation 2.2.14) now becomes

$$
\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma \begin{cases}=0 & \text { if } j \nprec i,  \tag{2.2.15}\\ \geq-\frac{1}{2 \pi}\left|D_{i}\right|\left|D_{j}\right| & \text { if } j \prec i \text { and } i \notin L, \\ \geq-\left(\frac{1}{2 \pi}-\epsilon\right)\left|D_{i}\right|\left|D_{j}\right| & \text { if } j \prec i \text { and } i \in L\end{cases}
$$

Now we are ready to estimate $F_{2}$. Let us break it into

$$
\begin{aligned}
F_{2} & =\sum_{\substack{j<i \\
(i, j) \notin L \times L}} \alpha_{i} \alpha_{j} \int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma+\sum_{\substack{j<i \\
(i, j) \in L \times L}} \alpha_{i} \alpha_{j} \int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma \\
& \geq-\sum_{\substack{j \nless i \\
(i, j) \notin L \times L}} \alpha_{i} \alpha_{j} \frac{1}{2 \pi}\left|D_{i}\right|\left|D_{j}\right|-\sum_{\substack{j \prec i \\
(i, j) \in L \times L}} \alpha_{i} \alpha_{j}\left(\frac{1}{2 \pi}-\epsilon\right)\left|D_{i}\right|\left|D_{j}\right|
\end{aligned}
$$

where the first equality follows from case 1 of (Equation 2.2.15), and the second inequality follows from case 2,3 of (Equation 2.2.15). Finally, by exchanging $i$ with $j$ and taking average with the original inequality, we have

$$
\begin{align*}
F_{2} & \geq-\frac{1}{4 \pi} \sum_{\substack{i \neq j \\
(i, j) \notin L \times L}}\left(\mathbb{1}_{i \prec j}+\mathbb{1}_{j \prec i}\right) \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|-\frac{1}{2} \sum_{\substack{i \neq j \\
(i, j) \in L \times L}}\left(\mathbb{1}_{i \prec j}+\mathbb{1}_{j \prec i}\right) \alpha_{i} \alpha_{j}\left(\frac{1}{2 \pi}-\epsilon\right)\left|D_{i}\right|\left|D_{j}\right| \\
& \geq-\frac{1}{4 \pi} \sum_{\substack{i \neq j \\
(i, j) \not L \times L}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|-\frac{1}{2} \sum_{\substack{i \neq j \\
(i, j) \in L \times L}} \alpha_{i} \alpha_{j}\left(\frac{1}{2 \pi}-\epsilon\right)\left|D_{i}\right|\left|D_{j}\right| \\
& =-\frac{1}{4 \pi} \sum_{i \neq j} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|+\frac{\epsilon}{2} \sum_{\substack{i \neq j \\
(i, j) \in L \times L}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right| \tag{2.2.16}
\end{align*}
$$

where the second inequality is due to the fact that for any $i \neq j$, at most one of $i \prec j$ and $j \prec i$ can be true, thus we always have $\mathbb{1}_{i \prec j}+\mathbb{1}_{j \prec i} \leq 1$.

Therefore, from (Equation 2.2.12) and (Equation 2.2.16) it follows that

$$
\begin{align*}
F_{1}+F_{2} & \geq-\frac{1}{4 \pi} \sum_{i, j=1}^{M_{n}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|+\frac{\epsilon}{2} \sum_{(i, j) \in L \times L} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|-\frac{\epsilon}{2} \sum_{i \in L} \alpha_{i}^{2}\left|D_{i}\right|^{2} \\
& =-\frac{1}{4 \pi}\left(\sum_{i=1}^{M_{n}} \alpha_{i}\left|D_{i}\right|\right)^{2}+\frac{\epsilon}{2}\left(\sum_{i \in L} \alpha_{i}\left|D_{i}\right|\right)^{2}-\frac{\epsilon}{2} \sum_{i \in L} \alpha_{i}^{2}\left|D_{i}\right|^{2} \tag{2.2.17}
\end{align*}
$$

Since we will send $n \rightarrow \infty$, in the rest of step 3 we will denote $L$ by $L^{n}$ to emphasize that $L$ depends on $\omega_{n}$. (In fact $\alpha_{i}$ and $D_{i}$ depend on $n$ as well, and we omit the $n$ dependence for them to
avoid overcomplicating the notations.)
Note that $\sum_{i \in L^{n}} \alpha_{i} 1_{D_{i}}$ converges to $\omega 1_{V}$ in $L^{1}\left(\mathbb{R}^{2}\right)$. Also if $i \in L^{n}$, then the nondegeneracy of $|\nabla \omega|$ on $V$ yields that $\lim _{n \rightarrow \infty} \sup _{i \in L^{n}}\left|D_{i}\right|=0$, consequently

$$
\lim _{n \rightarrow \infty} \sum_{i \in L^{n}} \alpha_{i}^{2}\left|D_{i}\right|^{2} \leq\|\omega\|_{L^{\infty}} \lim _{n \rightarrow \infty} \sup _{i \in L^{n}}\left|D_{i}\right| \int_{\mathbb{R}^{2}} \omega d x=0 .
$$

Therefore it follows that

$$
\begin{align*}
\liminf _{n \rightarrow \infty} \mathcal{I}_{2}^{n} & =\liminf _{n \rightarrow \infty}\left(F_{1}+F_{2}\right) \\
& \geq-\lim _{n \rightarrow \infty} \frac{1}{4 \pi}\left(\sum_{i=1}^{M_{n}} \alpha_{i}\left|D_{i}\right|\right)^{2}+\lim _{n \rightarrow \infty} \frac{\epsilon}{2}\left(\sum_{i \in L^{n}} \alpha_{i}\left|D_{i}\right|\right)^{2} \\
& =-\frac{1}{4 \pi}\left(\int_{\mathbb{R}^{2}} \omega(x) d x\right)^{2}+\frac{\epsilon}{2}\left(\int_{V} \omega(x) d x\right)^{2} \tag{2.2.18}
\end{align*}
$$

Note that $\omega$ is strictly positive in $V$, due to $|\nabla \omega|>0$ in $V$ and $\omega \geq 0$ in $\mathbb{R}^{2}$. Thus from (Equation 2.2.8), (Equation 2.2.9) and (Equation 2.2.18), it follows that

$$
\begin{equation*}
0=\lim _{n \rightarrow \infty} \mathcal{I}^{n} \geq \lim _{n \rightarrow \infty} \mathcal{I}_{1}^{n}+\liminf _{n \rightarrow \infty} \mathcal{I}_{2}^{n} \geq \frac{\epsilon}{2}\left(\int_{V} \omega(x) d x\right)^{2}>0 \tag{2.2.19}
\end{equation*}
$$

which is a contradiction and we have proved that any connected component of a regular level set is a circle.

Step 4. In this step we show that every pair of disjoint level set components are nested. Towards a contradiction, assume that there exist $\Gamma_{1}$ and $\Gamma_{2}$ that are connected components of level sets of regular values of $\omega$, such that $\Gamma_{1}$ and $\Gamma_{2}$ are not nested.

From step 3, we know that $\Gamma_{1}$ and $\Gamma_{2}$ are circles. Then the disks $\operatorname{int}\left(\Gamma_{1}\right)$ and $\operatorname{int}\left(\Gamma_{2}\right)$ are disjoint, and they must be separated by a positive distance since $\Gamma_{1}$ and $\Gamma_{2}$ are level sets of regular values of $\omega$. As in step 1 , using the flow map $\Phi_{t}$ originating from the two circles, we can find disjoint open
annuli $V_{1}$ and $V_{2}$ such that $\Gamma_{i} \subset V_{i}$ for $i=1,2$, and both $\partial_{\text {out }} V_{i}$ and $\partial_{\text {in }} V_{i}$ are level set components of $\omega$.

For any $n>1$, let $\omega_{n}(x)=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}(x)$ be constructed in step 2 , and let

$$
L_{1}^{n}:=\left\{i: D_{i} \subset V_{1}\right\} \quad \text { and } \quad L_{2}^{n}:=\left\{i: D_{i} \subset V_{2}\right\} .
$$

Let $p_{i}$ be defined in (Equation 2.2.6) of step 3, and $\mathcal{I}^{n}$ defined in (Equation 2.2.7). Then on the one hand, the same computations in step 3 give

$$
\begin{equation*}
\lim _{n \rightarrow \infty} \mathcal{I}^{n}=0 \quad \text { and } \quad \lim _{n \rightarrow \infty} \mathcal{I}_{1}^{n}=\frac{1}{4 \pi}\left(\int_{\mathbb{R}^{2}} \omega(x) d x\right)^{2} \tag{2.2.20}
\end{equation*}
$$

Let $F_{1}$ and $F_{2}$ be given by (Equation 2.2.10). For $F_{1}$, the estimate (Equation 2.2.12) still holds. For $F_{2}$, using (Equation 2.2.13) and Proposition 2.1.6, we have

$$
F_{2} \geq-\frac{1}{4 \pi} \sum_{i \prec j \text { or } j \prec i} \alpha_{i} \alpha_{j}\left|D_{i} \| D_{j}\right| .
$$

Since $V_{1}$ and $V_{2}$ are assumed to be not nested, if $(i, j) \in L_{1}^{n} \times L_{2}^{n}$ then neither $i \prec j$ nor $j \prec i$. Therefore it follows that

$$
F_{2} \geq-\frac{1}{4 \pi} \sum_{i \neq j} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|+\frac{1}{4 \pi} \sum_{(i, j) \in L_{1} \times L_{2}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|+\frac{1}{4 \pi} \sum_{(j, i) \in L_{1} \times L_{2}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{j}\right|
$$

Combining the estimates for $F_{1}$ and $F_{2}$ yields

$$
F_{1}+F_{2} \geq-\frac{1}{4 \pi} \sum_{i, j=1}^{M_{n}} \alpha_{i} \alpha_{j}\left|D_{i}\right|\left|D_{i}\right|+\frac{1}{2 \pi}\left(\sum_{i \in L_{1}^{n}} \alpha_{i}\left|D_{i}\right|\right)\left(\sum_{i \in L_{2}^{n}} \alpha_{i}\left|D_{i}\right|\right)
$$

As $n \rightarrow \infty$, since $\sum_{i \in L_{1}^{n}} \alpha_{i} 1_{D_{i}}$ and $\sum_{i \in L_{2}^{n}} \alpha_{i} 1_{D_{i}}$ converge to $\omega 1_{V_{1}}$ and $\omega 1_{V_{2}}$ respectively in
$L^{1}\left(\mathbb{R}^{2}\right)$, we have

$$
\begin{equation*}
\liminf _{n \rightarrow \infty} \mathcal{I}_{2}^{n} \geq \lim _{n \rightarrow \infty}\left(F_{1}+F_{2}\right)=-\frac{1}{4 \pi}\left(\int_{\mathbb{R}^{2}} \omega(x) d x\right)^{2}+\frac{1}{2 \pi}\left(\int_{V_{1}} \omega(x) d x\right)\left(\int_{V_{2}} \omega(x) d x\right) . \tag{2.2.21}
\end{equation*}
$$

Combining (Equation 2.2.20) and (Equation 2.2.21) gives us a similar contradiction as in (Equation 2.2.19), except that $\frac{\epsilon}{2}\left(\int_{V} \omega(x) d x\right)^{2}$ is now replaced by $\frac{1}{2 \pi}\left(\int_{V_{1}} \omega(x) d x\right)\left(\int_{V_{2}} \omega(x) d x\right)$. Thus we complete the proof that level sets are nested.

Step 5. In this step, we aim to show that all level set components are concentric within the same connected component of supp $\omega$. This immediately implies that each connected component of supp $\omega$ is an annulus or a disk, and $\omega$ is radially symmetric about its center.

Towards a contradiction, suppose that there are two level set components $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$ in the same connected component of supp $\omega$, such that they are nested circles, but their centers $O_{\text {in }}$ and $O_{\text {out }}$ do not coincide. We denote their radii by $r_{\text {in }}$ and $r_{\text {out }}$, and define

$$
U:=\overline{\operatorname{int}\left(\Gamma_{\text {out }}\right)} \backslash \operatorname{int}\left(\Gamma_{\text {in }}\right) .
$$

For an illustration of $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$ and $U$, see Figure Figure 2.2(a).
We claim that $\omega$ is uniformly positive in $U$. Recall that all level set components of $\omega$ are nested by step 4 . Thus if $\omega$ achieves zero in $U$, the zero level set must be also nested between $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$, since it can be taken as a limit of level set components whose value approaches 0 ; but this contradicts with the assumption that $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$ lie in the same connected component of supp $\omega$. As a result, we have $\omega_{\text {min }}:=\inf _{U} \omega>0$.

For a sufficiently large $n$, let $\omega_{n}=\sum_{i=1}^{M_{n}} \alpha_{i} 1_{D_{i}}(x)$ be given in step 2 , where we further require both $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$ coincide with some boundary of $D_{i}$. (This is allowed in our construction of $\omega_{n}$ in


Figure 2.2: (a) Illustration of the circles $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$, whose centers are $O_{\text {in }}$ and $O_{\text {out }}$. The set $U$ is colored in blue. (b) For a fixed $n$, each open set $\left\{D_{i}\right\}_{i \in B_{n}}$ is colored in yellow. Note that their union gives exactly the set $U$.
step 2 , since $\omega$ is regular along both $\Gamma_{\text {in }}$ and $\Gamma_{\text {out }}$.) Let us denote

$$
B_{n}:=\left\{1 \leq i \leq M_{n}: D_{i} \subset U\right\}
$$

and note that $U:=\cup_{i \in B_{n}} \bar{D}_{i}$. See Figure Figure 2.2(b) for an illustration of $\left\{D_{i}\right\}_{i \in B_{n}}$.
As before, we denote $i \prec j$ if $D_{i}$ is nested in $D_{j}$. For the integral $\mathcal{I}^{n}$ in (Equation 2.2.7), on the one hand, we have $\mathcal{I}^{n}=0$ for all $n>1$ by (Equation 2.2.8). On the other hand, following the same argument as in step 3 up to (Equation 2.2.13) (where we also use that each $D_{i}$ is already known to be doubly-connected, thus $\int_{\partial D_{i}} p_{i} \nabla\left(1_{D_{j}} * \mathcal{N}\right) \cdot \vec{n} d \sigma=-p_{i}| |_{\partial_{\text {in }} D_{i}}\left|D_{j}\right|$ if $j \prec i$, we have

$$
\begin{aligned}
\liminf _{n \rightarrow \infty} \mathcal{I}^{n} & =\liminf _{n \rightarrow \infty}\left(\frac{1}{4 \pi}\left(\sum_{i=1}^{M_{n}} \alpha_{i}\left|D_{i}\right|\right)^{2}-\sum_{i=1}^{M_{n}} \alpha_{i}^{2} \int_{D_{i}} p_{i} d x-\left.\sum_{1 \leq i, j \leq M_{n}, j \prec i} \alpha_{i} \alpha_{j} p_{i}\right|_{\partial_{\text {in }} D_{i}}\left|D_{j}\right|\right) \\
& \geq \liminf _{n \rightarrow \infty}(\underbrace{\sum_{1 \leq i, j \leq M_{n}, j \prec i} \alpha_{i} \alpha_{j}\left(\frac{1}{2 \pi}\left|D_{i}\right|-\left.p_{i}\right|_{\partial_{\text {in }} D_{i}}\right)\left|D_{j}\right|}_{=: T_{n}})
\end{aligned},
$$

where in the last step we used Proposition 2.1.6.
Note that Proposition 2.1.6 gives $T_{n}>0$, where we have strict positivity, since $O_{\text {in }} \neq O_{\text {out }}$ implies that some $\left\{D_{i}\right\}_{i \in B_{n}}$ must be non-radial. But since the area of these $D_{i}$ 's may approach 0 as
$n \rightarrow \infty$, in order to derive a contradiction after taking $\lim _{\inf }^{n \rightarrow \infty}$, we need to obtain a quantitative estimate for Proposition 2.1.6 for a thin tubular region $D_{i}$ between two circles, which is done in Lemma 2.2.3.

Next we show that the sets $\left\{D_{i}\right\}_{i \in B_{n}}$ that are "non-radial to some extent" must occupy a certain portion of $U$. For $i \in B_{n}$, denote by $o_{\text {in }}^{i}$ and $r_{\text {in }}^{i}$ the center and radius of $\partial_{\text {in }} D_{i}$, and likewise $o_{\text {out }}^{i}$ and $r_{\text {out }}^{i}$ the center and radius of $\partial_{\text {out }} D_{i}$. Note that if $D_{i}$ is the inner-most set in $\left\{D_{i}\right\}_{i \in B_{n}}$, then we have $o_{\mathrm{in}}^{i}=O_{\mathrm{in}}$, and the outmost $D_{i}$ satisfies $o_{\text {out }}^{i}=O_{\text {out }}$. In addition, if $\partial_{\text {out }} D_{i}=\partial_{\text {in }} D_{j}$ for some $i, j \in B_{n}$, then $o_{\text {out }}^{i}=o_{\mathrm{in}}^{j}$. Thus triangle inequality gives

$$
\begin{equation*}
\sum_{i \in B_{n}}\left|o_{\mathrm{in}}^{i}-o_{\text {out }}^{i}\right| \geq\left|O_{\mathrm{in}}-O_{\text {out }}\right|=: c_{0}>0 \quad \text { for all } n \geq n_{0} \tag{2.2.22}
\end{equation*}
$$

In order to apply Lemma 2.2 .3 (which requires the region to have inner radius 1), for each $i \in B_{n}$, consider the scaling

$$
\tilde{p}_{i}(x):=\left(r_{\text {in }}^{i}\right)^{-2} p_{i}\left(r_{\text {in }}^{i} x\right) .
$$

Then $\tilde{p}_{i}$ is defined in $\tilde{D}_{i}:=\left(r_{\text {in }}^{i}\right)^{-1} D_{i}$. Due to the scaling, $\tilde{D}_{i}$ has inner radius 1 (denote the hole by $\tilde{h}_{i}$ ), and outer radius $1+\epsilon_{i}$, where $\epsilon_{i}:=\frac{r_{\text {out }}^{i}-r_{\text {in }}^{i}}{r_{\text {in }}^{i}}>0$. In addition, the distance between the centers of $\partial_{\text {in }} \tilde{D}_{i}$ and $\partial_{\text {out }} \tilde{D}_{i}$ is $a_{i} \epsilon_{i}$, where

$$
a_{i}:=\frac{\left|o_{\text {in }}^{i}-o_{\text {out }}^{i}\right|}{\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right|} .
$$

One can also easily check that $\tilde{p}_{i}$ satisfies $\Delta \tilde{p}_{i}=-2$ in $\tilde{D}_{i}$, and $\int_{\partial \tilde{h}}$ $\nabla \tilde{p} \cdot \vec{n} d \sigma=-2\left|\tilde{h}_{i}\right|=-2 \pi$. By Lemma 2.2.3, if $0<\epsilon_{i}<\frac{\left(a_{i}\right)^{2}}{64}$, then $\left.\tilde{p}_{i}\right|_{\partial \tilde{h}_{i}} \leq \frac{\left|\tilde{D}_{i}\right|}{2 \pi}\left(1-\frac{a_{i}^{2}}{16}\right)$. Thus in terms of $p_{i}$, we have

$$
\begin{equation*}
\left.p_{i}\right|_{\partial_{\text {in }} D_{i}} \leq \frac{1}{2 \pi}\left|D_{i}\right|-c_{1} a_{i}^{2}\left|D_{i}\right| \quad \text { if } i \in B_{n} \text { satisfies } r_{\text {out }}^{i}-r_{\text {in }}^{i} \leq c_{2} a_{i}^{2} \tag{2.2.23}
\end{equation*}
$$

where $c_{1}:=\frac{1}{32 \pi}$ and $c_{2}:=\frac{r_{\text {in }}}{64}$ are independent of $n$ and $i$, due to the fact that $r_{\text {in }}^{i} \geq r_{\text {in }}>0$ for all
$i \in B_{n}$. Using the definition of $a_{i}$, (Equation 2.2.22) can be written as

$$
\begin{equation*}
\sum_{i \in B_{n}} a_{i}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right|>c_{0} \tag{2.2.24}
\end{equation*}
$$

Note that $\sum_{i \in B_{n}}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right|$ satisfies the upper bound

$$
\begin{equation*}
\sum_{i \in B_{n}}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right| \leq \frac{|U|}{2 \pi r_{\text {in }}}=: M \tag{2.2.25}
\end{equation*}
$$

which follows from

$$
|U|=\sum_{i \in B_{n}}\left|D_{i}\right|=\pi \sum_{i \in B_{n}}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right| \underbrace{\left(r_{\text {in }}^{i}+r_{\text {out }}^{i}\right)}_{>2 r_{\text {in }}} .
$$

Combining (Equation 2.2.24) and (Equation 2.2.25) gives

$$
\begin{equation*}
\sum_{i \in B_{n}} \mathbb{1}_{a_{i}>\frac{c_{0}}{2 M}}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right| \geq \sum_{i \in B_{n}}\left(a_{i}-\frac{c_{0}}{2 M}\right)\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right| \geq \frac{c_{0}}{2} \tag{2.2.26}
\end{equation*}
$$

where the first inequality follows from $\mathbb{1}_{a_{i}>\frac{c_{0}}{2 M}} \geq a_{i}-\frac{c_{0}}{2 M}$ (recall that $a_{i} \in(0,1)$ ), and the second inequality follows from subtracting $\frac{c_{0}}{2 M}$ times (Equation 2.2.25) from (Equation 2.2.24).

Let

$$
K_{n}:=\left\{i \in B_{n}: a_{i}>\frac{c_{0}}{2 M}\right\} .
$$

Using this definition and the fact that $\left|D_{i}\right|>2 \pi r_{\text {in }} \mid r_{\text {in }}^{i}-r_{\text {out }}^{i}$, (Equation 2.2.26) can be rewritten as

$$
\begin{equation*}
\sum_{i \in K_{n}}\left|D_{i}\right|>2 \pi r_{\text {in }} \sum_{i \in K_{n}}\left|r_{\text {in }}^{i}-r_{\text {out }}^{i}\right| \geq \pi r_{\text {in }} c_{0} \tag{2.2.27}
\end{equation*}
$$

Now we take a sufficiently large $n$, and discuss two cases (note that different $n$ may lead to different cases):

Case 1. Every $i \in K_{n}$ satisfies $r_{\text {out }}^{i}-r_{\text {in }}^{i} \leq \min \left\{c_{2}\left(\frac{c_{0}}{2 M}\right)^{2}, \frac{r_{\text {in }} c_{0}}{4 r_{\text {out }}}\right\}$. By definition of $K_{n}$, we have $r_{\text {out }}^{i}-r_{\text {in }}^{i} \leq c_{2}\left(\frac{c_{0}}{2 M}\right)^{2} \leq c_{2} a_{i}^{2}$ for $i \in K_{n}$. (The motivation of the second term in the min function
will be made clear later.) Then by (Equation 2.2.23), we have

$$
\frac{1}{2 \pi}\left|D_{i}\right|-\left.p_{i}\right|_{\partial_{\text {in }} D_{i}} \geq c_{1} a_{i}^{2}\left|D_{i}\right| \geq \frac{c_{1} c_{0}^{2}}{4 M^{2}}\left|D_{i}\right| \quad \text { for all } i \in K_{n} .
$$

Since $K_{n}$ is a subset of $B_{n}$ (and recall that $\alpha_{i} \geq \omega_{\min }>0$ for all $i \in B_{n}$ ), we have the following lower bound for $T_{n}$ :

$$
\begin{align*}
T_{n} & \geq \omega_{\min }^{2} \sum_{i, j \in K_{n}, j \prec i}\left(\frac{1}{2 \pi}\left|D_{i}\right|-\left.p_{i}\right|_{\partial_{\text {in }} D_{i}}\right)\left|D_{j}\right| \\
& \geq \omega_{\min }^{2} \sum_{i, j \in K_{n}, j \prec i} \frac{c_{1} c_{0}^{2}}{4 M^{2}}\left|D_{i}\right|\left|D_{j}\right|=\omega_{\min }^{2} \sum_{i, j \in K_{n}, i \neq j} \frac{c_{1} c_{0}^{2}}{8 M^{2}}\left|D_{i}\right|\left|D_{j}\right|  \tag{2.2.28}\\
& =\omega_{\min }^{2} \frac{c_{1} c_{0}^{2}}{16 M^{2}}\left(\left(\sum_{i \in K_{n}}\left|D_{i}\right|\right)^{2}-\sum_{i \in K_{n}}\left|D_{i}\right|^{2}\right)
\end{align*}
$$

Note that the second term in the min function in the assumption gives

$$
\max _{i \in K_{n}}\left|D_{i}\right|<2 \pi r_{\text {out }}\left(r_{\text {out }}^{i}-r_{\text {in }}^{i}\right) \leq \frac{\pi r_{\text {in }} c_{0}}{2} \leq \frac{1}{2} \sum_{i \in K_{n}}\left|D_{i}\right|,
$$

where we use (Equation 2.2.27) in the last inequality. Applying this to the right hand side of (Equation 2.2.28) gives

$$
T_{n} \geq \omega_{\min }^{2} \frac{c_{1} c_{0}^{2}}{16 M^{2}} \cdot \frac{1}{2}\left(\sum_{i \in K_{n}}\left|D_{i}\right|\right)^{2} \geq \omega_{\min }^{2} \frac{c_{1} c_{0}^{2}}{32 M^{2}}\left(\pi r_{\mathrm{in}} c_{0}\right)^{2}
$$

Case 2. If Case 1 is not true, then there must be some $i_{0} \in K_{n}$ satisfying $r_{\text {out }}^{i_{0}}-r_{\text {in }}^{i_{0}}>$ $\min \left\{c_{2}\left(\frac{c_{0}}{2 M}\right)^{2}, \frac{r_{\text {in }} c_{0}}{4 r_{\text {out }}}\right\}=: c_{3}$, which leads to

$$
\left|o_{\text {in }}^{i_{0}}-o_{\text {out }}^{i_{0}}\right|=a_{i_{0}}\left(r_{\text {out }}^{i_{0}}-r_{\text {in }}^{i_{0}}\right)>\frac{c_{0} c_{3}}{2 M}=: l .
$$

Although this set $D_{i_{0}}$ is likely not thin enough for us to apply Lemma 2.2.3, since $\left|o_{\mathrm{in}}^{i_{0}}-o_{\text {out }}^{i_{0}}\right|$ is
bounded below by a positive constant independent of $n$, we can still use Lemma 2.2.4 to conclude that $\frac{1}{2 \pi}\left|D_{i_{0}}\right|-\left.p_{i_{0}}\right|_{\partial_{\text {in }} D_{i_{0}}} \geq c_{4}$ for some $c_{4}>0$ only depending on $r_{\text {in }}, r_{\text {out }}$ and $l$. This leads to

$$
T_{n} \geq \sum_{i_{0} \succ j} \omega_{\min } \alpha_{j} c_{4}\left|D_{j}\right| \geq \omega_{\min } c_{4} \sum_{D_{j} \subset \operatorname{int}\left(\Gamma_{\text {in }}\right)} \alpha_{j}\left|D_{j}\right| \geq \omega_{\min } c_{4} \cdot \frac{1}{2} \int_{\mathrm{int}\left(\Gamma_{\mathrm{in}}\right)} \omega d x
$$

where the last inequality follows from the fact that for all sufficiently large $n$, the definition of $\omega_{n}$ gives $\sum_{D_{j} \subset \operatorname{int}\left(\Gamma_{\text {in }}\right)} \alpha_{j}\left|D_{j}\right|=\int_{\operatorname{int}\left(\Gamma_{\text {in }}\right)} \omega_{n} d x \geq \frac{1}{2} \int_{\text {int }\left(\Gamma_{\text {in }}\right)} \omega d x$. Note that the last integral is positive since $\omega>0$ on $\Gamma_{\mathrm{in}}$, and it is clearly independent of $n$.

From the above discussion, for all sufficiently large $n$, regardless of whether we are in Case 1 or 2 for this $n$, we always have that $T_{n}$ is bounded below by some uniformly positive constant independent of $n$. Therefore taking the $n \rightarrow \infty$ limit gives

$$
\liminf _{n \rightarrow \infty} \mathcal{I}^{n} \geq \liminf _{n \rightarrow \infty} T_{n}>0
$$

This contradicts $\mathcal{I}^{n}=0$, therefore finishing the proof of step 5 .
Step 6. It remains to show that all connected components of supp $\omega$ are concentric. If $\operatorname{supp} \omega$ has finitely many connected components, we could proceed similarly as the end of the proof of Corollary 2.1.8. But since supp $\omega$ may have countably many connected components, we need to use a different argument.

Let us denote the connected components of supp $\omega$ by $\left\{U_{i}\right\}_{i \in I}$, where $I$ may have countably many elements. Denote their centers by $\left\{o_{i}\right\}_{i \in I}$, their radii by $\left\{R_{i}\right\}_{i \in I}$, and their outer boundaries by $\left\{\partial_{\text {out }} U_{i}\right\}_{i \in I}$. Without loss of generality, suppose the $x$-coordinates of their centers $\left\{o_{i}^{1}\right\}_{i \in I}$ are not all identical.

Among the (possibly infinitely many) collection of circles $\left\{\partial_{\text {out }} U_{i}\right\}_{i \in I}$, let $\Gamma_{r}$ be the "circle with rightmost center" among them, in the following sense:

- If there exists some $i_{0} \in I$ such that $o_{i_{0}}^{1}=\sup _{i \in I} o_{i}^{1}$, we define $\Gamma_{r}:=\partial_{\text {out }} U_{i_{0}}$. (If the
supremum is achieved at more than one indices, we set $i_{0}$ to be any of them.)
- Otherwise, take any subsequence $\left\{i_{k}\right\}_{k \in \mathbb{N}} \subset I$ such that $\lim _{k \rightarrow \infty} o_{i_{k}}^{1}=\sup _{i \in I} o_{i}^{1}$. Since $\omega$ has compact support, we can extract a further subsequence (which we still denote by $\left\{i_{k}\right\}_{k \in \mathbb{N}}$ ), such that both $o_{i_{k}}$ and $r_{i_{k}}$ converge as $k \rightarrow \infty$, and denote their limit by $O_{r} \in \mathbb{R}^{2}$ and $R_{r} \in \mathbb{R}$. Finally let $\Gamma_{r}:=\partial B\left(O_{r}, R_{r}\right)$.

With the above definition, we point out that $f:=\omega * \mathcal{N}=$ const on $\Gamma_{r}$. Note that in both cases above, we can find a sequence of level set components of $\omega$ that converges to $\Gamma_{r}$, in the sense that the Hausdorff distance between the two sets goes to 0 . Since $f=$ const on each level set component of $\omega$, continuity of $f$ gives that $f=$ const on $\Gamma_{r}$.

Let $f_{i}(x):=\left(\omega 1_{U_{i}}\right) * \mathcal{N}$ for $i \in I$; note that by definition we have $f=\sum_{i \in I} f_{i}$. Lemma 2.1.9 gives the following:
(a) For all $x \in\left(\operatorname{int}\left(\partial_{\text {out }} U_{i}\right)\right)^{c}$, we have $f_{i}(x)=\frac{1}{2 \pi}\left(\int_{U_{i}} \omega d x\right) \ln \left|x-o_{i}\right|$.
(b) If $U_{i}$ is doubly-connected, then $f_{i}=\operatorname{const} \operatorname{in} \operatorname{int}\left(\partial_{\text {in }} U_{i}\right)$, where the constants are different for different $i$.

Note that for any $i \in I, U_{i}$ must be either nested inside $\Gamma_{r}$, or have $\Gamma_{r}$ nested in its hole. (By a slight abuse of notation, we use $i \prec \Gamma_{r}$ and $i \succ \Gamma_{r}$ to denote these two relations.) Let $\Gamma_{r}^{R}:=\left(O_{r}^{1}+R_{r}, O_{r}^{2}\right)$ and $\Gamma_{r}^{L}:=\left(O_{r}^{1}-R_{r}, O_{r}^{2}\right)$ be the rightmost/leftmost point of the circle $\Gamma_{r}$. Note that (b) implies $f_{i}\left(\Gamma_{r}^{R}\right)=f_{i}\left(\Gamma_{r}^{L}\right)$ for all $i \succ \Gamma_{r}$, whereas (a) gives the following for all $i \prec \Gamma_{r}$ :

$$
f_{i}\left(\Gamma_{r}^{R}\right)=\frac{\int_{U_{i}} \omega d x}{2 \pi} \ln \left|\Gamma_{r}^{R}-o_{i}\right| \geq \frac{\int_{U_{i}} \omega d x}{2 \pi} \ln \left|\Gamma_{r}^{L}-o_{i}\right|=f_{i}\left(\Gamma_{r}^{L}\right),
$$

where the inequality follows from that $\left|O_{r}^{1}+R_{r}-o_{i}^{1}\right| \geq\left|O_{r}^{1}-R_{r}-o_{i}^{1}\right|$, which is a consequence of $o_{i}^{1} \leq O_{r}^{1}$ due to our choice of $O_{r}$. (Also note that $\Gamma_{r}^{R}$ and $\Gamma_{r}^{L}$ have the same $y$-coordinate.)

As a result, summing over all $i \in I$ gives $f\left(\Gamma_{r}^{R}\right) \geq f\left(\Gamma_{r}^{L}\right)$, where the equality is achieved if and only if $o_{i}^{1}=O_{r}$ for all $i \prec \Gamma_{r}$. Now we discuss two cases:

Case 1. There is some $i \prec \Gamma_{r}$ with $o_{i}^{1}<O_{r}$. In this case the above discussion gives $f\left(\Gamma_{r}^{R}\right)>$
$f\left(\Gamma_{r}^{L}\right)$, which directly leads to a contradiction to $f=\mathrm{const}$ on $\Gamma_{r}$.
Case 2. If case 1 does not hold, then let us define $\Gamma_{l}$ as a "circle with leftmost center" among $\left\{\partial_{\text {out }} U_{i}\right\}_{i \in I}$ in the same way as $\Gamma_{r}$. Then we must have $O_{l}^{1}<O_{r}^{1}$, and since case 1 does not hold (i.e. all $i \prec \Gamma_{r}$ satisfy that $o_{i}^{1}=O_{r}$ ), we must have $\Gamma_{l} \succ \Gamma_{r}$. By definition of $\Gamma_{r}$, there exists some $U_{i_{0}}$ whose outer boundary is sufficiently close to $\Gamma_{r}$ and center sufficiently close to $O_{r}$. As a result, $i_{0} \prec \Gamma_{l}$ and $o_{i_{0}}^{1}>O_{l}^{1}$.

Let $\Gamma_{l}^{L}$ and $\Gamma_{l}^{R}$ be the leftmost/rightmost point of $\Gamma_{l}$. A parallel argument as above then gives that $f_{i}\left(\Gamma_{l}^{L}\right) \geq f_{i}\left(\Gamma_{l}^{R}\right)$ for all $i \in I$. Since we have found an $i_{0} \prec \Gamma_{l}$ with $o_{i_{0}}^{1}>O_{l}^{1}$, we have $f_{i_{0}}\left(\Gamma_{l}^{L}\right)>f_{i_{0}}\left(\Gamma_{l}^{R}\right)$, thus summing over all $i \in I$ gives the strict inequality $f\left(\Gamma_{l}^{L}\right)>f\left(\Gamma_{l}^{R}\right)$, contradicting with $f=$ const on $\Gamma_{l}$.

In both cases above we have obtained a contradiction, thus $\left\{o_{i}\right\}_{i \in I}$ must have the same $x$ coordinate. An identical argument shows that their $y$-coordinate must also be identical, thus $\left\{U_{i}\right\}_{i \in I}$ are concentric. Since $\omega$ is known to be radial within each $U_{i}$ (about its own center) in step 1-5, the proof is now finished.

In the next corollary, we show that the above proof for stationary smooth solutions can be extended (with some modifications) to show radial symmetry for rotating patches with $\Omega<0$.

Corollary 2.2.6. Let $\omega$ be a compactly supported smooth nonnegative rotating solution to the $2 D$ Euler equation, with angular velocity $\Omega<0$. Then $\omega$ is radially symmetric about the origin.

Proof. The proof is very similar to the proof of Theorem 2.2.5, and we only highlight the differences. Let $\left\{\omega_{n}\right\}$ be the same approximation for $\omega$ as in step 2 of Theorem 2.2.5. We consider the same setting as in (Equation 2.2.6) and (Equation 2.2.7), except with $f(x)$ replaced by $f_{\Omega}(x):=\omega * \mathcal{N}-\frac{\Omega}{2}|x|^{2}$. From the assumption on $\omega$, we have that $f_{\Omega}$ is a constant on each regular level set component of $\omega$. Thus the same computations in (Equation 2.2.8) give $\mathcal{I}^{n}=0$ for all $n>1$.

On the other hand, we have

$$
\begin{equation*}
\mathcal{I}^{n}=\int_{\mathbb{R}^{2}} \omega_{n} \nabla \varphi^{n} \cdot \nabla(\omega * \mathcal{N}) d x+(-\Omega) \int_{\mathbb{R}^{2}} \omega_{n} \nabla \varphi^{n} \cdot x d x=: \mathcal{I}_{1}^{n}+\underbrace{(-\Omega)}_{\geq 0} \mathcal{I}_{2}^{n} \tag{2.2.29}
\end{equation*}
$$

The same argument as in (Equation 2.1.33) of Theorem 2.1 .10 gives that $\mathcal{I}_{2}^{n} \geq 0$. As for $\mathcal{I}_{1}^{n}$, in step 3 - step 5 of the proof of Theorem 2.2.5, we have already shown that $\lim _{\inf }^{n \rightarrow \infty} \mathcal{I}_{1}^{n} \geq 0$, and the equality is achieved if and only if each connected component of $\{\omega>0\}$ is radially symmetric up to a translation, and they are all nested.

Let us decompose supp $\omega$ into (possibly infinitely many) connected components $\cup_{i \in I} U_{i}$, with centers $\left\{o_{i}\right\}_{i \in I}$. Our goal is to show $o_{i} \equiv(0,0)$ for $i \in I$. Note that it suffices to show that their $x$-coordinates satisfy $\sup _{i \in I} o_{i}^{1} \leq 0$. Once we prove this, a parallel argument gives $\inf _{i \in I} o_{i}^{1} \geq 0$, which implies $o_{i}^{1} \equiv 0$ for $i \in I$, and the same can be done for the $y$-coordinate.

Towards a contradiction, suppose $\sup _{i \in I} o_{i}^{1}>0$. We can then define $\Gamma_{r}$ in the same way as step 6 of the proof of Theorem 2.2.5, i.e. it is the "circle with rightmost center" among $\left\{\partial_{\text {out }} U_{i}\right\}_{i \in I}$, and its center $O_{r}$ satisfies $O_{r}^{1}=\sup _{i \in I} o_{i}^{1}>0$. Since the new $f$ function takes constant values along each level set component of $\omega$, we again have that $f=$ const on $\Gamma_{r}$. Let $\Gamma_{r}^{R}$ and $\Gamma_{r}^{L}$ be the rightmost/leftmost points on $\Gamma_{r}$. Note that their distances to the origin satisfy $\left|\Gamma_{r}^{R}\right|>\left|\Gamma_{r}^{L}\right|$, where the strict inequality is due to the assumption $O_{r}^{1}>0$.

Let us define $f_{i}(x)=\left(\omega 1_{U_{i}}\right) * \mathcal{N}$ for $i \in I$, and note that $f_{\Omega}=\left(\sum_{i \in I} f_{i}\right)-\Omega|x|^{2}$. The properties ( $\mathbf{a}, \mathbf{b}$ ) in step 6 of Theorem 2.2.5 still hold for $f_{i}$, thus we have $f_{i}\left(\Gamma_{r}^{R}\right) \geq f_{i}\left(\Gamma_{r}^{L}\right)$ for all $i \in I$. This leads to

$$
f_{\Omega}\left(\Gamma_{r}^{R}\right)=\left(\sum_{i \in I} f_{i}\left(\Gamma_{r}^{R}\right)\right)+\underbrace{(-\Omega)}_{>0}\left|\Gamma_{r}^{R}\right|^{2}>\left(\sum_{i \in I} f_{i}\left(\Gamma_{r}^{L}\right)\right)+(-\Omega)\left|\Gamma_{r}^{L}\right|^{2}=f_{\Omega}\left(\Gamma_{r}^{L}\right)
$$

contradicting the fact that $f_{\Omega} \equiv$ const on $\Gamma_{r}$.

### 2.2.1 Proofs of the quantitative lemmas

Let us start with the proof of Lemma 2.2.2. Let us begin by stating two lemmas that we will use in the proof. The first one is a quantitative version of the isoperimetric inequality obtained by Fusco, Maggi and Pratelli [42].

Lemma 2.2 .7 (c.f. [42, Section 1.2]). Let $E \subseteq \mathbb{R}^{2}$ be a bounded domain. Then there is some constant $c \in(0,1)$, such that

$$
P(E) \geq 2 \sqrt{\pi} \sqrt{|E|}\left(1+c \mathcal{A}(E)^{2}\right)
$$

where $P(E)=\mathcal{H}^{1}(\partial E)$ denotes the perimeter of $E$.

The second lemma is a simple result relating the Fraenkel asymmetry of a set $E$ with its subsets $U$.

Lemma 2.2.8 (c.f. [34, Lemma 4.4]). Let $E \subseteq \mathbb{R}^{2}$ be a bounded domain. For all $U \subseteq E$ satisfying $|U| \geq|E|\left(1-\frac{\mathcal{A}(E)}{4}\right)$, we have

$$
\mathcal{A}(U) \geq \frac{\mathcal{A}(E)}{4}
$$

Proof of Lemma 2.2.2. The proof of the Lemma 2.2.2 is similar to [34, Proposition 4.5] obtained by Craig, Kim and the last author. For the sake of completeness, we give a proof below. Let $g(k)$, $D_{k}$ and $\tilde{D}_{k}$ be defined as in Proposition 2.1.6, let $\tilde{D}=D \cup \bar{h}$ and define $p_{h}:=\left.p\right|_{\partial h}$. We start by following the proof of Proposition 2.1.6, except that after obtaining (Equation 2.1.18), instead of using the isoperimetric inequality, we use the stability version in Lemma 2.2 .7 to control $P\left(\tilde{D}_{k}\right)$.

This gives

$$
\begin{aligned}
g^{\prime}(k)\left(g(k)+|h| \mathbb{1}_{p_{h}>k}\right) & \leq-\frac{1}{2} P\left(\tilde{D}_{k}\right)^{2} \\
& \leq-2 \pi\left|\tilde{D}_{k}\right|\left(1+c \mathcal{A}\left(\tilde{D}_{k}\right)^{2}\right)^{2} \\
& \leq-2 \pi\left(g(k)+|h| \mathbb{1}_{p_{h}>k}\right)\left(1+c \mathcal{A}\left(\tilde{D}_{k}\right)^{2}\right) .
\end{aligned}
$$

Hence it follows from Lemma 2.2.8 that

$$
\begin{equation*}
g^{\prime}(k) \leq-2 \pi\left(1+c \frac{\mathcal{A}(\tilde{D})^{2}}{16}\right) \text { for all } k \text { such that }\left|\tilde{D}_{k}\right| \geq|\tilde{D}|\left(1-\frac{\mathcal{A}(\tilde{D})}{4}\right) \tag{2.2.30}
\end{equation*}
$$

We claim that

$$
\begin{equation*}
g(k) \leq|D|-2 \pi\left(1+c \frac{\mathcal{A}(\tilde{D})^{2}}{16}\right) k \quad \text { for } k<\min \left\{p_{h}, \frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{16 \pi}\right\} \tag{2.2.31}
\end{equation*}
$$

Towards a contradiction, suppose there is $k_{0} \leq \min \left(p_{h}, \frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{16 \pi}\right)$ such that (Equation 2.2.31) is violated. Since $1+c \frac{\mathcal{A}(\tilde{D})^{2}}{16} \leq 2$, we have

$$
g\left(k_{0}\right)>|D|-4 \pi k_{0} \geq|D|-\frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{4}
$$

therefore

$$
\begin{aligned}
\left|\tilde{D}_{k_{0}}\right| & =g\left(k_{0}\right)+|h| \\
& >|D|-\frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{4}+|h| \\
& =|\tilde{D}|-\frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{4}=|\tilde{D}|\left(1-\frac{\mathcal{A}(\tilde{D})}{4}\right)
\end{aligned}
$$

Hence for all $k \leq k_{0}, g^{\prime}(k)$ satisfies the inequality (Equation 2.2.30). Thus we have

$$
\begin{aligned}
g\left(k_{0}\right) & \leq \int_{0}^{k_{0}}-2 \pi\left(1+c \frac{\mathcal{A}(\tilde{D})^{2}}{16}\right) d k+|D| \\
& =|D|-2 \pi\left(1+c \frac{\mathcal{A}(\tilde{D})^{2}}{16}\right) k_{0}
\end{aligned}
$$

contradicting our assumption on $k_{0}$.

Finally, to control $p_{h}$, we discuss two cases below, depending on which one in the minimum function in (Equation 2.2.31) is smaller. For simplicity, we denote $A:=\frac{\mathcal{A}(\tilde{D})|\tilde{D}|}{16 \pi}$ and $B:=c \frac{\mathcal{A}(\tilde{D})^{2}}{16}$.

Case 1: $p_{h} \leq A$. In this case (Equation 2.2.31) holds for all $k \leq p_{h}$, thus

$$
0 \leq g\left(p_{h}\right) \leq|D|-2 \pi(1+B) p_{h}
$$

implying

$$
p_{h} \leq \frac{|D|}{2 \pi(1+B)} \leq \frac{|D|}{2 \pi}\left(1-c_{0}\right)
$$

for some constant $c_{0}$ which only depends on $\mathcal{A}(\tilde{D})$.

Case 2: $p_{h}>A$. In this case (Equation 2.2.31) gives $g(A) \leq|D|-2 \pi(1+B) A$ and we use a crude bound for $k \geq A$ that is $g^{\prime}(k) \leq-2 \pi$. Therefore for $k>A$,

$$
\begin{aligned}
g(k) & =\int_{A}^{k} g^{\prime}(k) d k+g(A) \leq-2 \pi(k-A)+|D|-2 \pi(1+B) A \\
& =|D|-2 \pi k-2 \pi A B \\
& \leq|D|\left(1-\frac{\mathcal{A}(\tilde{D})}{8} B\right)-2 \pi k,
\end{aligned}
$$

where the last inequality follows from $A>\frac{\mathcal{A}(\tilde{D})|D|}{16 \pi}$. Plugging in $k=p_{h}$ gives

$$
0 \leq g\left(p_{h}\right) \leq|D|\left(1-\frac{\mathcal{A}(\tilde{D})}{8} B\right)-2 \pi p_{h}
$$

leading to

$$
p_{h} \leq \frac{|D|}{2 \pi}\left(1-c_{0}\right)
$$

again $c_{0}$ only depends on $\mathcal{A}(\tilde{D})$.
Next we prove Lemma 2.2.3.

Proof of Lemma 2.2.3. Without loss of generality, we can assume that $o_{1}=(0,0)$ and $o_{2}=$ $(a \epsilon, 0)$. To estimate $\left.p\right|_{\partial B_{1}}$, we decompose $p$ into

$$
p=\left.p\right|_{\partial B_{1}} g+u
$$

where $g$ satisfies

$$
\begin{cases}\Delta g=0 & \text { in } D  \tag{2.2.32}\\ g=1 & \text { on } \partial B_{1} \\ g=0 & \text { on } \partial B_{2}\end{cases}
$$

and $u$ satisfies

$$
\begin{cases}\Delta u=-2 & \text { in } D  \tag{2.2.33}\\ u=0 & \text { on } \partial D\end{cases}
$$

Using this decomposition as well as the definition of $p$, we have

$$
-2\left|B_{1}\right|=\int_{\partial B_{1}} \nabla p \cdot \vec{n} d \sigma=\left.p\right|_{\partial B_{1}} \int_{\partial B_{1}} \nabla g \cdot \vec{n} d \sigma+\int_{\partial B_{1}} \nabla u \cdot \vec{n} d \sigma
$$

where $\vec{n}$ is the outer-normal of $B_{1}$ throughout this proof. Thus

$$
\begin{equation*}
\left.p\right|_{\partial B_{1}}=\frac{1}{\int_{\partial B_{1}} \nabla g \cdot \vec{n} d \sigma}\left(-2 \pi-\int_{\partial B_{1}} \nabla u \cdot \vec{n} d \sigma\right) . \tag{2.2.34}
\end{equation*}
$$

To estimate $\left.p\right|_{\partial B_{1}}$, it remains to estimate the two integrals in (Equation 2.2.34).
The function $g$ can be explicitly constructed using the conformal mapping from $D$ to a perfect annulus centered at 0 . Consider the Möbius map $h: \mathbb{C} \rightarrow \mathbb{C}$ given by

$$
h(z):=\frac{z+b}{1+b z},
$$

where $b \in \mathbb{R}$ will be fixed soon. Note that the unit circle and the real line are both invariant under $h$, and $\partial B_{2}$ is mapped to some circle centered on the real line. In order to make $h\left(\partial B_{2}\right)$ centered at 0 , since the left/right endpoints of $\partial B_{2}$ are $\pm(1+\epsilon)+a \epsilon$, we look for $b \in \mathbb{R}$ that solves

$$
\begin{equation*}
h(1+\epsilon+a \epsilon)=-h(-1-\epsilon+a \epsilon) \tag{2.2.35}
\end{equation*}
$$

Plugging the definition of $h$ into the above equation, we know that $b$ is a root of the quadratic polynomial

$$
f(b):=b^{2}-\frac{2+\left(1-a^{2}\right) \epsilon}{a} b+1 .
$$

Clearly, for $0<a<1$, $f$ has two positive roots whose product is 1 , thus one is in $(0,1)$ and the other in $(1,+\infty)$. We define $b$ to be the root in $(0,1)$. One can easily check that $f(a)<0$, and $f\left(\frac{a}{2}\right)>0$ if $a^{2}>2\left(1-a^{2}\right) \epsilon$, which is true due to our assumption $a^{2}>64 \epsilon$. Thus for all $\epsilon \in\left(0, \frac{a^{2}}{64}\right)$ we have

$$
0<\frac{a}{2}<b<a<1 .
$$

Note that $h$ is holomorphic in $\mathbb{C}$ except at the two singularity points $-b$ and $-\frac{1}{b}$. We have already shown that $-b \in B_{1}$, thus it is outside of $D$. Next we will show that $-\frac{1}{b} \in B_{2}^{c}$, thus is also
outside of $D$. To see this, note that

$$
\frac{-1-\epsilon+a \epsilon+b}{1+b(-1-\epsilon+a \epsilon)}=h(-1-\epsilon+a \epsilon)=-h(1+\epsilon+a \epsilon)=-\frac{1+\epsilon+a \epsilon+b}{1+b(1+\epsilon+a \epsilon)}<0
$$

where the inequality follows from the fact that $a, b, \epsilon>0$. Since the numerator of the left hand side is already known to be negative due to $a, b \in(0,1)$, its denominator must be positive, leading to $-\frac{1}{b}<-1-\epsilon+a \epsilon$, i.e. $-\frac{1}{b} \in B_{2}^{c}$.

Now we define $g: \mathbb{R}^{2} \backslash\{(-b, 0) \cup(-1 / b, 0)\} \rightarrow \mathbb{R}$ as

$$
g(x):=-\frac{1}{\log |h(1+\epsilon+a \epsilon)|} \log |h(z)|+1 \quad \text { for } z=x_{1}+i x_{2}
$$

Let us check that $g$ indeed satisfies (Equation 2.2.32): first note that $g$ satisfies the boundary conditions in (Equation 2.2.32), since $h$ maps $D$ to a perfect annulus centered at the origin, whose inner boundary is $\partial B_{1}$. In addition, $g$ is harmonic in $\mathbb{R}^{2} \backslash\{(-b, 0) \cup(-1 / b, 0)\}$, thus harmonic in $D$.

Using the explicit formula of $g$, we have

$$
\Delta g(x)=-\frac{2 \pi}{\log |h(1+\epsilon+a \epsilon)|}\left(\delta_{(-b, 0)}(x)-\delta_{\left(-\frac{1}{b}, 0\right)}(x)\right)
$$

in the distribution sense. We can then apply the divergence theorem to $g$ in $B_{1}$, and compute the integral containing $g$ in (Equation 2.2.34) explicitly as

$$
\begin{equation*}
\int_{\partial B_{1}} \nabla g \cdot \vec{n} d \sigma=-\frac{2 \pi}{\log |h(1+\epsilon+a \epsilon)|} \tag{2.2.36}
\end{equation*}
$$

As for the integral containing $u$ in (Equation 2.2.34), we compare $u$ with a radial barrier function

$$
w(x):=-2(|x|-1)(|x|-1-2 \epsilon),
$$

which satisfies $w=0$ on $\partial B_{1}$ and $w>0$ on $\partial B_{2}$. Note that

$$
\Delta w=\left(\partial_{r r}+\frac{1}{r} \partial_{r}\right) w=-8+\frac{4+4 \epsilon}{r} \leq-2 \quad \text { in } D
$$

where we used that $\epsilon \in\left(0, \frac{1}{2}\right)$ and $r>1$ in $D$ in the last inequality. Thus $w-u$ is superharmonic in $D$ and nonnegative on $\partial D$, which allows us to apply the classical maximum principle to obtain $u \leq w$ in $\bar{D}$. Combining this with the fact that $u=w=0$ on $\partial B_{1}$, we have

$$
\nabla u(x) \cdot \vec{n}(x) \leq \nabla w(x) \cdot \vec{n}(x)=\left.\frac{d}{d r} w(r)\right|_{r=1}=4 \epsilon \quad \text { for all } x \in \partial B_{1}
$$

hence

$$
\begin{equation*}
\int_{\partial B_{1}} \nabla u \cdot \vec{n} d \sigma \leq 8 \pi \epsilon \tag{2.2.37}
\end{equation*}
$$

Plugging (Equation 2.2.36) and (Equation 2.2.37) into (Equation 2.2.34), we obtain

$$
\left.p\right|_{\partial B_{1}} \leq \log (|h(1+\epsilon+a \epsilon)|)(1+4 \epsilon) .
$$

Since $\log s \leq s-1$ for $s>1$, it follows that

$$
\begin{aligned}
\log |h(1+\epsilon+a \epsilon)| & \leq h(1+\epsilon+a \epsilon)-1=\frac{1+\epsilon+a \epsilon+b}{1+b(1+\epsilon+a \epsilon)}-1 \\
& =\epsilon\left(1+\frac{a-2 b-a b-b \epsilon-a b \epsilon}{1+b(1+\epsilon+a \epsilon)}\right) \\
& \leq \epsilon\left(1-\frac{a b}{4}\right) \leq \epsilon\left(1-\frac{a^{2}}{8}\right)
\end{aligned}
$$

where we used $b>\frac{a}{2}$ to obtain the last two inequalities. Finally, using that $\epsilon<\frac{a^{2}}{64}$, we have

$$
\left.p\right|_{\partial B_{1}} \leq \epsilon\left(1-\frac{a^{2}}{8}\right)\left(1+\frac{a^{2}}{16}\right) \leq \epsilon\left(1-\frac{1}{16} a^{2}\right)<\frac{|D|}{2 \pi}\left(1-\frac{1}{16} a^{2}\right)
$$

where in the last step we use that $|D|=\pi(1+\epsilon)^{2}-\pi>2 \pi \epsilon$. This finishes the proof of the
lemma.

Finally we give the proof of Lemma 2.2.4.

Proof of Lemma 2.2.4. Without loss of generality, we can assume that $o_{2}$ is the origin. Let $\beta:=$ $\left.p\right|_{\partial B_{r}}$. From the proof of Proposition 2.1.6, we already know that $g^{\prime}(k) \leq-2 \pi$, where $g(k):=$ $|\{x \in D: p(x)>k\}|$. This implies that $g(k) \geq-2 \pi(k-\beta)$. Therefore we have

$$
\int_{D} p d x=\int_{0}^{\beta} g(k) d k \geq \int_{0}^{\beta}-2 \pi(k-\beta) d k=\pi \beta^{2}
$$

On the other hand, the same computation in the proof of Lemma 2.1.11 gives

$$
\beta\left|B_{r}\right|+\int_{D} p d x=\frac{1}{2} \int_{D}|\nabla p|^{2} d x \leq \frac{1}{2} \int_{D}|x|^{2} d x .
$$

Since

$$
\begin{aligned}
\frac{1}{2} \int_{D}|x|^{2} d x & =\frac{1}{2}\left(\int_{B_{R}}|x|^{2} d x-\int_{B_{r}}|x|^{2} d x\right) \\
& =\frac{|D|^{2}}{4 \pi}+\frac{|D|\left|B_{r}\right|}{2 \pi}+\frac{\left|B_{r}\right|^{2}}{4 \pi}-\frac{1}{2} \int_{B_{r}}|x|^{2} d x \\
& =\frac{|D|^{2}}{4 \pi}+\frac{|D|\left|B_{r}\right|}{2 \pi}-\frac{l^{2}\left|B_{r}\right|}{2}
\end{aligned}
$$

it follows that

$$
\begin{equation*}
\pi \beta^{2}+\beta\left|B_{r}\right| \leq \frac{|D|^{2}}{4 \pi}+\frac{|D|\left|B_{r}\right|}{2 \pi}-\frac{l^{2}\left|B_{r}\right|}{2} . \tag{2.2.38}
\end{equation*}
$$

By solving the quadratic inequality (Equation 2.2.38), we find that

$$
\beta \leq \frac{|D|}{2 \pi}\left(1-c_{0}\right),
$$

for some constant $c_{0}$ which only depends on $\delta_{1}, \delta_{2}$ and $l$.

## CHAPTER 3

## RIGIDITY RESULTS FOR GENERALIZED SQG

### 3.1 Radial symmetry for stationary/rotating gSQG solutions with $\Omega \leq 0$

In this section, we consider the family of gSQG equations with $0<\alpha<2$, and study the symmetry property for rotating patch/smooth solutions with angular velocity $\Omega \leq 0$.

Let us deal with patch solutions first. As we have discussed in the introduction, we cannot expect a non-simply-connected patch $D$ with $\Omega \leq 0$ to be radial, due to the non-radial examples in $[68,49]$ for $\alpha \in(0,2)$. For a simply-connected patch $D$, the constant on the right hand side of (Equation 1.1.6) is the same on $\partial D$, which motivates us to consider Question 2 in the introduction. The goal of this section is to prove Theorem C, which gives an affirmative answer to Question 2 for the whole range $\alpha \in[0,2)$.

Our results are not limited to the Riesz potentials $K_{\alpha, d}$ in (Equation 1.1.4); in fact, we only need the potential being radially increasing and not too singular at the origin. Below we state our assumption on the potential $K$, which covers the whole range of $K_{\alpha, d}$ with $\alpha \in[0,2)$.
(HK) Let $K \in C^{1}\left(\mathbb{R}^{d} \backslash\{0\}\right)$ be radially symmetric with $K^{\prime}(r)>0$ for all $r>0$. (Here we denote $K(x)=K(r)$ by a slight abuse of notation.) Also assume there is some $\delta>0$ such that $K^{\prime}(r) \leq r^{-d-1+\delta}$ for all $0<r \leq 1$.

Our proof is done by a variational approach, which relies on a continuous Steiner symmetrization argument in a similar spirit as [18].

### 3.1.1 Definition and properties of continuous Steiner symmetrization

Below we define the continuous Steiner symmetrization for a bounded open set $D \subset \mathbb{R}^{d}$ with respect to the direction $e_{1}=(1,0, \ldots, 0)$, which can be easily adapted to any other direction in $\mathbb{R}^{d}$.

The definition is the same as [18, Section 2.2.1], which we briefly outline below for completeness.
For a one-dimensional open set $U \subset \mathbb{R}$, we define its continuous Steiner symmetrization $M^{\tau}[U]$ as follows. If $U=(a, b)$ is an open interval, then $M^{\tau}[U]$ shifts the midpoint of this interval towards the origin with velocity 1 , while preserving the length of interval. That is,

$$
M^{\tau}[U]:= \begin{cases}\left(a-\tau \operatorname{sgn}\left(\frac{a+b}{2}\right), b-\tau \operatorname{sgn}\left(\frac{a+b}{2}\right)\right) & \text { for } 0 \leq \tau<\frac{|a+b|}{2} \\ \left(-\frac{b-a}{2}, \frac{b-a}{2}\right) & \text { for } \tau \geq \frac{|a+b|}{2}\end{cases}
$$

If $U=\cup_{i=1}^{N} U_{i}$ is a finite union of open intervals, then $M^{\tau}[U]$ is defined by $\cup_{i=1}^{N} M^{\tau}\left[U_{i}\right]$, and as soon as two intervals touch each other, we merge them into one interval as in [18, Definition 2.10(2)]. Finally, if $U=\cup_{i=1}^{\infty} U_{i}$ is a countable union of open intervals, we define $M^{\tau}[U]$ as a limit of $M^{\tau}\left[\cup_{i=1}^{N} U_{i}\right]$ as $N \rightarrow \infty$ as in [18, Definition 2.10(3)]. See [18, Figure 1] for an illustration of $M^{\tau}[U]$.

Next we move on to higher dimensions. We denote a point $x \in \mathbb{R}^{d}$ by $\left(x_{1}, x^{\prime}\right)$, where $x^{\prime}=$ $\left(x_{2}, \ldots, x_{d}\right) \in \mathbb{R}^{d-1}$. For a bounded domain $D \subset \mathbb{R}^{d}$ and any $x^{\prime} \in \mathbb{R}^{d-1}$, we define the section of $D$ with respect to the direction $x_{1}$ as

$$
D_{x^{\prime}}:=\left\{x_{1} \in \mathbb{R}:\left(x_{1}, x^{\prime}\right) \in D\right\},
$$

which is an open set in $\mathbb{R}$. If the section $D_{x^{\prime}}$ is a single open interval centered at 0 for all $x^{\prime} \in \mathbb{R}^{d-1}$, then we say the set $D$ is Steiner symmetric about the hyperplane $\left\{x_{1}=0\right\}$. Note that this definition is stronger than being symmetric about $\left\{x_{1}=0\right\}$. For example, an annulus in $\mathbb{R}^{2}$ is symmetric about $\left\{x_{1}=0\right\}$, but not Steiner symmetric about it.

Finally, for any $\tau>0$, the continuous Steiner symmetrization of $D \subset \mathbb{R}^{d}$ is defined as

$$
S^{\tau}[D]:=\left\{\left(x_{1}, x^{\prime}\right) \in \mathbb{R}^{d}: x_{1} \in M^{\tau}\left[D_{x^{\prime}}\right]\right\},
$$

with $M^{\tau}$ given above being the continuous Steiner symmetrization for one-dimensional open sets. See Figure Figure 3.1 for a comparison of the sets $D$ and $S^{\tau}[D]$ for small $\tau>0$.


Figure 3.1: Illustration of the continuous Steiner symmetrization $S^{\tau}[D]$ for a set $D \subset \mathbb{R}^{2}$. The left figure is the set $D$, with the midpoints of all subintervals of its 1 D section highlighted in red circles. The right figure shows the set $S^{\tau}[D]$ for some small $\tau>0$, with the new midpoints denoted by blue squares.

One can easily check that $S^{\tau}[D]$ satisfies the following properties.

Lemma 3.1.1. For any bounded open set $D \subset \mathbb{R}^{d}$, its continuous Steiner symmetrization $S^{\tau}[D]$ satisfies the following properties:
(a) $\left|S^{\tau}[D]\right|=|D|$ for any $\tau>0$, where $|\cdot|$ denotes the Lebesgue measure in $\mathbb{R}^{d}$.
(b) $\left(S^{\tau}[D]\right) \triangle D \subset B^{\tau}[D]$ for any $\tau>0$, where $\triangle$ is the symmetric difference between the two sets, and $B^{\tau}[D]$ is the $\tau$-neighborhood of $\partial D$, given by

$$
\begin{equation*}
B^{\tau}[D]:=\left\{x \in \mathbb{R}^{d}: \operatorname{dist}(x, \partial D) \leq \tau\right\} . \tag{3.1.1}
\end{equation*}
$$

Proof. (a) is a direct consequence of the fact that $\left|M^{\tau}[U]\right|=|U|$ for any open set $U \subset \mathbb{R}$ and $\tau>0$ [18, Lemma 2.11(b)]. To prove (b), one can start with the one-dimensional version: For any bounded open set $U \subset \mathbb{R}$, we have $M^{\tau}[U] \triangle U \subset\{x \in \mathbb{R}: \operatorname{dist}(x, \partial U) \leq \tau\}$, which follows from
the fact that the intervals move with velocity at most 1 . Thus for any bounded open set $D \subset \mathbb{R}^{d}$,

$$
\begin{aligned}
S^{\tau}[D] \triangle D & =\left\{\left(x_{1}, x^{\prime}\right) \in \mathbb{R}^{d}: x_{1} \in M^{\tau}\left[D_{x^{\prime}}\right] \triangle D_{x^{\prime}}\right\} \\
& \subset\left\{\left(x_{1}, x^{\prime}\right): \operatorname{dist}\left(x_{1}, \partial\left(D_{x^{\prime}}\right)\right) \leq \tau\right\} \\
& \subset B^{\tau}[D],
\end{aligned}
$$

finishing the proof.

### 3.1.2 Simply-connected patch solutions with $\Omega \leq 0$

We assume that $D \subset \mathbb{R}^{d}$ satisfies the following condition.
(HD) $D \subset \mathbb{R}^{d}$ is a bounded domain, and there exists some $M>0$ depending on $D$, such that $\left|B^{\tau}[D]\right| \leq M \tau$ for all sufficiently small $\tau>0$, where $B^{\tau}[D]$ is given in (item 3.1.1).

It can be easily checked that for $d \geq 2$, any bounded domain $D$ with Lipschitz continuous boundary satisfies condition (HD). In fact, for $d=2$, we will show any domain $D \subset \mathbb{R}^{2}$ with a rectifiable boundary satisfies (HD), with a precise bound

$$
\begin{equation*}
\left|B^{\tau}[D]\right| \leq 2|\partial D| \tau \quad \text { for all } \tau \geq 0 \tag{3.1.2}
\end{equation*}
$$

where $|\partial D|$ is the total length of $\partial D$. Let us first prove (Equation 3.1.2) holds for any polygon $P \subset \mathbb{R}^{2}$. Erect two polygons at distance $\tau$ from $P$ and the transversal sides being bisectors of the inner angles of $P$ (see Figure Figure 3.2). It is clear that $B^{\tau}[P]$ is contained in the trapezoidal region, which has area no more than $2|\partial P| \tau$. Finally, this can be extended to the general case by approximating any rectifiable curve by polygons.

Below we state our main theorem of this section, which is slightly more general than Theorem Theorem C.

Theorem 3.1.2. Let $D \subset \mathbb{R}^{d}$ and $K \in C^{1}\left(\mathbb{R}^{d} \backslash\{0\}\right)$ satisfy the conditions (HD) and (HK)


Figure 3.2: Illustration of the polygon $P$ and the underlying trapezoidal region (the whole colored region). Here the blue trapezoid has area $2 l_{1} \tau$ ( $l_{1}$ is the corresponding side length in $P$ ), and summing over all edges gives a total area $2|\partial P| \tau$. Since the trapezoids may intersect for large $\tau$, the whole trapezoidal region has area no more than $2|\partial P| \tau$.
respectively. Let $g \in C^{1}\left(\mathbb{R}^{d}\right)$ be a radial function with $g^{\prime}(r)>0$ for all $r>0$.
If $D$ satisfies that

$$
\begin{equation*}
1_{D} * K-\frac{\Omega}{2} g(x)=\text { const } \quad \text { on } \partial D \tag{3.1.3}
\end{equation*}
$$

for some $\Omega \leq 0$ (where the constant is the same on all connected components of $\partial D$ ), then $D$ is a ball. Moreover, the ball is centered at the origin if $\Omega<0$.

Remark 3.1.3. (1) Note that $D$ does not need to be simply-connected in Theorem 3.1.2. However, since the constant on the right hand side of (Equation 3.1.3) is assumed to be the same on all connected components of $\partial D$, comparing with (Equation 1.1.6), Theorem 3.1.2 only implies that all simply-connected patches with $\Omega \leq 0$ must be a disk.
(2) In the case $\Omega=0$, the problem is translation invariant, so in the proof we assume without loss of generality that the center of mass of $D$ is at the origin.

Proof. We prove it by contradiction. Without loss of generality, we assume $D$ is not Steiner symmetric about the hyperplane $\left\{x_{1}=0\right\}$. Let $D^{\tau}:=S^{\tau}[D]$ be the continuous Steiner symmetrization of $D$ at time $\tau>0$. By Lemma 3.1.1(b), we have

$$
\begin{equation*}
D^{\tau} \triangle D \subset B^{\tau}[D] \tag{3.1.4}
\end{equation*}
$$

where $B^{\tau}$ is defined in (item 3.1.1). Let us consider the functional

$$
\mathcal{E}[D]:=\underbrace{\int_{\mathbb{R}^{d}} \int_{\mathbb{R}^{d}} 1_{D}(x) 1_{D}(y) K(x-y) d x d y}_{=: \mathcal{I}[D]}+\underbrace{(-\Omega) \int_{\mathbb{R}^{d}} g(x) 1_{D}(x) d x}_{=: \mathcal{V}[D]} .
$$

We will use two different ways to compute $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}$, where $\frac{d^{+}}{d \tau}$ denotes the right derivative. On the one hand, using the equation (Equation 3.1.3) and the regularity assumptions on $D, K$ and $g$, we aim to show that

$$
\begin{equation*}
\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}=0 \tag{3.1.5}
\end{equation*}
$$

Instead of directly taking the derivative, we consider the finite difference

$$
\mathcal{E}\left[D^{\tau}\right]-\mathcal{E}[D]=\underbrace{\int_{\mathbb{R}^{d}} 2\left(1_{D^{\tau}}-1_{D}\right)\left(1_{D} * K-\frac{\Omega}{2} g(x)\right) d x}_{=: I_{1}}+\underbrace{\int_{\mathbb{R}^{d}}\left(1_{D^{\tau}}-1_{D}\right)\left(\left(1_{D^{\tau}}-1_{D}\right) * K\right) d x}_{=: I_{2}},
$$

where in the equality we used that $\int 1_{D}\left(1_{D^{\tau}} * K\right) d x=\int 1_{D^{\tau}}\left(1_{D} * K\right) d x$ for any radial kernel $K$.
Let us control the term $I_{1}$ first. First note that (Equation 3.1.4) implies that the integrand is supported in $B_{\tau}[D]$. Next we claim that $(\mathbf{H K})$ implies $1_{D} * K-\frac{\Omega}{2} g \in C^{0, \delta^{\prime}}\left(\mathbb{R}^{d}\right)$ for $\delta^{\prime}:=\min \{\delta, 1\}$, where $C^{0,1}$ stands for Lipschitz continuity. The proof is a simple potential theory estimate, which we provide below for completeness. For any $x, z \in \mathbb{R}^{d}$,

$$
\begin{aligned}
\mid\left(1_{D} * K\right)(x+z) & -\left(1_{D} * K\right)(x)\left|=\left|\int_{\mathbb{R}^{d}} 1_{D}(x-y)(K(y+z)-K(y)) d y\right|\right. \\
& \leq \int_{|y|<2|z|}|K(y+z)-K(y)| d y+\int_{|y|>2|z|} 1_{D}(x-y)|K(y+z)-K(y)| d y \\
& =: J_{1}+J_{2}
\end{aligned}
$$

For $J_{1}$, a crude estimate gives

$$
J_{1} \leq \int_{|y|<2|z|}|K(y+z)|+|K(y)| d y \leq 2 \int_{|y|<3|z|}|K(y)| d y \leq C(d)|z|^{\delta}
$$

where in the last step we used that (HK) implies $|K(y)| \leq C|y|^{-d+\delta}$ for $|y| \leq 1$. For $J_{2}$, note that (HK) and the mean-value theorem gives

$$
|K(y+z)-K(y)| \leq C|y|^{-d-1+\delta}|z| \quad \text { for all }|y|>2|z|
$$

and plugging it into the integral gives $J_{2} \leq C(d,|D|)|z|^{\delta}$. Putting the estimates for $J_{1}$ and $J_{2}$ together gives that $1_{D} * K \in C^{0, \delta^{\prime}}\left(\mathbb{R}^{d}\right)$ for $\delta^{\prime}=\min \{\delta, 1\}$, and combining this with the assumption $g \in C^{1}\left(\mathbb{R}^{d}\right)$ gives $1_{D} * K-\frac{\Omega}{2} g \in C^{0, \delta^{\prime}}\left(\mathbb{R}^{d}\right)$.

In addition, by (Equation 3.1.3), We have $1_{D} * K-\frac{\Omega}{2} g(x) \equiv C_{0}$ on $\partial D$ for some constant $C_{0}$. Thus we have

$$
\left|1_{D} * K-\frac{\Omega}{2} g(x)-C_{0}\right| \leq C\left(\delta^{\prime}, d,|D|\right) \tau^{\delta^{\prime}} \quad \text { in } B^{\tau}[D]
$$

for some constant $C>0$, where we used the Hölder continuity of $1_{D} * K-\frac{\Omega}{2} g$ and the definition of $B^{\tau}[D]$. This leads to

$$
\left|I_{1}\right| \leq 2\left|B^{\tau}[D]\right| \sup _{x \in B^{\top}[D]}\left|1_{D} * K-\frac{\Omega}{2} g(x)-C_{0}\right| \leq C\left(\delta^{\prime}, d,|D|\right) M \tau^{1+\delta^{\prime}}
$$

where in the first inequality we used that $\int_{B_{\tau}}\left(1_{D}-1_{D^{\tau}}\right) C_{0} d x=0$, which follows from Lemma 3.1.1(a); and in the second inequality we used (HD).

Next we control $I_{2}$ by the crude bound

$$
\left|I_{2}\right| \leq \int_{\mathbb{R}^{d}} 1_{B^{\tau}[D]}\left|\left(1_{B^{\tau}[D]} * K\right)\right| d x \leq\left|B^{\tau}[D]\right|\left\|1_{B^{\tau}[D]} * K\right\|_{\infty} \leq M \tau\left\|\left(1_{B^{\tau}[D]}\right)^{*} * K\right\|_{\infty}
$$

where the last step follows from the Hardy-Littlewood inequality, where $\left(1_{B^{\tau}[D]}\right)^{*}$ is the radial decreasing rearrangement of $1_{B^{\tau}[D]}$. $\mathrm{By}(\mathbf{H D}),\left(1_{B^{\tau}[D]}\right)^{*}$ is a characteristic function of a ball whose
radius is bounded by $C(d)(M \tau)^{1 / d}$, thus

$$
\left\|\left(1_{B^{\tau}[D]}\right)^{*} * K\right\|_{\infty} \leq \int_{0}^{C(d)(M \tau)^{1 / d}}|K(r)| \omega_{d} r^{d-1} d r \leq \int_{0}^{C(d)(M \tau)^{1 / d}} \omega_{d} r^{-1+\delta} d x \leq C(d)(M \tau)^{\frac{\delta}{d}}
$$

and plugging it into the $I_{2}$ estimate gives

$$
\left|I_{2}\right| \leq C(d) M^{\frac{d+\delta}{d}} \tau^{1+\frac{\delta}{d}}
$$

Putting the estimates of $I_{1}$ and $I_{2}$ together directly yields

$$
\frac{\left|\mathcal{E}\left[D^{\tau}\right]-\mathcal{E}[D]\right|}{\tau} \leq C\left(\delta^{\prime}, d, M,|D|\right) \tau^{\min \left\{\frac{\delta}{d}, \delta^{\prime}\right\}}
$$

and since $\delta>0$ we have $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}=0$.
Now, we use another way to calculate $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}$. Let us deal with the $\Omega<0$ case first. Since $K$ is radial and increasing in $r$, it has been shown in [11, Corollary 2] and [78, Theorem 3.7] that the interaction energy $\mathcal{I}\left[D^{\tau}\right]=\int_{D^{\tau}} \int_{D^{\tau}} K(x-y) d x d y$ is non-increasing along the continuous Steiner symmetrization, leading to

$$
\frac{d^{+}}{d \tau} \mathcal{I}\left[D^{\tau}\right] \leq 0 \quad \text { for all } \tau \geq 0
$$

For the other term $\mathcal{V}\left[D^{\tau}\right]=(-\Omega) \int_{D^{\tau}} g(x) d x$, by the assumptions that $g^{\prime}(r)>0$ for all $r>0$ and $D$ is not Steiner symmetric about $\left\{x_{1}=0\right\}$, we can use [18, Lemma 2.22] to show, for $\Omega<0$,

$$
\left.\frac{d^{+}}{d \tau} \mathcal{V}\left[D^{\tau}\right]\right|_{\tau=0}=\left.(-\Omega) \frac{d^{+}}{d \tau} \int_{D^{\tau}} g(x) d x\right|_{\tau=0}<0
$$

Adding them together gives

$$
\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}<0
$$

leading to a contradiction with (Equation 3.1.5).

In the $\Omega=0$ case, recall that we assume that the center of mass of $D$ is at the origin. Thus if $D$ is not Steiner symmetric about $\left\{x_{1}=0\right\}$, the same proof as [18, Proposition 2.15] gives that $\mathcal{I}[D]$ must be decreasing to the first order for a short time, leading to

$$
\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}=\left.\frac{d^{+}}{d \tau} \int_{D^{\tau}} \int_{D^{\tau}} K(|x-y|) d x d y\right|_{\tau=0}<0
$$

again contradicting (Equation 3.1.5). We point out that although the proposition was stated for continuous densities, the same proof works for the patch setting. In addition, although [18] only dealt with the kernels no more singular than Newtonian potential, the proof indeed holds for all kernels $K$ satisfying (HK): see [17, Theorem 6] for an extension to all Riesz potentials $K_{\alpha, d}$ with $\alpha \in(0,2)$.

The above theorem immediately leads to the following result concerning simply-connected stationary/rotating patch solution with $\Omega \leq 0$.

Theorem 3.1.4. Let $D \subset \mathbb{R}^{2}$ be a bounded, simply-connected domain with rectifiable boundary. If $1_{D}$ is a $V$-state for (Equation 1.1.2) for some $\alpha \in[0,2)$ with angular velocity $\Omega \leq 0$, then $D$ must be a disk. In addition, the disk must be centered at the origin if $\Omega<0$.

Proof. We have $1_{D} * K-\frac{\Omega}{2}|x|^{2}=C$ for some constant $C$ on $\partial D$. For the Euler equation, $K=\frac{1}{2 \pi} \ln |x|$. For the g -SQG equation. $K=-C_{\alpha}|x|^{-\alpha}$. In both cases, the proof follows from Theorem 3.1.2.

Remark 3.1.5. As we discussed in the beginning of this subsection, in the case of $g S Q G$ with $\alpha \in(0,2)$, Theorem 3.1.4 is not true if the simply connected assumption is dropped, due to the non-radial patches in [68, 49] bifurcating from annuli.
3.1.3 Smooth solutions with simply-connected level sets with $\Omega \leqq 0$

The rest of this section is devoted to the smooth setting. We will show that any nonnegative smooth rotating solution of the Euler or gSQG equation with angular velocity $\Omega \leq 0$ must be radial, under
the additional assumption that all the super level-sets $U^{h}$

$$
\begin{equation*}
U^{h}:=\left\{x \in \mathbb{R}^{d}: \omega(x)>h\right\} \tag{3.1.6}
\end{equation*}
$$

are simply-connected for any $h>0$. We believe that the simply-connected assumption is necessary, since it is likely that the bifurcation arguments from annuli in $[68,49]$ can be extended to the smooth setting as well, using a similar argument as in [24] or [23].

Theorem 3.1.6. Let $\omega(x) \in C^{1}\left(\mathbb{R}^{2}\right)$ be nonnegative and compactly supported. In addition, assume the super level-set $U^{h}$ as in (Equation 3.1.6) is simply connected for all $h \in(0, \sup \omega)$. Assume $K$ satisfies (HK). If for some $\Omega \leq 0$, we have

$$
\begin{equation*}
\omega * K-\frac{\Omega}{2}|x|^{2}=C_{0}(h) \quad \text { on } \partial U^{h} \text { for all } h \in(0, \sup \omega) \tag{3.1.7}
\end{equation*}
$$

then $\omega$ is radially decreasing up to a translation. Moreover, it is centered at the origin if $\Omega<0$.

Proof. We prove it by contradiction. For the $\Omega<0$ case, without loss of generality, we assume $\omega$ is not symmetric decreasing about the line $x_{1}=0$. For the $\Omega=0$ case, similar to Remark 3.1.3, without loss of generality we assume the center of mass is at the origin, and then we assume $\omega$ is not symmetric decreasing about the line $x_{1}=0$.

For any $\tau \geq 0$, we define the continuous Steiner symmetrization $\omega^{\tau}(x)$ in the same way as [18, Definition 2.12]:

$$
\omega^{\tau}(x):=\int_{0}^{h_{0}} 1_{S^{\tau}\left[U^{h}\right]}(x) d h,
$$

where $h_{0}:=\sup \omega$, and $S^{\tau}\left[U^{h}\right]$ is the continuous Steiner symmetrization of the super level set $U^{h}$ at time $\tau \geq 0$.

Consider the energy functional

$$
\mathcal{E}[\omega]:=\underbrace{\int_{\mathbb{R}^{2}} \int_{\mathbb{R}^{2}} \omega(x) \omega(y) K(x-y) d x d y}_{=: \mathcal{I}[\omega]}+\underbrace{(-\Omega) \int_{\mathbb{R}^{2}} \omega(x)|x|^{2} d x}_{=: \mathcal{V}[\omega]}
$$

We proceed similarly as in Theorem 3.1.2 to compute $\frac{d^{+}}{d \tau} \mathcal{E}\left[\omega^{\tau}\right]$ in two different ways. We first rewrite the finite difference $\mathcal{E}\left[\omega^{\tau}\right]-\mathcal{E}[\omega]$ as

$$
\begin{align*}
& \mathcal{E}\left[\omega^{\tau}\right]-\mathcal{E}[\omega] \\
= & \int_{\mathbb{R}^{2}} 2\left(\omega^{\tau}(x)-\omega(x)\right)\left(\omega * K-\frac{\Omega}{2}|x|^{2}\right) d x+\iint_{\mathbb{R}^{2} \times \mathbb{R}^{2}}\left(\omega^{\tau}(x)-\omega(x)\right)\left(\omega^{\tau}(y)-\omega(y)\right) K(x-y) d x d y \\
= & I_{1}+I_{2} \tag{3.1.8}
\end{align*}
$$

Since $\omega \in C_{c}^{1}\left(\mathbb{R}^{2}\right)$ and $K$ satisfies (HK) (hence is locally integrable), one can easily check that $\omega * K-\frac{\Omega}{2}|x|^{2}$ is Lipschitz in $\tilde{D}:=\left\{x \in \mathbb{R}^{2}: \operatorname{dist}(x, \operatorname{supp} \omega) \leq 1\right\}$. Note that we have supp $\omega^{\tau} \in \tilde{D}$ for all $\tau \in[0,1]$. Combining this fact with the assumption (Equation 3.1.7), there exists $C_{1}>0$ independent of $h$, such that

$$
\begin{equation*}
\left.\left.\left|(\omega * K)(x)-\frac{\Omega}{2}\right| x\right|^{2}-C_{0}(h) \right\rvert\, \leq C_{1} \tau \quad \text { on } S^{\tau}\left[U^{h}\right] \triangle U^{h} \quad \text { for all } h \in\left(0, h_{0}\right) \tag{3.1.9}
\end{equation*}
$$

Let us first rewrite $I_{1}$ as

$$
I_{1}=2 \int_{0}^{h_{0}} \int_{\mathbb{R}^{2}}\left(1_{S^{\tau}\left[U^{h}\right]}(x)-1_{U^{h}}(x)\right)\left((\omega * K)(x)-\frac{\Omega}{2}|x|^{2}\right) d x d h
$$

By Lemma 3.1.1(a), we have $\int_{\mathbb{R}^{2}}\left(1_{S^{\tau}\left[U^{h}\right]}(x)-1_{U^{h}}(x)\right) d x=0$ for all $h \in\left(0, h_{0}\right)$. Thus we can
control $I_{1}$ as

$$
\begin{align*}
\left|I_{1}\right| & =\left|2 \int_{0}^{h_{0}} \int_{\mathbb{R}^{2}}\left(1_{S^{\tau}\left[U^{h}\right]}(x)-1_{U^{h}}(x)\right)\left((\omega * K)(x)-\frac{\Omega}{2}|x|^{2}-C_{0}(h)\right) d x d h\right| \\
& \leq 2 C_{1} \tau \int_{0}^{h_{0}}\left|\left(S^{\tau}\left[U^{h}\right]\right) \triangle U^{h}\right| d h  \tag{3.1.10}\\
& \leq 2 C_{1} \tau \int_{0}^{h_{0}} 2\left|\partial U^{h}\right| \tau d h \\
& =4 C_{1} \tau^{2} \int_{\text {supp } \omega}|\nabla \omega| d x \leq C(\omega) \tau^{2} .
\end{align*}
$$

Here in the second line we used (Equation 3.1.9); in the third line we used Lemma 3.1.1(b) and the property (Equation 3.1.2) in two dimensions; and in the fourth line we used the co-area formula and the fact that $\omega \in C_{c}^{1}$.

We next move on to $I_{2}$. Since $|\nabla \omega|$ is bounded, Lemma 3.1.1(b) leads to

$$
\begin{aligned}
\left|\omega^{\tau}(x)-\omega(x)\right| & =\left|\int_{0}^{\infty} 1_{S^{\tau}\left[U^{h}\right]}(x)-1_{U^{h}}(x) d h\right| \\
& \leq\|\nabla \omega\|_{L^{\infty} \tau} \quad \text { for all } x \in \mathbb{R}^{2}
\end{aligned}
$$

and $\operatorname{supp} \omega^{\tau} \in \tilde{D}$ for all $\tau \in[0,1]$. Thus

$$
\begin{aligned}
\left|I_{2}\right| & \leq\left\|\omega^{\tau}-\omega\right\|_{L^{1}}\left\|\left(\omega^{\tau}-\omega\right) * K\right\|_{L^{\infty}} \\
& \leq\left\|\omega^{\tau}-\omega\right\|_{L^{1}}\left\|\omega^{\tau}-\omega\right\|_{L^{\infty}} \int_{\tilde{D}}|K(x)| d x \\
& \leq C(\omega) \tau^{2}
\end{aligned}
$$

Combining the estimates for $I_{1}$ and $I_{2}$ gives $\mathcal{E}\left[\omega^{\tau}\right]-\mathcal{E}[\omega] \leq C(\omega) \tau^{2}$ for all $\tau \in[0,1]$, thus

$$
\begin{equation*}
\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[\omega^{\tau}\right]\right|_{\tau=0}=\left.\frac{d^{+}}{d \tau}\left(I_{1}+I_{2}\right)\right|_{\tau=0}=0 \tag{3.1.11}
\end{equation*}
$$

On the other hand, we compute $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[\omega^{\tau}\right]\right|_{\tau=0}$ in a different way as

$$
\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[\omega^{\tau}\right]\right|_{\tau=0}=\left.\frac{d^{+}}{d \tau}\left(\mathcal{I}\left[\omega^{\tau}\right]+\mathcal{V}\left[\omega^{\tau}\right]\right)\right|_{\tau=0}
$$

In the $\Omega<0$ case, similarly as in Theorem 3.1.2, we have $\mathcal{I}\left[\omega^{\tau}\right]$ is non-increasing along the continuous Steiner symmetrization by [11, Corollary 2] and [78, Theorem 3.7], thus

$$
\frac{d^{+}}{d \tau} \mathcal{I}\left[\omega^{\tau}\right] \leq 0 \quad \text { for all } \tau>0
$$

For $\mathcal{V}\left[\omega^{\tau}\right]$, by the assumption that $\omega$ is not symmetric decreasing about $\left\{x_{1}=0\right\}$, we again use [18, Lemma 2.22] to show, for $\Omega<0$,

$$
\frac{d^{+}}{d \tau} \mathcal{V}\left[\omega^{\tau}\right]=\left.(-\Omega) \frac{d^{+}}{d \tau} \int_{\mathbb{R}^{2}} \omega(x)|x|^{2} d x\right|_{\tau=0}<0
$$

Adding them together gives $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}<0$, contradicting (Equation 3.1.11).
In the $\Omega=0$ case, we assume that the center of mass of $\omega$ is at the origin. Thus if $\omega$ is not symmetric decreasing about $\left\{x_{1}=0\right\}$, the same proof as [18, Proposition 2.15] gives that $\mathcal{I}[D]$ must be decreasing to the first order for a short time (again, the proof holds for all kernels $K$ satisfying (HK); see [17, Theorem 6] for extensions to Riesz kernels $K_{\alpha, d}$ with $\alpha \in(0,2)$ ). This gives $\left.\frac{d^{+}}{d \tau} \mathcal{E}\left[D^{\tau}\right]\right|_{\tau=0}<0$, again contradicting (Equation 3.1.11).

The above theorem immediately gives the following corollary concerning the V-states for the Euler and gSQG equations.

Corollary 3.1.7. Assume $\omega(x) \in C^{1}\left(\mathbb{R}^{2}\right)$ is a nonnegative, compactly supported $V$-state satisfying the Euler equation or the gSQG equation for some $\alpha \in(0,2)$ with $\Omega \leq 0$. In addition, assume the super level-set $U^{h}$ as in (Equation 3.1.6) is simply connected for all $h \in(0, \sup \omega)$. Then $\omega$ must be radially decreasing if $\Omega<0$, and radially decreasing up to a translation if $\Omega=0$.

Proof. For the Euler equation, $K=\frac{1}{2 \pi} \ln |x|$. For the gSQG equation. $K=-C_{\alpha}|x|^{-\alpha}$. In both cases, the proof follows from Theorem 3.1.6.

### 3.2 Radial symmetry of rotating gSQG solutions with $\Omega>\Omega_{\alpha}$

In this section, we focus on rotating gSQG patches with area $\pi$ and $\alpha \neq 0$. As we discussed in the introduction, for $\alpha \in[0,2)$, there exist rotating patches bifurcating from the unit disk at angular velocities $\Omega_{m}^{\alpha}=2^{\alpha-1} \frac{\Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}}\left(\frac{\Gamma\left(1+\frac{\alpha}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)}-\frac{\Gamma\left(m+\frac{\alpha}{2}\right)}{\Gamma\left(m+1-\frac{\alpha}{2}\right)}\right)$, where $\Omega_{m}^{\alpha}$ is increasing in $m$ for any fixed $\alpha \in[0,2)$. Let us denote $\Omega_{\alpha}:=\lim _{m \rightarrow \infty} \Omega_{m}^{\alpha}$. If $\alpha \in(0,1)$ we have that

$$
\begin{equation*}
\Omega_{\alpha}=2^{\alpha-1} \frac{\Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}} \frac{\Gamma\left(1+\frac{\alpha}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)} \tag{3.2.1}
\end{equation*}
$$

Note that $\Omega_{\alpha}$ is a continuous function of $\alpha$ for $\alpha \in(0,1)$, with $\Omega_{0}=\frac{1}{2}$, and $\Omega_{\alpha}=+\infty$ for all $\alpha \in[1,2)$.

A natural question is whether there can be rotating patches with area $\pi$ with $\Omega \geq \Omega_{\alpha}$ for $\alpha \in(0,1)$. Note that the area constraint is necessary for all $\alpha>0$ : if $D$ is a rotating gSQG patch for $\alpha \in(0,2)$ with angular velocity $\Omega$, then one can easily check that its scaling $\lambda D=\{\lambda x: x \in D\}$ is a rotating patch with angular velocity $\lambda^{-\alpha} \Omega$.

In Theorem 2.1.12, we showed that for the 2D Euler case ( $\alpha=0$ ), all rotating patches with $\Omega \geq \Omega_{0}=\frac{1}{2}$ must be a disk. In this section, our goal is to show that all simply-connected rotating patches with area $\pi$ with $\Omega \geq \Omega_{\alpha}$ for $\alpha \in(0,1)$ must be a disk. Whether there exist non-simplyconnected or disconnected rotating patches with $\Omega \geq \Omega_{\alpha}$ for $\alpha \in(0,1)$ is still an open question.

Below is the main theorem of this section. Recall that for $\alpha \in(0,2), K_{\alpha}=-C_{\alpha}|x|^{-\alpha}$ is the fundamental solution for $-(-\Delta)^{-1+\frac{\alpha}{2}}$, where $C_{\alpha}=\frac{1}{2 \pi} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{2^{1-\alpha} \Gamma\left(1-\frac{\alpha}{2}\right)}$.

Theorem 3.2.1. Let $D \subset \mathbb{R}^{2}$ be a bounded, simply-connected patch with $C^{1}$ boundary. Let us denote $R:=\max _{x \in D}|x|$. Assume that $D$ is a uniformly rotating patch with angular velocity $\Omega$ of
the gSQG equation with $\alpha \in(0,1)$, i.e.,

$$
\begin{equation*}
1_{D} * K_{\alpha}-\frac{\Omega}{2}|x|^{2}=C \quad \text { on } \partial D \tag{3.2.2}
\end{equation*}
$$

Let $\Omega_{c}(R):=R^{-\alpha} \Omega_{\alpha}$. If $\Omega \geq \Omega_{c}(R)$, then $D$ must coincide with $B(0, R)$.

Remark 3.2.2. (a) Note that all sets $D \subset \mathbb{R}^{2}$ with area $\pi$ must have $R \geq 1$. In this case we have $\Omega_{c}(R) \leq \Omega_{\alpha}$, thus Theorem 3.2.1 immediately implies that all simply-connected rotating patches with area $\pi$ and $\Omega \geq \Omega_{\alpha}$ must be a disk.
(b) Note that the constant $\Omega_{\alpha}$ is sharp, since there exist patches bifurcating from a disk of radius 1 at velocities $\Omega_{m}^{\alpha}$, which can get arbitrarily close to $\Omega_{\alpha}$ as $m \rightarrow \infty$ [57, Theorem 1.4].

Proof. Towards a contradiction, assume that $D \neq B(0, R)$. Let $x_{0} \in \partial D$ be the farthest point from 0 . Then we have that $D \subset B(0, R)$, and let $U:=B(0, R) \backslash D$. See Figure Figure 3.3 for an illustration of $U$ and $x_{0}$. Then (Equation 3.2.2) can be rewritten as

$$
\begin{equation*}
1_{U} * K_{\alpha}=1_{B(0, R)} * K_{\alpha}-\frac{\Omega}{2}|x|^{2}-C \quad \text { on } \partial D \tag{3.2.3}
\end{equation*}
$$



Figure 3.3: Illustration of the set $U$ and the point $x_{0}$.

The key idea of this proof is to use two different ways to compute $\nabla\left(1_{U} * K_{\alpha}\right)\left(x_{0}\right) \cdot x_{0}$, and
obtain a contradiction if $\Omega \geq \Omega_{c}(R)$. On the one hand,

$$
\begin{equation*}
\nabla\left(1_{U} * K_{\alpha}\right)\left(x_{0}\right) \cdot x_{0}=\alpha C_{\alpha} \int_{U} \frac{\left(x_{0}-y\right) \cdot x_{0}}{\left|x_{0}-y\right|^{\alpha+2}} d y>0 \tag{3.2.4}
\end{equation*}
$$

where we used the fact that $\left(x_{0}-y\right) \cdot x_{0}>0$ for all $y \in U \subset B(0, R)$ since the two vectors point to the same halfplane.

On the other hand, we claim the following properties hold for $1_{U} * K_{\alpha}$ :

1. $\Delta\left(1_{U} * K_{\alpha}\right)<0$ in $D$.
2. Along $\partial D$, the minimum of $1_{U} * K_{\alpha}$ is achieved at $x_{0}$.

To show property 1 , using the fact that $K_{\alpha}=-C_{\alpha}|x|^{-\alpha}$ is the fundamental solution for $-(-\Delta)^{-1+\frac{\alpha}{2}}$, we have $1_{U} * K_{\alpha}=-(-\Delta)^{-1+\frac{\alpha}{2}} 1_{U}$, thus $\Delta\left(1_{U} * K_{\alpha}\right)=(-\Delta)^{\alpha / 2} 1_{U}$. Thus for any $x \in D$, using the singular integral definition of the fractional Laplacian [77, Theorem 1.1, Definition (e)] and the fact that $1_{U} \equiv 0$ in $D$, we have

$$
(-\Delta)^{\alpha / 2} 1_{U}(x)=C_{1}(\alpha) \int_{\mathbb{R}^{2}} \frac{1_{U}(x)-1_{U}(y)}{|x-y|^{2+\alpha}} d y=C_{1}(\alpha) \int_{\mathbb{R}^{2}} \frac{0-1_{U}(y)}{|x-y|^{2+\alpha}} d y<0 \quad \text { for } x \in D
$$

for some constant $C_{1}(\alpha)>0$. Note that despite the denominator being singular, the integral indeed converges for all $x \in D$, due to the fact that $D$ is open and the integrand is identically zero in $D$ which yields

$$
\Delta\left(1_{U} * K_{\alpha}\right)(x)=(-\Delta)^{\alpha / 2} 1_{U}(x)<0 \text { in } D .
$$

We now move on to property 2. Due to (Equation 3.2.3) and the fact that $x_{0}$ is the outmost point on $\partial D$, it suffices to show that the radial function $1_{B(0, R)} * K_{\alpha}-\frac{\Omega}{2}|x|^{2}$ is non-increasing in $|x|$ for all $\Omega \geq \Omega_{c}(R)$. We prove this in Proposition 3.2.3 right after this theorem.

The above claims allow us to apply the maximum principle to $1_{U} * K_{\alpha}$, which yields that the
minimum of $1_{U} * K_{\alpha}$ in $\bar{D}$ is also achieved at $x_{0}$, thus

$$
\nabla\left(1_{U} * K_{\alpha}\right)\left(x_{0}\right) \cdot \vec{n}\left(x_{0}\right) \leq 0
$$

where $\vec{n}\left(x_{0}\right)$ is the outer normal of $D$ at $x_{0}$. Since $\vec{n}\left(x_{0}\right)=x_{0} /\left|x_{0}\right|$, the above inequality contradicts with (Equation 3.2.4). As a result, $D$ must coincide with $B(0, R)$.

Now we prove the proposition that was used in the proof of the above theorem.

Proposition 3.2.3. For a fixed $\alpha \in(0,1)$ and $R>0$, let $\Omega_{c}(R)$ be the smallest number such that

$$
g_{R}(x):=1_{B(0, R)} * K_{\alpha}-\frac{\Omega_{c}}{2}|x|^{2}
$$

is non-increasing in $|x|$. Then we have $\Omega_{c}(R)=R^{-\alpha} \Omega_{\alpha}$, with $\Omega_{\alpha}$ given in (Equation 3.2.1).
Proof. Recall that $K_{\alpha}=-C_{\alpha}|x|^{-\alpha}$ with $C_{\alpha}=\frac{1}{2 \pi} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{2^{1-\alpha} \Gamma\left(1-\frac{\alpha}{2}\right)}$. Since $|x|^{2}$ and $1_{B(0, R)} * K_{\alpha}$ are both radially symmetric and increasing in $|x|$, we have

$$
\Omega_{c}(R)=2 C_{\alpha} \sup _{\left|x_{1}\right| \neq\left|x_{2}\right|} \frac{\int_{B(0, R)}\left|x_{2}-y\right|^{-\alpha} d y-\int_{B(0, R)}\left|x_{1}-y\right|^{-\alpha} d y}{\left|x_{1}\right|^{2}-\left|x_{2}\right|^{2}} .
$$

Let us denote the fraction above by $F\left(x_{1}, x_{2}\right)$. We claim that the $\sup _{\left|x_{1}\right| \neq\left|x_{2}\right|} F\left(x_{1}, x_{2}\right)$ is attained when $\left|x_{1}\right|=R$, and $\left|x_{2}\right| \rightarrow R$.

To prove the claim, we first compute $I(x):=\int_{B(0, R)}|x-y|^{-\alpha} d y$. Taking the Fourier transform:

$$
I(x)=C R^{2-\alpha} \int_{0}^{\infty} r^{a-2} J_{1}(r) J_{0}\left(\frac{|x| r}{R}\right) d r
$$

where $C$ is some positive constant. By Sonine-Schafheitlin's formula [115, p. 401] and by conti-
nuity, we obtain

$$
I(x)=\left\{\begin{array}{cc}
C R^{2-\alpha} 2^{\alpha-2} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)}{ }_{2} F_{1}\left(\frac{\alpha}{2}-1, \frac{\alpha}{2}, 1, \frac{|x|^{2}}{R^{2}}\right) & \text { if }|x| \leq R \\
C R^{2-\alpha} 2^{\alpha-2}|x|^{-\alpha} R^{\alpha} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{\Gamma\left(1-\frac{\alpha}{2}\right)}{ }_{2} F_{1}\left(\frac{\alpha}{2}, \frac{\alpha}{2}, 2, \frac{R^{2}}{|x|^{2}}\right) & \text { if }|x|>R .
\end{array}\right.
$$

By the mean value theorem, it is enough to check that min $J(z)=J\left(R^{2}\right)$, where

$$
\begin{aligned}
J(z) & =\left\{\begin{array}{cl}
\frac{d}{d z}\left({ }_{2} F_{1}\left(\frac{\alpha}{2}-1, \frac{\alpha}{2}, 1, \frac{z}{R^{2}}\right)\right) & \text { if } z \leq R^{2} \\
\frac{d}{d z}\left(\left(1-\frac{\alpha}{2}\right) z^{-\frac{\alpha}{2}} R^{\alpha}{ }_{2} F_{1}\left(\frac{\alpha}{2}, \frac{\alpha}{2}, 2, \frac{R^{2}}{z}\right)\right) & \text { if } z>R^{2}
\end{array}\right. \\
& =\left\{\begin{array}{cl}
\frac{\alpha(\alpha-2)}{4} \frac{1}{R^{2}}{ }_{2} F_{1}\left(\frac{\alpha}{2}, 1+\frac{\alpha}{2}, 2, \frac{z}{R^{2}}\right) & \text { if } z \leq R^{2} \\
\frac{\alpha(\alpha-2)}{4} z^{-1-\frac{\alpha}{2}} R^{\alpha}{ }_{2} F_{1}\left(\frac{\alpha}{2}, 1+\frac{\alpha}{2}, 2, \frac{R^{2}}{z}\right) & \text { if } z>R^{2}
\end{array}\right.
\end{aligned}
$$

Writing the series expansion (respectively at $z=0$ and $z=\infty$ ) of the hypergeometric series:

$$
\begin{gathered}
\frac{\alpha(\alpha-2)}{4} \frac{1}{R^{2}}{ }_{2} F_{1}\left(\frac{\alpha}{2}, 1+\frac{\alpha}{2}, 2, z\right)=\frac{1}{R^{2}} \sum_{n=0}^{\infty} \frac{\Gamma\left(\frac{\alpha}{2}+n\right) \Gamma\left(n+1+\frac{\alpha}{2}\right)}{\Gamma\left(\frac{\alpha}{2}-1\right) \Gamma\left(\frac{\alpha}{2}\right) \Gamma(1+n) \Gamma(2+n)}\left(\frac{z}{R^{2}}\right)^{n} \\
\frac{\alpha(\alpha-2)}{4} z^{-1-\frac{\alpha}{2}} R^{\alpha}{ }_{2} F_{1}\left(\frac{\alpha}{2}, 1+\frac{\alpha}{2}, 2, \frac{1}{z}\right)=\left(\frac{1}{z}\right)^{\frac{\alpha}{2}} R^{\alpha-2} \sum_{n=1}^{\infty} \frac{\Gamma\left(\frac{\alpha}{2}+n\right) \Gamma\left(n-1+\frac{\alpha}{2}\right)}{\Gamma\left(\frac{\alpha}{2}-1\right) \Gamma\left(\frac{\alpha}{2}\right) \Gamma(n) \Gamma(n+1)}\left(\frac{R^{2}}{z}\right)^{n},
\end{gathered}
$$

which are both minimized at $z=R^{2}$ since every coefficient is negative. This proves the claim.
The claim immediately implies

$$
\begin{equation*}
\Omega_{c}(R)=-\left.\frac{C_{\alpha}}{R} \frac{d}{d|x|} \int_{B(0, R)}|x-y|^{-\alpha} d y\right|_{|x|=R}, \tag{3.2.5}
\end{equation*}
$$

Where $\frac{d}{d|x|}$ denotes the derivative in the radial variable (recall that $\int_{B(0, R)}|x-y|^{-\alpha} d y$ is radially symmetric). To compute the derivative at $|x|=R$, we can simply compute the partial derivative in
the $x_{1}$ direction at the point $(R, 0)$ :

$$
\begin{align*}
& \left.\frac{\partial}{\partial x_{1}} \int_{B(0, R)}|x-y|^{-\alpha} d y\right|_{x=(R, 0)}=-\alpha \int_{B(0, R)}\left(\left(R-y_{1}\right)^{2}+y_{2}^{2}\right)^{-\frac{\alpha}{2}-1}\left(R-y_{1}\right) d y_{1} d y_{2} \\
& =-\left.2 \int_{0}^{R}\left(\left(R-y_{1}\right)^{2}+y_{2}^{2}\right)^{-\frac{\alpha}{2}}\right|_{y_{1}=-\sqrt{y_{1}=\sqrt{R^{2}-y_{2}{ }^{2}}}} d y_{2} \\
& =-2^{1-\frac{\alpha}{2}} R^{1-\alpha}\left(\int_{0}^{1}\left(1-\sqrt{1-u^{2}}\right)^{-\frac{\alpha}{2}} d u-\int_{0}^{1}\left(1+\sqrt{1-u^{2}}\right)^{-\frac{\alpha}{2}} d u\right)  \tag{3.2.6}\\
& =-2^{1-\frac{\alpha}{2}} R^{1-\alpha}\left(\int_{0}^{\frac{\pi}{2}}(1-\cos \theta)^{-\frac{\alpha}{2}} \cos \theta d \theta-\int_{0}^{\frac{\pi}{2}}(1+\cos \theta)^{-\frac{\alpha}{2}} \cos \theta d \theta\right) \\
& =:-2^{1-\frac{\alpha}{2}} R^{1-\alpha}\left(I_{1}-I_{2}\right),
\end{align*}
$$

where in the third line we used the identity $\left(R \pm \sqrt{R^{2}-y_{2}^{2}}\right)^{2}+y_{2}^{2}=2 R^{2}\left(1 \pm \sqrt{1-\left(R^{-1} y_{2}\right)^{2}}\right)$, as well as the substitution $u=R^{-1} y_{2}$.

Using a substitution $\theta=2 \beta$, we rewrite $I_{1}$ as

$$
I_{1}=2 \int_{0}^{\frac{\pi}{4}}(1-\cos (2 \beta))^{-\frac{\alpha}{2}} \cos (2 \beta) d \beta=2^{1-\frac{\alpha}{2}} \int_{0}^{\frac{\pi}{4}}(\sin \beta)^{-\alpha}\left(1-2 \sin ^{2} \beta\right) d \beta .
$$

Likewise, the substitution $\theta=\pi-2 \beta$ allows us to rewrite $-I_{2}$ as

$$
-I_{2}=2 \int_{\frac{\pi}{4}}^{\frac{\pi}{2}}(1-\cos (2 \beta))^{-\frac{\alpha}{2}} \cos (2 \beta) d \beta=2^{1-\frac{\alpha}{2}} \int_{\frac{\pi}{4}}^{\frac{\pi}{2}}(\sin \beta)^{-\alpha}\left(1-2 \sin ^{2} \beta\right) d \beta
$$

Adding the above two identities for $I_{1}$ and $-I_{2}$ together gives

$$
\begin{aligned}
I_{1}-I_{2} & =2^{1-\frac{\alpha}{2}} \int_{0}^{\frac{\pi}{2}}(\sin \beta)^{-\alpha}\left(1-2 \sin ^{2} \beta\right) d \beta \\
& =2^{1-\frac{\alpha}{2}}\left(\frac{1}{2} B\left(\frac{1-\alpha}{2}, \frac{1}{2}\right)-B\left(\frac{3-\alpha}{2}, \frac{1}{2}\right)\right) \\
& =2^{-\frac{\alpha}{2}} \frac{\Gamma\left(\frac{1-\alpha}{2}\right) \Gamma\left(\frac{1}{2}\right)}{\Gamma\left(1-\frac{\alpha}{2}\right)}-2^{1-\frac{\alpha}{2}} \frac{\Gamma\left(\frac{3-\alpha}{2}\right) \Gamma\left(\frac{1}{2}\right)}{\Gamma\left(2-\frac{\alpha}{2}\right)},
\end{aligned}
$$

where $B$ stands for the beta function. Here the second identity follows from the property that
$B(x, y)=2 \int_{0}^{\pi / 2}(\sin \theta)^{2 x-1}(\cos \theta)^{2 y-1} d \theta$, and the third line follows from the property that $B(x, y)=\frac{\Gamma(x) \Gamma(y)}{\Gamma(x+y)}$. According to the properties of the gamma function $\Gamma(z+1)=z \Gamma(z)$ and $\Gamma(z) \Gamma\left(z+\frac{1}{2}\right)=2^{1-2 z} \sqrt{\pi} \Gamma(2 z)$, we have

$$
\begin{equation*}
I_{1}-I_{2}=2^{-1+\frac{\alpha}{2}} \frac{\alpha}{2-\alpha} \cdot \frac{2 \pi \Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}} \tag{3.2.7}
\end{equation*}
$$

Finally, plugging this into (Equation 3.2.6) and (Equation 3.2.5) gives

$$
\begin{aligned}
\Omega_{c}(R) & =R^{-\alpha} C_{\alpha} 2^{1-\frac{\alpha}{2}}\left(I_{1}-I_{2}\right) \\
& =R^{-\alpha} \frac{1}{2 \pi} \frac{\Gamma\left(\frac{\alpha}{2}\right)}{2^{1-\alpha} \Gamma\left(1-\frac{\alpha}{2}\right)} \frac{\alpha}{2-\alpha}\left(\frac{2 \pi \Gamma(1-\alpha)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2}}\right) \\
& =R^{-\alpha} \frac{2^{\alpha-1} \Gamma(1-\alpha) \Gamma\left(\frac{\alpha}{2}+1\right)}{\Gamma\left(1-\frac{\alpha}{2}\right)^{2} \Gamma\left(2-\frac{\alpha}{2}\right)}=R^{-\alpha} \Omega_{\alpha},
\end{aligned}
$$

finishing the proof.

At the end of this section, we point out that Theorem 3.2.1 directly gives the following quantitative estimate: if a simply-connected patch $D$ rotates with angular velocity $\Omega \in\left(0, \Omega_{\alpha}\right)$ that is very close to $\Omega_{\alpha}$, then $D$ must be very close to a disk in terms of symmetric difference.

Corollary 3.2.4. Assume $0<\alpha<1$. Let $D$ be a rotating patch with area $\pi$ and angular velocity $\Omega \in\left(0, \Omega_{\alpha}\right)$, and let $B$ be the unit disk. Then we have

$$
|D \triangle B| \leq 2 \pi\left(\left(\frac{\Omega_{\alpha}}{\Omega}\right)^{2 / \alpha}-1\right)
$$

Note that for a fixed $\alpha \in(0,1)$, the right hand side goes to 0 as $\Omega \nearrow \Omega_{\alpha}$.

Proof. Denote $R:=\max _{x \in D}|x|$. If $D$ is a rotating patch with angular velocity $\Omega$ and is not a disk, Theorem 3.2.1 gives that $\Omega \leq R^{-\alpha} \Omega_{\alpha}$, which gives that $R \leq\left(\frac{\Omega_{\alpha}}{\Omega}\right)^{1 / \alpha}$. Thus $D \subset B\left(0,\left(\frac{\Omega_{\alpha}}{\Omega}\right)^{1 / \alpha}\right)$,
which implies that the symmetric difference $D \triangle B$ satisfies

$$
|D \triangle B|=2|D \backslash B| \leq 2\left|B\left(0,\left(\frac{\Omega_{\alpha}}{\Omega}\right)^{1 / \alpha}\right) \backslash B\right|=2 \pi\left(\left(\frac{\Omega_{\alpha}}{\Omega}\right)^{2 / \alpha}-1\right) .
$$

## CHAPTER 4

## RIGIDITY RESULTS FOR VORTEX SHEETS

In this chapter, we study radial symmetry of stationary/uniformly-rotating vortex sheets. We fix the following notations in this chapter.

Notations For a bounded domain $D \subset \mathbb{R}^{2}$, we denote $|D|$ by its area (i.e. its Lebesgue measure). For $x \in \mathbb{R}^{2}$ and $r>0$, denote by $B(x, r)$ or $B_{r}(x)$ the open ball centered at $x$ with radius $r$.

Through section 4.2 -section 4.3 , we will desingularize the vortex sheet into a vortex layer with width $\sim \epsilon$, and obtain various quantitative estimates. In all these estimates, we say a term $f$ is $O(g(\epsilon))$ if $|f| \leq C g(\epsilon)$ for some constant $C$ independent of $\epsilon$.

For a domain $U \subset \mathbb{R}^{2}$, in the boundary integral $\int_{\partial U} \vec{f} \cdot n d \sigma, n$ denotes the outer normal of the domain $U$.

### 4.1 Equations for a stationary/rotating vortex sheet

Let $\omega(\cdot, t)=\omega_{0}\left(R_{\Omega t}\right)$ be a stationary/rotating vortex sheet solution to the incompressible 2D Euler equation, where $\omega_{0} \in \mathcal{M}\left(\mathbb{R}^{2}\right) \cap H^{-1}\left(\mathbb{R}^{2}\right)$ is a Radon measure. Here $\Omega=0$ corresponds to a stationary solution, and $\Omega \neq 0$ corresponds to a rotating solution. Assume $\omega_{0}$ is concentrated on $\Gamma$, which is a finite disjoint union of curves. Throughout this paper we assume $\Gamma$ satisfies the following:
(H1) Each connected component of $\Gamma$ is smooth and with finite length, and it is either a simple closed curve (denote them by $\Gamma_{1}, \ldots, \Gamma_{n}$ ), or a non-self-intersecting curve with two endpoints (denote them by $\Gamma_{n+1}, \ldots, \Gamma_{n+m}$ ). Here we require $n+m \geq 1$, but allow either $n$ or $m$ to be 0 .

Let us denote

$$
\begin{equation*}
d_{\Gamma}:=\min _{k \neq i} \operatorname{dist}\left(\Gamma_{i}, \Gamma_{k}\right), \tag{4.1.1}
\end{equation*}
$$

which is strictly positive since we assume the curves $\left\{\Gamma_{i}\right\}_{i=1}^{n+m}$ are disjoint. For $i=1, \ldots, n+m$, denote by $L_{i}$ the length of $\Gamma_{i}$. Let $z_{i}: S_{i} \rightarrow \Gamma_{i}$ denote a constant-speed parameterization of $\Gamma_{i}$ (in counter-clockwise direction if $\Gamma_{i}$ is a closed curve), where the parameter domain $S_{i}$ is given by

$$
S_{i}:= \begin{cases}\mathbb{R} / \mathbb{Z} & \text { for } i=1, \ldots, n \\ {[0,1] \quad \text { for } i=n+1, \ldots, n+m}\end{cases}
$$

Note that this gives $\left|z_{i}^{\prime}\right| \equiv L_{i}$, and the arc-chord constant

$$
\begin{equation*}
F_{\Gamma}:=\max _{i=1, \ldots, n+m} \sup _{\alpha \neq \beta \in S_{i}} \frac{|\alpha-\beta|}{\left|z_{i}(\alpha)-z_{i}(\beta)\right|} \tag{4.1.2}
\end{equation*}
$$

is finite, since $\Gamma$ is non-self-intersecting. Let $\mathrm{s}: \Gamma \rightarrow \mathbb{R}^{2}$ be the unit tangential vector on $\Gamma$, given by $\mathbf{s}\left(z_{i}(\alpha)\right):=\frac{z_{i}^{\prime}(\alpha)}{\left|z_{i}^{\prime}(\alpha)\right|}=\frac{z_{i}^{\prime}(\alpha)}{L_{i}}$, and $\mathbf{n}: \Gamma \rightarrow \mathbb{R}^{2}$ be the unit normal vector, given by $\mathbf{n}=\mathbf{s}^{\perp}$. See Figure Figure 4.1 for an illustration.

For $i=1, \ldots, n+m$, let us denote by $\gamma_{i}(\alpha)$ the vorticity strength at $z_{i}(\alpha)$ with respect to the arclength parametrization, which is related to $\varpi_{i}(\alpha)$ by

$$
\begin{equation*}
\gamma_{i}(\alpha)=\frac{\varpi_{i}(\alpha)}{\left|z_{i}^{\prime}(\alpha)\right|} \quad \text { for } \alpha \in S_{i} \tag{4.1.3}
\end{equation*}
$$

Throughout this paper we will be working with $\gamma_{i}$, instead of $\varpi_{i}$. We impose the following regularity and positivity assumptions on $\gamma_{i}$ :
(H2) Assume that $\gamma_{i} \in C^{2}\left(S_{i}\right)$ for $i=1, \ldots, n$ and $\gamma_{i} \in C^{b}\left(S_{i}\right) \cap C^{1}\left(S_{i}^{\circ}\right)$ for some $b \in(0,1)$ for $i=n+1, \ldots, n+m .{ }^{1}$
(H3) For $i=1, \ldots, n$, assume $\gamma_{i}>0$ in $S_{i}$. And for $i=n+1, \ldots, n+m$, assume $\gamma_{i}>0$ in $S_{i}^{\circ}$, and $\gamma_{i}(0)=\gamma_{i}(1)=0$.

[^0]Note that for a closed curve, (H3) implies that $\gamma_{i}$ is uniformly positive; whereas for an open curve, $\gamma_{i}$ is positive in the interior of $S_{i}$ but vanishes at its endpoints. This is because any stationary/rotating vortex sheet with continuous $\gamma_{i}$ must have it vanishing at the two endpoints of any open curve: if not, one can easily check that $\left|B R\left(z_{i}(\alpha)\right) \cdot \mathbf{n}\left(z_{i}(\alpha)\right)\right| \rightarrow \infty$ as $\alpha$ approaches the endpoint, thus such a vortex sheet cannot be stationary in the rotating frame.

With the above notations of $z_{i}$ and $\gamma_{i}$, the Birkhoff-Rott integral (Equation 1.1.10) along the sheet can now be expressed as

$$
\begin{equation*}
B R\left(z_{i}(\alpha)\right)=\sum_{k=1}^{n+m} B R_{k}\left(z_{i}(\alpha)\right):=\sum_{k=1}^{n+m} P V \int_{S_{k}} K_{2}\left(z_{i}(\alpha)-z_{k}\left(\alpha^{\prime}\right)\right) \gamma_{k}\left(\alpha^{\prime}\right)\left|z_{k}^{\prime}\left(\alpha^{\prime}\right)\right| d \alpha^{\prime} \tag{4.1.4}
\end{equation*}
$$

with the kernel $K_{2}$ given by

$$
\begin{equation*}
K_{2}(x):=(2 \pi)^{-1} \nabla^{\perp} \log |x|=\frac{x^{\perp}}{2 \pi|x|^{2}}, \tag{4.1.5}
\end{equation*}
$$

and the principal value in (Equation 4.1.4) is only needed for the integral with $k=i$.
Let $\mathbf{v}: \mathbb{R}^{2} \rightarrow \mathbb{R}^{2}$ be the velocity field generated by $\omega_{0}$, given by $\mathbf{v}:=\nabla^{\perp}\left(\omega_{0} * \mathcal{N}\right)$. Note that $\mathbf{v} \in C^{\infty}\left(\mathbb{R}^{2} \backslash \Gamma\right)$, but $\mathbf{v}$ is discontinuous across $\Gamma$. Let $\mathbf{v}^{+}, \mathbf{v}^{-}: \Gamma \rightarrow \mathbb{R}^{2}$ denote the two limits of $\mathbf{v}$ on the two sides of $\Gamma$ (with $\mathbf{v}^{+}$being the limit on the side that $\mathbf{n}$ points into - see Figure Figure 4.1 for an illustration), and $[\mathbf{v}]:=\mathbf{v}^{-}-\mathbf{v}^{+}$the jump in $\mathbf{v}$ across the sheets. $[\mathbf{v}]$ is related to the vortex-sheet strength $\gamma$ as follows (see [87, Eq. (9.8)] for a derivation): $[\mathbf{v}] \cdot \mathbf{n}=0$, and

$$
[\mathbf{v}] \times \mathbf{n}=[\mathbf{v}] \cdot \mathbf{s}=\gamma
$$

In addition, the Birkhoff-Rott integral (Equation 4.1.4) is the the average of $\mathbf{v}^{+}$and $\mathbf{v}^{-}$, namely

$$
B R\left(z_{i}(\alpha)\right)=\frac{1}{2}\left(\mathbf{v}^{+}\left(z_{i}(\alpha)\right)+\mathbf{v}^{-}\left(z_{i}(\alpha)\right)\right) \quad \text { for all } \alpha \in S_{i}, i=1, \ldots, n+m
$$



Figure 4.1: Illustration of the closed curves $\Gamma_{1}, \ldots, \Gamma_{n}$ and the open curves $\Gamma_{n+1}, \ldots, \Gamma_{n+m}$, and the definitions of $\mathbf{n}, \mathbf{s}, \mathbf{v}^{+}$and $\mathbf{v}^{-}$.

In the following lemma, we derive the equation that the Birkhoff-Rott integral satisfies for a stationary/rotating vortex sheet.

Lemma 4.1.1. Assume $\omega(\cdot, t)=\omega_{0}\left(R_{\Omega t} x\right)$ is a stationary/uniformly-rotating vortex sheet with angular velocity $\Omega \in \mathbb{R}$, and $\omega_{0}$ is concentrated on $\cup_{i=1}^{n+m} \Gamma_{i}$, with $z_{i}$ and $\gamma_{i}$ defined as above. Then the Birkhoff-Rott integral BR (Equation 4.1.4) and the strength $\gamma_{i}$ satisfy the following two equations:

$$
\begin{equation*}
\left(B R-\Omega x^{\perp}\right) \cdot \mathbf{n}=\mathbf{v}^{+} \cdot \mathbf{n}=\mathbf{v}^{-} \cdot \mathbf{n}=0 \quad \text { on } \Gamma \tag{4.1.6}
\end{equation*}
$$

and

$$
\left(B R\left(z_{i}(\alpha)\right)-\Omega z_{i}^{\perp}(\alpha)\right) \cdot \mathbf{s}\left(z_{i}(\alpha)\right) \gamma_{i}(\alpha)= \begin{cases}C_{i} & \text { on } S_{i} \text { for } i=1, \ldots, n  \tag{4.1.7}\\ 0 & \text { on } S_{i} \text { for } i=n+1, \ldots, n+m\end{cases}
$$

In particular, the above two equations imply that $B R\left(z_{i}(\alpha)\right)-\Omega z_{i}^{\perp}(\alpha) \equiv \mathbf{0}$ for $i=n+1, \ldots, n+$ $m$.

Proof. By definition of the stationary/uniformly-rotating solutions, $\omega_{0}$ is a stationary vortex sheet in the rotating frame with angular velocity $\Omega$. In this rotating frame, an extra velocity $-\Omega z_{i}^{\perp}$
should be added to the right hand side of (Equation 1.1.11). Therefore the evolution equations (Equation 1.1.11)-(Equation 1.1.12) become the following in the rotating frame (where we also use (Equation 4.1.4)):

$$
\begin{align*}
\partial_{t} z_{i}(\alpha, t) & =B R\left(z_{i}(\alpha, t)\right)-\Omega z_{i}^{\perp}(\alpha, t)+c_{i}(\alpha, t) \partial_{\alpha} z_{i}(\alpha, t)  \tag{4.1.8}\\
\partial_{t} \varpi_{i}(\alpha, t) & =\partial_{\alpha}\left(c_{i}(\alpha, t) \varpi_{i}(\alpha, t)\right), \tag{4.1.9}
\end{align*}
$$

where the term $c_{i}(\alpha, t)$ accounts for the reparametrization freedom of the curves. Since $\omega_{0}$ is stationary in the rotating frame, $z_{i}(\cdot, t)$ parametrizes the same curve as $z_{i}(\cdot, 0)$. Therefore $\partial_{t} z_{i}(\alpha, t)$ is tangent to the curve $\Gamma_{i}$, and multiplying $\mathbf{n}\left(z_{i}(\alpha, t)\right)$ to (Equation 4.1.8) gives

$$
\begin{equation*}
0=\partial_{t} z_{i}(\alpha, t) \cdot \mathbf{n}\left(z_{i}(\alpha, t)\right)=\left(B R\left(z_{i}(\alpha, t)\right)-\Omega z_{i}^{\perp}(\alpha, t)\right) \cdot \mathbf{n}\left(z_{i}(\alpha, t)\right), \tag{4.1.10}
\end{equation*}
$$

where we use that $\mathbf{n}\left(z_{i}(\alpha, t)\right) \cdot \partial_{\alpha} z_{i}(\alpha, t)=0$. This proves (Equation 5.1.1).
Now we prove (Equation 5.1.2). Towards this end, let us choose

$$
c_{i}(\alpha, t):=-\frac{\left(B R\left(z_{i}(\alpha, t)\right)-\Omega z_{i}^{\perp}(\alpha, t)\right) \cdot \mathbf{s}\left(z_{i}(\alpha, t)\right)}{\left|\partial_{\alpha} z_{i}(\alpha, t)\right|},
$$

so that multiplying $\mathbf{s}\left(z_{i}(\alpha, t)\right)$ to (Equation 4.1.8) gives $\partial_{t} z_{i}(\alpha, t) \cdot \mathbf{s}\left(z_{i}(\alpha, t)\right)=0$, and combining it with (Equation 4.1.10) gives $\partial_{t} z_{i}(\alpha, t)=0$. In other words, with such choice of $c_{i}$, the parametrization $z_{i}(\alpha, t)$ remains fixed in time. Since $\omega_{0}$ is stationary in the rotating frame, we know that with a fixed parametrization $z_{i}(\alpha, t)=z_{i}(\alpha, 0)$, the strength $\varpi_{i}(\alpha, t)$ must also remain invariant in time. Thus (Equation 4.1.9) becomes

$$
c_{i}(\alpha, t) \varpi_{i}(\alpha, t) \equiv C_{i} .
$$

Plugging the definition of $c_{i}$ into the equation above and using the fact that $z_{i}$ is invariant in $t$, we
have

$$
\frac{\left(B R\left(z_{i}(\alpha)\right)-\Omega z_{i}^{\perp}(\alpha)\right) \cdot \mathbf{s}\left(z_{i}(\alpha)\right) \varpi_{i}(\alpha)}{\left|\partial_{\alpha} z_{i}(\alpha)\right|} \equiv-C_{i} \quad \text { for all } \alpha \in S_{i},
$$

and finally the relationship between $\gamma_{i}$ and $\varpi_{i}$ in (Equation 4.1.3) yields (Equation 5.1.2) for $i=$ $1, \ldots, n$.

And for the open curves $i=n+1, \ldots, n+m$, note that we do not have any reparametrization freedom at the two endpoints $\alpha=0,1$, therefore the endpoint velocity $B R\left(z_{i}(0, t)\right)-\Omega z_{i}^{\perp}(0, t)$ must be 0 to ensure that $\omega_{0}$ is stationary in the rotating frame. This immediately leads to $C_{i}=0$ for $i=n+1, \ldots, n+m$, finishing the proof of (Equation 5.1.2).

### 4.2 Approximation by a thin vortex layer

Our aim in this section is to desingularize the vortex sheet $\omega_{0}$. Namely, for $0<\epsilon \ll 1$, we will construct a vorticity $\omega^{\epsilon} \in L^{\infty}\left(\mathbb{R}^{2}\right) \cap L^{1}\left(\mathbb{R}^{2}\right)$ that only takes values 0 and $\epsilon^{-1}$, and is supported in an $O(\epsilon)$ neighborhood of $\Gamma$, such that $\omega^{\epsilon}$ weakly converges to $\omega_{0}$ as $\epsilon \rightarrow 0^{+}$.

For each $i=1, \ldots, n+m$, we will describe a neighborhood of $\Gamma_{i}$ using the following change of coordinates: let $R_{i}^{\epsilon}: S_{i} \times \mathbb{R} \rightarrow \mathbb{R}^{2}$ be given by

$$
\begin{equation*}
R_{i}^{\epsilon}(\alpha, \eta):=z_{i}(\alpha)+\epsilon \gamma_{i}(\alpha) \mathbf{n}\left(z_{i}(\alpha)\right) \eta \tag{4.2.1}
\end{equation*}
$$

and let

$$
D_{i}^{\epsilon}:=\left\{R_{i}^{\epsilon}(\alpha, \eta): \alpha \in S_{i}^{\circ}, \eta \in(-1,0)\right\} .
$$

Note that each $D_{i}^{\epsilon}$ is a connected open set, and for all $\epsilon>0$ sufficiently small, the sets $\left(D_{i}^{\epsilon}\right)_{i=1}^{n+m}$ are disjoint. For $i=1, \ldots, n$, the domains $D_{i}^{\epsilon}$ are doubly-connected with smooth boundary, and its inner boundary coincides with $\Gamma_{i}$; see the left of Figure Figure 4.2 for an illustration. And for $i=n+1, \ldots, n+m$, the domains $D_{i}^{\epsilon}$ are simply-connected, and its boundary is smooth except at at most two points; see the right of Figure Figure 4.2 for an illustration.


Figure 4.2: Illustration of the definitions of $R_{i}^{\epsilon}$ and $D_{i}^{\epsilon}$ for a closed curve (left) and an open curve (right).

In addition, for $\epsilon>0$ that is sufficiently small, one can check that $R_{i}^{\epsilon}: S_{i}^{\circ} \times(-1,0) \rightarrow D_{i}^{\epsilon}$ is a diffeomorphism. Since $\gamma_{i} \in C^{1}\left(S_{i}\right)$ and $z_{i} \in C^{2}\left(S_{i}\right)$, we only need to show $R_{i}^{\epsilon}: S_{i}^{\circ} \times(-1,0) \rightarrow$ $D_{i}^{\epsilon}$ is injective. Below we prove this fact in a stronger quantitative version, which will be used later.

Lemma 4.2.1. For any $i=1, \ldots, n+m$, assume $\Gamma_{i}$ and $\gamma_{i}$ satisfy $\mathbf{( H 1 ) - ( \mathbf { H } 2 ) . ~ T h e n ~ t h e ~ m a p ~}$ $R_{i}^{\epsilon}: S_{i}^{\circ} \times(-1,0) \rightarrow D_{i}^{\epsilon}$ given by (Equation 4.2.1) is injective. In addition, there exist some $c_{0}, \epsilon_{0}>0$ depending on $\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{L^{\infty}\left(S_{i}\right)}$ and $F_{\Gamma}$, such that for all $\epsilon \in\left(0, \epsilon_{0}\right)$ we have

$$
\begin{equation*}
\left|R_{i}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)-R_{i}^{\epsilon}(\alpha, \eta)\right| \geq c_{0}\left(\left|\alpha^{\prime}-\alpha\right|+\epsilon\left|\gamma_{i}(\alpha) \eta-\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}\right|\right), \tag{4.2.2}
\end{equation*}
$$

for all $\alpha, \alpha^{\prime} \in S_{i}^{\circ}, \eta, \eta^{\prime} \in(-1,0) .{ }^{2}$

Proof. To begin with, note that (Equation 4.2.2) immediately implies that $R_{i}^{\epsilon}: S_{i}^{\circ} \times(-1,0) \rightarrow D_{i}^{\epsilon}$ is injective, where we used the positivity assumption $\gamma_{i}>0$ in $S_{i}^{\circ}$ in (H2). Thus it suffices to prove (Equation 4.2.2). Throughout the proof, we fix any $i \in\{1, \ldots, n+m\}$, and we will omit the subscript $i$ for notational simplicity. Using the definition (Equation 4.2.1), let us break

[^1]$R^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)-R^{\epsilon}(\alpha, \eta)$ into
\[

$$
\begin{equation*}
R^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)-R^{\epsilon}(\alpha, \eta)=\underbrace{z\left(\alpha^{\prime}\right)-z(\alpha)}_{=: T_{1}}+\underbrace{\epsilon\left(\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right) \mathbf{n}\left(z\left(\alpha^{\prime}\right)\right)}_{=: T_{2}}+\underbrace{\epsilon \gamma(\alpha) \eta\left(\mathbf{n}\left(z\left(\alpha^{\prime}\right)\right)-\mathbf{n}(z(\alpha))\right)}_{=: T_{3}} \tag{4.2.3}
\end{equation*}
$$

\]

For $T_{1}$ and $T_{3}$, we have

$$
\begin{align*}
\left|T_{1}-z^{\prime}\left(\alpha^{\prime}\right)\left(\alpha^{\prime}-\alpha\right)\right| & \leq\|z\|_{C^{2}(S)}\left|\alpha-\alpha^{\prime}\right|^{2}  \tag{4.2.4}\\
\left|T_{3}\right| & \leq \epsilon \gamma(\alpha)\|z\|_{C^{2}(S)}\left|\alpha-\alpha^{\prime}\right|
\end{align*}
$$

Also, using that $z^{\prime}\left(\alpha^{\prime}\right)=L \mathbf{s}\left(z\left(\alpha^{\prime}\right)\right)$ is perpendicular to $\mathbf{n}\left(z\left(\alpha^{\prime}\right)\right)$, we have

$$
\begin{aligned}
\left|z^{\prime}\left(\alpha^{\prime}\right)\left(\alpha^{\prime}-\alpha\right)+T_{2}\right| & =\left|L\left(\alpha^{\prime}-\alpha\right) \mathbf{s}\left(z\left(\alpha^{\prime}\right)\right)+\epsilon\left(\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right) \mathbf{n}\left(z\left(\alpha^{\prime}\right)\right)\right| \\
& \geq \frac{1}{2} L\left|\alpha^{\prime}-\alpha\right|+\frac{1}{2} \epsilon\left|\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right|
\end{aligned}
$$

where we use that $\sqrt{x^{2}+y^{2}} \geq \frac{1}{2}(|x|+|y|)$. Combining this with (Equation 4.2.4) gives

$$
\left|T_{1}+T_{2}+T_{3}\right| \geq\left|\alpha-\alpha^{\prime}\right|\left(\frac{L}{2}-\|z\|_{C^{2}(S)}\left(\left|\alpha-\alpha^{\prime}\right|+\epsilon \gamma(\alpha)\right)\right)+\frac{1}{2} \epsilon\left|\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right|
$$

thus

$$
\begin{equation*}
\left|R^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)-R^{\epsilon}(\alpha, \eta)\right| \geq \frac{L}{4}\left|\alpha-\alpha^{\prime}\right|+\frac{1}{2} \epsilon\left|\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right| \tag{4.2.5}
\end{equation*}
$$

for all $0<\epsilon<L\left(8\|z\|_{C^{2}}\|\gamma\|_{L^{\infty}}\right)^{-1}$ and $\left|\alpha-\alpha^{\prime}\right| \leq \frac{L}{8\|z\|_{C^{2}}}$.
For $\left|\alpha-\alpha^{\prime}\right|>\frac{L}{8\|z\|_{C^{2}}}$, recall that the definition of $F_{\Gamma}$ in (Equation 4.1.2) gives $\left|z\left(\alpha^{\prime}\right)-z(\alpha)\right| \geq$ $F_{\Gamma}^{-1}\left|\alpha^{\prime}-\alpha\right|$. Thus a crude estimate gives

$$
\begin{equation*}
\left|R^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)-R^{\epsilon}(\alpha, \eta)\right| \geq\left|z\left(\alpha^{\prime}\right)-z(\alpha)\right|-2 \epsilon\|\gamma\|_{L^{\infty}(S)} \geq \frac{1}{2 F_{\Gamma}}\left|\alpha^{\prime}-\alpha\right|+\epsilon\left|\gamma\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma(\alpha) \eta\right| \tag{4.2.6}
\end{equation*}
$$

for $0<\epsilon<L\left(64 F_{\Gamma}\|z\|_{C^{2}}\|\gamma\|_{L^{\infty}}\right)^{-1}$. (Note that for such $\epsilon$ we have $4 \epsilon\|\gamma\|_{L^{\infty}} \leq \frac{1}{2 F_{\Gamma}}\left|\alpha^{\prime}-\alpha\right|$ due
to our assumption that $\left.\left|\alpha-\alpha^{\prime}\right|>\frac{L}{8\|z\|_{C^{2}}}\right)$.
Finally, combining (Equation 4.2.5) and (Equation 4.2.6), it follows that (Equation 4.2.2) holds for $c_{0}=\min \left\{\frac{L}{4}, \frac{1}{2 F_{\Gamma}}, \frac{1}{2}\right\}$ and $\epsilon_{0}=\min \left\{L\left(8\|z\|_{C^{2}}\|\gamma\|_{L^{\infty}}\right)^{-1}, L\left(64 F_{\Gamma}\|z\|_{C^{2}}\|\gamma\|_{L^{\infty}}\right)^{-1}\right\}$. This finishes the proof.

In the next lemma we compute the partial derivatives and Jacobian of $R_{i}^{\epsilon}(\alpha, \eta)$, which will be useful later.

Lemma 4.2.2. For any $i=1, \ldots, n+m$, let $z_{i}$ be a constant-speed parameterization of the curve $\Gamma_{i}$ (with length $L_{i}$ ), and let $R_{i}^{\epsilon}$ be given by (Equation 4.2.1). Then its partial derivatives are

$$
\begin{align*}
& \partial_{\alpha} R_{i}^{\epsilon}(\alpha, \eta)=z_{i}^{\prime}(\alpha)+\epsilon\left(\gamma_{i}^{\prime}(\alpha) \frac{z_{i}^{\prime}(\alpha)^{\perp}}{L_{i}} \eta+\gamma_{i}(\alpha) \frac{z_{i}^{\prime \prime}(\alpha)^{\perp}}{L_{i}} \eta\right),  \tag{4.2.7}\\
& \partial_{\eta} R_{i}^{\epsilon}(\alpha, \eta)=\epsilon \gamma_{i}(\alpha) \frac{z_{i}^{\prime}(\alpha)^{\perp}}{L_{i}} .
\end{align*}
$$

Moreover, its Jacobian is given by

$$
\begin{equation*}
\operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}\right)=\epsilon L_{i} \gamma_{i}(\alpha)-\epsilon^{2} L_{i} \gamma_{i}^{2}(\alpha) \kappa_{i}(\alpha) \eta, \tag{4.2.8}
\end{equation*}
$$

where $\kappa_{i}(\alpha)$ denotes the signed curvature of $\Gamma_{i}$ at $z_{i}(\alpha)$.

Proof. Since $z_{i}$ is the constant-speed parameterization of $\Gamma_{i}$ (which has length $L_{i}$ ), we have $\left|z_{i}^{\prime}\right| \equiv$ $L_{i}$ and $\mathbf{n}\left(z_{i}(\alpha)\right)=z_{i}^{\prime}(\alpha)^{\perp} / L_{i}$. Taking the $\alpha$ and $\eta$ partial derivatives of (Equation 4.2.1) directly yields (Equation 4.2.7).

Putting the two partial derivatives into columns of a $2 \times 2$ matrix and computing the determinant, we have

$$
\begin{aligned}
\operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}\right) & =\epsilon \gamma_{i}(\alpha) \frac{\left|z_{i}^{\prime}(\alpha)\right|^{2}}{L_{i}}+\epsilon^{2} \gamma_{i}^{2}(\alpha) \frac{z_{i}^{\prime \prime}(\alpha)^{\perp} \cdot z_{i}^{\prime}(\alpha)}{L_{i}^{2}} \eta \\
& =\epsilon L_{i} \gamma_{i}(\alpha)-\epsilon^{2} L_{i} \gamma_{i}^{2}(\alpha) \kappa_{i}(\alpha) \eta,
\end{aligned}
$$

where in the second equality we used that $z_{i}^{\prime \prime}(\alpha)=\kappa_{i}(\alpha) \mathbf{n}\left(z_{i}(\alpha)\right) L_{i}^{2}$ (recall that $z_{i}$ has constant speed $L_{i}$ ). This finishes the proof.

Remark 4.2.3. We point out that for each $i=1, \ldots, n+m$, the determinant formula (Equation 4.2.8) immediately gives the following approximation of $\left|D_{i}^{\epsilon}\right|$, which will be helpful in the proofs later:

$$
\begin{equation*}
\frac{\left|D_{i}^{\epsilon}\right|}{\epsilon}=\frac{1}{\epsilon} \int_{D_{i}^{\epsilon}} 1 d x=\frac{1}{\epsilon} \int_{S_{i}} \int_{-1}^{0} \operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}(\alpha, \eta)\right) d \eta d \alpha=L_{i} \int_{S_{i}} \gamma_{i}(\alpha) d \alpha+O(\epsilon) \tag{4.2.9}
\end{equation*}
$$

where the $O(\epsilon)$ error term has its absolute value bounded by $C \epsilon$, with $C$ only depending on $\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)}$ and $\left\|\gamma_{i}\right\|_{L^{\infty}\left(S_{i}\right)}$.

Finally, let $D^{\epsilon}:=\cup_{i=1}^{n+m} D_{i}^{\epsilon}$, and $\omega^{\epsilon}: \mathbb{R}^{2} \rightarrow \mathbb{R}$ is defined as

$$
\omega^{\epsilon}(x):=\epsilon^{-1} 1_{D^{\epsilon}}(x)=\epsilon^{-1} \sum_{i=1}^{n+m} 1_{D_{i}^{\epsilon}}(x),
$$

and let

$$
\begin{equation*}
\mathbf{v}^{\epsilon}=\nabla^{\perp}\left(\omega^{\epsilon} * \mathcal{N}\right) \tag{4.2.10}
\end{equation*}
$$

be the velocity field generated by $\omega^{\epsilon}$.
In the next lemma we aim to obtain some fine estimate of $\mathbf{v}^{\epsilon}$ in the thin vortex layer $D^{\epsilon}$. Our goal is to show that along each cross section of the thin layer (i.e. fix $i$ and $\alpha$, and let $\eta$ vary in $[-1,0]$ ), the function $\eta \mapsto \mathbf{v}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)$ is almost a linear function in $\eta$, with the endpoint values (at $\eta=-1$ and 0 ) being almost $\mathbf{v}^{-}\left(z_{i}(\alpha)\right)$ and $\mathbf{v}^{+}\left(z_{i}(\alpha)\right)$ respectively.

Lemma 4.2.4. For $i=1, \ldots, n+m$, assume $\Gamma_{i}$ and $\gamma_{i}$ satisfy (H1)-(H3). Let

$$
g_{i}(\alpha, \eta):=B R\left(z_{i}(\alpha)\right)-\left(\eta+\frac{1}{2}\right)[\mathbf{v}]\left(z_{i}(\alpha)\right) \quad \text { for } \alpha \in S_{i}
$$

and note that $g_{i}(\alpha, 0)=\mathbf{v}^{+}\left(z_{i}(\alpha)\right)$ and $g_{i}(\alpha,-1)=\mathbf{v}^{-}\left(z_{i}(\alpha)\right)$ (see Figure Figure 4.3 for an
illustration of $g_{i}(\alpha, \eta)$ ). Then for all sufficiently small $\epsilon>0$, for all $i=1, \ldots, n+m$ we have

$$
\begin{equation*}
\left|\mathbf{v}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-g_{i}(\alpha, \eta)\right| \leq C \epsilon^{b}|\log \epsilon| \quad \text { for all } \alpha \in S_{i}, \eta \in[-1,0], \tag{4.2.11}
\end{equation*}
$$

where $b \in(0,1)$ is as in (H2), and C depends on $b, \max _{i}\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)}, \max _{i}\left\|\gamma_{i}\right\|_{C^{b}\left(S_{i}\right)}, d_{\Gamma}$ and $F_{\Gamma}$.


Figure 4.3: Illustration of the definition of $g_{i}(\alpha, \cdot)$ (the orange arrows).

Proof. Let $i$ be any fixed index in $1, \ldots, n+m$. We begin with breaking $\mathbf{v}^{\epsilon}$ into contributions from different components $\left\{D_{k}^{\epsilon}\right\}_{k=1}^{n+m}$, namely

$$
\mathbf{v}^{\epsilon}(x)=\sum_{k=1}^{n+m} \mathbf{v}_{k}^{\epsilon}(x):=\sum_{k=1}^{n+m} \epsilon^{-1} \int_{D_{i}^{\epsilon}} K_{2}(x-y) d y
$$

where the kernel $K_{2}$ is given by (Equation 4.1.5). Similarly, we can break $B R\left(z_{i}(\alpha)\right)$ into $B R\left(z_{i}(\alpha)\right)=\sum_{k=1}^{n+m} B R_{k}\left(z_{i}(\alpha)\right)$, where $B R_{k}$ is the contribution from the $k$-th integral in (Equation 4.1.4), and note that the PV symbol is only needed for $k=i$.

- Estimates for $k \neq i$ terms. For any $k \neq i$, we aim to show that

$$
\begin{equation*}
\left|\mathbf{v}_{k}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-B R_{k}\left(z_{i}(\alpha)\right)\right| \leq C \epsilon \tag{4.2.12}
\end{equation*}
$$

where $C$ depends on $d_{\Gamma}, \max _{k}\left\|z_{k}\right\|_{C^{2}}$ and $\max _{k}\left\|\gamma_{k}\right\|_{L^{\infty}}$. Applying a change of variable $y=$
$R_{k}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)$, we can rewrite $\mathbf{v}_{k}^{\epsilon}$ as

$$
\begin{align*}
\mathbf{v}_{k}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right) & =\epsilon^{-1} \int_{D_{k}^{\epsilon}} K_{2}\left(R_{i}^{\epsilon}(\alpha, \eta)-y\right) d y \\
& =\int_{S_{k}} \int_{-1}^{0} \underbrace{K_{2}\left(R_{i}^{\epsilon}(\alpha, \eta)-R_{k}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)\right)}_{=: T_{1}} \underbrace{\epsilon^{-1} \operatorname{det}\left(\nabla_{\alpha^{\prime}, \eta^{\prime}} R_{k}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)\right)}_{=: T_{2}} d \eta^{\prime} d \alpha^{\prime} \tag{4.2.13}
\end{align*}
$$

Using the facts that $R_{i}^{\epsilon}(\alpha, \eta)-R_{k}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)=z_{i}(\alpha)-z_{k}\left(\alpha^{\prime}\right)+O(\epsilon)$ as well as $\left|z_{i}(\alpha)-z_{k}\left(\alpha^{\prime}\right)\right| \geq$ $d_{\Gamma}>0$ (recall that $d_{\Gamma}$ is as given in (Equation 4.1.1)), for all sufficiently small $\epsilon>0$ we have $T_{1}=K_{2}\left(z_{i}(\alpha)-z_{k}\left(\alpha^{\prime}\right)\right)+O(\epsilon)$. For $T_{2}$, the explicit formula (Equation 4.2.8) for the determinant gives $T_{2}=L_{k} \gamma_{k}\left(\alpha^{\prime}\right)+O(\epsilon)$. Plugging these into the above integral yields

$$
\mathbf{v}_{k}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)=\int_{S_{k}} K_{2}\left(z_{i}(\alpha)-z_{k}\left(\alpha^{\prime}\right)\right) L_{k} \gamma_{k}\left(\alpha^{\prime}\right) d \alpha^{\prime}+O(\epsilon)=B R_{k}\left(z_{i}(\alpha)\right)+O(\epsilon)
$$

finishing the proof of (Equation 4.2.12).

- Estimates for the $k=i$ term. It will be more involved to control the $k=i$ term, and our goal is to show that

$$
\begin{equation*}
\left|\mathbf{v}_{i}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-B R_{i}\left(z_{i}(\alpha)\right)+\left(\eta+\frac{1}{2}\right)[\mathbf{v}]\left(z_{i}(\alpha)\right)\right| \leq C \epsilon^{b}|\log \epsilon| . \tag{4.2.14}
\end{equation*}
$$

To begin with, we again rewrite $\mathbf{v}_{i}^{\epsilon}$ as in (Equation 4.2.13) with $k=i$, and plug in the formula (Equation 4.2.8) for the determinant. This leads to

$$
\begin{aligned}
\mathbf{v}_{i}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right) & =\int_{S_{k}} \int_{-1}^{0} K_{2}\left(R_{i}^{\epsilon}(\alpha, \eta)-R_{i}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)\right)\left(L_{i} \gamma_{i}\left(\alpha^{\prime}\right)-\epsilon L_{i} \gamma_{i}^{2}\left(\alpha^{\prime}\right) \kappa_{i}\left(\alpha^{\prime}\right) \eta^{\prime}\right) d \eta^{\prime} d \alpha^{\prime} \\
& =: I_{1}+I_{2}
\end{aligned}
$$

where $I_{1}, I_{2}$ are the contributions from the two terms in the last parenthesis respectively. Let us control $I_{2}$ first, and we claim that

$$
\begin{equation*}
\left|I_{2}\right| \leq C \epsilon|\log \epsilon| \tag{4.2.15}
\end{equation*}
$$

Using (Equation 4.2.2) of Lemma 4.2.1 and the fact that $\left|K_{2}(x)\right| \leq|x|^{-1}$, we can bound $I_{2}$ as

$$
\begin{align*}
\left|I_{2}\right| & =\left|\int_{S_{k}} \int_{-1}^{0} K_{2}\left(R_{i}^{\epsilon}(\alpha, \eta)-R_{i}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)\right) \epsilon L_{i} \gamma_{i}^{2}\left(\alpha^{\prime}\right) \kappa_{i}\left(\alpha^{\prime}\right) \eta^{\prime} d \eta^{\prime} d \alpha^{\prime}\right| \\
& \leq C \epsilon \int_{S_{k}} \int_{-1}^{0} \frac{\gamma_{i}\left(\alpha^{\prime}\right)}{\left|\alpha^{\prime}-\alpha\right|+\epsilon\left|\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta\right|} d \eta^{\prime} d \alpha^{\prime} \\
& \leq C \epsilon \int_{S_{k}} \int_{-\left\|\gamma_{i}\right\|_{\infty}}^{\left\|\gamma_{i}\right\|_{\infty}} \frac{1}{\left|\alpha^{\prime}-\alpha\right|+\epsilon\left|\theta^{\prime}\right|} d \theta^{\prime} d \alpha^{\prime} \quad\left(\theta^{\prime}:=\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta\right)  \tag{4.2.16}\\
& \leq C \epsilon \int_{-1 / \epsilon}^{1 / \epsilon} \int_{-\left\|\gamma_{i}\right\|_{\infty}}^{\left\|\gamma_{i}\right\|_{\infty}} \frac{1}{\left|\beta^{\prime}\right|+\left|\theta^{\prime}\right|} d \theta^{\prime} d \beta^{\prime} \quad\left(\beta^{\prime}:=\epsilon^{-1}\left(\alpha^{\prime}-\alpha\right)\right) \\
& \leq C \epsilon|\log \epsilon|
\end{align*}
$$

where $C$ depends on $\left\|z_{i}\right\|_{C^{2}}$ and $\left\|\gamma_{i}\right\|_{L^{\infty}}$.
In the rest of the proof we focus on estimating $I_{1}=\int_{S_{k}} \int_{-1}^{0} K_{2}\left(R_{i}^{\epsilon}(\alpha, \eta)-\right.$ $\left.R_{i}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}\right)\right) L_{i} \gamma_{i}\left(\alpha^{\prime}\right) d \eta^{\prime} d \alpha^{\prime}$. For $t \in[0,1]$, let us define

$$
\begin{align*}
& f\left(\alpha, \alpha^{\prime}, \eta, \eta^{\prime} ; t\right):=R_{i}^{\epsilon}\left(\alpha, \eta-t \eta^{\prime}\right)-R_{i}^{\epsilon}\left(\alpha^{\prime}, \eta^{\prime}-t \eta^{\prime}\right) \\
& J(t):=\int_{S_{k}} \int_{-1}^{0} K_{2}\left(f\left(\alpha, \alpha^{\prime}, \eta, \eta^{\prime} ; t\right)\right) L_{i} \gamma_{i}\left(\alpha^{\prime}\right) d \eta^{\prime} d \alpha^{\prime} \tag{4.2.17}
\end{align*}
$$

Note that in the definition of $f$, the argument $\eta-t \eta^{\prime}$ of $R_{i}^{\epsilon}$ belongs to $[-1,1]$, instead of $[-1,0]$ as in the original definition of (Equation 4.2.1). Here $R_{i}^{\epsilon}\left(\alpha, \eta-t \eta^{\prime}\right)$ is defined as in the formula (Equation 4.2.1), even though it might not belong to $D_{i}^{\epsilon}$. Clearly, $J(0)=I_{1}$. The motivation for us to define such $f$ and $J(t)$ is that at $t=1$, we have

$$
\begin{equation*}
J(1)=\int_{S_{k}} \int_{-1}^{0} K_{2}\left(R_{i}^{\epsilon}\left(\alpha, \eta-\eta^{\prime}\right)-z_{i}\left(\alpha^{\prime}\right)\right) L_{i} \gamma_{i}\left(\alpha^{\prime}\right) d \eta^{\prime} d \alpha^{\prime}=\int_{-1}^{0} \mathbf{v}_{i}\left(R_{i}^{\epsilon}\left(\alpha, \eta-\eta^{\prime}\right)\right) d \eta^{\prime} \tag{4.2.18}
\end{equation*}
$$

where $\mathbf{v}_{i}$ is the velocity field generated by the sheet $\Gamma_{i}$. Recall that $\mathbf{v}_{i}$ has a jump across $\Gamma_{i}$, where we denote its limits on two sides by $\mathbf{v}_{i}^{ \pm}$. Using Lemma 4.2.5, which we will prove momentarily,
we have

$$
\mathbf{v}_{i}\left(R_{i}^{\epsilon}\left(\alpha, \eta-\eta^{\prime}\right)\right)= \begin{cases}\mathbf{v}_{i}^{+}\left(z_{i}(\alpha)\right)+O\left(\epsilon^{b}|\log \epsilon|\right) & \text { if } \eta-\eta^{\prime} \in(0,2)  \tag{4.2.19}\\ \mathbf{v}_{i}^{-}\left(z_{i}(\alpha)\right)+O\left(\epsilon^{b}|\log \epsilon|\right) & \text { if } \eta-\eta^{\prime} \in(-2,0)\end{cases}
$$

We can then split the integration domain on the right hand side of (Equation 6.2.56) into $\eta^{\prime} \in$ $(-1, \eta)$ and $\eta^{\prime} \in(\eta, 0)$, and use (Equation 4.2.19) to approximate the integrand in each interval. This gives

$$
\begin{align*}
J(1) & =(\eta+1) \mathbf{v}_{i}^{+}\left(z_{i}(\alpha)\right)-\eta \mathbf{v}_{i}^{-}\left(z_{i}(\alpha)\right)+O\left(\epsilon^{b}|\log \epsilon|\right) \\
& =B R_{i}\left(z_{i}(\alpha)\right)-\left(\eta+\frac{1}{2}\right)[\mathbf{v}]\left(z_{i}(\alpha)\right)+O\left(\epsilon^{b}|\log \epsilon|\right), \tag{4.2.20}
\end{align*}
$$

where in the last step we used that $[\mathbf{v}]\left(z_{i}(\alpha)\right)=\left[\mathbf{v}_{i}\right]\left(z_{i}(\alpha)\right)$, since all other $\mathbf{v}_{k}$ with $k \neq i$ are continuous across $\Gamma_{i}$.

Finally, it remains to control $|J(0)-J(1)|$. Note that by (Equation 4.2.2), we have

$$
f\left(\alpha, \alpha^{\prime}, \eta, \eta^{\prime} ; t\right) \geq c_{0}\left(\left|\alpha-\alpha^{\prime}\right|+\epsilon\left|\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta\right|\right) .
$$

In addition, we have

$$
\left|\frac{\partial}{\partial t} f\left(\alpha, \alpha^{\prime}, \eta, \eta^{\prime} ; t\right)\right|=\left|\epsilon\left(\gamma_{i}(\alpha) \mathbf{n}\left(z_{i}(\alpha)\right)-\gamma_{i}\left(\alpha^{\prime}\right) \mathbf{n}\left(z_{i}\left(\alpha^{\prime}\right)\right)\right) \eta^{\prime}\right| \leq C \epsilon\left|\alpha-\alpha^{\prime}\right|^{b}
$$

where the last inequality follows from (H2) and the fact that $\mathbf{n}\left(z_{i}(\alpha)\right) \in C^{1}\left(S_{i}\right)$. Therefore, for any $t \in(0,1)$, taking the $t$ derivative of (Equation 4.2.17) and using that $\left|\nabla K_{2}(x)\right| \leq|x|^{-2}$, we
have

$$
\begin{aligned}
\left|J^{\prime}(t)\right| & \leq C \int_{S_{k}} \int_{-1}^{0} \frac{\epsilon\left|\alpha-\alpha^{\prime}\right|^{b} \gamma_{i}\left(\alpha^{\prime}\right)}{\left(\left|\alpha-\alpha^{\prime}\right|+\epsilon\left|\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta\right|\right)^{2}} d \eta^{\prime} d \alpha^{\prime} \\
& \leq C \epsilon \int_{S_{k}} \int_{-1}^{0} \frac{\gamma_{i}\left(\alpha^{\prime}\right)}{\left|\alpha-\alpha^{\prime}\right|^{1-b}\left(\left|\alpha-\alpha^{\prime}\right|+\epsilon\left|\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta\right|\right)} d \eta^{\prime} d \alpha^{\prime} \\
& \leq C \epsilon^{b} \int_{-1 / \epsilon}^{1 / \epsilon} \int_{-\left\|\gamma_{i}\right\|_{\infty}}^{\left\|\gamma_{i}\right\| \infty} \frac{1}{\left|\beta^{\prime}\right|^{1-b}\left(\left|\beta^{\prime}\right|+\left|\theta^{\prime}\right|\right)} d \theta^{\prime} d \beta^{\prime} \quad\left(\theta^{\prime}:=\gamma_{i}\left(\alpha^{\prime}\right) \eta^{\prime}-\gamma_{i}(\alpha) \eta, \beta^{\prime}:=\epsilon^{-1}\left(\alpha^{\prime}-\alpha\right)\right) \\
& \leq C \epsilon^{b} \int_{-1 / \epsilon}^{1 / \epsilon}\left|\beta^{\prime}\right|^{b-1} \log \left(1+\frac{\left\|\gamma_{i}\right\|_{L^{\infty}}}{\left|\beta^{\prime}\right|}\right) d \beta^{\prime} \\
& \leq C \epsilon^{b},
\end{aligned}
$$

where $C$ depends on $b,\left\|\gamma_{i}\right\|_{C^{b}\left(S_{i}\right)},\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)}$ and $F_{\Gamma}$. This leads to

$$
\left|J(1)-I_{1}\right|=|J(1)-J(0)| \leq C \epsilon^{b}|\log \epsilon| .
$$

Finally, combining this with (Equation 4.2.20) and (Equation 4.2.15) yields (Equation 4.2.14), finishing the proof of the $k=i$ case. We can then conclude the proof by taking the sum of this estimate with all the $k \neq i$ estimates in (Equation 4.2.12).

The following lemma proves (Equation 4.2.19). Let $\mathbf{v}_{i}$ be the velocity field generated by the sheet $\Gamma_{i}$, which is smooth in $\mathbb{R}^{2} \backslash \Gamma_{i}$, and has a discontinuity across $\Gamma_{i}$. It is known that $\mathbf{v}_{i}$ converges to $\mathbf{v}_{i}^{ \pm}$respectively on the two sides of $\Gamma_{i}$ [87]. However, we were unable to find a quantitative convergence rate (in terms of the distance from the point to $\Gamma_{i}$ ) in the literature, especially under the assumption that $\gamma_{i}$ is only in $C^{b}\left(S_{i}\right)$ for the open curves. Below we prove such an estimate.

Lemma 4.2.5. For $i=1, \ldots, n+m$, let $\mathbf{v}_{i}$ be the velocity field generated by the sheet $\Gamma_{i}$, given by

$$
\mathbf{v}_{i}(x):=\int_{S_{i}} K_{2}\left(x-z_{i}\left(\alpha^{\prime}\right)\right) \gamma_{i}\left(\alpha^{\prime}\right)\left|z_{i}^{\prime}\left(\alpha^{\prime}\right)\right| d \alpha^{\prime} \quad \text { for } x \in \mathbb{R}^{2} \backslash \Gamma_{i} .
$$

Then there exist constants $C, \epsilon_{0}>0$ depending on on b(as in (H2)), $\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{C^{b}\left(S_{i}\right)}$ and $F_{\Gamma}$,
such that for all $\epsilon \in\left(0, \epsilon_{0}\right)$ and $\eta \in(-2,2)$ we have

$$
\begin{array}{ll}
\left|\mathbf{v}_{i}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-\mathbf{v}_{i}^{+}\left(z_{i}(\alpha)\right)\right| \leq C \epsilon^{b}|\log \epsilon| & \text { if } \eta \in(0,2) \\
\left|\mathbf{v}_{i}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-\mathbf{v}_{i}^{-}\left(z_{i}(\alpha)\right)\right| \leq C \epsilon^{b}|\log \epsilon| & \text { if } \eta \in(-2,0) \tag{4.2.22}
\end{array}
$$

where

$$
\mathbf{v}_{i}^{+}=B R_{i}\left(z_{i}(\alpha)\right)+\frac{\mathbf{n}\left(z_{i}(\alpha)\right)^{\perp} \gamma_{i}(\alpha)}{2}, \quad \mathbf{v}_{i}^{-}=B R_{i}\left(z_{i}(\alpha)\right)-\frac{\mathbf{n}\left(z_{i}(\alpha)\right)^{\perp} \gamma_{i}(\alpha)}{2}
$$

and $B R_{i}$ is the contribution from the $i$-th integral in (Equation 4.1.4).

Proof. We will show (Equation 4.2.21) only since (Equation 4.2.22) can be treated in the same way. From the definition of $R_{i}^{\epsilon}$ in (Equation 4.2.1), we have

$$
\begin{aligned}
\mathbf{v}_{i}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)= & \frac{L_{i}}{2 \pi} \int_{S_{i}} \frac{\left(z_{i}(\alpha)-z_{i}\left(\alpha^{\prime}\right)\right)^{\perp} \gamma_{i}\left(\alpha^{\prime}\right)}{\left|z_{i}(\alpha)-z_{i}\left(\alpha^{\prime}\right)+\epsilon \eta \mathbf{n}\left(z_{i}(\alpha)\right) \gamma_{i}(\alpha)\right|^{2}} d \alpha^{\prime} \\
& +\frac{L_{i}}{2 \pi} \int_{S_{i}} \frac{\epsilon \eta \mathbf{n}\left(z_{i}(\alpha)\right)^{\perp} \gamma_{i}(\alpha) \gamma_{i}\left(\alpha^{\prime}\right)}{\left|z_{i}(\alpha)-z_{i}\left(\alpha^{\prime}\right)+\epsilon \eta \mathbf{n}\left(z_{i}(\alpha)\right) \gamma_{i}(\alpha)\right|^{2}} d \alpha^{\prime} \\
= & A_{1}+A_{2} .
\end{aligned}
$$

We claim that for all $\epsilon>0$ sufficiently small and $\eta \in[0,2)$, we have

$$
\begin{align*}
& \left|A_{1}-B R_{i}(z(\alpha))\right| \leq C \epsilon^{b}|\log \epsilon|,  \tag{4.2.23}\\
& \left|A_{2}-\frac{\mathbf{n}(z(\alpha))^{\perp} \gamma(\alpha)}{2}\right| \leq C \epsilon^{b}, \tag{4.2.24}
\end{align*}
$$

and note that these two claims immediately yield (Equation 4.2.21). From now on, let us fix $i \in\{1, \ldots, n+m\}$ and omit it in the notation for simplicity. Throughout this proof, let us denote

$$
\mathbf{y}\left(\alpha, \alpha^{\prime}\right):=z(\alpha)-z\left(\alpha^{\prime}\right) \quad \text { and } \quad \mathbf{c}(\alpha):=\epsilon \eta \mathbf{n}(z(\alpha)) \gamma(\alpha),
$$

so that

$$
A_{1}=\frac{L}{2 \pi} \int_{S} \frac{\mathbf{y}^{\perp}\left(\alpha, \alpha^{\prime}\right) \gamma\left(\alpha^{\prime}\right)}{\left|\mathbf{y}\left(\alpha, \alpha^{\prime}\right)+\mathbf{c}(\alpha)\right|^{2}} d \alpha^{\prime}, \quad A_{2}=\frac{L}{2 \pi} \int_{S} \frac{\mathbf{c}^{\perp}(\alpha) \gamma\left(\alpha^{\prime}\right)}{\left|\mathbf{y}\left(\alpha, \alpha^{\prime}\right)+\mathbf{c}(\alpha)\right|^{2}} d \alpha^{\prime}
$$

Note that

$$
\begin{equation*}
F_{\Gamma}^{-1}\left|\alpha-\alpha^{\prime}\right| \leq\left|\mathbf{y}\left(\alpha, \alpha^{\prime}\right)\right| \leq\|z\|_{C^{1}}\left|\alpha-\alpha^{\prime}\right| . \tag{4.2.25}
\end{equation*}
$$

For the closed curves with $i=1, \ldots, n$, since $z$ has period 1 , we can always set $\alpha-\alpha^{\prime} \in\left[-\frac{1}{2}, \frac{1}{2}\right)$ in this proof.

Applying (Equation 4.2.2) (with $\eta^{\prime}=0$ ), we have

$$
\begin{equation*}
\left|\mathbf{y}\left(\alpha, \alpha^{\prime}\right)+\mathbf{c}(\alpha)\right|^{2} \geq c_{0}\left(\left|\alpha-\alpha^{\prime}\right|^{2}+\epsilon^{2} \eta^{2} \gamma^{2}(\alpha)\right)=c_{0}\left(\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}(\alpha)|^{2}\right) . \tag{4.2.26}
\end{equation*}
$$

Since $z^{\prime}(\alpha)=L \mathbf{s}(z(\alpha))$, let us define

$$
\tilde{\mathbf{y}}\left(\alpha, \alpha^{\prime}\right):=L \mathbf{s}(z(\alpha))\left(\alpha-\alpha^{\prime}\right),
$$

which is a close approximation of $\mathbf{y}$ in the sense that

$$
\begin{equation*}
\left|\mathbf{y}\left(\alpha, \alpha^{\prime}\right)-\tilde{\mathbf{y}}\left(\alpha, \alpha^{\prime}\right)\right| \leq\|z\|_{C^{2}}\left(\alpha-\alpha^{\prime}\right)^{2} \tag{4.2.27}
\end{equation*}
$$

Using $\mathbf{s}(z(\alpha)) \perp \mathbf{n}(z(\alpha))$, we have

$$
\begin{equation*}
\left|\tilde{\mathbf{y}}\left(\alpha, \alpha^{\prime}\right)+\mathbf{c}(\alpha)\right|^{2}=L^{2}\left|\alpha-\alpha^{\prime}\right|^{2}+\epsilon^{2} \eta^{2} \gamma^{2}(\alpha)=L^{2}\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}(\alpha)|^{2} . \tag{4.2.28}
\end{equation*}
$$

From now on, for notational simplicity, we compress the dependence of $\mathbf{y}\left(\alpha, \alpha^{\prime}\right), \tilde{\mathbf{y}}\left(\alpha, \alpha^{\prime}\right), \mathbf{c}(\alpha)$ on $\alpha$ and $\alpha^{\prime}$ in the rest of the proof.

- Estimate (Equation 4.2.23). Note that $B R_{i}(z(\alpha))$ can also be written using the above nota-
tions as

$$
B R_{i}(z(\alpha))=\frac{L}{2 \pi} P V \int_{S} \frac{\mathbf{y}^{\perp}}{|\mathbf{y}|^{2}} \gamma\left(\alpha^{\prime}\right) d \alpha
$$

thus $A_{1}-B R_{i}(z(\alpha))$ can be written as follows:

$$
\begin{aligned}
A_{1}-B R_{i}(z(\alpha)) & =\frac{L}{2 \pi} P V \int_{S} \underbrace{\left(\frac{\mathbf{y}^{\perp}}{|\mathbf{y}+\mathbf{c}|^{2}}-\frac{\mathbf{y}^{\perp}}{|\mathbf{y}|^{2}}\right)}_{=: \mathbf{f}(\mathbf{y}, \mathbf{c})} \gamma\left(\alpha^{\prime}\right) d \alpha^{\prime} \\
& =\frac{L}{2 \pi} \int_{S} \mathbf{f}(\mathbf{y}, \mathbf{c})\left(\gamma\left(\alpha^{\prime}\right)-\gamma(\alpha)\right) d \alpha^{\prime}+\frac{L \gamma(\alpha)}{2 \pi} P V \int_{S} \mathbf{f}(\mathbf{y}, \mathbf{c}) d \alpha^{\prime} \\
& =: A_{11}+A_{12} .
\end{aligned}
$$

A direct computation gives

$$
\begin{equation*}
\mathbf{f}(\mathbf{y}, \mathbf{c})=-\frac{\mathbf{y}^{\perp}}{|\mathbf{y}|^{2}} \frac{2 \mathbf{y} \cdot \mathbf{c}+|\mathbf{c}|^{2}}{|\mathbf{y}+\mathbf{c}|^{2}} \tag{4.2.29}
\end{equation*}
$$

Since $\mathbf{y} \cdot \mathbf{c}=(\mathbf{y}-\tilde{\mathbf{y}}) \cdot \mathbf{c} \leq C\left|\alpha-\alpha^{\prime}\right|^{2}|\mathbf{c}|$, (where we use $\tilde{\mathbf{y}} \perp \mathbf{n}(z(\alpha))$ and (Equation 4.2.27)), combining this with (Equation 4.2.25) and (Equation 4.2.26) gives a crude bound

$$
|\mathbf{f}(\mathbf{y}, \mathbf{c})| \lesssim \frac{\left|\alpha-\alpha^{\prime}\right|^{2}|\mathbf{c}|+|\mathbf{c}|^{2}}{\left|\alpha-\alpha^{\prime}\right|\left(\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}|^{2}\right)}
$$

Plugging this into $A_{11}$ and using the Hölder continuity of $\gamma$, we have

$$
\begin{aligned}
\left|A_{11}\right| & \lesssim \int_{S} \frac{\left|\alpha-\alpha^{\prime}\right|^{2}|\mathbf{c}|+|\mathbf{c}|^{2}}{\left|\alpha-\alpha^{\prime}\right|\left(\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}|^{2}\right)}\left|\alpha-\alpha^{\prime}\right|^{b} d \alpha^{\prime} \\
& \lesssim \int_{|\theta|<|\mathbf{c}|}\left(|\theta|^{1+b}|\mathbf{c}|^{-1}+|\theta|^{b-1}\right) d \theta+\int_{|\mathbf{c}| \leq|\theta| \leq 1}\left(|\mathbf{c}||\theta|^{b-1}+|\mathbf{c}|^{2}|\theta|^{b-3}\right) d \theta \quad\left(\theta:=\alpha^{\prime}-\alpha\right) \\
& \lesssim|\mathbf{c}|^{b} \leq C \epsilon^{b},
\end{aligned}
$$

where the last step follows from the fact that $|\mathbf{c}| \leq 2 \epsilon\|\gamma\|_{\infty}$. Now let us turn to $A_{12}$, which requires a more delicate estimate of $\mathbf{f}(\mathbf{y}, \mathbf{c})$. Let us break $A_{12}$ as

$$
A_{12}=\frac{L \gamma(\alpha)}{2 \pi} \int_{S}(\mathbf{f}(\mathbf{y}, \mathbf{c})-\mathbf{f}(\tilde{\mathbf{y}}, \mathbf{c})) d \alpha^{\prime}+\frac{L \gamma(\alpha)}{2 \pi} P V \int_{S} \mathbf{f}(\tilde{\mathbf{y}}, \mathbf{c}) d \alpha^{\prime}=: B_{1}+B_{2}
$$

For $B_{1}$, let us take the gradient of $\mathbf{f}(\mathbf{y}, \mathbf{c})$ (as in (Equation 4.2.29)) in the first variable. An elementary computation yields that

$$
\begin{equation*}
\left|\nabla_{\mathbf{x}} \mathbf{f}(\mathbf{x}, \mathbf{c})\right| \leq C|\mathbf{x}|^{-2} \min \left\{1, \frac{|\mathbf{c}|}{|\mathbf{x}|}\right\} \tag{4.2.30}
\end{equation*}
$$

as long as $\mathbf{x}$ satisfies

$$
\begin{equation*}
|\mathbf{x}+\mathbf{c}|^{2} \geq c_{0}\left(|\mathbf{x}|^{2}+|\mathbf{c}|^{2}\right) \tag{4.2.31}
\end{equation*}
$$

We point out that $\mathbf{x}=\xi \mathbf{y}+(1-\xi) \tilde{\mathbf{y}}$ indeed satisfies (Equation 4.2.31) for all $\xi \in[0,1]$ : to see this, in the proof of Lemma 4.2.1, if we replace $T_{1}$ in (Equation 4.2.3) by $\xi \mathbf{y}+(1-\xi) \tilde{\mathbf{y}}$, one can easily check the proof still goes through for $\xi \in[0,1]$. In addition, for any $\xi \in[0,1]$ we also have

$$
\begin{equation*}
|\xi \mathbf{y}+(1-\xi) \tilde{\mathbf{y}}| \geq c_{0}\left|\alpha-\alpha^{\prime}\right| . \tag{4.2.32}
\end{equation*}
$$

Thus the gradient estimate (Equation 4.2.30) together with (Equation 4.2.27) and (Equation 4.2.32) yields

$$
|f(\mathbf{y}, \mathbf{c})-f(\tilde{\mathbf{y}}, \mathbf{c})| \lesssim \min \left\{1,|\mathbf{c}|\left|\alpha-\alpha^{\prime}\right|^{-1}\right\} \lesssim \min \left\{1, \epsilon\left|\alpha-\alpha^{\prime}\right|^{-1}\right\}
$$

and plugging this into $B_{1}$ gives

$$
\left|B_{1}\right| \lesssim \epsilon+\int_{\epsilon<\left|\alpha-\alpha^{\prime}\right|<1} \epsilon\left|\alpha-\alpha^{\prime}\right|^{-1} d \alpha^{\prime} \lesssim \epsilon|\log \epsilon|
$$

As for $B_{2}$, using the definition of $\tilde{\mathbf{y}}$, the identity (Equation 4.2.28) and the fact that $\tilde{\mathbf{y}} \cdot \mathbf{c}=0$, we have

$$
\begin{aligned}
B_{2} & =\frac{L \gamma(\alpha)}{2 \pi} P V \int_{S}-\frac{\tilde{\mathbf{y}}^{\perp}}{|\tilde{\mathbf{y}}|^{2}} \frac{|\mathbf{c}|^{2}}{|\tilde{\mathbf{y}}+\mathbf{c}|^{2}} d \alpha^{\prime} \\
& =\frac{L \gamma(\alpha)|\mathbf{c}|^{2} \mathbf{n}(z(\alpha))}{2 \pi L} P V \int_{S} \frac{\alpha^{\prime}-\alpha}{\left|\alpha^{\prime}-\alpha\right|^{2}\left(L^{2}\left|\alpha^{\prime}-\alpha\right|^{2}+|\mathbf{c}|^{2}\right)} d \alpha^{\prime} .
\end{aligned}
$$

For the closed curves $i=1, \ldots, n$, we immediately have $B_{2}=0$ since $\alpha-\alpha^{\prime} \in\left[-\frac{1}{2}, \frac{1}{2}\right)$, and the integrand is an odd function of $\alpha^{\prime}-\alpha$.

For the open curves $i=n+1, \ldots, n+m$, the above integral becomes

$$
\begin{aligned}
B_{2} & =\frac{L \gamma(\alpha)|\mathbf{c}|^{2} \mathbf{n}(z(\alpha))}{2 \pi L} P V \int_{-\alpha}^{1-\alpha} \frac{\theta}{|\theta|^{2}\left(L^{2}|\theta|^{2}+|\mathbf{c}|^{2}\right)} d \theta \quad\left(\theta:=\alpha^{\prime}-\alpha\right) \\
& =\frac{L \gamma(\alpha)|\mathbf{c}|^{2} \mathbf{n}(z(\alpha))}{2 \pi L} \int_{\alpha}^{1-\alpha} \frac{\theta}{\theta^{2}\left(L^{2} \theta^{2}+|\mathbf{c}|^{2}\right)} d \theta
\end{aligned}
$$

where in the second inequality we used that the integral in $[-\alpha, \alpha]$ gives zero contribution to the principal value, since the integrand is odd.

Next we discuss two cases. If $\alpha>|\mathbf{c}|$, we bound the integrand by $C \theta^{-3}$, which gives

$$
\left|B_{2}\right| \leq C \gamma(\alpha)|\mathbf{c}|^{2} \alpha^{-2} \leq C|\mathbf{c}|^{2} \alpha^{b-2} \leq C|\mathbf{c}|^{b} \leq C \epsilon^{b}
$$

where the second inequality follows from the assumption $\gamma(0)=0$ for an open curve in (H3), as well as the Hölder continuity of $\gamma$. And if $0<\alpha \leq|\mathbf{c}|$, the integrand can be bounded above by $\theta^{-1}|\mathbf{c}|^{-2}$, which immediately leads to

$$
\left|B_{2}\right| \leq C \gamma(\alpha)|\log \alpha| \leq C|\mathbf{c}|^{b}|\log | \mathbf{c}| | \leq C \epsilon^{b}|\log \epsilon|
$$

In both cases we have $\left|B_{2}\right| \leq C \epsilon^{b}|\log \epsilon|$, and combining it with the $B_{1}$ and $A_{11}$ estimates gives (Equation 4.2.23).

- Estimate (Equation 4.2.24). We break $A_{2}$ into

$$
\begin{aligned}
A_{2} & =\frac{L \mathbf{c}^{\perp}}{2 \pi} \int_{S} \frac{\gamma\left(\alpha^{\prime}\right)-\gamma(\alpha)}{|\mathbf{y}+\mathbf{c}|^{2}} d \alpha^{\prime}+\frac{L \mathbf{c}^{\perp} \gamma(\alpha)}{2 \pi} \int_{S}\left(\frac{1}{|\mathbf{y}+\mathbf{c}|^{2}}-\frac{1}{|\tilde{\mathbf{y}}+\mathbf{c}|^{2}}\right) d \alpha^{\prime}+\frac{L \mathbf{c}^{\perp} \gamma(\alpha)}{2 \pi} \int_{S} \frac{1}{|\tilde{\mathbf{y}}+\mathbf{c}|^{2}} d \alpha^{\prime} \\
& =: A_{21}+A_{22}+A_{23} .
\end{aligned}
$$

For $A_{21}$, (Equation 4.2.26) and the Hölder continuity of $\gamma$ immediately lead to

$$
\begin{equation*}
\left|A_{21}\right| \leq C|\mathbf{c}| \int_{S} \frac{\left|\alpha-\alpha^{\prime}\right|^{b}}{\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}|^{2}} d \alpha^{\prime} \leq|\mathbf{c}|^{b} \leq C \epsilon^{b} \tag{4.2.33}
\end{equation*}
$$

For $A_{22}$, its integrand can be controlled as

$$
\left|\frac{1}{|\mathbf{y}+\mathbf{c}|^{2}}-\frac{1}{|\tilde{\mathbf{y}}+\mathbf{c}|^{2}}\right| \leq \frac{|\mathbf{y}-\tilde{\mathbf{y}}|(|\mathbf{y}+\mathbf{c}|+|\tilde{\mathbf{y}}+\mathbf{c}|)}{|\mathbf{y}+\mathbf{c}|^{2}|\tilde{\mathbf{y}}+\mathbf{c}|^{2}} \leq \frac{C\left|\alpha-\alpha^{\prime}\right|^{2}}{\left(\left|\alpha-\alpha^{\prime}\right|^{2}+|\mathbf{c}|^{2}\right)^{3 / 2}},
$$

where the last step follows from (Equation 4.2.26), (Equation 4.2.27) and (Equation 4.2.28). This allows us to control $A_{22}$ as

$$
\begin{equation*}
\left|A_{22}\right| \leq C|\mathbf{c}| \int_{-1}^{1} \frac{\theta^{2}}{\left(\theta^{2}+|\mathbf{c}|^{2}\right)^{3 / 2}} d \theta \leq C|\mathbf{c}||\log | \mathbf{c}| | \leq C \epsilon|\log \epsilon| \tag{4.2.34}
\end{equation*}
$$

Finally, for the $A_{23}$ term, (Equation 4.2.28) gives

$$
A_{23}=\frac{L \mathbf{c}^{\perp} \gamma(\alpha)}{2 \pi} \int_{S} \frac{1}{L^{2}\left|\alpha^{\prime}-\alpha\right|^{2}+|\mathbf{c}|^{2}} d \alpha^{\prime}=\frac{\mathbf{n}^{\perp}(\alpha) \gamma(\alpha)}{2 \pi} \int_{I} \frac{1}{\theta^{2}+1} d \theta \quad\left(\operatorname{set} \theta:=\frac{L\left(\alpha^{\prime}-\alpha\right)}{|\mathbf{c}|}\right)
$$

where the integration interval $I=\left(-\frac{L}{2|\mathbf{c}|}, \frac{L}{2|\mathbf{c}|}\right)$ for $i=1, \ldots, n$, and $I=\left(-\frac{L \alpha}{|\mathbf{c}|}, \frac{L(1-\alpha)}{|\mathbf{c}|}\right)$ for $i=n+1, \ldots, n+m$, and in the last equality we also used that $\frac{\mathbf{c}^{\perp}}{|\mathbf{c}|}=\mathbf{n}^{\perp}$. For $i=1, \ldots, n$, one can easily check that

$$
\left|\int_{I} \frac{1}{\theta^{2}+1} d \theta-\pi\right|=2 \int_{\frac{L}{2|\mathbf{c}|}}^{\infty} \frac{1}{\theta^{2}+1} d \theta \leq C|\mathbf{c}| \leq C \epsilon
$$

which immediately leads to

$$
\left|A_{23}-\frac{\mathbf{n}(z(\alpha))^{\perp} \gamma(\alpha)}{2}\right|=\left|\frac{\mathbf{n}^{\perp}(\alpha) \gamma(\alpha)}{2 \pi}\left(\int_{I} \frac{1}{\theta^{2}+1} d \theta-\pi\right)\right| \leq C \epsilon
$$

for $i=1, \ldots, n$. Next we turn to the open curves $i=n+1, \ldots, n+m$, and let us assume $\alpha \in\left[0, \frac{1}{2}\right]$
without loss of generality. In this case we have

$$
\left|\int_{I} \frac{1}{\theta^{2}+1} d \theta-\pi\right|=\int_{-\infty}^{-\frac{L \alpha}{|c|}} \frac{1}{\theta^{2}+1} d \theta+\int_{\frac{L(1-\alpha)}{|c|}}^{\infty} \frac{1}{\theta^{2}+1} d \theta \leq \min \left\{C \frac{|\mathbf{c}|}{\alpha}, \frac{\pi}{2}\right\}+C \epsilon
$$

where we used $1-\alpha>\frac{1}{2}$ to control the second integral by $C \epsilon$. Using the above inequality as well as the fact that $\gamma(\alpha) \leq C \alpha^{b}$ due to (H3), we have
$\left|A_{23}-\frac{\mathbf{n}(z(\alpha))^{\perp} \gamma(\alpha)}{2}\right|=\frac{\gamma(\alpha)}{2 \pi}\left|\int_{I} \frac{1}{\theta^{2}+1} d \theta-\pi\right| \leq C \alpha^{b} \min \left\{\frac{|\mathbf{c}|}{\alpha}, 1\right\}+C \epsilon \leq C\left(|\mathbf{c}|^{b}+\epsilon\right) \leq C \epsilon^{b}$
for $i=n+1, \ldots, n+m$. Finally, combining the $A_{23}$ estimates together with (Equation 4.2.33) and (Equation 4.2.34) yields (Equation 4.2.24).

### 4.3 Constructing a divergence-free perturbation

In this section, we aim to construct a divergence-free velocity field $\mathbf{u}^{\epsilon}: D^{\epsilon} \rightarrow \mathbb{R}^{2}$, such that $-\mathbf{u}^{\epsilon}$ tends to make each $D_{i}^{\epsilon}$ "more symmetric". Let $\mathbf{u}^{\epsilon}: D^{\epsilon} \rightarrow \mathbb{R}^{2}$ be given by

$$
\begin{equation*}
\mathbf{u}^{\epsilon}:=x+\nabla p^{\epsilon} \quad \text { in } D^{\epsilon}, \tag{4.3.1}
\end{equation*}
$$

where the function $p^{\epsilon}: \overline{D^{\epsilon}} \rightarrow \mathbb{R}$ is chosen such that

$$
\begin{equation*}
\nabla \cdot \mathbf{u}^{\epsilon}=0 \quad \text { in } D^{\epsilon}, \tag{4.3.2}
\end{equation*}
$$

and on each connected component $l$ of $\partial D^{\epsilon}, u^{\epsilon}$ satisfies

$$
\begin{equation*}
\int_{l} \mathbf{u}^{\epsilon} \cdot n d \sigma=0 \tag{4.3.3}
\end{equation*}
$$

where $n$ is the unit normal of $l$ pointing outwards of $D^{\epsilon}$. Note that $\partial D^{\epsilon}$ has a total of $2 n+$ $m$ connected components: $D_{i}^{\epsilon}$ is doubly-connected for $i=1, \ldots, n$ (denote its outer and inner boundaries by $\partial D_{i, \text { out }}^{\epsilon}$ and $\partial D_{i, \text { in }}^{\epsilon}$; note that $\partial D_{i, \text { in }}^{\epsilon}$ coincides with $\Gamma_{i}$, whereas it is simply-connected for $i=n+1, \ldots, n+m$ (denote its boundary by $\partial D_{i}^{\epsilon}$ ).

Next we show that there indeed exists a function $p^{\epsilon}$ so that $\mathbf{u}^{\epsilon}$ satisfies (Equation 4.3.2)(Equation 4.3.3). Clearly, (Equation 4.3.2) requires that $p^{\epsilon}$ satisfies

$$
\begin{equation*}
\Delta p^{\epsilon}=-2 \quad \text { in } D^{\epsilon} \tag{4.3.4}
\end{equation*}
$$

As for the boundary conditions, we let

$$
\begin{equation*}
\left.p^{\epsilon}\right|_{\partial D_{i}^{\epsilon}}=0 \quad \text { for } i=n+1, \ldots, n+m, \tag{4.3.5}
\end{equation*}
$$

so the divergence theorem yields that (Equation 4.3.3) is satisfied for each $l=\partial D_{i}^{\epsilon}$ for $i=n+$ $1, \ldots, n+m$. As for $i=1, \ldots, n$, we define

$$
p^{\epsilon}=\left\{\begin{array}{ll}
0 & \text { on } \partial D_{i, \mathrm{out}}^{\epsilon}  \tag{4.3.6}\\
c_{i}^{\epsilon} & \text { on } \partial D_{i, \text { in }}^{\epsilon}=\Gamma_{i}
\end{array} \quad \text { for } i=1, \ldots, n\right.
$$

where $c_{i}^{\epsilon}>0$ is the unique constant such that

$$
\begin{equation*}
\int_{\partial U_{i}} \nabla p^{\epsilon} \cdot n d \sigma=-2\left|U_{i}\right| \quad \text { for } i=1, \ldots, n \tag{4.3.7}
\end{equation*}
$$

where $U_{i}$ is the domain enclosed by $\partial D_{i, \text { in }}^{\epsilon}=\Gamma_{i}$ (thus $U_{i}$ is independent of $\epsilon$ ), and $n$ is the outer normal of $U_{i}$ (thus the inner normal of $D_{i}^{\epsilon}$ ). The existence of $c_{i}^{\epsilon}$ is guaranteed by [50, Lemma 2.5]. One can then check that $\int_{\partial U_{i}} \mathbf{u}^{\epsilon} \cdot n d \sigma=0$. Applying the divergence theorem in $D_{i}^{\epsilon}$ then gives us that $\int_{\partial D_{i, \text { out }}^{\epsilon}} \mathbf{u}^{\epsilon} \cdot n d \sigma=0$ as well.

In [50] we proved a rearrangement inequality for such $p^{\epsilon}$ in a similar spirit of Talenti's rearrangement inequality for elliptic equations [109], which we state below.

Lemma 4.3.1 ([50, Proposition 2.6]). The function $p^{\epsilon}: \overline{D^{\epsilon}} \rightarrow \mathbb{R}$ defined in (Equation 4.3.4)(Equation 4.3.7) satisfies the following in each $D_{i}^{\epsilon}$ for $i=1, \ldots, n+m$ :

$$
\begin{equation*}
\sup _{D_{i}^{\epsilon}} p^{\epsilon} \leq \frac{\left|D_{i}^{\epsilon}\right|}{2 \pi} \tag{4.3.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{D_{i}^{\epsilon}} p^{\epsilon}(x) d x \leq \frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi} \tag{4.3.9}
\end{equation*}
$$

Moreover, each inequality above achieves equality if and only $D_{i}^{\epsilon}$ is either a disk or an annulus.
Note that the inequalities (Equation 4.3.8)-(Equation 4.3.9) hold for any domain with $C^{1, \alpha}$ boundary. Even though the inequalities are strict when $D_{i}^{\epsilon}$ is non-radial, they are not strong enough to rule out non-radial vortex sheets, as we need quantitative versions of strict inequalities that are still valid in the $\epsilon \rightarrow 0^{+}$limit. As we will see in the proof of Proposition 4.4.2, the key step is to show that if some $\Gamma_{i}$ is either not a circle or does not have a constant $\gamma_{i}$, then the following quantitative version of (Equation 4.3.9) holds: $\epsilon^{-2}\left(\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi}-\int_{D_{i}^{\epsilon}} p^{\epsilon}(x) d x\right) \geq c_{0}>0$, where $c_{0}$ is independent of $\epsilon$.

In order to upgrade (Equation 4.3.9) into a quantitative version, we need to obtain some fine estimates for $p^{\epsilon}$ that take into account the shape of the thin domains $D_{i}^{\epsilon}$. For $i=n+1, \ldots, n+m$, since $p^{\epsilon}=0$ on $\partial D_{i}^{\epsilon}$, and the domain $D_{i}^{\epsilon}$ is a thin simply-connected domain with width $\epsilon \ll 1$, intuitively one would expect that $\left|p^{\epsilon}\right| \leq C \epsilon^{2}$. The next proposition shows that this crude estimate is indeed true, and its proof is postponed to Section subsection 4.3.1.

Proposition 4.3.2. For any $i=n+1, \ldots, n+m$, let $p^{\epsilon}: \overline{D_{i}^{\epsilon}} \rightarrow \mathbb{R}$ be given by (Equation 4.3.4)(Equation 4.3.5). Then there exist $\epsilon_{1}$ and $C$ only depending on $\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{L^{\infty}\left(S_{i}\right)}$ and $F_{\Gamma}$, such
that

$$
\left|p^{\epsilon}\right| \leq C \epsilon^{2} \quad \text { in } D_{i}^{\epsilon}
$$

for all $\epsilon \in\left(0, \epsilon_{1}\right)$.

For $i=1, \ldots, n$, the estimate is more involved, since $p^{\epsilon}$ takes different values $c_{i}^{\epsilon}$ and 0 on the inner and outer boundaries of $D_{i}^{\epsilon}$. Heuristically speaking, since $D_{i}^{\epsilon}$ is a doubly-connected thin tubular domain with width $\sim \epsilon$, we would expect that $p_{i}^{\epsilon}$ (in $\alpha, \eta$ coordinate) changes almost linearly from 0 to $c_{i}^{\epsilon}$ as $\eta$ goes from -1 (outer boundary) to 0 (inner boundary). Next we will show that the error between $p^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)$ and the linear-in- $\eta$ function $c_{i}^{\epsilon}(1+\eta)$ is indeed controlled by $O\left(\epsilon^{2}\right)$. We will also obtain fine estimates of the gradient of the function $c_{i}^{\epsilon}(1+\eta)$, as well as the boundary value $c_{i}^{\epsilon}$. Again, its proof is postponed to Section subsection 4.3.1.

Proposition 4.3.3. For any $i=1, \ldots, n$, let $p^{\epsilon}: \overline{D_{i}^{\epsilon}} \rightarrow \mathbb{R}$ and $c_{i}^{\epsilon} \in \mathbb{R}$ be given by (Equation 4.3.4) and (Equation 4.3.6)-(Equation 4.3.7). For such $p^{\epsilon}$, let us define $\tilde{p}^{\epsilon}, q^{\epsilon}: \overline{D_{i}^{\epsilon}} \mapsto \mathbb{R}$ as follows:

$$
\begin{align*}
\tilde{p}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right) & :=c_{i}^{\epsilon}(1+\eta) \quad \text { for } \alpha \in S_{i}, \eta \in[0,-1],  \tag{4.3.10}\\
q^{\epsilon} & :=p^{\epsilon}-\tilde{p}^{\epsilon} \quad \text { in } \overline{D_{i}^{\epsilon}} .
\end{align*}
$$

Also let

$$
\begin{equation*}
\beta_{i}:=\frac{2\left|U_{i}\right|}{L_{i} \int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha} . \tag{4.3.11}
\end{equation*}
$$

Then there exist $\epsilon_{1}$ and $C$ only depending on $\left\|z_{i}\right\|_{C^{3}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{C^{2}\left(S_{i}\right)}$ and $F_{\Gamma}$, such that for all $\epsilon \in$
$\left(0, \epsilon_{1}\right)$ we have the following:

$$
\begin{align*}
& \begin{cases}\left|q^{\epsilon}\right| \leq C \epsilon^{2} & \text { in } D_{i}^{\epsilon}, \\
q^{\epsilon}=0 & \text { on } \partial D_{i}^{\epsilon},\end{cases}  \tag{4.3.12}\\
& \left|\frac{c_{i}^{\epsilon}}{\epsilon}-\beta_{i}\right| \leq C \epsilon,  \tag{4.3.13}\\
& \left|\nabla \tilde{p}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-\frac{\beta_{i}}{\gamma_{i}(\alpha)} \mathbf{n}\left(z_{i}(\alpha)\right)\right| \leq C \epsilon \quad \text { for } \alpha \in S_{i}, \eta \in[0,-1] \text {. } \tag{4.3.14}
\end{align*}
$$

### 4.3.1 Proof of the quantitative lemmas for $p^{\epsilon}$

In this subsection we aim to prove Propositions 4.3.2 and 4.3.3. We start with a technical lemma on estimating the solution of Poisson's equation (with zero boundary condition) in the domain $D_{i}^{\epsilon}$.

Lemma 4.3.4. For any $i=1, \ldots, n+m$, assume $\Gamma_{i}$ and $\gamma_{i}$ satisfy (H1)-(H3). Let $v^{\epsilon} \in C^{2}\left(D_{i}^{\epsilon}\right) \cap$ $C\left(\overline{D_{i}^{\epsilon}}\right)$ solve the Poisson's equation with zero boundary condition:

$$
\begin{cases}\Delta v^{\epsilon}=-1 & \text { in } D_{i}^{\epsilon}  \tag{4.3.15}\\ v^{\epsilon}=0 & \text { on } \partial D_{i}^{\epsilon}\end{cases}
$$

Then there exist positive constants $\epsilon_{0}=C\left(\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{L^{\infty}\left(S_{i}\right)}, F_{\Gamma}\right)$ and $C_{1}, C_{2}=$ $C\left(\left\|\gamma_{i}\right\|_{L^{\infty}\left(S_{i}\right)}\right)$, such that for all $\epsilon \in\left(0, \epsilon_{0}\right)$ we have

$$
\begin{equation*}
0 \leq v^{\epsilon} \leq C_{1} \epsilon^{2} \quad \text { in } D_{i}^{\epsilon} \tag{4.3.16}
\end{equation*}
$$

and

$$
\begin{equation*}
\left\|\nabla v^{\epsilon}\right\|_{L^{\infty}\left(\Gamma_{i}\right)} \leq C_{2} \epsilon \quad \text { for } i=1, \ldots, n . \tag{4.3.17}
\end{equation*}
$$

Proof. Throughout the proof, let $i \in\{1, \ldots, n+m\}$ be fixed. For notational simplicity, in the rest of the proof we omit the subscript $i$ in $R_{i}^{\epsilon}, D_{i}^{\epsilon}, S_{i}, z_{i}$ and $\gamma_{i}$.

Step 1. We start with a simple geometric result that $D^{\epsilon}$ is "flat" in a small neighborhood of any $z(\alpha)$. For any $\alpha \in S$, let $V^{\epsilon}(\alpha):=D^{\epsilon} \cap B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$, where $\|\cdot\|_{\infty}$ denotes $\|\cdot\|_{L^{\infty}(S)}$. We will show that any $y \in V^{\epsilon}(\alpha)$ satisfies

$$
\begin{equation*}
|(z(\alpha)-y) \cdot \mathbf{n}(z(\alpha))| \leq 2 \epsilon\|\gamma\|_{\infty} \tag{4.3.18}
\end{equation*}
$$

for all sufficiently small $\epsilon>0$ (to be quantified in (Equation 4.3.23)). See Figure Figure 4.4(a) for an illustration.

Since $y \in V^{\epsilon}(\alpha) \subset D^{\epsilon}$, there exist $\beta \in S$ and $\eta \in(-1,0)$ such that $y=R^{\epsilon}(\beta, \eta)=$ $z(\beta)+\epsilon \gamma(\beta) \mathbf{n}(z(\beta)) \eta$. It follows that

$$
\begin{align*}
|(z(\alpha)-y) \cdot \mathbf{n}(z(\alpha))| & \leq|(z(\alpha)-z(\beta)) \cdot \mathbf{n}(z(\alpha))|+\epsilon\|\gamma\|_{\infty}  \tag{4.3.19}\\
& \leq\left\|z^{\prime \prime}\right\|_{\infty}(\alpha-\beta)^{2}+\epsilon\|\gamma\|_{\infty}
\end{align*}
$$

where in the second inequality we used

$$
\begin{equation*}
\left|(z(\alpha)-z(\beta))-z^{\prime}(\alpha)(\alpha-\beta)\right| \leq\left\|z^{\prime \prime}\right\|_{\infty}(\alpha-\beta)^{2} \tag{4.3.20}
\end{equation*}
$$

and $z^{\prime}(\alpha) \cdot \mathbf{n}(z(\alpha))=0$. To bound $\alpha-\beta$ on the right hand side of (Equation 4.3.19), the fact that $y \in B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$ gives

$$
\begin{equation*}
6 \epsilon\|\gamma\|_{\infty} \geq|z(\alpha)-y| \geq|z(\alpha)-z(\beta)|-\epsilon \gamma(\beta) \tag{4.3.21}
\end{equation*}
$$

which implies $|z(\alpha)-z(\beta)| \leq 7 \epsilon\|\gamma\|_{\infty}$. Since the arc-chord constant $F_{\Gamma}$ given in (Equation 4.1.2) is finite, this implies

$$
\begin{equation*}
|\alpha-\beta| \leq 7 F_{\Gamma}\|\gamma\|_{\infty} \epsilon \tag{4.3.22}
\end{equation*}
$$

Plugging this into the right hand side of (Equation 4.3.19), we know (Equation 4.3.18) holds for


Figure 4.4: (a) In Step 1, $V^{\epsilon}(\alpha)$ (the yellow set) must lie between the two dashed lines for small $\epsilon$. (b) In Step 2, $\partial V^{\epsilon}\left(\alpha_{0}\right)$ is decomposed into $\partial V_{1}^{\epsilon}\left(\alpha_{0}\right)$ (in dark green) and $\partial V_{2}^{\epsilon}\left(\alpha_{0}\right)$ (in purple).
all

$$
\begin{equation*}
0<\epsilon \leq\left(49\left\|z^{\prime \prime}\right\|_{\infty} F_{\Gamma}^{2}\|\gamma\|_{\infty}\right)^{-1} \tag{4.3.23}
\end{equation*}
$$

Step 2. Next we prove (Equation 4.3.16). Note that $v^{\epsilon}$ is superharmonic in $D^{\epsilon}$ and vanishes on the boundary, thus it follows from the maximum principle that $v^{\epsilon} \geq 0$ in $D^{\epsilon}$. Denote $M:=$ $\max _{x \in D^{\epsilon}} v(x)$, and pick $x_{0}=R\left(\alpha_{0}, \eta_{0}\right) \in D^{\epsilon}$ such that $v\left(x_{0}\right)=M$. Without loss of generality, we can assume that $z\left(\alpha_{0}\right)=(0,0)$ and $\mathbf{s}\left(z\left(\alpha_{0}\right)\right)=\mathbf{e}_{1}:=(1,0)$, so that $\mathbf{n}\left(z\left(\alpha_{0}\right)\right)=(0,1)$ and $x_{0}=\left(0, \epsilon \gamma\left(\alpha_{0}\right) \eta_{0}\right)$. Let us consider a barrier function $b_{1}: \mathbb{R}^{2} \mapsto \mathbb{R}$ given by

$$
b_{1}\left(x_{1}, x_{2}\right)=x_{2}^{2}-\frac{x_{1}^{2}}{2}
$$

Clearly $\Delta b_{1}=1$, so $v^{\epsilon}+b_{1}$ is harmonic in $D^{\epsilon}$. It then follows from the maximum principle that $\max _{\overline{V^{\epsilon}\left(\alpha_{0}\right)}}\left(v^{\epsilon}+b_{1}\right)$ is achieved at some boundary point $\tilde{x}_{0} \in \partial V^{\epsilon}\left(\alpha_{0}\right)$. Let us break $\partial V^{\epsilon}\left(\alpha_{0}\right)$ into $\partial V_{1}^{\epsilon}\left(\alpha_{0}\right) \cup \partial V_{2}^{\epsilon}\left(\alpha_{0}\right)$ (see Figure Figure 4.4(b) for an illustration), given by

$$
\begin{equation*}
\partial V_{1}^{\epsilon}\left(\alpha_{0}\right):=\partial D^{\epsilon} \cap B_{6 \epsilon\|\gamma\|_{\infty}}\left(z\left(\alpha_{0}\right)\right), \quad \partial V_{2}^{\epsilon}\left(\alpha_{0}\right):=\overline{D^{\epsilon}} \cap \partial B_{6 \epsilon\|\gamma\|_{\infty}}\left(z\left(\alpha_{0}\right)\right) . \tag{4.3.24}
\end{equation*}
$$

We claim that $\tilde{x}_{0} \in \partial V_{1}^{\epsilon}\left(\alpha_{0}\right)$. To see this, note that any $y=\left(y_{1}, y_{2}\right) \in \partial V_{2}^{\epsilon}\left(\alpha_{0}\right)$ satisfies $|y|=$ $6 \epsilon\|\gamma\|_{\infty}$ and $\left|y_{2}\right| \leq 2 \epsilon\|\gamma\|_{\infty}$, where the latter follows from (Equation 4.3.18) and our assumption that $\mathbf{s}\left(z\left(\alpha_{0}\right)\right)=\mathbf{e}_{1}$. This implies that $\left|y_{1}\right| \geq 4 \epsilon\|\gamma\|_{\infty}>\left|y_{2}\right|$, thus $b_{1}(y)<0$. Using that $v^{\epsilon}\left(x_{0}\right)=$ $M \geq v^{\epsilon}(y)$ and $b_{1}\left(x_{0}\right)=b_{1}\left(0, \epsilon \gamma\left(\alpha_{0}\right) \eta_{0}\right) \geq 0$, we have $\left(v^{\epsilon}+b_{1}\right)(y)<\left(v^{\epsilon}+b_{1}\right)\left(x_{0}\right)$. This shows that $\max _{\overline{V^{\epsilon}\left(\alpha_{0}\right)}}\left(v^{\epsilon}+b_{1}\right)$ cannot be achieved on $\partial V_{2}^{\epsilon}\left(\alpha_{0}\right)$, finishing the proof of the claim.

Since $\tilde{x}_{0} \in \partial V_{1}^{\epsilon}\left(\alpha_{0}\right) \subset \partial D^{\epsilon}$, the boundary condition in (Equation 4.3.15) yields that $v^{\epsilon}\left(\tilde{x}_{0}\right)=$ 0 . Thus

$$
M+b_{1}\left(x_{0}\right)=v^{\epsilon}\left(x_{0}\right)+b_{1}\left(x_{0}\right) \leq v^{\epsilon}\left(\tilde{x}_{0}\right)+b_{1}\left(\tilde{x}_{0}\right)=b_{1}\left(\tilde{x}_{0}\right) .
$$

Using $b_{1}\left(x_{0}\right)=b_{1}\left(0, \epsilon \gamma\left(\alpha_{0}\right) \eta_{0}\right) \geq 0$, the above inequality becomes

$$
\begin{equation*}
M \leq b_{1}\left(\tilde{x}_{0}\right) \leq\left|\tilde{x}_{0}\right|^{2} \leq 36\|\gamma\|_{\infty}^{2} \epsilon^{2}, \tag{4.3.25}
\end{equation*}
$$

where the second inequality follows from the definition of $b_{1}$. This proves (Equation 4.3.16) for $C_{1}=36\|\gamma\|_{\infty}^{2}$.

Step 3. It remains to prove (Equation 4.3.17). First note that for $i \in\{1, \ldots, n\}$, the assumptions (H1)-(H3) yield that $D_{i}^{\epsilon}$ has $C^{2}$ boundary, therefore $v^{\epsilon} \in C^{2}\left(D_{i}^{\epsilon}\right) \cap C^{1}\left(\overline{D_{i}^{\epsilon}}\right)$. Let us fix $i \in\{1, \ldots, n\}$ and any $\alpha \in S$, and we aim to show that $\left|\nabla v^{\epsilon}(z(\alpha))\right| \leq C_{2} \epsilon$. Again, without loss of generality we can assume that $z(\alpha)=(0,0)$ and $\mathbf{s}(z(\alpha))=\mathbf{e}_{1}$. Let us consider a new barrier function $b_{2}: \mathbb{R}^{2} \rightarrow \mathbb{R}$

$$
\begin{equation*}
b_{2}\left(x_{1}, x_{2}\right):=x_{2}^{2}+4 \epsilon\|\gamma\|_{\infty} x_{2}-\frac{x_{1}^{2}}{2} \tag{4.3.26}
\end{equation*}
$$

which satisfies $b_{2}(0,0)=0$, and one can easily check that its zero level set has horizontal tangent at $(0,0)$ (thus tangent to $\partial D^{\epsilon}$ at $z(\alpha)$ ).

Again, let us decompose $\partial V^{\epsilon}(\alpha)$ as $\partial V_{1}^{\epsilon}(\alpha) \cup \partial V_{2}^{\epsilon}(\alpha)$ as in (Equation 4.3.24) (except that $\alpha_{0}$
now becomes $\alpha$ ). We claim that for all sufficiently small $\epsilon>0$, the new barrier function $b_{2}$ satisfies

$$
\begin{array}{ll}
\Delta b_{2}=1 & \text { in } V^{\epsilon}(\alpha), \\
b_{2} \leq 0 & \text { on } \partial V_{1}^{\epsilon} \\
b_{2} \leq-\epsilon^{2} & \text { on } \partial V_{2}^{\epsilon} \tag{4.3.29}
\end{array}
$$

Let us assume for a moment that (Equation 4.3.27)-(Equation 4.3.29) are true. Then it follows that

$$
\begin{equation*}
v^{\epsilon}+C_{2} b_{2} \leq 0 \text { in } V^{\epsilon}(\alpha), \tag{4.3.30}
\end{equation*}
$$

where $C_{2}:=\max \left\{1, C_{1}\right\}$ and $C_{1}$ is as in (Equation 4.3.16) (in the end of step 2 we have $C_{1}=36\|\gamma\|_{\infty}^{2}$ ). To show (Equation 4.3.30), note that $v^{\epsilon}+C_{2} b_{2}$ is subharmonic in $V^{\epsilon}(\alpha)$ due to (Equation 4.3.27) and the definition of $C_{2}$, thus its maximum is attained on its boundary. The boundary conditions in (Equation 4.3.15) and (Equation 4.3.28) yield that $v^{\epsilon}+C_{2} b_{2} \leq 0$ on $\partial V_{1}^{\epsilon}(\alpha)$; whereas (Equation 4.3.16), (Equation 4.3.29) and the definition of $C_{2}$ yield that $v^{\epsilon}+C_{2} b_{2} \leq 0$ on $\partial V_{2}^{\epsilon}(\alpha)$. Thus $v^{\epsilon}+C_{2} b_{2} \leq 0$ on $\partial V_{1}^{\epsilon}(\alpha) \cup \partial V_{2}^{\epsilon}(\alpha)$, implying (Equation 4.3.30).

However, $v^{\epsilon}+C_{2} b_{2}$ is actually zero at $z(\alpha) \in \partial V^{\epsilon}(\alpha)$, therefore Hopf's Lemma implies that $\nabla\left(v^{\epsilon}+C_{2} b_{2}\right)(z(\alpha)) \cdot \vec{n}(z(\alpha))>0$, where $\vec{n}(z(\alpha))$ is the outer normal of $\partial D^{\epsilon}$ at $z(\alpha)$. Hence

$$
\begin{equation*}
\left|\nabla v^{\epsilon}(z(\alpha))\right|=-\nabla v^{\epsilon}(z(\alpha)) \cdot \vec{n}(z(\alpha))<C_{2} \nabla b_{2}(z(\alpha)) \cdot \vec{n}(z(\alpha))=4 C_{2}\|\gamma\|_{\infty} \epsilon, \tag{4.3.31}
\end{equation*}
$$

where the first equality follows from the fact that $v^{\epsilon}$ is superharmonic in $D^{\epsilon}$ and constant on $\partial D^{\epsilon}$, and the second equality is a direct computation of $\nabla b_{2}$. Thus (Equation 4.3.31) proves (Equation 4.3.17).

To complete the proof, we only need to prove (Equation 4.3.27)-(Equation 4.3.29) for small $\epsilon>0$. Note that (Equation 4.3.27) follows immediately from computing the Laplacian of $b_{2}$. For (Equation 4.3.28), let us pick $y \in \partial V_{1}^{\epsilon}(\alpha)$, and we aim to show that $b_{2}(y) \leq 0$. Note that
$y=R^{\epsilon}(\beta, 0)$ or $R^{\epsilon}(\beta,-1)$ for some $\beta \in S$. We first deal with the first case.
Let us denote $y=\left(y_{1}, y_{2}\right)$. Rewriting (Equation 4.3.20) into two inequalities for the two components, and using that $z(\alpha)=(0,0)$ and $z^{\prime}(\alpha)=L \mathbf{e}_{1}$ ( $L$ is the length of the curve $\Gamma_{i}$ ), we have

$$
\begin{align*}
& \left|0-y_{1}-L(\alpha-\beta)\right| \leq\left\|z^{\prime \prime}\right\|_{\infty}(\alpha-\beta)^{2}  \tag{4.3.32}\\
& \left|y_{2}\right|=\left|0-y_{2}\right| \leq\left\|z^{\prime \prime}\right\|_{\infty}(\alpha-\beta)^{2} . \tag{4.3.33}
\end{align*}
$$

Also, (Equation 4.3.22) gives $|\alpha-\beta| \leq 7 F_{\Gamma}\|\gamma\|_{\infty} \epsilon$. Applying it to (Equation 4.3.32), for all $\epsilon>0$ sufficiently small we have that

$$
\begin{equation*}
\left|y_{1}\right| \geq \frac{L}{2}|\alpha-\beta| . \tag{4.3.34}
\end{equation*}
$$

Plugging (Equation 4.3.34) and (Equation 4.3.33) into $b_{2}(y)=-\frac{1}{2} y_{1}^{2}+y_{2}^{2}+4 \epsilon\|\gamma\|_{\infty} y_{2}$, we have

$$
\begin{aligned}
b_{2}(y) & \leq-\frac{L^{2}}{8}(\alpha-\beta)^{2}+\left\|z^{\prime \prime}\right\|_{\infty}^{2}(\alpha-\beta)^{4}+4 \epsilon\|\gamma\|_{\infty}\left\|z^{\prime \prime}\right\|_{\infty}(\alpha-\beta)^{2} \\
& \leq\left(-\frac{L^{2}}{8}+C \epsilon^{2}+C \epsilon\right)(\alpha-\beta)^{2} \leq 0
\end{aligned}
$$

for all $\epsilon>0$ sufficiently small, where the second inequality follows from (Equation 4.3.22). This finishes the proof of (Equation 4.3.28) for the case $y=R^{\epsilon}(\beta, 0)$.

Before we deal with the case $y=R^{\epsilon}(\beta,-1)$, let us prove (Equation 4.3.29) first. For any $y=$ $\left(y_{1}, y_{2}\right) \in \partial V_{2}^{\epsilon}(\alpha)$, (Equation 4.3.18) gives $\left|y_{2}\right| \leq 2 \epsilon\|\gamma\|_{\infty}$. Combining this with $|y|=6 \epsilon\|\gamma\|_{\infty}$ yields $\left|y_{1}\right| \geq \sqrt{32} \epsilon\|\gamma\|_{\infty}$. Thus

$$
b_{2}(y) \leq\left(2 \epsilon\|\gamma\|_{\infty}\right)^{2}+4 \epsilon\|\gamma\|_{\infty}\left(2 \epsilon\|\gamma\|_{\infty}\right)-\frac{\left(\sqrt{32} \epsilon\|\gamma\|_{\infty}\right)^{2}}{2} \leq-4 \epsilon^{2}\|\gamma\|_{\infty}^{2}
$$

Finally we turn to the proof of (Equation 4.3.28) for the case $y=R^{\epsilon}(\beta,-1)$. Note that the curve $\left\{R^{\epsilon}(\beta,-1): \beta \in S\right\} \cap B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$ lies in the interior of the region bounded by $\Gamma \cap$
$B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$ on the top, $\partial B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$ on the sides, and $y_{2}=-2 \epsilon\|\gamma\|_{\infty}$ on the bottom. (The last one follows from (Equation 4.3.18) and our assumption that $\mathbf{s}(z(\alpha))=\mathbf{e}_{1}$ ). We have already shown $b_{2} \leq 0$ on $\Gamma \cap B_{6 \epsilon\|\gamma\|_{\infty}}(z(\alpha))$ and the lateral boundaries, and it is easy to check that $b_{2} \leq 0$ on $y_{2}=-2 \epsilon\|\gamma\|_{\infty}$. Since the set $\left\{b_{2} \leq 0\right\}$ is simply-connected, it implies that $b_{2} \leq 0$ in the interior of this region, finishing the proof.

Note that (Equation 4.3.16) of Lemma 4.3.4 immediately implies Proposition 4.3.2. (The only difference is that $\Delta v^{\epsilon}=-1$ in Lemma 4.3.4 whereas $\Delta p^{\epsilon}=-2$ in Proposition 4.3.2, so the constant $C$ in Proposition 4.3 .2 is twice of that in (Equation 4.3.16)). The lemma also implies the following corollary, which will be helpful in the proof of Proposition 4.3.3.

Corollary 4.3.5. For any $i=1, \ldots, n+m$, assume $\Gamma_{i}$ and $\gamma_{i}$ satisfy (H1)-(H3). Assume $v^{\epsilon} \in$ $C^{2}\left(D_{i}^{\epsilon}\right) \cap C\left(\overline{\overline{D_{i}}}\right)$ satisfies that

$$
\begin{cases}\left|\Delta v^{\epsilon}\right| \leq C_{0} & \text { in } D_{i}^{\epsilon} \\ v^{\epsilon}=0 & \text { on } \partial D_{i}^{\epsilon}\end{cases}
$$

for some constant $C_{0}>0$. Then for the same constants $\epsilon_{0}, C_{1}, C_{2}$ as in Lemma 4.3.4, the following holds for all $\epsilon \in\left(0, \epsilon_{0}\right)$ :

$$
\begin{equation*}
\left|v^{\epsilon}\right| \leq C_{0} C_{1} \epsilon^{2} \quad \text { in } D_{i}^{\epsilon}, \tag{4.3.35}
\end{equation*}
$$

and if $v^{\epsilon} \in C^{2}\left(D_{i}^{\epsilon}\right) \cap C^{1}\left(\overline{D_{i}^{\epsilon}}\right)$, we have

$$
\begin{equation*}
\left\|\nabla v^{\epsilon}\right\|_{L^{\infty}\left(\Gamma_{i}\right)} \leq C_{0} C_{2} \epsilon \quad \text { for } \quad i=1, \ldots, n . \tag{4.3.36}
\end{equation*}
$$

Proof. Let $\tilde{v}$ be a solution to

$$
\begin{cases}\Delta \tilde{v}=-C_{0} & \text { in } D_{i}^{\epsilon} \\ \tilde{v}=0 & \text { on } \partial D_{i}^{\epsilon}\end{cases}
$$

It is clear that $v^{\epsilon}+\tilde{v}$ is super-harmonic and $v^{\epsilon}-\tilde{v}$ is sub-harmonic in $D_{i}^{\epsilon}$, and they both vanish on
the boundary. Thus the maximum principle implies that

$$
\begin{equation*}
-\tilde{v} \leq v^{\epsilon} \leq \tilde{v} \quad \text { in } D_{i}^{\epsilon} \tag{4.3.37}
\end{equation*}
$$

Applying (Equation 4.3.16) of Lemma 4.3 .4 to $\frac{\tilde{v}}{C_{0}}$, we obtain $0 \leq \tilde{v} \leq C_{0} C_{1} \epsilon^{2}$ in $D_{i}^{\epsilon}$ for all $\epsilon \in\left(0, \epsilon_{0}\right)$, leading to (Equation 4.3.35). Furthermore, (Equation 4.3.37) and the fact that $v^{\epsilon}$ and $v$ both have zero boundary condition imply that

$$
\left|\nabla v^{\epsilon}\right| \leq|\nabla \tilde{v}| \quad \text { on } \partial D_{i}^{\epsilon}
$$

We then apply (Equation 4.3.17) of Lemma 4.3 .4 to $\frac{\tilde{v}}{C_{0}}$ and obtain $\left\|\nabla v^{\epsilon}\right\|_{L^{\infty}\left(\Gamma_{i}\right)} \leq C_{0} C_{2} \epsilon$, which proves (Equation 4.3.36).

Now we are ready to prove Proposition 4.3.3.

Proof of Proposition 4.3.3. Throughout the proof, let $i \in\{1, \ldots, n\}$ be fixed. For notational simplicity, in the rest of the proof we omit the subscript $i$ from all terms.

We claim that

$$
\begin{align*}
& \left|\nabla \tilde{p}^{\epsilon}\left(R^{\epsilon}(\alpha, \eta)\right)-\frac{c^{\epsilon}}{\epsilon \gamma(\alpha)} \mathbf{n}(z(\alpha))\right| \leq C \epsilon \quad \text { for all } \alpha \in S, \eta \in[0,-1]  \tag{4.3.38}\\
& \left\|\Delta q^{\epsilon}\right\|_{L^{\infty}\left(D^{\epsilon}\right)} \leq C \tag{4.3.39}
\end{align*}
$$

for some constant $C>0$ only depending on $\left\|z_{i}\right\|_{C^{3}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{C^{2}\left(S_{i}\right)}$ and $F_{\Gamma}$. Assuming these are true, let us explain how they lead to (Equation 4.3.12)-(Equation 4.3.14). By (Equation 4.3.6) and (Equation 4.3.10), $p^{\epsilon}$ and $\tilde{p}^{\epsilon}$ have the same boundary condition, thus $q^{\epsilon}=0$ on $\partial D^{\epsilon}$. This and (Equation 4.3.39) allow us to apply Corollary 4.3 .5 to $q^{\epsilon}$ to obtain the estimate (Equation 4.3.35), implying (Equation 4.3.12).

Due to (Equation 4.3.36) of Corollary 4.3.5, we also have

$$
\begin{equation*}
\left\|\nabla q^{\epsilon}\right\|_{L^{\infty}(\Gamma)} \leq C \epsilon \tag{4.3.40}
\end{equation*}
$$

Using (Equation 4.3.7) and $p^{\epsilon}=\tilde{p}^{\epsilon}+q^{\epsilon}$, we have

$$
\begin{aligned}
-2|U| & =\int_{\partial U} \nabla \tilde{p}^{\epsilon} \cdot n d \sigma+\int_{\partial U} \nabla q^{\epsilon} \cdot n d \sigma \\
& =-\frac{c^{\epsilon} L}{\epsilon} \int_{S} \gamma^{-1}(\alpha) d \alpha+O(\epsilon)
\end{aligned}
$$

where the second equality follows from (Equation 4.3.38) for $\eta=0, n(z(\alpha))=-\mathbf{n}(z(\alpha))$ and $d \sigma=L d \alpha$, as well as (Equation 4.3.40). Rearranging the terms and using the definition of $\beta$ in (Equation 4.3.11) yields (Equation 4.3.13).

Finally, note that (Equation 4.3.13) and (Equation 4.3.38) directly lead to (Equation 4.3.14), where we are using the fact that $\gamma_{i}$ is uniformly positive for $i=1, \ldots, n$, due to (H3).

The rest of the proof is devoted to proving the claims (Equation 4.3.38) and (Equation 4.3.39). For (Equation 4.3.38), we compute the gradient of $\tilde{p}^{\epsilon}$. Differentiating (Equation 4.3.10) with respect to $\alpha$ and $\eta$, we obtain

$$
\begin{equation*}
\left(\nabla_{\alpha, \eta} R^{\epsilon}(\alpha, \eta)\right)^{t} \nabla \tilde{p}\left(R^{\epsilon}(\alpha, \eta)\right)=\binom{0}{c^{\epsilon}} \tag{4.3.41}
\end{equation*}
$$

where $\left(\nabla_{\alpha, \eta} R^{\epsilon}\right)^{t}$ denotes the transpose of the Jacobian matrix of $R^{\epsilon}$. Since $\nabla_{\alpha, \eta} R^{\epsilon}=$ $\left(\partial_{\alpha} R^{\epsilon}, \partial_{\eta} R^{\epsilon}\right)$, using the formula for inverses of $2 \times 2$ matrices, we have

$$
\begin{equation*}
\left(\left(\nabla_{\alpha, \eta} R^{\epsilon}\right)^{t}\right)^{-1}=\frac{1}{J(\alpha, \eta)}\left(-\left(\partial_{\eta} R^{\epsilon}\right)^{\perp},\left(\partial_{\alpha} R^{\epsilon}\right)^{\perp}\right) \tag{4.3.42}
\end{equation*}
$$

where $J(\alpha, \eta):=\operatorname{det}\left(\nabla_{\alpha, \eta} R^{\epsilon}\right)$. Multiplying the inverse matrix on both sides of (Equation 4.3.41),
we have

$$
\begin{equation*}
\nabla \tilde{p}^{\epsilon}\left(R^{\epsilon}(\alpha, \eta)\right)=\frac{1}{J}\left(-\left(\partial_{\eta} R^{\epsilon}\right)^{\perp},\left(\partial_{\alpha} R^{\epsilon}\right)^{\perp}\right)\binom{0}{c^{\epsilon}}=\frac{c^{\epsilon}}{J}\left(\partial_{\alpha} R^{\epsilon}\right)^{\perp} \tag{4.3.43}
\end{equation*}
$$

Recall that Lemma 4.2.2 gives $\left(\partial_{\alpha} R^{\epsilon}\right)^{\perp}=z^{\prime}(\alpha)^{\perp}+O(\epsilon)=\operatorname{Ln}(z(\alpha))+O(\epsilon)$, and $J=\epsilon L \gamma+$ $O\left(\epsilon^{2}\right)$. Plugging these into (Equation 4.3.43) gives

$$
\begin{equation*}
\nabla \tilde{p}^{\epsilon}(R(\alpha, \eta))=\frac{c^{\epsilon}}{\epsilon}\left(\frac{\mathbf{n}(z(\alpha))}{\gamma}+O(\epsilon)\right) . \tag{4.3.44}
\end{equation*}
$$

Note that it follows from (Equation 4.3.8) that $c^{\epsilon} \leq \frac{\left|D^{\epsilon}\right|}{2 \pi}$, where $\left|D^{\epsilon}\right| \leq C \epsilon$ due to (Equation 4.2.9). These imply

$$
\begin{equation*}
\frac{c_{\epsilon}}{\epsilon} \leq C, \tag{4.3.45}
\end{equation*}
$$

and applying it to (Equation 4.3.44) yields (Equation 4.3.38).
To prove (Equation 4.3.39), since $q^{\epsilon}=p^{\epsilon}-\tilde{p}^{\epsilon}$ and $\Delta p^{\epsilon}=-2$ in $D^{\epsilon}$, it suffices to show that

$$
\begin{equation*}
\left|\Delta \tilde{p}^{\epsilon}\right| \leq C \quad \text { in } D^{\epsilon}, \tag{4.3.46}
\end{equation*}
$$

and we will begin with an explicit computation of $\partial_{x_{1} x_{1}} \tilde{p}^{\epsilon}$ and $\partial_{x_{2} x_{2}} \tilde{p}^{\epsilon}$. Let us denote $R^{\epsilon}=$ : ( $R^{1}, R^{2}$ ). For notational simplicity, in the rest of the proof we will use subscripts on $R^{\epsilon}, R^{1}, R^{2}$ and $J$ to denote their partial derivative, e.g. $R_{\alpha}^{1}:=\partial_{\alpha} R^{1}$.

From (Equation 4.3.43), it follows that

$$
\partial_{x_{1}} \tilde{p}^{\epsilon}\left(R^{\epsilon}(\alpha, \eta)\right)=-\frac{c^{\epsilon}}{J} R_{\alpha}^{2} .
$$

Differentiating in $\alpha$ and $\eta$, we get

$$
\begin{aligned}
\nabla\left(\partial_{x_{1}} \tilde{p}^{\epsilon}\right)\left(R^{\epsilon}(\alpha, \eta)\right) & =\left(\left(\nabla_{\alpha, \eta} R^{\epsilon}\right)^{t}\right)^{-1} \nabla_{\alpha, \eta}\left(-\frac{c^{\epsilon}}{J} R_{\alpha}^{2}\right) \\
& =\frac{c^{\epsilon}}{J}\left(\begin{array}{cc}
R_{\eta}^{2} & -R_{\alpha}^{2} \\
-R_{\eta}^{1} & R_{\alpha}^{1}
\end{array}\right)\binom{\frac{J_{\alpha}}{J^{2}} R_{\alpha}^{2}-\frac{1}{J} R_{\alpha \alpha}^{2}}{\frac{J_{\eta}}{J^{2}} R_{\alpha}^{2}-\frac{1}{J} R_{\alpha \eta}^{2}}
\end{aligned}
$$

thus

$$
\partial_{x_{1} x_{1}} \tilde{p}^{\epsilon}(R(\alpha, \eta))=\frac{c^{\epsilon}}{J}\left(\frac{J_{\alpha}}{J^{2}} R_{\eta}^{2} R_{\alpha}^{2}-\frac{1}{J} R_{\eta}^{2} R_{\alpha \alpha}^{2}-\frac{J_{\eta}}{J^{2}}\left(R_{\alpha}^{2}\right)^{2}+\frac{1}{J} R_{\alpha}^{2} R_{\alpha \eta}^{2}\right) .
$$

Likewise, $\partial_{x_{2} x_{2}} \tilde{p}(R(\alpha, \eta))$ takes the same expression except every $R^{2}$ is changed into $R^{1}$. Adding them together gives

$$
\begin{equation*}
\Delta \tilde{p}^{\epsilon}(R(\alpha, \eta))=\frac{c^{\epsilon}}{J}\left(\frac{J_{\alpha}}{J^{2}} R_{\eta}^{\epsilon} \cdot R_{\alpha}^{\epsilon}-\frac{1}{J} R_{\eta}^{\epsilon} \cdot R_{\alpha \alpha}^{\epsilon}-\frac{J_{\eta}}{J^{2}} R_{\alpha}^{\epsilon} \cdot R_{\alpha}^{\epsilon}+\frac{1}{J} R_{\alpha}^{\epsilon} \cdot R_{\alpha \eta}^{\epsilon}\right) . \tag{4.3.47}
\end{equation*}
$$

Using the explicit formulae of $R_{\alpha}, R_{\eta}$ and $J$ in Lemma 4.2.2, we directly obtain $\left|R_{\alpha}^{\epsilon}\right|,\left|R_{\alpha \alpha}^{\epsilon}\right| \leq C$; $\left|R_{\eta}^{\epsilon}\right|,\left|R_{\alpha \eta}^{\epsilon}\right|,\left|J_{\alpha}\right| \leq C \epsilon ; \quad\left|J_{\eta}\right| \leq C \epsilon^{2} ;$ and $J^{-1} \leq C \epsilon^{-1}$ when $\epsilon$ is sufficiently small, where $C$ depends on $\left\|z_{i}\right\|_{C^{3}\left(S_{i}\right)}$ and $\left\|\gamma_{i}\right\|_{C^{2}\left(S_{i}\right)}$. As a result, all the four terms in the parenthesis of (Equation 4.3.47) are bounded by some constant $C$ independent of $\epsilon$. Finally, (Equation 4.3.45) yields $\frac{c_{\epsilon}}{J} \leq C$ as well, thus $\left|\Delta \tilde{p}^{\epsilon}\right| \leq C$, and this proves the second claim (Equation 4.3.39).

### 4.4 Proof of the symmetry result

In this section we prove that a stationary vortex sheet with positive vorticity must be radially symmetric up to a translation, and a rotating vortex sheet with positive vorticity and angular velocity $\Omega<0$ must be radially symmetric. The key idea of the proof is to define the integral

$$
\begin{align*}
I^{\epsilon} & :=\int_{D^{\epsilon}} \epsilon^{-1} \mathbf{u}^{\epsilon} \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x \\
& =\int_{D^{\epsilon}} \epsilon^{-1}\left(x+\nabla p^{\epsilon}\right) \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x \tag{4.4.1}
\end{align*}
$$

and compute it in two different ways. The motivation of the definition is as follows. As discussed in [50, Section 2.1], $I^{\epsilon}$ can be thought of as a first variation of an "energy functional"

$$
\mathcal{E}\left[\omega^{\epsilon}\right]:=\int \frac{1}{2} \omega^{\epsilon}\left(\omega^{\epsilon} * \mathcal{N}\right)-\frac{\Omega}{2} \omega^{\epsilon}|x|^{2} d x
$$

when we perturb $\omega^{\epsilon}$ by a divergence free vector $\mathbf{u}^{\epsilon}$ in $D^{\epsilon}$. (This functional $\mathcal{E}$ only serves as our motivation, and will not appear in the proof.) On the one hand, using that $\omega_{0}$ is stationary in the rotating frame with angular velocity $\Omega$ and $\omega^{\epsilon}$ is a close approximation of $\omega_{0}$, we will show in Proposition 4.4.1 that $I^{\epsilon}$ is of order $O(\epsilon|\log \epsilon|)$, thus goes to zero as $\epsilon \rightarrow 0$. On the other hand, using the particular $\mathbf{u}^{\epsilon}$ that we constructed in Section section 4.3, we will prove in Proposition 4.4.2 that if $\Omega=0, I^{\epsilon}$ is strictly positive independently of $\epsilon$ unless all the vortex sheets are nested circles with constant density; and also prove a similar result in Corollary 4.4.3 for $\Omega<0$.

Proposition 4.4.1. Assume $\omega(\cdot, t)=\omega_{0}\left(R_{\Omega t} \cdot\right)$ is a stationary/uniformly-rotating vortex sheet with angular velocity $\Omega \in \mathbb{R}$, where $\omega_{0}$ satisfies $(\mathbf{H} 1)-(\mathbf{H} 3)$. Then there exists some $C>0$ only depending on $b$ (as in (H2)), $\max _{i}\left\|z_{i}\right\|_{C^{3}\left(S_{i}\right)}, \max _{i \leq n}\left\|\gamma_{i}\right\|_{C^{2}\left(S_{i}\right)}, \max _{i>n}\left\|\gamma_{i}\right\|_{C^{b}\left(S_{i}\right)}, d_{\Gamma}$ and $F_{\Gamma}$, such that $\left|I^{\epsilon}\right|<C \epsilon^{b}|\log \epsilon|$ for all sufficiently small $\epsilon>0$.

Proof. Let us decompose $I^{\epsilon}=: \sum_{i=1}^{n+m} I_{i}^{\epsilon}$, where $I_{i}^{\epsilon}:=\int_{D_{i}^{\epsilon}} \epsilon^{-1}\left(x+\nabla p^{\epsilon}\right) \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x$.
We start with showing that $\left|I_{i}^{\epsilon}\right| \leq C \epsilon^{b}|\log \epsilon|$ for $i=n+1, \ldots, n+m$. For such $i, p^{\epsilon}=0$ on $\partial D_{i}^{\epsilon}$, thus the divergence theorem (and the fact that $\omega^{\epsilon}=\epsilon^{-1}$ in $D_{i}^{\epsilon}$ ) gives

$$
I_{i}^{\epsilon}=\underbrace{\int_{D_{i}^{\epsilon}} \epsilon^{-1} x \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x}_{=: T_{i}^{\epsilon}}-\int_{D_{i}^{\epsilon}} \epsilon^{-1}\left(\epsilon^{-1}-2 \Omega\right) p^{\epsilon}(x) d x
$$

Using the estimate $\left|p^{\epsilon}\right| \leq C \epsilon^{2}$ in Proposition 4.3 .2 and the fact that $\left|D_{i}^{\epsilon}\right| \leq C \epsilon$ from (Equation 4.2.9), we easily bound the second integral by $C \epsilon$. To control the first integral $T_{i}^{\epsilon}$,
we rewrite it using the change of variables $x=R_{i}^{\epsilon}(\alpha, \eta)$ and the definition $\mathbf{v}^{\epsilon}:=\nabla^{\perp}\left(\omega^{\epsilon} * \mathcal{N}\right)$ in (Equation 4.2.10): (also note that on the right hand side we group $\epsilon^{-1}$ with the determinant)

$$
T_{i}^{\epsilon}=\int_{S_{i}} \int_{-1}^{0} R_{i}^{\epsilon}(\alpha, \eta) \cdot \underbrace{\left(-\left(\mathbf{v}^{\epsilon}\right)^{\perp}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-\Omega R_{i}^{\epsilon}(\alpha, \eta)\right)}_{=: J_{i}^{\epsilon}} \underbrace{\epsilon^{-1} \operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}(\alpha, \eta)\right)}_{=: K_{i}^{\epsilon}} d \eta d \alpha .
$$

Let us take a closer look at the integrand, which is a product of 3 terms. Clearly, the definition of $R_{i}^{\epsilon}$ gives $R_{i}^{\epsilon}(\alpha, \eta)=z_{i}(\alpha)+O(\epsilon)$. As for the middle term $J_{i}^{\epsilon}$, Lemma 4.2.4 yields

$$
\begin{equation*}
J_{i}^{\epsilon}(\alpha, \eta)=-B R^{\perp}\left(z_{i}(\alpha)\right)+\left(\eta+\frac{1}{2}\right)[\mathbf{v}]^{\perp}\left(z_{i}(\alpha)\right)-\Omega z_{i}(\alpha)+O\left(\epsilon^{b}|\log \epsilon|\right) \tag{4.4.2}
\end{equation*}
$$

Using the fact that $B R\left(z_{i}(\alpha)\right)=\Omega z_{i}^{\perp}(\alpha)$ for $i=n+1, \ldots, n+m$ (which follows from (Equation 5.1.1) and (Equation 5.1.2)), it becomes

$$
\begin{equation*}
J_{i}^{\epsilon}(\alpha, \eta)=\left(\eta+\frac{1}{2}\right)[\mathbf{v}]^{\perp}\left(z_{i}(\alpha)\right)+O\left(\epsilon^{b}|\log \epsilon|\right) \tag{4.4.3}
\end{equation*}
$$

Also it follows from (Equation 4.2.8) that $K_{i}^{\epsilon}(\alpha, \eta)=L_{i} \gamma_{i}(\alpha)+O(\epsilon)$. Plugging these three estimates into the above integral gives

$$
T_{i}^{\epsilon}=\int_{S_{i}} \int_{-1}^{0} z_{i}(\alpha) \cdot\left(\eta+\frac{1}{2}\right)[\mathbf{v}]^{\perp}\left(z_{i}(\alpha)\right) L_{i} \gamma_{i}(\alpha) d \eta d \alpha+O\left(\epsilon^{b}|\log \epsilon|\right)=O\left(\epsilon^{b}|\log \epsilon|\right)
$$

where the last step follows from the fact that $\int_{-1}^{0}\left(\eta+\frac{1}{2}\right) d \eta=0$. This finishes the proof that $\left|I_{i}^{\epsilon}\right| \leq C \epsilon^{b}|\log \epsilon|$ for $i=n+1, \ldots, n+m$, where $C$ depends on $b, \max _{i}\left\|z_{i}\right\|_{C^{2}\left(S_{i}\right)}, \max _{i}\left\|\gamma_{i}\right\|_{C^{b}\left(S_{i}\right)}$, $d_{\Gamma}$ and $F_{\Gamma}$.

In the rest of the proof we aim to show $\left|I_{i}^{\epsilon}\right| \leq C \epsilon^{b}|\log \epsilon|$ for $i=1, \ldots, n$, which is slightly more involved. Recall that in Proposition 4.3 .3 we defined $\tilde{p}^{\epsilon}$ and $q^{\epsilon}$ in $D_{i}^{\epsilon}$ for $i=1, \ldots, n$, where they satisfy $p^{\epsilon}=\tilde{p}^{\epsilon}+q^{\epsilon}$ in $D_{i}^{\epsilon}$, and $q^{\epsilon}=0$ on $\partial D_{i}^{\epsilon}$. This allows us to apply the divergence theorem
(to the $q^{\epsilon}$ term only) and decompose $I_{i}^{\epsilon}$ as

$$
I_{i}^{\epsilon}=\int_{D_{i}^{\epsilon}} \epsilon^{-1}\left(x+\nabla \tilde{p}_{\epsilon}\right) \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x-\int_{D_{i}^{\epsilon}} \epsilon^{-1}\left(\epsilon^{-1}-2 \Omega\right) q^{\epsilon}(x) d x=: I_{i, 1}^{\epsilon}+I_{i, 2}^{\epsilon} .
$$

We can easily show that $I_{i, 2}^{\epsilon}=O(\epsilon)$ : (Equation 4.3.12) of Proposition 4.3.3 gives $\left|q^{\epsilon}\right| \leq C \epsilon^{2}$, and combining it with $\left|D_{i}^{\epsilon}\right| \leq C \epsilon$ in (Equation 4.2.9) immediately yields the desired estimate.

Next we turn to $I_{i, 1}^{\epsilon}$. Again, the change of variables $x=R_{i}^{\epsilon}(\alpha, \eta)$ and the definition $\mathbf{v}^{\epsilon}:=$ $\nabla^{\perp}\left(\omega^{\epsilon} * \mathcal{N}\right)$ gives

$$
I_{i, 1}^{\epsilon}=\int_{S_{i}} \int_{-1}^{0}\left(R_{i}^{\epsilon}(\alpha, \eta)+\nabla \tilde{p}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)\right) \cdot \underbrace{\left(-\left(\mathbf{v}^{\epsilon}\right)^{\perp}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)-\Omega R_{i}^{\epsilon}(\alpha, \eta)\right)}_{=: J_{i}^{\epsilon}} \underbrace{\epsilon^{-1} \operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}(\alpha, \eta)\right)}_{=: K_{i}^{\epsilon}} d \eta d \alpha
$$

For the three terms in the product of the integrand, we will approximate the first term using the definition of $R_{i}^{\epsilon}$ and (Equation 4.3.14) of Proposition 4.3.3:

$$
R_{i}^{\epsilon}(\alpha, \eta)+\nabla \tilde{p}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)=z_{i}(\alpha)+\frac{\beta_{i}}{\gamma_{i}(\alpha)} \mathbf{n}\left(z_{i}(\alpha)\right)+O(\epsilon)
$$

where $\beta_{i}:=\frac{2\left|U_{i}\right|}{L_{i} \int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha}$ is given by (Equation 4.3.11). Lemma 4.2.4 allows us to approximate the middle term $J_{i}^{\epsilon}$ as (Equation 4.4.2), however (Equation 4.4.3) no longer holds since for $i=$ $1, \ldots, n$ we do not have $B R\left(z_{i}(\alpha)\right)=\Omega z_{i}^{\perp}(\alpha)$. As for $K_{i}^{\epsilon}$, we again use (Equation 4.2.8) to approximate it by $K_{i}^{\epsilon}(\alpha, \eta)=L_{i} \gamma_{i}(\alpha)+O(\epsilon)$. Plugging these three estimates into the integrand of $I_{i, 1}^{\epsilon}$ gives

$$
I_{i, 1}^{\epsilon}=\int_{S_{i}}\left(z_{i}(\alpha)+\frac{\beta_{i}}{\gamma_{i}(\alpha)} \mathbf{n}\left(z_{i}(\alpha)\right)\right) \cdot\left(-B R^{\perp}\left(z_{i}(\alpha)\right)-\Omega z_{i}(\alpha)\right) L_{i} \gamma_{i}(\alpha) d \alpha+O\left(\epsilon^{b}|\log \epsilon|\right)
$$

where we again use the fact that the $\left(\eta+\frac{1}{2}\right)$ term gives zero contribution since $\int_{-1}^{0}\left(\eta+\frac{1}{2}\right) d \eta=0$. Next we will show the integral on the right hand side is in fact 0 . Since $\omega$ is a rotating solution
with angular velocity $\Omega$, the conditions (Equation 5.1.1) and (Equation 5.1.2) yield that

$$
-B R^{\perp}\left(z_{i}(\alpha)\right)-\Omega z_{i}(\alpha)=C_{i} \gamma_{i}^{-1}(\alpha) \mathbf{n}\left(z_{i}(\alpha)\right)
$$

for some constant $C_{i}$. Plugging this into the above integral gives

$$
\begin{aligned}
I_{i, 1}^{\epsilon} & =C_{i} L_{i} \int_{S_{i}}\left(z_{i}(\alpha) \cdot \mathbf{n}\left(z_{i}(\alpha)\right)+\frac{\beta_{i}}{\gamma_{i}(\alpha)}\right) d \alpha+O\left(\epsilon^{b}|\log \epsilon|\right) \\
& =C_{i} L_{i}\left(\int_{S_{i}} z_{i}(\alpha) \cdot \mathbf{n}\left(z_{i}(\alpha)\right) d \alpha+\frac{2\left|U_{i}\right|}{L_{i}}\right)+O\left(\epsilon^{b}|\log \epsilon|\right),
\end{aligned}
$$

where the second step follows from the definition of $\beta_{i}$ in (Equation 4.3.11). Let us compute the integral on the right hand side by changing to arclength parametrization and applying the divergence theorem:

$$
\int_{S_{i}} z_{i}(\alpha) \cdot \mathbf{n}\left(z_{i}(\alpha)\right) d \alpha=-\frac{1}{L_{i}} \int_{\partial U_{i}} x \cdot n d \sigma=-\frac{2\left|U_{i}\right|}{L_{i}},
$$

which yields $I_{i, 1}^{\epsilon}=O\left(\epsilon^{b}|\log \epsilon|\right)$, and finishes the proof that $\left|I_{i}^{\epsilon}\right| \leq C \epsilon^{b}|\log \epsilon|$ for $i=1, \ldots, n$, where $C$ depends on $b,\left\|z_{i}\right\|_{C^{3}\left(S_{i}\right)},\left\|\gamma_{i}\right\|_{C^{2}\left(S_{i}\right)}, d_{\Gamma}$ and $F_{\Gamma}$.

Finally, summing the $I_{i}^{\epsilon}$ estimates for $i=1, \ldots, n+m$ gives $\left|I^{\epsilon}\right| \leq C \epsilon^{b}|\log \epsilon|$ for all sufficiently small $\epsilon>0$, thus we can conclude.

Now we will use a different way to compute $I^{\epsilon}$. Let us first define a new integral $\tilde{I}^{\epsilon}$ that is the same as $I^{\epsilon}$ except with $\Omega$ set to zero:

$$
\begin{equation*}
\tilde{I}^{\epsilon}:=\int_{D^{\epsilon}} \epsilon^{-1}\left(x+\nabla p^{\epsilon}\right) \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}\right) d x . \tag{4.4.4}
\end{equation*}
$$

Next we will prove that $\tilde{I}^{\epsilon}$ is strictly positive independently of $\epsilon$ unless all the vortex sheets are nested circles with constant density. As we will see in the proof, the key step is to show that if some $\Gamma_{i}$ is either not a circle or does not have a constant $\gamma_{i}$, then the estimates on $p^{\epsilon}$ in Propositions 4.3.2-4.3.3 lead to the following quantitative version of (Equation 4.3.9):
$\epsilon^{-2}\left(\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi}-\int_{D_{i}^{\epsilon}} p^{\epsilon}(x) d x\right) \geq c_{0}>0$, where $c_{0}$ is independent of $\epsilon$.
Proposition 4.4.2. Let $\tilde{I}^{\epsilon}$ be defined as in (Equation 4.4.4). Assume that $\Gamma_{i}$ and $\gamma_{i}$ satisfy (H1)(H3) for $i=1, \ldots, n+m$. Then we have $\tilde{I}^{\epsilon} \geq 0$ for all sufficiently small $\epsilon>0$.

In addition, if $\Gamma$ is not a union of nested circles with constant $\gamma_{i}$ 's on each connected component, there exists some $c_{0}>0$ independent of $\epsilon$, such that $\tilde{I}^{\epsilon}>c_{0}>0$ for all sufficiently small $\epsilon>0$.

Proof. We start by decomposing $\tilde{I}^{\epsilon}$ as

$$
\tilde{I}^{\epsilon}=\int_{D^{\epsilon}} \epsilon^{-1} x \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}\right) d x+\int_{D^{\epsilon}} \epsilon^{-1} \nabla p^{\epsilon} \cdot \nabla\left(\omega^{\epsilon} * \mathcal{N}\right) d x=: I_{1}^{\epsilon}+I_{2}^{\epsilon}
$$

$I_{1}^{\epsilon}$ can be easily computed as

$$
\begin{equation*}
I_{1}^{\epsilon}=\frac{1}{2 \pi \epsilon^{2}} \int_{D^{\epsilon}} \int_{D^{\epsilon}} \frac{x \cdot(x-y)}{|x-y|^{2}} d x d y=\frac{\left|D^{\epsilon}\right|^{2}}{4 \pi \epsilon^{2}}=\frac{1}{4 \pi \epsilon^{2}}\left(\sum_{i=1}^{n+m}\left|D_{i}^{\epsilon}\right|\right)^{2} \tag{4.4.5}
\end{equation*}
$$

where the second equality is obtained by exchanging $x$ with $y$ and taking the average with the original integral. As for $I_{2}^{\epsilon}$, we have

$$
\begin{align*}
I_{2}^{\epsilon} & =\frac{1}{\epsilon} \int_{\partial D^{\epsilon}} p^{\epsilon} \nabla\left(\omega^{\epsilon} * \mathcal{N}\right) \cdot n d \sigma-\frac{1}{\epsilon} \int_{D^{\epsilon}} p^{\epsilon} \omega^{\epsilon} d x \\
& =-\frac{1}{\epsilon} \sum_{i=1}^{n} c_{i}^{\epsilon} \int_{\partial U_{i}} \nabla\left(\omega^{\epsilon} * \mathcal{N}\right) \cdot n d \sigma-\frac{1}{\epsilon^{2}} \int_{D^{\epsilon}} p^{\epsilon} d x  \tag{4.4.6}\\
& \geq-\frac{1}{\epsilon^{2}} \sum_{i=1}^{n} \sum_{j=1}^{n+m} \frac{\left|D_{i}^{\epsilon}\right|}{2 \pi} \int_{U_{i}} 1_{D_{j}^{\epsilon}} d x-\frac{1}{\epsilon^{2}} \sum_{i=1}^{n+m} \int_{D_{i}^{\epsilon}} p^{\epsilon} d x
\end{align*}
$$

where the first equality follows from the divergence theorem, the second equality follows from the boundary conditions (Equation 4.3.5) and (Equation 4.3.6) for $p^{\epsilon}$ (as well as the fact that $\partial U_{i}$ and $\partial D_{i}^{\epsilon}$ have opposite outer normals), and the last inequality follows from the divergence theorem as well as the inequality $c_{i}^{\epsilon} \leq \sup _{D_{i}^{\epsilon}} p \leq \frac{\left|D_{i}^{\epsilon}\right|}{2 \pi}$ due to (Equation 4.3.8).

Let us denote $j \prec i$ if $i \in\{1, \ldots, n\}, j \in\{1, \ldots, n+m\}, j \neq i$ and $\Gamma_{j}$ lies in the interior of the domain enclosed by $\Gamma_{i}$ (that is, $\Gamma_{j} \subset U_{i}$ ). If not, we denote $j \nprec i$. Note that for sufficiently small $\epsilon>0$, we have

$$
\int_{U_{i}} 1_{D_{j}^{\epsilon}} d x= \begin{cases}\left|D_{j}^{\epsilon}\right| & \text { if } j \prec i  \tag{4.4.7}\\ 0 & \text { otherwise }\end{cases}
$$

Applying this to (Equation 4.4.6) yields

$$
\begin{align*}
I_{2}^{\epsilon} & \geq-\frac{1}{2 \pi \epsilon^{2}} \sum_{i, j=1}^{n+m} \mathbb{1}_{j \prec i}\left|D_{i}^{\epsilon}\right|\left|D_{j}^{\epsilon}\right|-\frac{1}{\epsilon^{2}} \sum_{i=1}^{n+m} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x  \tag{4.4.8}\\
& =-\frac{1}{4 \pi \epsilon^{2}} \sum_{i, j=1}^{n+m}\left(\mathbb{1}_{j \prec i}+\mathbb{1}_{i \prec j}\right)\left|D_{i}^{\epsilon}\right|\left|D_{j}^{\epsilon}\right|-\frac{1}{\epsilon^{2}} \sum_{i=1}^{n+m} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x
\end{align*}
$$

where in the first step we used that the $i=n+1, \ldots, n+m$ terms have zero contribution in the first sum, due to the definition of $j \prec i$.

Adding (Equation 4.4.5) and (Equation 4.4.8) together, we obtain

$$
\begin{equation*}
\tilde{I}^{\epsilon} \geq \sum_{i=1}^{n+m} \underbrace{\frac{1}{\epsilon^{2}}\left(\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi}-\int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x\right)}_{=: A_{i}^{\epsilon}}+\sum_{i, j=1}^{n+m} \underbrace{\frac{1}{\epsilon^{2}}\left(\mathbb{1}_{i \neq j}-\left(\mathbb{1}_{j \prec i}+\mathbb{1}_{i \prec j}\right)\right) \frac{\left|D_{i}^{\epsilon}\right|\left|D_{j}^{\epsilon}\right|}{4 \pi}}_{=: B_{i, j}^{\epsilon}}, \tag{4.4.9}
\end{equation*}
$$

From (Equation 4.3.9), it follows that $A_{i}^{\epsilon} \geq 0$ for all $i=1, \ldots, n+m$, with equality achieved if and only if each $D_{i}^{\epsilon}$ is a disk or an annulus. Note that $B_{i, j}^{\epsilon} \geq 0$ as well for all $i$ and $j$, since for any $i \neq j$, at most one of $i \prec j$ and $j \prec i$ can hold. Putting these together yields that $\tilde{I}^{\epsilon} \geq 0$ for any sufficiently small $\epsilon>0$.

In the rest of the proof, we assume $\Gamma$ is not a union of nested circles with constant $\gamma_{i}$ 's on each connected component. Therefore at least one of the following 3 cases must be true. In each case we aim to show that $\tilde{I}_{\epsilon} \geq c_{0}>0$, where $c_{0}$ is independent of $\epsilon$ for all sufficiently small $\epsilon>0$.

Case 1. There exists some open curve $\Gamma_{i}$ that is not a loop. In this case $D_{i}^{\epsilon}$ is simplyconnected, and $p^{\epsilon}=0$ on $\partial D_{i}^{\epsilon}$ by (Equation 4.3.5). Applying Proposition 4.3.2 to $p^{\epsilon}$ in $D_{i}^{\epsilon}$,
we have $\sup _{D_{i}^{\epsilon}} p^{\epsilon} \leq C \epsilon^{2}$, where $C$ is independent of $\epsilon$. This leads to $\int_{D_{i}^{\epsilon}} p^{\epsilon} d x \leq C \epsilon^{3}$, since $\left|D_{i}^{\epsilon}\right|=O(\epsilon)$ by (Equation 4.2.9). As a result, for the index $i$ we have

$$
A_{i}^{\epsilon}=\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi \epsilon^{2}}-\epsilon^{-2} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x \geq \frac{L_{i}^{2}}{4 \pi}\left(\int_{S_{i}} \gamma_{i}(\alpha) d \alpha\right)^{2}-C \epsilon,
$$

where we again used (Equation 4.2.9) in the second inequality. This gives that $A_{i}^{\epsilon} \geq$ $\frac{L_{i}^{2}}{8 \pi}\left(\int_{S_{i}} \gamma_{i}(\alpha) d \alpha\right)^{2}>0$ for all sufficiently small $\epsilon>0$.

Case 2. There exists some closed curve $\Gamma_{i}$ that is either not a circle, or $\gamma_{i}$ is not a constant. In this case we aim to show that $A_{i}^{\epsilon} \geq c_{0}>0$, and this will be done by finding good approximations (independent of $\epsilon$ ) for both terms in $A_{i}^{\epsilon}$. For the first term $\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi \epsilon^{2}}$, using (Equation 4.2.9) we again have

$$
\begin{equation*}
\frac{\left|D_{i}^{\epsilon}\right|^{2}}{4 \pi \epsilon^{2}} \geq \frac{L_{i}^{2}}{4 \pi}\left(\int_{S_{i}} \gamma_{i}(\alpha) d \alpha\right)^{2}-C \epsilon=: J_{i}-C \epsilon \tag{4.4.10}
\end{equation*}
$$

where $J_{i}>0$ is independent of $\epsilon$. For the second term $\epsilon^{-2} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x$, rewriting the integral using the change of variables $x=R_{i}^{\epsilon}(a, \eta)$ gives

$$
\epsilon^{-2} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x=\int_{S_{i}} \int_{-1}^{0} \frac{p^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)}{\epsilon} \frac{\operatorname{det}\left(\nabla_{\alpha, \eta} R_{i}^{\epsilon}\right)}{\epsilon} d \eta d \alpha
$$

Recall that in Proposition 4.3 .3 we defined $\tilde{p}^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right):=c_{i}^{\epsilon}(1+\eta)$ and $q_{\epsilon}$ such that $p^{\epsilon}-\tilde{p}_{\epsilon}=q_{\epsilon}$. By (Equation 4.3.12) and (Equation 4.3.13), for all $\alpha \in S_{i}$ and $\eta \in(-1,0)$ we have

$$
\left|\frac{p^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)}{\epsilon}-\beta_{i}(1+\eta)\right| \leq\left|\frac{p^{\epsilon}\left(R_{i}^{\epsilon}(\alpha, \eta)\right)}{\epsilon}-\frac{c_{i}^{\epsilon}}{\epsilon}(1+\eta)\right|+\left|\frac{c_{i}^{\epsilon}}{\epsilon}-\beta_{i}\right| \leq C \epsilon,
$$

where $\beta_{i}:=\frac{2\left|U_{i}\right|}{L_{i} \int_{S_{i}}^{\gamma_{i}^{-1}}(\alpha) d \alpha}$ is defined in (Equation 4.3.11). Combining this with the expression of
the determinant in (Equation 4.2.8), we have

$$
\begin{aligned}
\epsilon^{-2} \int_{D_{i}^{\epsilon}} p_{i}^{\epsilon} d x & =\int_{S_{i}} \int_{-1}^{0}\left(\beta_{i}(1+\eta)+O(\epsilon)\right)\left(L_{i} \gamma_{i}(\alpha)+O(\epsilon)\right) d \eta d \alpha \\
& \leq \frac{\left|U_{i}\right|}{\int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha} \int_{S_{i}} \gamma_{i}(\alpha) d \alpha+C \epsilon=: K_{i}+C \epsilon,
\end{aligned}
$$

where $K_{i}$ is independent of $\epsilon$. Putting this together with (Equation 4.4.10) yields the following:

$$
\begin{align*}
A_{i}^{\epsilon} & \geq J_{i}-K_{i}-C \epsilon \\
& =\frac{L_{i}^{2}}{4 \pi} \frac{\int_{S_{i}} \gamma_{i}(\alpha) d \alpha}{\int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha}\left(\int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha \int_{S_{i}} \gamma_{i}(\alpha) d \alpha-\frac{4 \pi\left|U_{i}\right|}{L_{i}^{2}}\right)-C \epsilon . \tag{4.4.11}
\end{align*}
$$

Let us take a closer look at the two terms inside the parenthesis. For the first term, Cauchy-Schwarz inequality gives

$$
\int_{S_{i}} \gamma_{i}^{-1}(\alpha) d \alpha \int_{S_{i}} \gamma_{i}(\alpha) d \alpha \geq 1
$$

with equality achieved if and only if $\gamma_{i}$ is a constant. For the second term, the isoperimetric inequality yields

$$
\frac{4 \pi\left|U_{i}\right|}{L_{i}^{2}} \leq 1
$$

(recall that $L_{i}=\left|\partial U_{i}\right|$ ), with equality achieved if and only $U_{i}$ is a disk. By the assumption of Case 2, at least one of the inequalities must be strict, thus the parenthesis on the right hand side of (Equation 4.4.11) is strictly positive (and independent of $\epsilon$ ). Therefore there exists some constant $c_{0}>0$ such that $\tilde{I}^{\epsilon} \geq A_{i}^{\epsilon} \geq c_{0}$ for all sufficiently small $\epsilon$.

Case 3. There exist $i \neq j$ such that $i \nprec j$ and $j \nprec i$. Then it is clear that for such $i, j, B_{i, j}^{\epsilon}$ in (Equation 4.4.9) is given by $B_{i, j}^{\epsilon}=\frac{\left|D_{i}^{\epsilon}\right|\left|D_{j}^{\epsilon}\right|}{4 \pi \epsilon^{2}}$. Hence (Equation 4.2.9) gives

$$
B_{i, j}^{\epsilon} \geq L_{i} L_{j}\left(\int_{S_{i}} \gamma_{i}(\alpha) d \alpha\right)\left(\int_{S_{j}} \gamma_{j}(\alpha) d \alpha\right)-C \epsilon
$$

which yields $\tilde{I}^{\epsilon} \geq \frac{1}{2} L_{i} L_{j}\left(\int_{S_{i}} \gamma_{i} d \alpha\right)\left(\int_{S_{j}} \gamma_{j} d \alpha\right)>0$ for all sufficiently small $\epsilon>0$.
This finishes our discussion on all 3 cases. To conclude, since $\Gamma$ is not a union of nested circles with constant $\gamma_{i}$ 's on each connected component, at least one of the 3 cases must hold, and all of them lead to $\tilde{I}^{\epsilon} \geq c_{0}>0$.

The above proposition immediately leads to the following corollary for the $\Omega<0$ case.

Corollary 4.4.3. Assume that $\Gamma_{i}$ and $\gamma_{i}$ satisfy $(\mathbf{H 1})-(\mathbf{H 3})$ for $i=1, \ldots, n+m$. Let $I^{\epsilon}$ be defined as in (Equation 4.4.1), and assume $\Omega<0$. Then we have $I^{\epsilon} \geq 0$ for all sufficiently small $\epsilon>0$. In addition, if $\Gamma$ is not a union of concentric circles all centered at the origin with constant $\gamma_{i}$ 's, there exists some $c_{0}>0$ independent of $\epsilon$, such that $I^{\epsilon}>c_{0}>0$ for all sufficiently small $\epsilon>0$.

Proof. Let us decompose $I^{\epsilon}$ as follows (recall the definition of $\tilde{I}^{\epsilon}$ in (Equation 4.4.4))

$$
\begin{equation*}
I^{\epsilon}=\tilde{I}^{\epsilon}+(-\Omega)\left(\epsilon^{-1} \int_{D^{\epsilon}}\left(|x|^{2}+\nabla p^{\epsilon} \cdot x\right) d x\right)=: \tilde{I}^{\epsilon}+\underbrace{(-\Omega)}_{>0} J^{\epsilon} . \tag{4.4.12}
\end{equation*}
$$

Recall that Proposition 4.4.2 gives $\tilde{I}_{\epsilon} \geq c_{0}>0$ as long as $\Gamma$ is not a union of nested circles with constant $\gamma_{i}$ 's. By [50, Lemma 2.11], we have

$$
\int_{D_{i}^{\epsilon}}\left(|x|^{2}+\nabla p^{\epsilon} \cdot x\right) d x \geq 0 \quad \text { for any } i=1, \ldots, n+m
$$

thus $J^{\epsilon} \geq 0$. Putting them together, and using the fact that $\Omega<0$, we know $I^{\epsilon} \geq c_{0}>0$ if $\Gamma$ is not a union of nested circles with constant $\gamma_{i}$ 's.

To finish the proof, we only need to focus on the case that the $\Gamma_{i}$ 's are nested circles with constant $\gamma_{i}$ 's, but not all of them are centered at the origin. Assume that there exists $k \in\{1, \ldots, n\}$ such that $\Gamma_{k}$ is a circle with radius $r_{k}$ centered at $x_{k} \neq 0$. Since $\gamma_{k}$ is a constant, $D_{k}^{\epsilon}$ is an annulus given by $B\left(x_{k}, r_{k}+\epsilon \gamma_{k}\right) \backslash B\left(x_{k}, r_{k}\right)$. The symmetry of $D_{k}^{\epsilon}$ about $x_{k}$ immediately leads to
$\left.p^{\epsilon}\right|_{D_{k}^{\epsilon}}=-\frac{1}{2}\left|x-x_{k}\right|^{2}+\frac{1}{2}\left(r_{k}+\epsilon \gamma_{k}\right)^{2}$. An elementary computation gives

$$
\epsilon^{-1} \int_{D_{k}^{\epsilon}}\left(|x|^{2}+\nabla p^{\epsilon} \cdot x\right) d x=\epsilon^{-1} \int_{D_{k}^{\epsilon}}|x|^{2}-\left(x-x_{k}\right) \cdot x d x=\epsilon^{-1}\left|x_{k}\right|^{2}\left|D_{k}^{\epsilon}\right| \geq 2 \pi r_{k} \gamma_{k}\left|x_{k}\right|^{2}>0
$$

where in the second-to-last step we used that $\left|D_{k}^{\epsilon}\right|=2 \pi \epsilon r_{k} \gamma_{k}+\pi \epsilon^{2} \gamma_{k}^{2}$. Setting $c_{0}:=2 \pi r_{k} \gamma_{k}\left|x_{k}\right|^{2}$ gives $I^{\epsilon} \geq c_{0}>0$, thus we can conclude.

Now we are ready to prove Theorem A. Note that for $\Omega<0$, the symmetry result immediately follows from Proposition 4.4.1 and Corollary 4.4.3. For $\Omega=0$, Proposition 4.4.1-4.4.2 already imply that a stationary vortex sheet with positive strength must be a union of nested circles with constant strength on each of them. To finish the proof, we only need to show that these nested circles must be concentric.

Proof of Theorem A. For a uniformly-rotating vortex sheet with $\Omega<0$, the symmetry result for $\Omega<0$ is a direct consequence of Proposition 4.4.1 and Corollary 4.4.3. Next we focus on the stationary (i.e. $\Omega=0$ ) case.

Combining Propsitions 4.4 .1-4.4.2, we obtain that $\Gamma$ is a union of nested circles, and $\gamma_{i}$ is constant on $\Gamma_{i}$ for all $i=1 \ldots, n$. It remains to show that all $\Gamma_{i}$ 's are concentric. Let us denote by $\mathbf{v}_{i}$ the contribution to the velocity field by $\Gamma_{i}$. Since $\Gamma_{i}$ is a circle with constant strength $\gamma_{i}$, a quick application of the divergence theorem yields that $\mathbf{v}_{i} \equiv 0$ in the open disk enclosed by $\Gamma_{i}$, whereas $\mathbf{v}_{i}(x)=\frac{\gamma_{i} L_{i}\left(x-x_{i}^{0}\right)^{\perp}}{2 \pi\left|x-x_{i}^{0}\right|^{2}}$ in the open set outside $\Gamma_{i}$, where $x_{i}^{0}$ is the center of the circle $\Gamma_{i}$.

Without loss of generality, let us reorder the indices such that $\Gamma_{i}$ is nested inside $\Gamma_{j}$ for $i<j$. Towards a contradiction, let $k>1$ be such that $\Gamma_{k}$ is the first circle that is not concentric with $\Gamma_{1}$. From the above discussion, we know that $\mathbf{v}_{i}=0$ on $\Gamma_{k}$ for $i=k+1, \ldots, n$ (since $\Gamma_{k}$ is nested inside $\Gamma_{i}$ ), whereas for $i=1, \ldots, k-1$ we have $\mathbf{v}_{i}=\frac{\gamma_{i} L_{i}\left(x-x_{1}^{0}\right)^{\perp}}{2 \pi\left|x-x_{1}^{0}\right|^{2}}$ on $\Gamma_{k}$, since all these $\Gamma_{i}$ 's have the same center $x_{1}^{0}$ and are nested inside $\Gamma_{k}$. Summing them up (and also using the fact that
$\Gamma_{k}$ contributes zero normal velocity on itself, since it is a circle with constant strength), we have

$$
B R(x) \cdot \mathbf{n}=\sum_{i=1}^{n} \mathbf{v}_{i}(x) \cdot \mathbf{n}=\left(\sum_{i=1}^{k-1} \gamma_{i} L_{i}\right) \frac{\left(x-x_{1}^{0}\right)^{\perp} \cdot \mathbf{n}}{2 \pi\left|x-x_{1}^{0}\right|^{2}} \quad \text { on } \Gamma_{k},
$$

where the right hand side is not a zero function since $\Gamma_{k}$ has a different center from $x_{1}^{0}$. This causes a contradiction with the fact that $\omega=\omega_{0}$ is stationary. As a result, all $\Gamma_{1}, \ldots, \Gamma_{n}$ must be concentric circles, finishing the proof.

## CHAPTER 5

## FLEXIBILITY RESULTS FOR VORTEX SHEETS

### 5.1 The equations and the functional spaces

Let $\omega(\cdot, t)=\omega_{0}\left(R_{\Omega t}\right)$ be a stationary/rotating vortex sheet solution to the incompressible 2D Euler equation, where $\omega_{0} \in \mathcal{M}\left(\mathbb{R}^{2}\right) \cap H^{-1}\left(\mathbb{R}^{2}\right)$. Here $\Omega=0$ corresponds to a stationary solution, and $\Omega \neq 0$ corresponds to a rotating solution. Assume that $\omega_{0}$ is concentrated on $\Gamma$. Throughout this paper we will assume that $\Gamma$ is a simple closed curve and $\Omega>0$. Following [52, Lemma 2.1], we have that:

Lemma 5.1.1. Assume $\omega(\cdot, t)=\omega_{0}\left(R_{\Omega t} x\right)$ is a stationary/uniformly-rotating vortex sheet with angular velocity $\Omega \in \mathbb{R}$, and $\omega_{0}$ is concentrated on $\Gamma$, with $z$ and $\gamma$ defined as above. Then the Birkhoff-Rott integral $B R$ (Equation 4.1.4) and the strength $\gamma$ satisfy the following two equations:

$$
\begin{equation*}
\left(B R-\Omega x^{\perp}\right) \cdot \mathbf{n}=\mathbf{v}^{+} \cdot \mathbf{n}=\mathbf{v}^{-} \cdot \mathbf{n}=0 \quad \text { on } \Gamma \tag{5.1.1}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(B R(z(\alpha))-\Omega z^{\perp}(\alpha)\right) \cdot \mathbf{s}(z(\alpha)) \frac{\gamma(\alpha)}{\left|z^{\prime}(\alpha)\right|}=C . \tag{5.1.2}
\end{equation*}
$$

Note that (Equation 5.1.2) can be written as

$$
\begin{equation*}
\left(I-P_{0}\right)\left[\left(B R(z, \Gamma)(z(\theta))-\Omega z(\theta)^{\perp}\right) \cdot \frac{z^{\prime}(\theta) \gamma(\theta)}{\left|z^{\prime}(\theta)\right|^{2}}\right]=0 \tag{5.1.3}
\end{equation*}
$$

where $P_{0}$ is a projection to the mean, that is, $P_{0} f:=\frac{1}{2 \pi} \int_{-\pi}^{\pi} f(\theta) d \theta$. For simplicity, we also denote $f f(\theta) d \theta:=\frac{1}{2 \pi} \int_{-\pi}^{\pi} f(\theta) d \theta$. Now plugging $z(\theta)=(1+r(\theta))(\cos (\theta), \sin (\theta))$ and $\gamma(\theta):=b+g(\theta)$
into (Equation 5.1.1), (Equation 5.1.2) and (Equation 5.1.3) yields that

$$
\begin{equation*}
\mathcal{F}(b, g, r):=\left(\mathcal{F}_{1}, \mathcal{F}_{2}\right)=(0,0), \tag{5.1.4}
\end{equation*}
$$

where

$$
\begin{aligned}
\mathcal{F}_{1}(b, g, r):= & f_{-\pi}^{\pi}(b+g(\eta)) \frac{\left(r^{\prime}(\theta) \cos (\theta-\eta)-(1+r(\theta)) \sin (\theta-\eta)\right)(1+r(\eta))-(1+r(\theta)) r^{\prime}(\theta)}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)} d \eta \\
& +\Omega r^{\prime}(\theta)(1+r(\theta)), \\
\mathcal{F}_{2}(b, g, r):= & \left(I-P_{0}\right) \tilde{\mathcal{F}}_{2}(b, g, r), \\
\tilde{\mathcal{F}}_{2}(b, g, r):= & f_{-\pi}^{\pi}(b+g(\eta)) \frac{(1+r(\theta))^{2}-\left(r^{\prime}(\theta) \sin (\theta-\eta)+(1+r(\theta)) \cos (\theta-\eta)\right)(1+r(\eta))}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)} d \eta \\
& \times \frac{(b+g(\theta))}{r^{\prime}(\theta)^{2}+(1+r(\theta))^{2}}-\Omega(1+r(\theta))^{2} \frac{b+g(\theta)}{r^{\prime}(\theta)^{2}+(1+r(\theta))^{2}} .
\end{aligned}
$$

Throughout the paper we will work with the following analytic function spaces. Let $c>0$ be a sufficiently small parameter and let $\mathcal{C}_{w}(c)$ be the space of analytic functions in the strip $|\Im(z)| \leq c$. For $k \in \mathbb{N}$, denote

$$
\begin{aligned}
& X_{c}^{k}:=\left\{f(\theta) \in \mathcal{C}_{w}(c), \quad f(\theta)=\sum_{n=1}^{\infty} a_{n} \cos (2 n \theta),\right. \\
& Y_{c}^{k}:=\left\{f(\theta) \in \mathcal{C}_{w}(c), \quad f(\theta)=\sum_{-\pi}^{\pi}|f(\theta \pm i c)|^{2}+\left|\partial^{k} f(\theta \pm i c)\right|^{2} d \theta<\infty\right\} \\
&\left.a_{n} \sin (2 n \theta), \quad \sum_{ \pm} \int_{-\pi}^{\pi}|f(\theta \pm i c)|^{2}+\left|\partial^{k} f(\theta \pm i c)\right|^{2} d \theta<\infty\right\},
\end{aligned}
$$

From now on, due to scaling considerations, we will fix $\Omega=1$ and $b$ will play the role as bifurcation parameter. It is clear that $\mathcal{F}(b, 0,0)=(0,0)$ for all $b \in \mathbb{R}$ since $\mathcal{F}_{1}(b, 0,0)=0$ and $\tilde{\mathcal{F}}_{2}(b, 0,0)$ is constant. Our main theorem in this paper is the following:

Theorem 5.1.2. Let $k \geq 3$, and let $c>0$ be sufficiently small. Then, there exists a curve of solutions $(b, g, r)$ of $\mathcal{F}=(0,0)$, belonging to $\mathbb{R} \times X_{c}^{k} \times X_{c}^{k+1}$ and a neighbourhood of $(b, g, r)=$ $(2,0,0)$, bifurcating from $(b, g, r)=(2,0,0)$ such that $(g, r) \neq(0,0)$.

### 5.2 Proof of Theorem 5.1.2

The goal of this section is to prove the existence of non-radial uniformly-rotating vortex sheets. To do so, we will split the proof into the following steps: first we will prove that the functional $\mathcal{F}$ is $C^{3}$, next we will study $D \mathcal{F}$ to show that, as mentioned in the introduction, it is a Fredholm operator of index 0 , with $\operatorname{dim}(\operatorname{Ker}(D \mathcal{F}))=1$. The next step is to apply Lyapunov-Schmidt theory and reduce the problem to a finite (2) dimensional one. In those coordinates, linear expansions fail to be conclusive (all the linear terms vanish) since 2 nontrivial branches emanate from the bifurcation point (as opposed to 1). Instead, we perform a quadratic expansion to determine that locally the bifurcation branches look like two pairs of straight lines (specifically as $x^{2}-y^{2}=0$ in some well-chosen coordinates) and hence the bifurcation does not trivialize (as if it had been of the type $x^{2}+y^{2}=0$ ). We conclude the proof by handling the higher order terms and showing that they don't alter the quadratic behaviour in a sufficiently small neighbourhood of the bifurcation point.

### 5.2.1 Continuity of the functional

In this subsection, we will check the regularity of $\mathcal{F}$. As explained above, we will reduce the infinite dimensional problem to a finite dimensional problem and investigate its Taylor expansion up to quadratic order. Hence, we need to check if the functional is regular enough to do so. To this end, we have the following proposition:

Proposition 5.2.1. Let $k \geq 3$. Then there exists a neighborhood $U$ of $(2,0,0) \in \mathbb{R} \times X_{c}^{k} \times X_{c}^{k+1}$ such that $\mathcal{F} \in C^{3}\left(U ; Y_{c}^{k} \times X_{c}^{k}\right)$.

Proof. Since the stream function, $\omega * \mathcal{N}$, is invariant under rotations, it follows immediately that $\mathcal{F}$ is also invariant under rotation by $\pi$-radians, hence $\mathcal{F}$ has only even Fourier modes. Also the oddness of $\mathcal{F}_{1}$ and evenness of $\mathcal{F}_{2}$ follow from the invariance under reflection.

To prove the regularity, we briefly sketch the idea. We impose $k \geq 3$ to ensure that $H^{k}$ is a Banach algebra. It is clear that $\mathcal{F}$ is smooth in $b$. It is also straightforward that, for example, for all $(g, r)$ near $(0,0) \in X_{c}^{k} \times X_{c}^{k+1}$,

$$
\begin{aligned}
&\left.\frac{d}{d t} \mathcal{F}_{1}\left(b, g+t h_{1}, r+t h_{2}\right)\right|_{t=0} \\
&= P V f h_{1}(\eta) \frac{\left(r^{\prime}(\theta) \cos (\theta-\eta)-(1+r(\theta)) \sin (\theta-\eta)\right)(1+r(\eta))-(1+r(\theta)) r^{\prime}(\theta)}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)} \\
&+(b+g(\eta))\left[\frac{\left(h_{2}^{\prime}(\theta) \cos (\theta-\eta)-h_{2}(\theta) \sin (\theta-\eta)\right)(1+r(\eta))-h_{2}(\theta) r^{\prime}(\theta)}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)}\right. \\
&+\frac{\left(r^{\prime}(\theta) \cos (\theta-\eta)-(1+r(\theta)) \sin (\theta-\eta)\right) h_{2}(\eta)-(1+r(\theta)) h_{2}^{\prime}(\theta)}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)} \\
&-\left[\left(r^{\prime}(\theta) \cos (\theta-\eta)-(1+r(\theta)) \sin (\theta-\eta)\right)(1+r(\eta))-(1+r(\theta)) r^{\prime}(\theta)\right] \\
&\left.\times \frac{\left[2 \left(1+r(\theta) h_{2}(\theta)+2\left(1+r(\eta) h_{2}(\eta)-2 \cos (\theta-\eta)\left(h_{2}(\theta)(1+r(\eta))+h_{2}(\eta)(1+r(\theta))\right)\right]\right.\right.}{\left((1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)\right)^{2}}\right] d \eta \\
&+\Omega h_{2}^{\prime}(\theta) \\
&= D \mathcal{F}_{1}(b, g, r)\left[h_{1}, h_{2}\right],
\end{aligned}
$$

and $D \mathcal{F}_{1}: \mathbb{R} \times X_{c}^{k} \times X_{c}^{k+1} \mapsto \mathcal{L}\left(X_{c}^{k} \times X_{c}^{k+1} ; Y_{c}^{k} \times X_{c}^{k}\right)$ is continuous. A similar derivation can be performed for $D \mathcal{F}_{2}$. For the higher derivatives, we refer to [21, 22, 49, 63, 100] for the method to deal with the singular integrals arising throughout the calculations.

### 5.2.2 Fredholm index of the linearized operator $D \mathcal{F}$

This subsection is devoted to show that $D \mathcal{F}$ is Fredholm of index zero. We can make all the calculations explicit, moreover the operator diagonalizes in Fourier modes. We have the following lemmas:

Lemma 5.2.2. Let $g(\theta)=\sum_{n=1}^{\infty} a_{n} \cos (2 n \theta)$ and $r(\theta)=\sum_{n=1}^{\infty} b_{n} \cos (2 n \theta)$. Then we have that

$$
D \mathcal{F}(b, 0,0)[g, r]=\binom{\hat{g}(\theta)}{\hat{r}(\theta)}
$$

where

$$
\hat{g}(\theta)=\sum_{n=1}^{\infty} \hat{a}_{n} \sin (2 n \theta), \quad \hat{r}(\theta)=\sum_{n=1}^{\infty} \hat{b}_{n} \cos (2 n \theta)
$$

and the coefficients satisfy, for any $n \geq 1$ :

$$
M_{n}\binom{a_{n}}{b_{n}}:=\left(\begin{array}{cc}
-\frac{1}{2} & -2 n\left(\Omega-\frac{b}{2}\right) \\
\frac{b}{2}-\Omega & b^{2}(n-1)
\end{array}\right)\binom{a_{n}}{b_{n}}=\binom{\hat{a}_{n}}{\hat{b}_{n}} .
$$

Proof. We use (Equation A.1.3) in Lemma A.1.2 and obtain

$$
\begin{aligned}
\hat{g}(\theta)=\frac{d}{d t} \mathcal{F}_{1}(b, t g, t r) & =-f \frac{g(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta) \\
& =-\sum_{n=1}^{\infty} a_{n} f \frac{\cos (2 n \eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+\sum_{n=1}^{\infty}(-2 n)\left(\Omega-\frac{b}{2}\right) b_{n} \sin (2 n \theta) \\
& =\sum_{n=1}^{\infty}\left(-\frac{a_{n}}{2}+(-2 n)\left(\Omega-\frac{b}{2}\right) b_{n}\right) \sin (2 n \theta)
\end{aligned}
$$

where the last equality follows from (Equation A.1.29). Similarly, we apply (Equation A.1.4) in Lemma A.1.2 and (Equation A.1.30) to obtain

$$
\begin{aligned}
\hat{r}(\theta) & =\left(\frac{b}{2}-\Omega\right) g(\theta)+b^{2}\left(f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)} d \eta-r(\theta)\right) \\
& =\sum_{n=1}^{\infty}\left(\frac{b}{2}-\Omega\right) a_{n} \cos (2 n \theta)+b^{2} \sum_{n=1}^{\infty} b_{n}\left(f \frac{\cos (2 n \theta)-\cos (2 n \eta)}{2-2 \cos (\theta-\eta)} d \eta-\cos (2 n \theta)\right) \\
& =\sum_{n=1}^{\infty}\left(\left(\frac{b}{2}-\Omega\right) a_{n}+b^{2}(n-1) b_{n}\right) \cos (2 n \theta) .
\end{aligned}
$$

This proves the lemma.

Lemma 5.2.3. Let us fix $b=2$ and $\Omega=1$. We also denote $v:=(0, \cos (2 \theta)) \in X_{c}^{k} \times X_{c}^{k+1}$ and $w:=(0, \cos (2 \theta)) \in Y_{c}^{k} \times X_{c}^{k}$. Then it holds that

$$
\begin{aligned}
& \operatorname{Ker}(D \mathcal{F}(2,0,0))=\operatorname{span}\{v\} \subset X_{c}^{k} \times X_{c}^{k+1}, \\
& \operatorname{Im}(D \mathcal{F}(2,0,0))^{\perp}=\operatorname{span}\{w\} \subset Y_{c}^{k} \times X_{c}^{k}
\end{aligned}
$$

Proof. From Lemma 5.2.2, we have

$$
M_{n}=\left(\begin{array}{cc}
-\frac{1}{2} & 0 \\
0 & 4(n-1)
\end{array}\right)
$$

for all $n \geq 1$. For all $n \geq 2, M_{n}$ is clearly an isomorphism, while $\operatorname{Ker}\left(M_{1}\right)=\operatorname{Im}\left(M_{1}\right)^{\perp}=\binom{0}{1}$. By orthogonality of Fourier modes, this proves the lemma.

### 5.2.3 Lyapunov-Schmidt reduction

In this subsection, we will aim to derive a finite dimensional system which is equivalent to (Equation 5.1.4). From Lemma 5.2.3, we have the following orthogonal decompositions of the function spaces:

$$
\begin{gathered}
X:=X_{c}^{k} \times X_{c}^{k+1}=\operatorname{span}\{v\} \oplus \operatorname{Ker}(D \mathcal{F}(2,0,0))^{\perp}=: \operatorname{span}\{v\} \oplus \mathcal{X}, \quad v \in \operatorname{Ker}(D \mathcal{F}(2,0,0)), \\
Y:=Y_{c}^{k} \times X_{c}^{k}=\operatorname{span}\{w\} \oplus \operatorname{Im}(D \mathcal{F}(2,0,0))=: \operatorname{span}\{w\} \oplus \mathcal{Y}, \quad w \in \operatorname{Im}(D \mathcal{F}(2,0,0))^{\perp},
\end{gathered}
$$

where $v$ and $w$ are as defined in Lemma 5.2.3. Let us consider the orthogonal projections

$$
P: X \rightarrow \operatorname{span}\{v\}, \quad Q: Y \rightarrow \operatorname{span}\{w\} .
$$

More precisely, we have

$$
\begin{align*}
& P(g(\theta), r(\theta))=\left(0,\left(\frac{1}{\pi} \int r(\eta) \cos (2 \eta) d \eta\right) \cos (2 \theta)\right) \quad \text { for all }(g, r) \in X_{c}^{k} \times X_{c}^{k+1}  \tag{5.2.1}\\
& Q(G(\theta), R(\theta))=\left(0,\left(\frac{1}{\pi} \int R(\eta) \cos (2 \eta) d \eta\right) \cos (2 \theta)\right) \quad \text { for all }(G, R) \in Y_{c}^{k} \times X_{c}^{k} \tag{5.2.2}
\end{align*}
$$

We remark that we will sometimes abuse notation and identify $\mathcal{F}(b, g, r)$ with $\mathcal{F}(b,(g, r))$, where $(g, r) \in X$. Let us define $G: \mathbb{R} \times \operatorname{span}\{v\} \times \mathcal{X} \mapsto Y$ as follows:

$$
G(b, f, x):=\mathcal{F}(b, f+x), \quad \text { for } \quad b \in \mathbb{R}, \quad f \in \operatorname{span}\{v\}, \quad x \in \mathcal{X} .
$$

Then (Equation 5.1.4) is equivalent to $($ for $(g, r)=f+x)$

$$
\begin{equation*}
Q G(b, f, x)=0 \quad \text { and } \quad(I-Q) G(b, f, x)=0 \tag{5.2.3}
\end{equation*}
$$

However, it follows from Lemma 5.2.3 that

$$
\begin{equation*}
D_{x}((I-Q) G)(b, 0,0)=(I-Q) D \mathcal{F}(b, 0) P: \mathcal{X} \mapsto \mathcal{Y} \tag{5.2.4}
\end{equation*}
$$

is an isomorphism, consequently, the implicit function theorem yields that there is an open set $U \subset \mathbb{R} \times \operatorname{span}\{v\}$ near $(b, 0)$ and a function $\varphi: U \mapsto \mathcal{X}$ such that

$$
(I-Q) G(b, f, \varphi(b, f))=(I-Q) \mathcal{F}(b, f+\varphi(b, f))=0 .
$$

Note that from $\mathcal{F}(b, 0)=(0,0)$ for any $b \in \mathbb{R}$, we have

$$
\begin{equation*}
\varphi(b, 0)=0 \tag{5.2.5}
\end{equation*}
$$

and thus (Equation 5.2.3) is equivalent to

$$
\begin{equation*}
0=Q G(b, f, \varphi(b, f))=Q \mathcal{F}(b, f+\varphi(b, f)), \quad(b, f) \in U \tag{5.2.6}
\end{equation*}
$$

Since span $\{v\}$ is one dimensional, we have $f=t v$ for some $t \in \mathbb{R}$, therefore the system (Equation 5.2.6) can be written in terms of the variables $b$ and $t$ as

$$
\begin{aligned}
0=Q \mathcal{F}(b, t v+\varphi(b, t v)) & =\int_{0}^{1} \frac{d}{d s}(Q \mathcal{F}(b, s t v+\varphi(b, s t v))) d s \\
& =\int_{0}^{1} Q D \mathcal{F}(b, s t v+\varphi(b, s t v))\left(t v+t \partial_{f} \varphi(b, s t v) v\right) d s
\end{aligned}
$$

where we used (Equation 5.2.5) to obtain the second equality. Dividing the right-hand side by $t$ to get rid of the trivial solutions, we are led to solve the following two dimensional problem:

$$
\begin{equation*}
0=F_{r e d}(b, t):=\int_{0}^{1} Q D \mathcal{F}(b, s t v+\varphi(b, s t v))\left(v+\partial_{f} \varphi(b, s t v) v\right) d s, \quad(b, t v) \in U \tag{5.2.7}
\end{equation*}
$$

### 5.2.4 Quadratic expansion of the reduced functional

The main idea is to expand the reduced functional $F_{r e d}$ up to quadratic terms. To this end, we recall the following proposition for the derivatives of $F_{\text {red }}$.

Proposition 5.2.4. ([61, Proposition 3], [65, Proposition 3.1]) Let $F_{\text {red }}$ be defined as in (Equation 5.2.7). Then the following hold:
(a) First derivatives:

$$
\begin{aligned}
\partial_{b} F_{r e d}(2,0) & =Q \partial_{b} D \mathcal{F}(2,0) v, \\
\partial_{t} F_{r e d}(2,0) & =\left.\frac{1}{2} \frac{d^{2}}{d t^{2}} Q \mathcal{F}(2, t v)\right|_{t=0}
\end{aligned}
$$

(b) Second derivatives:

$$
\begin{aligned}
& \partial_{b b} F_{r e d}(2,0)= 2 Q \partial_{b} D \mathcal{F}(2,0) \tilde{v}, \\
& \partial_{t t} F_{r e d}(2,0)=\left.\frac{1}{3} \frac{d^{3}}{d t^{3}}[Q \mathcal{F}(2, t v)]\right|_{t=0}+\left.Q \frac{d^{2}}{d t d s} \mathcal{F}(2, t v+s \hat{v})\right|_{t=s=0} \\
& \partial_{t b} F_{r e d}(2,0)=\left.\frac{1}{2} \partial_{b} Q \frac{d^{2}}{d t^{2}} \mathcal{F}(b, t v)\right|_{b=2, t=0} \\
& \quad+\frac{1}{2} Q \partial_{b} D \mathcal{F}(2,0) \hat{v}+\left.Q \frac{d^{2}}{d t d s} \mathcal{F}(2, t v+s \tilde{v})\right|_{t=s=0}
\end{aligned}
$$

where

$$
\begin{aligned}
& \hat{v}:=-\left.[D \mathcal{F}(2,0)]^{-1} \frac{d^{2}}{d t^{2}}[(I-Q) \mathcal{F}(2, t v)]\right|_{t=0} \\
& \tilde{v}:=-[D \mathcal{F}(2,0)]^{-1}(I-Q) \partial_{b} D \mathcal{F}(2,0) v
\end{aligned}
$$

Now using the values found in Lemma A.1.7, we can obtain the derivatives of $F_{\text {red }}$.

Proposition 5.2.5. Let $F_{\text {red }}$ be defined as in (Equation 5.2.7). Then it holds that

$$
\begin{align*}
& \partial_{b} F_{r e d}(2,0)=0  \tag{5.2.8}\\
& \partial_{t} F_{r e d}(2,0)=0  \tag{5.2.9}\\
& \partial_{b b} F_{r e d}(2,0)=2 w,  \tag{5.2.10}\\
& \partial_{t t} F_{r e d}(2,0)=-8 w,  \tag{5.2.11}\\
& \partial_{t b} F_{r e d}(2,0)=0 \tag{5.2.12}
\end{align*}
$$

Proof. (Equation 5.2.8) follows immediately from (Equation 5.2.2) and (Equation A.1.14). For (Equation 5.2.9), we use (Equation A.1.15) and the orthogonality of the Fourier modes. (Equation 5.2.10) follows from (Equation A.1.23). (Equation 5.2.11) follows from (Equation A.1.19) and (Equation A.1.20). Lastly, (Equation 5.2.12) follows from (Equation A.1.16), (Equation A.1.21) and (Equation A.1.22).

### 5.2.5 Proof of Theorem 5.1.2

Now we are ready to prove the main theorem of this section.

Proof. From (Equation 5.2.7), it suffices to show that there exist $(b, t)$ such that $t \neq 0$ and $F_{r e d}(b, t)=0$. To do so, we expand $F_{r e d}$ up to quadratic terms and obtain that for all $(b, t)$ near $(2,0)$,

$$
\begin{aligned}
F_{r e d}(b, t) & =\left[F_{r e d}(2,0)+\partial_{b} F_{r e d}(2,0)(b-2)+\partial_{t} F_{r e d}(2,0) t\right. \\
& \left.+\frac{1}{2} \partial_{b b} F_{r e d}(2,0)(b-2)^{2}+\frac{1}{2} \partial_{t t} F_{r e d}(2,0) t^{2}+\partial_{t b} F_{r e d}(2,0)(b-2) t+\left((b-2)^{2}+t^{2}\right) \epsilon(b, t)\right] w,
\end{aligned}
$$

where $\epsilon(b, t)$ is a continuous function such that $\lim _{(b, t) \rightarrow(2,0)} \epsilon(b, t)=\epsilon(2,0)=0$. From Proposition 5.2.5, it follows that (we drop $w$ for simplicity)

$$
F_{r e d}(b, t)=(b-2)^{2}-4 t^{2}+\left((b-2)^{2}+t^{2}\right) \epsilon(b, t) .
$$

Now we use the change of variables $b:=x+2$ and $t=x y$, so that

$$
\begin{equation*}
\hat{F}(x, y):=\frac{F_{r e d}(x+2, x y)}{x^{2}}=\left(1-4 y^{2}\right)+\left(1+y^{2}\right) \epsilon(x+2, x y) . \tag{5.2.13}
\end{equation*}
$$

Clearly, $\hat{F}\left(0, \frac{1}{2}\right)=0$ and $\partial_{y} \hat{F}\left(0, \frac{1}{2}\right)=-4 \neq 0$. Therefore the implicit function theorem implies that there exists a continuous function $\phi$ near 0 such that $\hat{F}(x, \phi(x))=0$ and $\phi(0)=\frac{1}{2}$. Therefore it follows from (Equation 5.2.13) that there exists a pair $(b, t)$ such that $t \neq 0$ and $F_{r e d}(b, t)=0$. This finishes the proof.


Figure 5.1: Plot of the bifurcation diagram of the solutions given by Theorem 5.1.2. The dotted red lines correspond to the linear expansion (Equation 5.2.13) around the bifurcation point $(2,0)$. See Figure Figure 5.2 for a numerical plot of the solutions at the points $A, B, C, D$. The branches continue beyond what is calculated.

### 5.3 Numerical results

In this section, we describe how to compute numerically the branches of solutions emanating from the disk, previously proved (locally) in Theorem 5.1.2. See Figure Figure 5.1. To do so, we calculate solutions of the form

$$
R(\theta)=1+\sum_{k=1}^{N} r_{k} \cos (2 k \theta), \quad \gamma(\theta)=\sum_{k=0}^{N} \gamma_{k} \cos (2 k \theta)
$$

with $\gamma_{0}=b$. We first employed continuation in $b$, in increments of $\Delta b=0.001$, starting from $b=1.8$ and $b=2.1$ and using as initial guess for the starting $b$ the solution given by the linear theory and for the subsequent $b$ the solution found in the previous iteration. After discovering a fold at approximately $b \sim 1.68$, we switched variables and instead we recalculated using continuation in $r_{1}$, which appears to be monotonic along the branches. As before, we start at $r_{1}= \pm 0.125$ and take an increment $\Delta r_{1}=0.001$.

To compute a solution for a fixed $r_{1}$ we use the Levenberg-Marquardt algorithm. We aim to find a zero of the system of equations $\mathcal{F}(b, g, r)\left(\theta_{j}\right)$, with $\theta_{j}=\frac{j \pi}{N_{\theta}}, \quad j=1, \ldots, N_{\theta}$ and $N_{\theta}=1024$ with variables $r_{k}, k \neq 1$ (recall that $r_{1}$ is fixed at each iteration since it is the continuation parameter) and $\gamma_{k}$. We take $N=160$. In order to perform the integration in space, we desingularize the principal value at $\eta=\theta$ by subtracting $\frac{1}{2} \mathcal{H}(\gamma)$ to $\mathcal{F}_{1}$, where $\mathcal{H}$ denotes the Hilbert transform, computed explicitly since we have the Fourier expansion of $\gamma$, and perform a trapezoidal integration on the rest (for which the integrand is smooth), with step $h=\frac{2 \pi}{N_{\theta}}$. We remark that the integrand of $\mathcal{F}_{2}$ has a removable singularity (thus no principal value integration is needed) and can be integrated using the trapezoidal integration if the limit at $\eta=\theta$ is taken properly.


(a) $\gamma(\theta)$ and $z(\theta)$ for $r_{1}=0.362, b \sim 1.6799$


(c) $\gamma(\theta)$ and $z(\theta)$ for $r_{1}=0.525, b \sim 4.0954$


(b) $\gamma(\theta)$ and $z(\theta)$ for $r_{1}=0.825, b \sim 1.7779$


(d) $\gamma(\theta)$ and $z(\theta)$ for $r_{1}=0.925, b \sim 9.3439$

Figure 5.2: Panels (a)-(d): $\gamma(\theta)$ and $z(\theta)$ at the points $A-D$ highlighted in Figure Figure 5.1. In panel (b), $\gamma$ appears to tend to be concentrated only on the horizontal parts of $z$, leading to a possible solution consisting only of two symmetric curves (cf. [93, Figure 1]) and a change of topology.

## CHAPTER 6

## QUANTITATIVE ESTIMATES FOR UNIFORMLY ROTATING VORTEX PATCHES

### 6.1 Quantitative estimates for small $\Omega$

This section is devoted to the proof of Theorem G. Throughout this section, we will always assume that $|D|=|B|=\pi$. We begin this section by proving two identities for simply-connected rotating patches.

Lemma 6.1.1. Let $(D, \Omega)$ be a solution to (Equation 1.2.4). Then it holds that

$$
\begin{align*}
& \Omega\left(\int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}\right)=(1-2 \Omega)\left(\frac{|D|^{2}}{4 \pi}-\int_{D} p d x\right)  \tag{6.1.1}\\
& \left(\frac{1}{2}-\Omega\right)\left(\int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}\right)=\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \tag{6.1.2}
\end{align*}
$$

where $p$ is the unique solution to (Equation 1.2.7).

Proof. The proof of (Equation 6.1.1) can be found in [50, Theorem 2.2]. For the sake of completeness, we give a proof below.

In order to prove (Equation 6.1.1), we plug $\vec{v}=x+\nabla p$ into (Equation 1.2.6) to get

$$
\begin{align*}
0 & =\int_{D}(x+\nabla p) \cdot \nabla\left(1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x \\
& =\int_{D} x \cdot \nabla\left(1_{D} * \mathcal{N}\right)-\Omega \int_{D}|x|^{2} d x+\int_{D} \nabla p \cdot \nabla\left(1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x \\
& =\int_{D} x \cdot \nabla\left(1_{D} * \mathcal{N}\right)-\Omega \int_{D}|x|^{2} d x-(1-2 \Omega) \int_{D} p d x \tag{6.1.3}
\end{align*}
$$

where we used divergence theorem for the last equality. Note that the first integral can be computed

$$
\begin{equation*}
\int_{D} x \cdot \nabla\left(1_{D} * \mathcal{N}\right) d x=\frac{1}{2 \pi} \int_{D} \int_{D} \frac{x \cdot(x-y)}{|x-y|^{2}} d y d x=\frac{|D|^{2}}{4 \pi}, \tag{6.1.4}
\end{equation*}
$$

where the last equality is obtained by exchanging $x$ and $y$ in the double integral, and then taking the average with the original integral. Therefore (Equation 6.1.3) and (Equation 6.1.4) yield

$$
0=\frac{|D|^{2}}{4 \pi}-\Omega \int_{D}|x|^{2} d x-(1-2 \Omega) \int_{D} p d x
$$

which is equivalent to (Equation 6.1.1).
For (Equation 6.1.2), we choose $\vec{v}=x-2 \nabla\left(1_{D} * \mathcal{N}\right)$ in (Equation 1.2.6) and obtain

$$
\begin{align*}
0 & =\int_{D}\left(x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right) \cdot \nabla\left(1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}\right) d x \\
& =(1+2 \Omega) \frac{|D|^{2}}{4 \pi}-2 \int_{D}\left|\nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x-\Omega \int_{D}|x|^{2} d x \tag{6.1.5}
\end{align*}
$$

where we used (Equation 6.1.4). Since $\int_{D}\left|\nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x$ can be computed as

$$
\begin{aligned}
2 \int_{D}\left|\nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x & =\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x+2 \int_{D} x \cdot \nabla\left(1_{D} * \mathcal{N}\right) d x-\frac{1}{2} \int_{D}|x|^{2} d x \\
& =\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x+\frac{|D|^{2}}{2 \pi}-\frac{1}{2} \int_{D}|x|^{2} d x
\end{aligned}
$$

plugging it into (Equation 6.1.5) yields (Equation 6.1.2).

Thanks to Lemma 6.1.1, the angular velocity can be estimated by comparing the quantities, $\int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}, \frac{|D|^{2}}{4 \pi}-\int_{D} p d x$ and $\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x$, which vanish if and only if $D=B$. To estimate those quantities for non-radial patches, we use the following notion of asymmetry.

Definition 6.1.2. [42, Section 1.1] For a bounded domain $D \subseteq \mathbb{R}^{2}$, the Fraenkel asymmetry $\mathcal{A}(D)$
is defined by

$$
\mathcal{A}(D):=\inf _{x \in \mathbb{R}^{2}}\left\{\frac{\left|D \triangle B_{r}(x)\right|}{|D|}: \pi r^{2}=|D|\right\} .
$$

If $\mathcal{A}(D)$ is not small, then we can find a lower bound of the right-hand side in (Equation 6.1.1) by using the following result:

Proposition 6.1.3. [10, Proposition 2.1] Let p be as in (Equation 1.2.7) and $|D|=\pi$. Then there exists a constant $\sigma>0$ such that

$$
\begin{equation*}
\frac{|D|^{2}}{4 \pi}-\int_{D} p d x \geq \sigma \mathcal{A}(D)^{2} \tag{6.1.6}
\end{equation*}
$$

Using the above proposition and the identity (Equation 6.1.1), one can easily show that $\sup _{x \in \partial D}|x| \gtrsim \sqrt{\mathcal{A}(D)} \Omega^{-\frac{1}{2}}$. Therefore Theorem $G$ can be proved if we can show $\mathcal{A}(D)$ is always bounded below by a strict positive constant. In other words, we will aim to prove in the next lemmas that if $\mathcal{A}(D)$ and $\Omega$ are sufficiently small, then $D$ must be a disk.

In the following lemma, we will estimate the boundedness of rotating patches in a crude way but this will be improved later.

Lemma 6.1.4. There exist positive constants $\Omega_{1}$ and $\alpha_{1}<\frac{1}{2}$ such that if $\Omega<\Omega_{1}$ and $\mathcal{A}(D)<\alpha_{1}$, then

$$
\begin{align*}
D & \subset B_{2}\left(x_{0}\right)  \tag{6.1.7}\\
\left|x_{0}\right| & \leq 4 \mathcal{A}(D) \tag{6.1.8}
\end{align*}
$$

where $x_{0}$ is a point such that $\frac{\left|D \Delta B\left(x_{0}\right)\right|}{\pi}=\mathcal{A}(D)$.

Proof. Let us pick $\Omega_{1}$ and $\alpha_{1}$ so that for all $\Omega<\Omega_{1}$ and $\alpha<\alpha_{1}<\frac{1}{2}$, it holds that

$$
\begin{equation*}
\frac{1}{2} \log 2 \geq \Omega\left(10+\frac{4 \alpha}{(1-\sqrt{2 \alpha})^{2}}\left(1+\sqrt{\frac{1}{4 \Omega}}\right)^{2}\right)+3 \sqrt{2 \alpha} \tag{6.1.9}
\end{equation*}
$$

We will first show that if $(D, \Omega)$ satisfies $\Omega<\Omega_{1}$ and $\mathcal{A}(D)<\alpha_{1}$, then (Equation 6.1.7) holds.
Note that the center of mass of $D$ is necessarily the origin ([64, Proposition 3]). Therefore we have

$$
\begin{equation*}
0=\frac{1}{\pi} \int x 1_{D}(x) d x=\frac{1}{\pi} \int x\left(1_{D}(x)-1_{B\left(x_{0}\right)}(x)\right) d x+x_{0} \tag{6.1.10}
\end{equation*}
$$

Hence it follows from Cauchy-Schwarz inequality that

$$
\begin{align*}
\left|x_{0}\right| \leq \frac{1}{\pi} \int|x|\left|1_{D}(x)-1_{B\left(x_{0}\right)}(x)\right| d x & \leq \frac{1}{\pi}\left(\sqrt{\left|D \triangle B\left(x_{0}\right)\right|} \sqrt{\int_{D}|x|^{2} d x+\int_{B\left(x_{0}\right)}|x|^{2} d x}\right) \\
& \leq \sqrt{\frac{\mathcal{A}(D)}{\pi}} \sqrt{\int_{D}|x|^{2} d x}+\sqrt{\frac{\mathcal{A}(D)}{\pi}} \sqrt{\int_{B\left(x_{0}\right)}|x|^{2} d x} \tag{6.1.11}
\end{align*}
$$

Since $\sqrt{\int_{B\left(x_{0}\right)}|x|^{2} d x} \leq \sqrt{\int_{B\left(x_{0}\right)} 2\left|x-x_{0}\right|^{2} d x+2 \int_{B\left(x_{0}\right)}\left|x_{0}\right|^{2} d x}=\sqrt{\pi+2\left|x_{0}\right|^{2} \pi} \leq \sqrt{\pi}+$ $\sqrt{2 \pi}\left|x_{0}\right|$, (Equation 6.1.11) yields that

$$
\begin{equation*}
(1-\sqrt{2 \mathcal{A}(D)})\left|x_{0}\right|<\sqrt{\mathcal{A}(D)}+\sqrt{\frac{\mathcal{A}(D)}{\pi}} \sqrt{\int_{D}|x|^{2} d x} . \tag{6.1.12}
\end{equation*}
$$

In addition, it follows from (Equation 6.1.1) that

$$
\Omega \int_{D}|x|^{2} d x=\frac{|D|^{2}}{4 \pi}-(1-2 \Omega) \int_{D} p d x<\frac{\pi}{4}
$$

where we used $\Omega<\frac{1}{2}, p \geq 0$ in $D$ and $|D|=\pi$ to get the last inequality. Plugging this into
(Equation 6.1.12), we obtain

$$
(1-\sqrt{2 \mathcal{A}(D)})\left|x_{0}\right|<\sqrt{\mathcal{A}(D)}\left(1+\sqrt{\frac{1}{4 \Omega}}\right)
$$

hence,

$$
\begin{equation*}
\left|x_{0}\right|<\frac{\sqrt{\mathcal{A}(D)}}{1-\sqrt{2 \mathcal{A}(D)}}\left(1+\sqrt{\frac{1}{4 \Omega}}\right) . \tag{6.1.13}
\end{equation*}
$$

To prove (Equation 6.1.7), let us suppose to the contrary that there exist $x_{1} \in \partial B\left(x_{0}\right) \cap \partial D$ and $x_{2} \in \partial B_{2}\left(x_{0}\right) \cap \partial D$. Then it follows from (Equation 1.2.4) that $0=\Psi\left(x_{1}\right)-\Psi\left(x_{2}\right)$, therefore

$$
\begin{equation*}
1_{B\left(x_{0}\right)} * \mathcal{N}\left(x_{2}\right)-1_{B\left(x_{0}\right)} * \mathcal{N}\left(x_{1}\right)=\Omega\left(\left|x_{2}\right|^{2}-\left|x_{1}\right|^{2}\right)+h\left(x_{1}\right)-h\left(x_{2}\right), \tag{6.1.14}
\end{equation*}
$$

where $h(x):=\left(1_{D}-1_{B\left(x_{0}\right)}\right) * \mathcal{N}$. For the left-hand side, we use

$$
1_{B\left(x_{0}\right)} * \mathcal{N}= \begin{cases}\frac{1}{4}\left|x-x_{0}\right|^{2}-\frac{1}{4} & \text { if }\left|x-x_{0}\right|<1 \\ \frac{1}{2} \log \left|x-x_{0}\right| & \text { otherwise }\end{cases}
$$

and obtain

$$
\begin{equation*}
1_{B\left(x_{0}\right)} * \mathcal{N}\left(x_{2}\right)-1_{B\left(x_{0}\right)} * \mathcal{N}\left(x_{1}\right)=\frac{1}{2} \log 2 . \tag{6.1.15}
\end{equation*}
$$

For $\Omega\left(\left|x_{2}\right|^{2}-\left|x_{1}\right|^{2}\right)$ in the right-hand side of (Equation 6.1.14), we use the triangular inequality
and (Equation 6.1.13) to obtain

$$
\begin{align*}
& \left|x_{1}\right|^{2}<2\left|x_{1}-x_{0}\right|^{2}+2\left|x_{0}\right|^{2}<2+\frac{2 \mathcal{A}(D)}{(1-\sqrt{2 \mathcal{A}(D)})^{2}}\left(1+\sqrt{\frac{1}{4 \Omega}}\right)^{2}  \tag{6.1.16}\\
& \left|x_{2}\right|^{2}<2\left|x_{2}-x_{0}\right|^{2}+2\left|x_{0}\right|^{2}<8+\frac{2 \mathcal{A}(D)}{(1-\sqrt{2 \mathcal{A}(D)})^{2}}\left(1+\sqrt{\frac{1}{4 \Omega}}\right)^{2} \tag{6.1.17}
\end{align*}
$$

To estimate $h\left(x_{1}\right)-h\left(x_{2}\right)$, we use the fact that

$$
\begin{equation*}
\|\nabla f * \mathcal{N}\|_{L^{\infty}} \leq \sqrt{\frac{2}{\pi}} \sqrt{\|f\|_{L^{1}}\|f\|_{L^{\infty}}} \quad \text { for any } f \in L^{1} \cap L^{\infty}\left(\mathbb{R}^{2}\right) \tag{6.1.18}
\end{equation*}
$$

Indeed, one can compute with $a:=\sqrt{\frac{\|f\|_{L^{1}}\left(\mathbb{R}^{2}\right)}{2 \pi\|f\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}}}$,

$$
|\nabla f * \mathcal{N}(x)| \leq \frac{1}{2 \pi} \int_{|x-y|>a} \frac{1}{|x-y|}|f(y)| d y+\frac{1}{2 \pi} \int_{|x-y| \leq a} \frac{1}{|x-y|}|f(y)| d y \leq \frac{\|f\|_{L^{1}\left(\mathbb{R}^{2}\right)}}{2 \pi a}+a\|f\|_{L^{\infty}\left(\mathbb{R}^{2}\right)},
$$

which yields (Equation 6.1.18). Thus we have

$$
\begin{equation*}
\left|h\left(x_{1}\right)-h\left(x_{2}\right)\right|<|\nabla h|_{L^{\infty}}\left|x_{1}-x_{2}\right|<\sqrt{\frac{2}{\pi}} \sqrt{\left|D \triangle B\left(x_{0}\right)\right|}\left|x_{1}-x_{2}\right|=3 \sqrt{2 \mathcal{A}(D)} \tag{6.1.19}
\end{equation*}
$$

Hence it follows from (Equation 6.1.14), (Equation 6.1.15), (Equation 6.1.16), (Equation 6.1.17) and (Equation 6.1.19) that

$$
\begin{equation*}
\frac{1}{2} \log 2<\Omega\left(10+\frac{4 \mathcal{A}(D)}{(1-\sqrt{2 \mathcal{A}(D)})^{2}}\left(1+\sqrt{\frac{1}{4 \Omega}}\right)^{2}\right)+3 \sqrt{2 \mathcal{A}(D)} \tag{6.1.20}
\end{equation*}
$$

which contradicts our choice of $\Omega_{1}$ and $\alpha_{1}$ for (Equation 6.1.9). This proves the claim (Equation 6.1.7).

To prove (Equation 6.1.8), let us fix $\Omega_{1}$ and $\alpha_{1}<\frac{1}{2}$ so that the claim (Equation 6.1.7)
holds. Then for all uniformly-rotating patches with $\Omega<\Omega_{1}$ and $\mathcal{A}(D)<\alpha_{1}$, it follows from (Equation 6.1.10) that

$$
\left|x_{0}\right| \leq \frac{1}{\pi}\left|D \triangle B\left(x_{0}\right)\right| \sup _{x \in D \cup B\left(x_{0}\right)}|x| \leq \mathcal{A}(D)\left(\left|x_{0}\right|+2\right)
$$

where we used (Equation 6.1.7) to get the last inequality. Therefore

$$
\left|x_{0}\right|<\frac{2 \mathcal{A}(D)}{1-\mathcal{A}(D)}<4 \mathcal{A}(D)
$$

where we used $\mathcal{A}(D) \leq \alpha_{1}<\frac{1}{2}$ for the last inequality. Hence (Equation 6.1.8) is proved.

Since we are interested in patches that rotate about the origin, let us consider the asymmetry between $D$ and the unit disk centered at the origin:

$$
\mathcal{A}_{0}(D):=\frac{|D \triangle B|}{|D|}=\frac{|D \triangle B|}{\pi} .
$$

Tautologically, it holds that $\mathcal{A}(D) \leq \mathcal{A}_{0}(D)$. For rotating patches, we have the following lemma:

Lemma 6.1.5. There exist positive constants $\Omega_{1}$ and $c_{1}$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{1}$, then

$$
\begin{equation*}
\mathcal{A}_{0}(D) \leq c_{1} \mathcal{A}(D) \tag{6.1.21}
\end{equation*}
$$

Proof. Let $x_{0}$ be a point in $\mathbb{R}^{2}$ such that $\frac{\left|D \cap B\left(x_{0}\right)\right|}{\pi}=\mathcal{A}(D)$. By Lemma 6.1.4, we can pick $\Omega_{1}$ and $\alpha_{1}<\frac{1}{2}$ such that if $\Omega<\Omega_{1}$ and $\mathcal{A}(D)<\alpha_{1}$, then

$$
\begin{equation*}
\left|x_{0}\right| \leq 4 \mathcal{A}(D) \tag{6.1.22}
\end{equation*}
$$

Let us assume for a moment that the claim is true. If $\mathcal{A}(D) \geq \alpha_{1}$, then it follows from the definition of $\mathcal{A}_{0}$ that

$$
\begin{equation*}
\mathcal{A}_{0}(D) \leq \frac{|D|+|B|}{\pi}=2 \leq \frac{2}{\alpha_{1}} \mathcal{A}(D) . \tag{6.1.23}
\end{equation*}
$$

Now let us assume that $\mathcal{A}(D)<\alpha_{1}$. For a constant $c>0$ such that $\left|B \triangle B\left(x_{0}\right)\right| \leq c\left|x_{0}\right|$, we can compute

$$
\begin{equation*}
\mathcal{A}_{0}(D)=\frac{|D \triangle B|}{\pi} \leq \frac{\left|D \triangle B\left(x_{0}\right)\right|+\left|B \triangle B\left(x_{0}\right)\right|}{\pi} \leq \mathcal{A}(D)+\frac{c\left|x_{0}\right|}{\pi} \leq\left(1+\frac{4 c}{\pi}\right) \mathcal{A}(D), \tag{6.1.24}
\end{equation*}
$$

where the last inequality follows from (Equation 6.1.22). Therefore (Equation 6.1.21) follows from (Equation 6.1.23) and (Equation 6.1.24) by choosing $c_{1}:=\max \left\{\frac{2}{\alpha_{1}},\left(1+\frac{4 c}{\pi}\right)\right\}$.

In the next lemma, we will prove that if $\mathcal{A}_{0}(D)$ is sufficiently small, then $D$ is necessarily star-shaped.

Lemma 6.1.6. There exist positive constants $\Omega_{2}, \alpha_{2}$ and $c_{2}$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{2}$ and $\mathcal{A}_{0}(D)<\alpha_{2}$, then there exists $u \in C^{1}(\mathbb{T})$ such that

$$
\begin{equation*}
\partial D=\{(1+u(\theta))(\cos \theta, \sin \theta): \theta \in \mathbb{T}\} \tag{6.1.25}
\end{equation*}
$$

and

$$
\begin{equation*}
\|u\|_{L^{\infty}} \leq c_{2} \mathcal{A}_{0}(D)\left|\log \mathcal{A}_{0}(D)\right| \tag{6.1.26}
\end{equation*}
$$

Proof. Without loss of generality, we assume that $D \neq B$. The key observation is that if $\Omega$ and $\mathcal{A}_{0}(D)$ are sufficiently small, then the radial derivative of the relative stream function $\Psi$ is strictly positive near $\partial B$, while $\partial D$ is a connected level set of $\Psi$.

To prove the lemma, let us consider the following decomposition of $\Psi$ :

$$
\Psi(x):=1_{D} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}=\underbrace{1_{B} * \mathcal{N}-\frac{\Omega}{2}|x|^{2}}_{=: \Psi^{\text {rad }}(x)}+\underbrace{\left(1_{D}-1_{B}\right) * \mathcal{N}(x)}_{=: \Psi^{e}(x)} .
$$

We claim that there exist positive constants $\Omega_{1}$ and $\alpha_{1}$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{1}$ and $\mathcal{A}_{0}(D)<\alpha_{1}$, then it holds for some $c, C>0$ that

$$
\begin{array}{r}
\partial_{r} \Psi^{r a d}(x)>c \text { for }|x| \in\left(\frac{7}{8}, \frac{9}{8}\right) \\
\left|\Psi^{e}(x)\right|<C \mathcal{A}_{0}(D)\left|\log \mathcal{A}_{0}(D)\right| \text { for } x \in \mathbb{R}^{2} . \tag{6.1.28}
\end{array}
$$

Let us assume for a moment that (Equation 6.1.27) and (Equation 6.1.28) are true. Then we set

$$
\Omega_{2}:=\Omega_{1} \quad \text { and } \quad \alpha_{2}:=\left\{\alpha_{1}, \alpha^{*}\right\}
$$

where $\alpha^{*}=\min \left\{\alpha>0: \frac{C}{c} \alpha \log \frac{1}{\alpha}=\frac{1}{16}\right\}$. If $\Omega<\Omega_{2}$ and $\mathcal{A}_{0}(D)<\alpha_{2}$, then for any $x_{1}$ and $x_{2}$ such that

$$
x_{1} \in \partial D \cap \partial B, \quad \text { and } \quad\left|x_{2}\right|=1-\frac{2 C}{c} \mathcal{A}_{0}(D)\left|\log \mathcal{A}_{0}(D)\right|>\frac{7}{8}
$$

we have

$$
\begin{align*}
\Psi\left(x_{1}\right)-\Psi\left(x_{2}\right) & =\left(\Psi^{r a d}\left(x_{1}\right)-\Psi^{r a d}\left(x_{2}\right)\right)+\left(\Psi^{e}\left(x_{1}\right)-\Psi^{e}\left(x_{2}\right)\right) \\
& >c\left(\left|x_{1}\right|-\left|x_{2}\right|\right)+\left(\Psi^{e}\left(x_{2}\right)-\Psi^{e}\left(x_{1}\right)\right) \\
& >0, \tag{6.1.29}
\end{align*}
$$

where the first and the second inequalities follow from (Equation 6.1.27) and (Equation 6.1.28)
respectively. In the same way, one can easily show that for any $x_{3}$ such that $\left|x_{3}\right|=1+$ $\frac{2 C}{c} \mathcal{A}_{0}(D) \log \frac{1}{\mathcal{A}_{0}(D)}<\frac{9}{8}$, we have $\Psi\left(x_{3}\right)-\Psi\left(x_{1}\right)>0$. Since $\partial D$ is a connected level set of $\Psi$ and $\partial B \cap \partial D \neq \emptyset$, we get

$$
\begin{equation*}
\partial D \subset\left\{x \in \mathbb{R}^{2}: 1-\frac{2 C}{c} \mathcal{A}_{0}(D) \log \frac{1}{\mathcal{A}_{0}(D)}<|x|<1+\frac{2 C}{c} \mathcal{A}_{0}(D) \log \frac{1}{\mathcal{A}_{0}(D)}\right\} \tag{6.1.30}
\end{equation*}
$$

Hence the implicit function theorem with (Equation 6.1.27) and (Equation 6.1.30) yields that there exists $u \in C^{1}(\mathbb{T})$ such that (Equation 6.1.25) holds. Furthermore, (Equation 6.1.30) immediately implies (Equation 6.1.26).

To complete the proof, we need to prove the claims. To prove (Equation 6.1.27), note that $\partial_{r} \Psi^{r a d}(r)$ is explicit and given by

$$
\partial_{r} \Psi^{r}(r)= \begin{cases}\left(\frac{1}{2}-\Omega\right) r & \text { if } r \leq 1 \\ \frac{1}{2 r}-\Omega r & \text { if } r>1\end{cases}
$$

Then (Equation 6.1.27) follows immediately by choosing sufficiently small $\Omega_{1}$. For (Equation 6.1.28), note that Lemma 6.1.4 implies that we can choose $\Omega_{1}$ and $\alpha_{1}$ so that $D \subset B_{3}$. Then we have for any $x \in \mathbb{R}^{2}$ that

$$
\begin{aligned}
\left|\Psi^{e}(x)\right| & =\left|\int_{y \in B_{3}}\left(1_{D}(y)-1_{B}(y)\right) \log \right| x-y|d y| \\
& \lesssim\left|\int_{y \in B_{3},|x-y|<10}\left(1_{D}(y)-1_{B}(y)\right) \log \right| x-y|d y|+\left|\int_{y \in B_{3},|x-y|>10}\left(1_{D}(y)-1_{B}(y)\right) \log \right| x-y|d y| \\
& \lesssim \int_{y \in B_{3},|x-y|<10}\left|1_{D}(y)-1_{B}(y)\right||\log | x-y| | d y+\left|\int_{y \in B_{3},|x-y|>10}\left(1_{D}(y)-1_{B}(y)\right) \frac{\log |x-y|}{\log x} d y\right| \\
& \lesssim|D \triangle B||\log | D \triangle B|\mid \\
& \lesssim \mathcal{A}_{0}(D)\left|\log \mathcal{A}_{0}(D)\right|
\end{aligned}
$$

where we used $\int 1_{D}(y)-1_{B}(y) d x=0$ to get the second inequality. This proves (Equation 6.1.28).

The proof of the following proposition will be postponed to the next subsection.

Proposition 6.1.7. There exist positive constants $\Omega_{3}$ and $\alpha_{3}$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{3}$ and $D$ is a star-shaped domain with $\mathcal{A}_{0}(D) \leq \alpha_{3}$, then $D=B$.

Now we are ready to prove Theorem G.

Proof of Theorem $G$ : We will choose $\Omega_{0}$ and $\alpha_{0}$ so small that all the previous lemmas are applicable. Let us set $\Omega_{0}:=\min \left\{\Omega_{1}, \Omega_{2}, \Omega_{3}, \frac{1}{4}\right\}$ and $\alpha_{0}:=\min \left\{\frac{\alpha_{2}}{c_{1}}, \frac{\alpha_{3}}{c_{1}}\right\}$, where $\alpha_{i}^{\prime} s$ and $c_{1}$ are as in Lemma 6.1.5, 6.1.6 and Proposition 6.1.7. Moreover, let $\sigma$ be as described in Proposition 6.1.3. We assume that $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{0}$ and $D \neq B$. Then we will prove

$$
\begin{equation*}
\sqrt{\frac{\sigma}{2 \pi}} \alpha_{0} \Omega^{-\frac{1}{2}}<\sup _{x \in \partial D}|x| \tag{6.1.31}
\end{equation*}
$$

Since $D \neq B$, we have $\mathcal{A}(D) \geq \alpha_{0}$. Indeed, if $\mathcal{A}(D)<\alpha_{0}$, then Lemma 6.1.5 and Lemma 6.1 .6 imply that $D$ is star-shaped and $\mathcal{A}_{0}(D)<c_{1} \alpha_{0}<\alpha_{3}$. Therefore, Proposition 6.1.7 yields that $D=B$, which is a contradiction. Thus it follows from (Equation 6.1.1) and (Equation 6.1.6) that

$$
\Omega \int_{D}|x|^{2} d x \geq(1-2 \Omega)\left(\frac{|D|^{2}}{4 \pi}-\int_{D} p d x\right)>\frac{1}{2} \sigma \alpha_{0}^{2}
$$

where we used $\Omega<\frac{1}{4}$. It is clear that $\Omega \int_{D}|x|^{2} d x \leq \pi \Omega\left(\sup _{x \in \partial D}|x|\right)^{2}$, hence the above inequality yields (Equation 6.1.31).

### 6.1.1 Proof of Proposition 6.1.7

In this subsection, we aim to prove Proposition 6.1.7. We say a simply-connected bounded domain is star-shaped if there exist $u: \mathbb{T} \mapsto(-1, \infty)$ such that

$$
\partial D=\left\{(1+u(\theta))(\cos \theta, \sin \theta) \in \mathbb{R}^{2}: \theta \in \mathbb{T}\right\}
$$

If $|D|=\pi$, we have that

$$
\pi=\int_{\mathbb{R}^{2}} 1_{D}(x) d x=\int_{\mathbb{T}} \int_{0}^{(1+u(\theta))} r d r d \theta=\pi+\frac{1}{2} \int_{\mathbb{T}} u(\theta)^{2}+2 u(\theta) d \theta
$$

thus

$$
\begin{equation*}
\int_{\mathbb{T}} u(\theta)^{2} d \theta=-\int_{\mathbb{T}} 2 u(\theta) d \theta \tag{6.1.32}
\end{equation*}
$$

Furthermore $\mathcal{A}_{0}(D)$ and the difference of second moments of $1_{D} d x$ and $1_{B} d x$ can be written in terms of $u$ as

$$
\begin{align*}
& \mathcal{A}_{0}(D)=\frac{1}{\pi} \int_{\mathbb{T}}|u(\theta)|+\operatorname{sgn}(u(\theta)) \frac{u(\theta)^{2}}{2} d \theta  \tag{6.1.33}\\
& \int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}=\int_{\mathbb{T}}|u(\theta)|^{2}+u(\theta)^{3}+\frac{1}{4} u(\theta)^{4} d \theta \tag{6.1.34}
\end{align*}
$$

where

$$
\operatorname{sgn}(x)= \begin{cases}1 & \text { if } x \geq 0 \\ -1 & \text { otherwise }\end{cases}
$$

Note that if $\|u\|_{L^{\infty}(\mathbb{T})}<\frac{1}{2}$, then (Equation 6.1.33) and (Equation 6.1.34) imply that there exists $c_{3}>0$ such that

$$
\begin{align*}
& \frac{1}{c_{3}} \int_{\mathbb{T}}|u(\theta)| d \theta \leq \mathcal{A}_{0}(D) \leq c_{3} \int_{\mathbb{T}}|u(\theta)| d \theta,  \tag{6.1.35}\\
& \frac{1}{c_{3}} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta \leq \int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi} \leq c_{3} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta . \tag{6.1.36}
\end{align*}
$$

The proof of Proposition 6.1.7 is based on the identity (Equation 6.1.2). We will estimate the right-hand side of (Equation 6.1.2) in the following proposition.

Proposition 6.1.8. Let $D$ be a star-shaped domain parametrized by $u: \mathbb{T} \mapsto \mathbb{R}$ with $\|u\|_{L^{\infty}}<\frac{1}{2}$. Then there exists $\delta>0$ such that for any $a \in\left(2\|u\|_{L^{\infty}(\mathbb{T})}, 1\right)$, it holds that

$$
\begin{equation*}
\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \delta\left(a \int_{\mathbb{T}}|u|^{2} d \theta+\frac{1}{a} \int_{\mathbb{T}} f(\theta)^{2} d \theta\right) \tag{6.1.37}
\end{equation*}
$$

where $f(\theta):=\int_{0}^{\theta} u(s)^{2}+2 u(s) d s$.

The above proposition will play a key role in the proofs of Proposition 6.1.7 and Theorem Theorem H. In the proof of Proposition 6.1.7, we simply use $|f(\theta)| \lesssim\|u\|_{L^{1}(\mathbb{T})}$, so that the left-hand side can be almost bounded by $L^{1}$-norm of $u$. Note that if we can choose $a$ small enough, then the proposition, together with (Equation 6.1.1) and (Equation 6.1.36) will give $\|u\|_{L^{2}(\mathbb{T})} \lesssim\|u\|_{L^{1}(\mathbb{T})}$.

In section section 6.2, we will use the fact that if $u(\theta)$ is $\frac{2 \pi}{m}$ periodic, then $f(\theta)$ is also $\frac{2 \pi}{m}$ periodic, which follows from (Equation 6.1.32). This will be used for the proof of Theorem Theorem H .

Proof. Using Cauchy-Schwarz inequality, we obtain that

$$
\begin{equation*}
\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \lesssim \int_{D}\left|x-2 \nabla\left(1_{B} * \mathcal{N}\right)\right|^{2} d x+\int_{D}\left|\nabla \mathcal{N} *\left(1_{B}-1_{D}\right)\right|^{2} d x=: H_{1}+H_{2} \tag{6.1.38}
\end{equation*}
$$

To estimate $H_{1}$, note that

$$
\nabla\left(1_{B} * \mathcal{N}\right)= \begin{cases}\frac{x}{2} & \text { if }|x| \leq 1 \\ \frac{x}{2|x|^{2}} & \text { if }|x|>1\end{cases}
$$

Therefore we can compute

$$
\begin{aligned}
\int_{D}\left|x-2 \nabla\left(1_{B} * \mathcal{N}\right)\right|^{2} d x & =\int_{D \backslash B}\left|x-\frac{x}{|x|^{2}}\right|^{2} d x \\
& =\int_{D \backslash B}|x|^{2}-2+\frac{1}{|x|^{2}} d x \\
& =\int_{\mathbb{T} \cap\{u>0\}} \int_{1}^{1+u(\theta)} \cdot\left(r^{2}-2+\frac{1}{r^{2}}\right) r d r d \theta
\end{aligned}
$$

However, we have that for $u(\theta)>0$,

$$
\begin{aligned}
\int_{1}^{1+u(\theta)} r^{3}-2 r+\frac{1}{r} d r & =\frac{1}{4} u(\theta)^{4}+u(\theta)^{3}+\frac{1}{2} u(\theta)^{2}-u(\theta)+\log (1+u(\theta)) \\
& \leq \frac{1}{4} u(\theta)^{4}+\frac{4}{3} u(\theta)^{3} \\
& \lesssim\|u\|_{L^{\infty}(\mathbb{T})}|u(\theta)|^{2} \\
& \lesssim a|u(\theta)|^{2}
\end{aligned}
$$

where we used $\log (1+x) \leq x-\frac{1}{2} x^{2}+\frac{1}{3} x^{3}$ for $x \geq 0$ and $0 \leq u(\theta)<\frac{1}{2}$. Hence it follows that

$$
\begin{equation*}
H_{1} \lesssim a \int_{\mathbb{T}}|u|^{2} d \theta \tag{6.1.39}
\end{equation*}
$$

In order to estimate $H_{2}$, we recall the following result:
Proposition 6.1.9. [79, Proposition 3.1] Let $\rho_{1}$ and $\rho_{2}$ be two probability measures on $\mathbb{R}^{d}$ with $L^{\infty}$ densities with respect to Lebesgue measure. Then

$$
\left\|\nabla\left(\mathcal{N} *\left(\rho_{1}-\rho_{2}\right)\right)\right\|_{L^{2}\left(\mathbb{R}^{d}\right)}^{2} \leq \max \left(\left\|\rho_{1}\right\|_{L^{\infty}},\left\|\rho_{2}\right\|_{L^{\infty}}\right) W_{2}^{2}\left(\rho_{1}, \rho_{2}\right)
$$

where $W_{2}\left(\rho_{1}, \rho_{2}\right)$ denotes 2 -Wasserstein distance between $\rho_{1}$ and $\rho_{2}$ defined by

$$
W_{2}^{2}\left(\rho_{1}, \rho_{2}\right):=\inf \left\{\int|T(x)-x|^{2} d \rho_{1}(x): T_{\#} \rho_{1}=\rho_{2}\right\} .
$$

Thanks to Proposition 6.1.9, it follows that

$$
\begin{equation*}
H_{2}=\left\|\nabla\left(\mathcal{N} *\left(1_{D}-1_{B}\right)\right)\right\|_{L^{2}\left(\mathbb{R}^{2}\right)}^{2} \leq \int_{D}|T(x)-x|^{2} d x \tag{6.1.40}
\end{equation*}
$$

for any $T: D \mapsto B$ such that

$$
\begin{equation*}
T_{\#}\left(1_{D}(x) d x\right)=1_{B}(x) d x \tag{6.1.41}
\end{equation*}
$$

where $T_{\#} \rho$ denotes the pushforward measure of $\rho$ by $T$. Note that in polar coordinates, (Equation 6.1.41) is equivalent to

$$
\begin{equation*}
T_{\#}\left(1_{\tilde{D}}(r, \theta) r d r d \theta\right)=1_{\tilde{B}}(r, \theta) r d r d \theta, \tag{6.1.42}
\end{equation*}
$$

where $\tilde{D}:=\{(r, \theta) \in[0,1) \times \mathbb{T}: 0 \leq r<1+u(\theta)\}$ and $\tilde{B}:=\{(r, \theta) \in[0,1) \times \mathbb{T}: 0 \leq r<1\}$. Hence it suffices to find a transport map $T$ which gives the desired estimate.

Let us define $T: \tilde{D} \mapsto \tilde{B}$ by,

$$
T(r, \theta):=\left(T^{r}(r, \theta), T^{\theta}(\theta)\right):= \begin{cases}\left(\sqrt{\frac{a(2-a)\left(r^{2}-(1+u(\theta))^{2}\right)}{(u(\theta)+a)(u(\theta)+2-a)}+1}, \frac{f(\theta)}{a(2-a)}+\theta\right) & \text { if } r>1-a  \tag{6.1.43}\\ (r, \theta) & \text { if } r \leq 1-a\end{cases}
$$

for $a \in\left(2\|u\|_{L^{\infty}}, 1\right)$, where $f(\theta):=\int_{0}^{\theta} u(\eta)^{2}+2 u(\eta) d \eta$.


Figure 6.1: Illustration of the transport map $T$ that pushes forwards $D$ to $B$.
Our motivation for the transport map $T$ is the following: We first choose $T^{\theta}$ so that $T^{\theta}$ is independent of $r$ and preserves the area in the sense that (see Figure Figure 6.1 for the illustration)

$$
\int_{0}^{\theta} \int_{1-a}^{1+u(s)} r d r d s=\int_{0}^{T^{\theta}(\theta)} \int_{1-a}^{1} r d r d s
$$

And then, we choose $T^{r}(r, \theta)$ so that (Equation 6.1.42) is satisfied. Note that in order to check the condition (Equation 6.1.42) for $T$, it suffices to show that

$$
\begin{equation*}
1_{\tilde{D}}(r, \theta) r=1_{\tilde{B}}(T(r, \theta)) T^{r}(r, \theta)|\operatorname{det}(\nabla T)|, \tag{6.1.44}
\end{equation*}
$$

almost everywhere with respect to the measure $1_{\tilde{D}} r d r d \theta$ (see [104]). Then it is clear that $\theta \mapsto$
$T^{\theta}(\theta)$ and $r \mapsto T^{r}(r, \theta)$ are increasing for fixed $r$ and $\theta$ respectively. Indeed,

$$
\frac{d}{d \theta} T^{\theta}(\theta)=1+\frac{u(\theta)^{2}+2 u(\theta)}{a(2-a)} \geq \frac{2 a-a^{2}+\frac{a^{2}}{4}-a}{a(2-a)} \geq \frac{a-\frac{3}{4} a^{2}}{a(2-a)}>0
$$

where the first inequality follows from that $\|u\|_{L^{\infty}(\mathbb{T})}<\frac{1}{2} a$ and $x \mapsto x^{2}+2 x$ is increasing for $x \geq-1$ thus $u(\theta)^{2}+2 u(\theta) \geq \frac{a^{2}}{4}-a$. Since $T$ maps $\{(r, \theta): r=1-a$ or $1+u(\theta)\}$ to $\{(r, \theta): r=1-a$ or $r=1\}$ continuously, $T$ is bijective and therefore $1_{\tilde{D}}(r, \theta)=1_{\tilde{B}} \circ T(r, \theta)$. Furthermore, the Jacobian matrix of $T$ can be computed as

$$
\nabla T(r, \theta)=\left\{\begin{array}{lc}
\left(\begin{array}{cc}
\frac{1}{T^{r}(r, \theta)} \frac{a(2-a) r}{(u(\theta)+a)(u(\theta)+2-a)} & \partial_{\theta} T^{r}(r, \theta) \\
0 & \frac{(u(\theta)+a)(u(\theta)+2-a)}{a(2-a)}
\end{array}\right) & \text { if } 1-a<r<1+u(\theta), \\
\left(\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right) & \text { otherwise },
\end{array}\right.
$$

therefore

$$
T^{r}(r, \theta)|\operatorname{det}(\nabla T)|=r,
$$

almost everywhere. This implies that $T$ satisfies (Equation 6.1.44) and thus (Equation 6.1.42) holds. Then it follows from (Equation 6.1.40) that

$$
H_{2} \leq \int_{\mathbb{T}} \int_{1-a}^{1+u(\theta)}\left|T^{r}(r, \theta) \cos \left(T^{\theta}\right)-r \cos \theta\right|^{2}+\left|T^{r}(r, \theta) \sin \left(T^{\theta}\right)-r \sin \theta\right|^{2} r d r d \theta
$$

The cosine term in the integrand can be estimated as

$$
\begin{aligned}
\left|T^{r}(r, \theta) \cos \left(T^{\theta}\right)-r \cos \theta\right|^{2} & =\left|\left(T^{r}(r, \theta)-r\right) \cos \left(T^{\theta}(\theta)\right)+r\left(\cos \left(T^{\theta}(\theta)\right)-\cos \theta\right)\right|^{2} \\
& \lesssim\left|T^{r}(r, \theta)-r\right|^{2}+\left|\cos \left(T^{\theta}(\theta)\right)-\cos \theta\right|^{2} \\
& \lesssim\left|T^{r}(r, \theta)-r\right|^{2}+\left|T^{\theta}(\theta)-\theta\right|^{2}
\end{aligned}
$$

In the same way, the sine term can be bounded as $\left|T^{r} \sin \left(T^{\theta}\right)-r \sin \theta\right|^{2} \lesssim\left|T^{r}-r\right|^{2}+\left|T^{\theta}-\theta\right|^{2}$, thus we have

$$
\begin{equation*}
H_{2} \lesssim \int_{\mathbb{T}} \int_{1-a}^{1+u(\theta)}\left|T^{r}(r, \theta)-r\right|^{2} r d r d \theta+\int_{\mathbb{T}} \int_{1-a}^{1+u(\theta)}\left|T^{\theta}(\theta)-\theta\right|^{2} r d r d \theta=: A_{1}+A_{2} \tag{6.1.45}
\end{equation*}
$$

$A_{2}$ is bounded by

$$
\begin{equation*}
A_{2} \leq \int_{\mathbb{T}} \int_{1-a}^{1+u(\theta)} \frac{f(\theta)^{2}}{a^{2}} r d r d \theta \lesssim \int_{\mathbb{T}} f(\theta)^{2} \frac{|u(\theta)|+a}{a^{2}} \lesssim \frac{1}{a} \int_{\mathbb{T}} f(\theta)^{2} d \theta \tag{6.1.46}
\end{equation*}
$$

where we used $\|u\|_{L^{\infty}(\mathbb{T})}<a$ to get the first and the last inequalities.
For $A_{1}$, we assume for a momoent that for $r \in(1-a, 1+u(\theta))$,

$$
\begin{equation*}
\left|T^{r}(r, \theta)-r\right| \lesssim|u(\theta)| . \tag{6.1.47}
\end{equation*}
$$

From (Equation 6.1.47), we obtain

$$
\begin{equation*}
A_{1} \lesssim \int_{\mathbb{T}} \int_{1-a}^{1+u(\theta)}|u(\theta)|^{2} r d r d \theta=\int_{\mathbb{T}}|u(\theta)|^{2} \int_{1-a}^{1+u(\theta)} r d r d \theta \lesssim a \int_{\mathbb{T}}|u(\theta)|^{2} d \theta \tag{6.1.48}
\end{equation*}
$$

where the last inequality follows from $a>\|u\|_{L^{\infty}(\mathbb{T})}$. Therefore, it follows from (Equation 6.1.45),
(Equation 6.1.46) and (Equation 6.1.48) that

$$
\begin{equation*}
H_{2} \lesssim a \int_{\mathbb{T}}|u|^{2} d \theta+\frac{1}{a} \int_{\mathbb{T}}|f|^{2} d \theta \tag{6.1.49}
\end{equation*}
$$

Thus (Equation 6.1.37) follows from (Equation 6.1.38), (Equation 6.1.39) and (Equation 6.1.49).
To check (Equation 6.1.47), let $g(a, r, x) \quad:=\frac{\sqrt{\frac{a(2-a)\left(r^{2}-(1+x)^{2}\right)}{(x+a)(x+2-a)}+1}}{x}$ so that $\frac{T^{r}(r, \theta)-r}{u(\theta)}=g(a, r, u(\theta))$. Then it suffices to show that $|g(a, r, x)| \lesssim 1$ in $\{(a, r, x):(1-a)<r<1+x, 2|x|<a<1\}$. Since $g(a, r, x)$ is continuous everywhere except for $x=0$, we only need to check $|g(a, r, x)| \lesssim 1$ when $0<x \ll 1$. Taking the limit, we obtain

$$
\lim _{x \rightarrow 0^{+}} g(a, r, x)=\frac{\left.\frac{\partial}{\partial x}\left(\sqrt{\frac{a(2-a)\left(r^{2}-(1+x)^{2}\right)}{(x+a)(x+2-a)}+1}-r\right)\right|_{x=0}}{1}=\frac{\left(1-r^{2}\right)-a(2-a)}{r a(2-a)} .
$$

If $r<\frac{1}{2}$, then $a>\frac{1}{2}$ therefore it follows from $r>1-a>0$ that

$$
\lim _{x \rightarrow 0}|g(a, r, x)|=\frac{\left|r^{2}-(a-1)^{2}\right|}{r a(2-a)} \leq \frac{r}{a(2-a)}+\frac{(a-1)^{2}}{r a(2-a)} \leq \frac{2 r}{a(2-a)} \lesssim 1
$$

where the second inequality follows from $(1-a)<r$. If $r>\frac{1}{2}$, then it follows from $|r-1|<a$ that

$$
\lim _{x \rightarrow 0}|g(a, r, x)| \leq \frac{\left|1-r^{2}\right|}{r a(2-a)}+\frac{(2-a)}{r(2-a)}<\frac{1+r}{r(2-a)}+\frac{2-a}{r(2-a)} \lesssim 1 .
$$

This proves (Equation 6.1.47) and finishes the proof.

Now we are ready to prove Proposition 6.1.7.

Proof of Proposition 6.1.7: We will fix $\Omega_{3}$ and $\alpha_{3}$ so small that all the lemmas are applicable. To do so, let us denote $h(x):=-x \log x$. Also we denote by $\alpha^{*}>0$ the smallest positive number
such that

$$
\begin{equation*}
\left(h\left(\alpha^{*}\right)\right)^{3}>\left(\alpha^{*}\right)^{2} . \tag{6.1.50}
\end{equation*}
$$

Furthermore, let $\Omega_{i}^{\prime} s, \alpha_{i}^{\prime} s$ and $c_{i}^{\prime} s$ for $i=1,2$ be as in Lemma 6.1.5 and Lemma 6.1.6, let $c_{3}$ be as in (Equation 6.1.35) and (Equation 6.1.36) and let $\delta$ be as in Proposition 6.1.8. Lastly, let $c_{4}:=18 \pi c_{3}^{3} \delta^{2}$. Then let us fix

$$
\begin{equation*}
\Omega_{3}:=\min \left\{\Omega_{1}, \Omega_{2}, \frac{1}{4}, \frac{\sigma}{16 c_{1}^{2} c_{3}^{2} c_{4}}\right\} \quad \text { and } \quad \alpha_{3}:=\left\{\alpha^{*}, \alpha_{1}, \alpha_{2},\left(\frac{1}{2 c_{2}}\right)^{\frac{3}{2}},\left(\frac{1}{4 c_{2} c_{3} \delta}\right)^{\frac{3}{2}}\right\} . \tag{6.1.51}
\end{equation*}
$$

Then our goal is to show that if $(D, \Omega)$ is a solution to (Equation 1.2.4) with $\Omega<\Omega_{3}$ and $\mathcal{A}_{0}(D)<$ $\alpha_{3}$, then $D=B$.

Step 1. Let us claim that

$$
\begin{align*}
& \mathcal{A}_{0}(D) \leq c_{1} \mathcal{A}(D) .  \tag{6.1.52}\\
& \|u\|_{L^{\infty}(\mathbb{T})} \leq c_{2} \mathcal{A}_{0}(D)^{\frac{2}{3}} \leq \frac{1}{2} . \tag{6.1.53}
\end{align*}
$$

Since $\Omega_{3}<\Omega_{1}$ and $\mathcal{A}(D) \leq \mathcal{A}_{0}(D)<\alpha_{3} \leq \alpha_{1}$, it follows from Lemma 6.1.5 that $\mathcal{A}_{0}(D)<$ $c_{1} \mathcal{A}(D)$. In addition, $\Omega_{3}<\Omega_{2}, \mathcal{A}_{0}(D)<\alpha_{3} \leq \alpha_{2}$ and Lemma 6.1.6 imply that

$$
\|u\|_{L^{\infty}(\mathbb{T})} \leq c_{2} h\left(\mathcal{A}_{0}(D)\right) \leq c_{2} \mathcal{A}_{0}(D)^{\frac{2}{3}},
$$

where the last inequality follows from $\alpha_{3} \leq \alpha^{*}$. Since $\mathcal{A}_{0}(D)<\alpha_{3} \leq\left(\frac{1}{2 c_{2}}\right)^{\frac{3}{2}}$, we have $c_{2} \mathcal{A}_{0}(D)^{\frac{2}{3}} \leq \frac{1}{2}$, which proves (Equation 6.1.53).

Step 2. In this step, we will show that

$$
\begin{equation*}
\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \frac{1}{4 c_{3}} \int_{\mathbb{T}}|u|^{2} d \theta+c_{4} \mathcal{A}_{0}(D)^{2} \tag{6.1.54}
\end{equation*}
$$

where $c_{4}:=18 \pi c_{3}^{3} \delta^{2}$. Since $\|u\|_{L^{\infty}((T))}<\frac{1}{2}$, we will apply Proposition 6.1 .8 with $a:=\frac{1}{2 c_{3} \delta}$. Note that

$$
2\|u\|_{L^{\infty}(\mathbb{T})} \leq 2 c_{2} \mathcal{A}_{0}(D)^{\frac{2}{3}}<2 c_{2} \alpha_{3}^{\frac{2}{3}} \leq a
$$

where the first inequality follows from (Equation 6.1.53), the second follows from the assumption that $\mathcal{A}_{0}(D)<\alpha_{3}$ and the last inequality follows from (Equation 6.1.51), which says $\alpha_{3} \leq$ $\left(\frac{1}{4 c_{2} c_{3} \delta}\right)^{\frac{3}{2}}$. Thus we can obtain by using Proposition 6.1 .8 that

$$
\begin{equation*}
\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \frac{1}{4 c_{3}} \int_{\mathbb{T}}|u|^{2} d \theta+c_{3} \delta^{2} \int_{\mathbb{T}} f(\theta)^{2} d \theta \tag{6.1.55}
\end{equation*}
$$

where $f(\theta)=\int_{0}^{\theta} u(s)^{2}+2 u(s) d s$. Moreover, we have

$$
\begin{equation*}
|f(\theta)| \leq \int_{0}^{\theta} 3|u(s)| d s<\int_{\mathbb{T}} 3|u(s)| d s \leq 3 c_{3} \mathcal{A}_{0}(D) \tag{6.1.56}
\end{equation*}
$$

where the last inequality follows from (Equation 6.1.35). Therefore it follows from (Equation 6.1.55) and (Equation 6.1.56) that

$$
\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \frac{1}{4 c_{3}} \int_{\mathbb{T}}|u|^{2} d \theta+18 \pi c_{3}^{3} \delta^{2} \mathcal{A}_{0}(D)^{2}
$$

which proves the claim (Equation 6.1.54).

Step 3. Now we will prove that

$$
\begin{equation*}
\int_{\mathbb{T}}|u(\theta)|^{2} d \theta \leq 4 c_{3} c_{4} \mathcal{A}_{0}(D)^{2} \tag{6.1.57}
\end{equation*}
$$

Since $\Omega<\Omega_{3} \leq \frac{1}{4}$, it follows from (Equation 6.1.36) that

$$
\begin{equation*}
(1-2 \Omega)\left(\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}\right)>\frac{1}{2 c_{3}} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta \tag{6.1.58}
\end{equation*}
$$

Thus it follows from (Equation 6.1.2) and (Equation 6.1.54) that

$$
\frac{1}{2 c_{3}} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta<\frac{1}{4 c_{3}} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta+c_{4} \mathcal{A}_{0}(D)^{2}
$$

which proves (Equation 6.1.57).
Step 4. Finally, we will prove $D=B$ by showing that $\mathcal{A}_{0}(D)=0$. This will be done by estimating the left/right-hand side in (Equation 6.1.1). It follows from Proposition 6.1.3 and (Equation 6.1.52) that

$$
\begin{equation*}
(1-2 \Omega)\left(\frac{|D|^{2}}{4 \pi}-\int_{D} p d x\right) \geq \frac{1}{2} \sigma \mathcal{A}(D)^{2} \geq \frac{1}{2 c_{1}^{2}} \sigma \mathcal{A}_{0}(D)^{2}, \tag{6.1.59}
\end{equation*}
$$

where we used $\Omega<\frac{1}{4}$. Moreover, it follows from (Equation 6.1.36) and (Equation 6.1.57) that

$$
\begin{equation*}
2 \Omega\left(\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}\right) \leq 2 \Omega c_{3} \int_{\mathbb{T}}|u(\theta)|^{2} d \theta \leq 8 \Omega c_{3}^{2} c_{4} \mathcal{A}_{0}(D)^{2} \tag{6.1.60}
\end{equation*}
$$

Therefore (Equation 6.1.1) yields that

$$
\left(8 \Omega c_{3}^{2} c_{4}-\frac{\sigma}{2 c_{1}^{2}}\right) \mathcal{A}_{0}(D)^{2} \geq 0
$$

This implies $\mathcal{A}_{0}(D)=0$, since $8 \Omega c_{3}^{2} c_{4}-\frac{\sigma}{2 c_{1}^{2}}<8 \Omega_{3} c_{3}^{2} c_{4}-\frac{\sigma}{2 c_{1}^{2}} \leq 0$, which follows from
(Equation 6.1.51) and $\Omega<\Omega_{3}$. This proves that $D=B$.

### 6.2 Rotating patches with $m$-fold symmetry

We now move on to the quantitative estimates for $m$-fold symmetric rotating patches. We say a domain $D$ is $m$-fold symmetric, if $D$ is invariant under rotation by $\frac{2 \pi}{m}$. We divide this section into two subsections: The first subsection is devoted to the proof of Theorem H and the second subsection is devoted to the proof of Theorem I.

### 6.2.1 Proof of Theorem H

The goal of this subsection is to prove Theorem H. As explained in Remark 1.2.1, angular velocity $\Omega$ is independent of radial dilation, thus we will assume that $|D|=|B|=\pi$ throughout this subsection.

For a simply-connected and $m$-fold symmetric patch $D$, we denote $r_{\text {min }}:=\inf _{x \in \partial D}|x|$, and $r_{\text {max }}:=\sup _{x \in \partial D}|x|$. Note that the origin is necessarily contained in $D$ since $D$ is simplyconnected and $m$-fold symmetric, therefore $r_{\min }>0$. Furthermore, since we are assuming $|D|=\pi$, it is necessarily $r_{\text {min }}<1$ and $r_{\max }>1$ if $D$ is not a disk.

We will prove the theorem by contrapositive. We suppose to the contrary that $(D, \Omega)$ is an $m$-fold symmetric solution with sufficiently large $m$ and $\lambda:=\frac{1}{2}-\Omega$ is sufficient large compared to $\frac{1}{m}$. Then Lemma 6.2 .2 tells us that the patch is necessarily star-shaped and the polar graph that parametrizes $\partial D$ must be small. With this fact, we will apply the identity (Equation 6.1.2) and Proposition 6.1 .8 to derive an upper bound of $\lambda$, which we expect to contradict our initial assumption on $\lambda$.

Now we introduce a decomposition of the stream function $1_{D} * \mathcal{N}$. We define a radial function
$g: \mathbb{R}^{2} \mapsto \mathbb{R}$ as follows (where we denote it by $g(r)$ by slight abuse of nontation):

$$
g(r):=\frac{1}{2 \pi r} \mathcal{H}^{1}\left(\partial B_{r} \cap D\right),
$$

where $\mathcal{H}^{1}$ denotes the 1-dimensional Hausdorff measure. Then we shall write, in polar coordinates,

$$
\begin{equation*}
\left(1_{D} * \mathcal{N}\right)(r, \theta)=(g * \mathcal{N})(r)+\left(1_{D}-g\right) * \mathcal{N}(r, \theta)=: \varphi^{r}(r)+\varphi_{m}(r, \theta) . \tag{6.2.1}
\end{equation*}
$$

Therefore the relative stream function can be written as $\Psi(r, \theta)=\varphi^{r}(r)-\frac{\Omega}{2} r^{2}+\varphi_{m}(r, \theta)$.
Note that $g$ is a radial function with the same integral as $1_{D}$ on each $\partial B_{r}$. If $D$ is $m$-fold symmetric for large $m$, we would expect that the velocity field generated by the vorticity $1_{D}$ must be very close to the velocity field generated by $g$, that is, we expect that $\left|\nabla \varphi_{m}\right| \ll 1$ if $m \gg 1$. Below we will give a quantitative proof of this fact in Lemma 6.2.1.

Lemma 6.2.1. Let $D$ be an $m$-fold symmetric bounded domain for $m \geq 3$. Then

$$
\begin{gather*}
\partial_{r} \varphi^{r}(r)=\frac{\left|D \cap B_{r}\right|}{2 \pi r}  \tag{6.2.2}\\
\left|\nabla \varphi_{m}(r, \theta)\right| \lesssim \frac{r}{m} \tag{6.2.3}
\end{gather*}
$$

Proof. Let us prove (Equation 6.2.2) first. Obviously, (Equation 6.2.2) is equivalent to

$$
\begin{equation*}
2 \pi r \partial_{r} \varphi^{r}(r)=\left|D \cap B_{r}\right| \tag{6.2.4}
\end{equation*}
$$

Clearly both sides of (Equation 6.2.4) are zero at $r=0$. Also we have that

$$
\partial_{r}\left(\left|D \cap B_{r}\right|\right)=\mathcal{H}^{1}\left(D \cap \partial B_{r}\right)=2 \pi r g(r)=2 \pi r \Delta\left(\varphi^{r}(r)\right)=\partial_{r}\left(2 \pi r \partial_{r} \varphi^{r}(r)\right),
$$

where we used $\Delta=\frac{1}{r} \partial_{r}\left(r \partial_{r}\right)+\frac{1}{r^{2}} \partial_{\theta \theta}$. This proves (Equation 6.2.4), thus (Equation 6.2.2).

We will prove (Equation 6.2.3) by using the formula for the stream function given in Lemma B.1.2. Let $h(r, \theta):=1_{D}(r \cos \theta, r \sin \theta)-g(r)$. We apply (Equation B.2.3) and (Equation B.2.4) to (Equation B.1.2) and (Equation B.1.3) respectively, and obtain

$$
\begin{align*}
\partial_{r} \varphi_{m}(r, \theta) & =\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\frac{\left(\frac{\rho}{r}\right)^{m+1}\left(\cos (m \eta)-\left(\frac{\rho}{r}\right)^{m}\right)}{\left(1-\left(\frac{\rho}{r}\right)^{m}\right)^{2}+2\left(\frac{\rho}{r}\right)^{m}(1-\cos (m \eta))}\right) d \rho d \eta \\
& -\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\frac{\left(\frac{r}{\rho}\right)^{m-1}\left(\cos (m \eta)-\left(\frac{r}{\rho}\right)^{m}\right)}{\left(1-\left(\frac{r}{\rho}\right)^{m}\right)^{2}+2\left(\frac{r}{\rho}\right)^{m}(1-\cos (m \eta))}\right) d \rho d \eta \\
& =: A_{1}-A_{2}  \tag{6.2.5}\\
\partial_{\theta} \varphi_{m}(r, \theta) & =-r \frac{1}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\frac{\left(\frac{\rho}{r}\right)^{m+1} \sin (m \eta)}{\left(1-\left(\frac{\rho}{r}\right)^{m}\right)^{2}+2\left(\frac{\rho}{r}\right)^{m}(1-\cos (m \eta))}\right) d \rho d \eta \\
& -r \frac{1}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\frac{\left(\frac{r}{\rho}\right)^{m-1} \sin (m \eta)}{\left(1-\left(\frac{r}{\rho}\right)^{m}\right)^{2}+2\left(\frac{r}{\rho}\right)^{m}(1-\cos (m \eta))}\right) d \rho d \eta \\
& =:-r A_{3}-r A_{4} \tag{6.2.6}
\end{align*}
$$

We claim that

$$
\begin{equation*}
\left|A_{i}\right| \lesssim \frac{r}{m} \text { for } i=1,2,3 \text { and } 4 \tag{6.2.7}
\end{equation*}
$$

Let us assume for a moment that the claim is true. Then (Equation 6.2.5) and (Equation 6.2.6) yield that $\left|\nabla \varphi_{m}(r, \theta)\right| \sim\left|\partial_{r} \varphi_{m}(r, \theta)\right|+\left|\frac{\partial_{\theta} \varphi_{m}(r, \theta)}{r}\right| \lesssim \frac{r}{m}$, which finishes the proof. We give a proof of (Equation 6.2.7) for only $A_{2}$ since the other terms can be proved in the same way. Note that in the proof, we will see that the assumption $m \geq 3$ is crucial to estimate $A_{2}$ and $A_{4}$.

From the change of the variables, $\left(\frac{r}{\rho}\right)^{m} \mapsto x$ and $\frac{2 \pi}{m}$-periodicity of the integrand in the angular
variable, it follows that

$$
\begin{aligned}
\left|A_{2}\right| & \leq \int_{\mathbb{T}} \int_{0}^{1}\left|h\left(r x^{-\frac{1}{m}}, \eta+\theta\right)\left(\frac{x^{1-\frac{1}{m}}(\cos (m \eta)-x)}{(1-x)^{2}+2 x(1-\cos (m \eta))}\right) \frac{r}{m} x^{-1-\frac{1}{m}}\right| d x d \eta \\
& \leq \frac{r}{m} \int_{\mathbb{T}} \int_{0}^{1}\left|\frac{x^{-\frac{2}{m}}(\cos (m \eta)-x)}{(1-x)^{2}+2 x(1-\cos (m \eta))}\right| d x d \eta \\
& \leq r \int_{0}^{\frac{2 \pi}{m}} \int_{0}^{1} \frac{x^{-\frac{2}{m}}((1-x)+(1-\cos (m \eta)))}{(1-x)^{2}+2 x(1-\cos (m \eta))} d x d \eta \\
& =\frac{r}{m} \int_{\mathbb{T}} \int_{0}^{1} \frac{x^{-\frac{2}{m}}((1-x)+(1-\cos \eta))}{(1-x)^{2}+2 x(1-\cos \eta)} d x d \eta \\
& =\frac{2 r}{m} \int_{0}^{\pi} \int_{0}^{1} \frac{x^{-\frac{2}{m}}((1-x)+(1-\cos \eta))}{(1-x)^{2}+2 x(1-\cos \eta)} d x d \eta \\
& =\frac{2 r}{m}\left(\int_{0}^{\pi} \int_{0}^{\frac{1}{2}} \frac{x^{-\frac{2}{m}}((1-x)+(1-\cos \eta))}{(1-x)^{2}+2 x(1-\cos \eta)} d x d \eta+\int_{0}^{\pi} \int_{\frac{1}{2}}^{1} \frac{x^{-\frac{2}{m}}((1-x)+(1-\cos \eta))}{(1-x)^{2}+2 x(1-\cos \eta)} d x d \eta\right) \\
& =\frac{2 r}{m}\left(A_{21}+A_{22}\right)
\end{aligned}
$$

where we used $\frac{2 \pi}{m}$-periodicity of the integrand to get the third inequality, the change of variables, $\eta \mapsto \frac{1}{m} \eta$ to get the first equality, and the evenness of the integrand in $\eta$ to get the second equality. Note that the denominator of the integrand $A_{21}$ is bounded from below by a strictly positive number, therefore

$$
A_{21} \lesssim \int_{0}^{\pi} \int_{0}^{\frac{1}{2}} x^{-\frac{2}{m}} d x d \eta \lesssim \frac{m}{m-2} \lesssim 1
$$

for $m \geq 3$. For $A_{22}$, we use that $(1-\cos \eta) \sim \eta^{2}$ for $\eta \in(0, \pi)$ and the change of variables, $x \mapsto 1-x$, to obtain

$$
\begin{aligned}
A_{22} \lesssim \int_{0}^{\pi} \int_{0}^{\frac{1}{2}} \frac{x+\eta^{2}}{x^{2}+\eta^{2}} d x d \eta & =\int_{0}^{\pi} \int_{0}^{\frac{1}{2}} 1_{\{x<\eta\}} \frac{x+\eta^{2}}{x^{2}+\eta^{2}} d x d \eta+\int_{0}^{\pi} \int_{0}^{\frac{1}{2}} 1_{\{x \geq \eta\}} \frac{x+\eta^{2}}{x^{2}+\eta^{2}} d x d \eta \\
& \leq \int_{0}^{\pi} \int_{0}^{\eta} \frac{\eta+\eta^{2}}{\eta^{2}} d x d \eta+\int_{0}^{\frac{1}{2}} \int_{0}^{x} \frac{x+x^{2}}{x^{2}} d \eta d x \\
& \lesssim 1
\end{aligned}
$$

This proves $\left|A_{2}\right| \lesssim \frac{r}{m}$. As mentioned, the same argument applies to $A_{1}, A_{3}$ and $A_{4}$ to prove (Equation 6.2.7). This completes the proof.

From (Equation 6.2.2) and (Equation 6.2.3) in the above lemma, it is clear that $\partial_{r} \Psi\left(r_{\text {min }}, \theta\right)=$ $r_{\min }\left(\frac{1}{2}-\Omega-\frac{\partial_{r} \varphi_{m}\left(r_{\min }, \theta\right)}{r_{\min }}\right)$ and $\left|\frac{\partial_{r} \varphi_{m}\left(r_{\min }, \theta\right)}{r_{\text {min }}}\right| \sim \frac{1}{m}$. Thus one can expect that if $\frac{1}{2}-\Omega$ is sufficiently large compared to $\frac{1}{m}$, then the level set $\partial D$ cannot be too far from the a circle. We give a detailed proof for this in the following lemma.

Lemma 6.2.2. Assume that $(D, \Omega)$ is a solution to (Equation 1.2.4). Then there exist constants $c_{1}, c_{2}>0$ and $m_{1} \geq 3$ such that if $D$ is $m$-fold symmetric for some $m \geq m_{1}$ and $\lambda=\frac{1}{2}-\Omega>\frac{c_{1}}{m}$, then $D$ is star-shaped and $\left|r_{\max }-r_{\min }\right|<\frac{c_{2}}{m}$. Hence there exist $u \in C^{1}(\mathbb{T})$ such that

$$
\partial D=\{(1+u(\theta))(\cos \theta, \sin \theta): \theta \in \mathbb{T}\} \quad \text { and } \quad\|u\|_{L^{\infty}(\mathbb{T})}<\frac{c_{2}}{m} .
$$

Proof. Thanks to (Equation 6.2.3) in Lemma 6.2.1, we can find a constant $C>0$ (which we can also assume to be larger then 1) such that

$$
\begin{equation*}
\left|\nabla_{x} \varphi_{m}(r, \theta)\right|<C \frac{r}{m} \tag{6.2.8}
\end{equation*}
$$

where $\nabla_{x}$ denotes the gradient in Cartesian coordinates, that is, $\nabla_{x}:=\partial_{r}+\frac{1}{r} \partial_{\theta}$. We will first prove the bound for $r_{\max }-r_{\text {min }}$ and show star-shapeness of $\partial D$ afterwards. Let

$$
\begin{equation*}
c_{1}:=\max \{6 C, \sqrt{48 \pi C}\}+1, \quad c_{2}:=\frac{c_{1}}{4} \quad \text { and } \quad m_{1}:=\max \left\{\frac{3 C}{2}, \frac{c_{1}}{4}, 3\right\}+1 \tag{6.2.9}
\end{equation*}
$$

We will show that if $\lambda>\frac{c_{1}}{m}$ and $m \geq m_{1}$, then $r_{\text {max }}-r_{\text {min }}<\frac{c_{2}}{m}$.
Let $q(r):=\frac{\left|D \cap B_{r}\right|}{2 \pi r^{2}}-\Omega$. Since $\frac{1}{r^{2}}>\frac{1}{r_{\text {min }}^{2}}-\frac{2}{r_{\text {min }}^{3}}\left(r-r_{\text {min }}\right)$ for $r>r_{\text {min }}$, and $\left|D \cap B_{r}\right|$ is
increasing in $r$, we have that

$$
q(r)>\frac{1}{2} \lambda, \text { for } r \in\left(r_{\min }, r_{\min }\left(1+\frac{1}{4} \lambda\right)\right)
$$

which implies that

$$
\begin{equation*}
\partial_{r}\left(\varphi^{r}(r)-\frac{\Omega}{2} r^{2}\right)=r q(r)>\frac{r_{\min }}{2} \lambda, \text { for } r \in\left(r_{\min }, r_{\min }\left(1+\frac{1}{4} \lambda\right)\right), \tag{6.2.10}
\end{equation*}
$$

where the equality follows from (Equation 6.2.2) in Lemma 6.2.1. Let $\varepsilon:=\frac{r_{\text {min }} c_{1}}{4 m}$. By the assumption $\lambda>\frac{c_{1}}{m}$, we have

$$
\begin{equation*}
\varepsilon<\frac{r_{\min }}{4} \lambda . \tag{6.2.11}
\end{equation*}
$$

We choose $x_{1}, x_{2} \in \mathbb{R}^{2}$ such that for some $\theta_{1}, \theta_{2} \in \mathbb{T}$,

$$
x_{1}=r_{\min }\left(\cos \left(\theta_{1}\right), \sin \left(\theta_{1}\right)\right), \quad x_{2}=\left(r_{\min }+\varepsilon\right)\left(\cos \left(\theta_{2}\right), \sin \left(\theta_{2}\right)\right) \quad \text { and } \quad\left|\theta_{1}-\theta_{2}\right| \leq \frac{2 \pi}{m}
$$

We claim that

$$
\begin{equation*}
\Psi\left(x_{2}\right)-\Psi\left(x_{1}\right)>0 \tag{6.2.12}
\end{equation*}
$$

Let us assume that the claim is true for a moment. Then from $m$-fold symmetry of $D$ and the fact that $\partial D$ is a level set of $\Psi$, it follows that $r_{\max } \leq r_{\min }+\varepsilon$. Thus it follows from (Equation 6.2.9), (Equation 6.2.11) and $r_{\text {min }}<1$ that

$$
\begin{equation*}
r_{\max }-r_{\min } \leq \varepsilon=\frac{c_{2} r_{\min }}{m}<\frac{c_{2}}{m} \tag{6.2.13}
\end{equation*}
$$

Furthermore, for all $x \in \partial D$, it follows from (Equation 6.2.8) and (Equation 6.2.10) that $\partial_{r} \Psi(x)>$ $\frac{r_{\text {min }}}{2} \lambda-\frac{C r_{\text {max }}}{m}$. Hence
$\partial_{r} \Psi(x)>\frac{r_{\min } c_{1}}{2 m}-\frac{C r_{\max }}{m} \geq \frac{c_{1}}{2 m}\left(r_{\max }-\frac{c_{2}}{m}\right)-\frac{C r_{\max }}{m}=\left(\frac{c_{1} r_{\max }}{4 m}-\frac{C r_{\max }}{m}\right)+\frac{c_{1}}{2 m}\left(\frac{r_{\max }}{2}-\frac{c_{2}}{m}\right)>0$,
where the first inequality follows from $\lambda>\frac{c_{1}}{m}$, the second inequality follows from (Equation 6.2.13) and the last inequality follows from (Equation 6.2.9) and $r_{\max } \geq 1$, which say $\frac{c_{1}}{4}>C$ and $\frac{r_{\max }}{2}>\frac{1}{2}>\frac{c_{2}}{m}$. Therefore the implicit function theorem yields that there exists $u \in C^{1}(\mathbb{T})$ such that $\partial D=\{(1+u(\theta))(\cos \theta, \sin \theta): \theta \in \mathbb{T}\}$. This proved star-shapeness of $D$ and the desired $L^{\infty}$-norm bound for $u$.

Now it suffices to prove (Equation 6.2.12). We compute

$$
\Psi\left(x_{2}\right)-\Psi\left(x_{1}\right)=\underbrace{\left(\varphi^{r}\left(\left|x_{2}\right|\right)-\frac{\Omega}{2}\left|x_{2}\right|^{2}\right)-\left(\varphi^{r}\left(\left|x_{1}\right|\right)-\frac{\Omega}{2}\left|x_{1}\right|^{2}\right)}_{=: L_{1}}+\underbrace{\varphi_{m}\left(x_{2}\right)-\varphi_{m}\left(x_{1}\right)}_{L_{2}} .
$$

Thanks to (Equation 6.2.10), we have

$$
\begin{equation*}
L_{1}>\frac{r_{\min }}{2} \lambda\left(\left|x_{2}\right|-\left|x_{1}\right|\right)=\frac{r_{\min } \lambda \varepsilon}{2} . \tag{6.2.14}
\end{equation*}
$$

To estimate $L_{2}$, let us pick $x_{1}^{\prime}=r_{\min }\left(\cos \left(\theta_{2}\right), \sin \left(\theta_{2}\right)\right)$. Then it follows from (Equation 6.2.8) that

$$
\begin{align*}
L_{2} & =\left(\varphi_{m}\left(x_{2}\right)-\varphi_{m}\left(x_{1}^{\prime}\right)\right)+\left(\varphi_{m}\left(x_{1}^{\prime}\right)-\varphi_{m}\left(x_{1}\right)\right) \\
& >-C \frac{\left|x_{2}\right|}{m}\left(\left|x_{2}\right|-\left|x_{1}\right|\right)-C \frac{r_{\min }^{2}}{m} \frac{2 \pi}{m} \\
& =-\frac{C r_{\min } \varepsilon}{m}-\frac{C \varepsilon^{2}}{m}-\frac{2 \pi C r_{\min }^{2}}{m^{2}} . \tag{6.2.15}
\end{align*}
$$

Hence it follows from (Equation 6.2.14) and (Equation 6.2.15) that (we split $\frac{r_{\text {min }} \lambda}{2}$ into three
pieces evenly)
$\Psi\left(x_{2}\right)-\Psi\left(x_{1}\right)>\left(\frac{r_{\min } \lambda \varepsilon}{6}-\frac{C r_{\min } \varepsilon}{m}\right)+\left(\frac{r_{\min } \lambda \varepsilon}{6}-\frac{C \varepsilon^{2}}{m}\right)+\left(\frac{r_{\min } \lambda \varepsilon}{6}-\frac{2 \pi C r_{\min }^{2}}{m^{2}}\right)=: L_{3}+L_{4}+L_{5}$.

From $\lambda>\frac{c_{1}}{m}$ and (Equation 6.2.9), which says $c_{1} \geq 6 C$, we have $L_{3}=\frac{r_{\text {min }} \varepsilon}{6}\left(\lambda-\frac{6 C}{m}\right) \geq 0$. For $L_{4}$, it follows from that

$$
L_{4}=\varepsilon\left(\frac{r_{\min } \lambda}{6}-\frac{C \varepsilon}{m}\right)>\varepsilon\left(\frac{r_{\min } \lambda}{6}-\frac{C r_{\min } \lambda}{4 m}\right)=\varepsilon \frac{r_{\min } \lambda}{6}\left(1-\frac{3 C}{2 m}\right)>0,
$$

where the first inequality follows from (Equation 6.2.11) and the last inequality follows from (Equation 6.2 .9 ), which says $m \geq m_{1}>\frac{3 C}{2}$. Finally,

$$
L_{5}=\frac{r_{\min }^{2} c_{1} \lambda}{24 m}-\frac{2 \pi C r_{\min }^{2}}{m^{2}}=\frac{r_{\min }^{2}}{24 m}\left(c_{1} \lambda-\frac{48 \pi C}{m}\right)>\frac{r_{\min }^{2}}{24 m}\left(\frac{c_{1}^{2}}{m}-\frac{48 \pi C}{m}\right)>0
$$

where the first equality follows from the definition of $\varepsilon$, the first inequality follows from $\lambda>\frac{c_{1}}{m}$ and the last inequality follows from (Equation 6.2 .9 ), which says $c_{1} \geq \sqrt{48 \pi C}$. Therefore it follows from (Equation 6.2.16) that

$$
\begin{equation*}
\Psi\left(x_{2}\right)-\Psi\left(x_{1}\right)>0, \tag{6.2.17}
\end{equation*}
$$

which finishes the proof.

Now we are ready to prove Theorem H.

Proof of Theorem $H$ : Let $c_{1}, c_{2}$ and $m_{1}$ be constants in Lemma 6.2.2 and $\delta$ be as in Proposi-
tion 6.1.8. Lastly, let $c_{3}$ be the constant in (Equation 6.1.36). Now we set

$$
\begin{equation*}
c:=\max \left\{c_{1}, \frac{c_{3}}{2}\left(c_{2} \delta+\frac{9 \pi^{2} \delta}{c_{2}}\right)\right\} \quad \text { and } \quad m_{0}:=\max \left\{2 c_{2}, m_{1}\right\}+1 \tag{6.2.18}
\end{equation*}
$$

We will prove that if $(D, \Omega)$ is a solution to (Equation 1.2.4) such that $D$ is $m$-fold symmetric for $m \geq m_{0}$ and simply-connected, then

$$
\begin{equation*}
\lambda:=\frac{1}{2}-\Omega \leq \frac{c}{m} \tag{6.2.19}
\end{equation*}
$$

Towards a contradiction, let us suppose that there exists $(D, \Omega)$ such that

$$
\begin{equation*}
\lambda>\frac{c}{m} . \tag{6.2.20}
\end{equation*}
$$

It is clear that (Equation 6.2.18) implies $\lambda>\frac{c_{1}}{m}$ and $m \geq m_{1}$. Thus Lemma 6.2.2 implies that there exists $u \in C^{1}(\mathbb{T})$ such that

$$
\begin{equation*}
\partial D=\{(1+u(\theta))(\cos \theta, \sin \theta): \theta \in \mathbb{T}\} \quad \text { and } \quad\|u\|_{L^{\infty}(\mathbb{T})}<\frac{c_{2}}{m} \tag{6.2.21}
\end{equation*}
$$

Since $m \geq m_{0}>2 c_{2}$, which follows from (Equation 6.2.18), we have that $\|u\|_{L^{\infty}(\mathbb{T})}<\frac{1}{2}$.
To derive a contradiction, we will use the identity (Equation 6.1.2). To estimate the right-hand side of it, we apply Proposition 6.1 .8 with $a:=\frac{2 c_{2}}{m} \in\left(2\|u\|_{L^{\infty}(\mathbb{T})}, 1\right)$ and obtain

$$
\begin{align*}
\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x & \leq \frac{2 c_{2} \delta}{m} \int_{\mathbb{T}}|u|^{2} d \theta+\frac{\delta m}{2 c_{2}} \int_{\mathbb{T}} f(\theta)^{2} d \theta \\
& \leq \frac{2 c_{2} \delta}{m} \int_{\mathbb{T}}|u|^{2} d \theta+\frac{\pi \delta m}{c_{2}}\|f\|_{L^{\infty}(\mathbb{T})}^{2} \tag{6.2.22}
\end{align*}
$$

where $f(\theta)=\int_{0}^{\theta} u(s)^{2}+2 u(s) d s$. Using (Equation 6.1.32) and $\frac{2 \pi}{m}$-periodicity of $u$, it is clear that
$f$ is also $\frac{2 \pi}{m}$-periodic. Furthermore, for $\theta \in\left(0, \frac{2 \pi}{m}\right)$, we have that (recall that $\left.\|u\|_{L^{\infty}(\mathbb{T})}<\frac{1}{2}\right)$,

$$
\begin{aligned}
|f(\theta)| & =\left|\int_{0}^{\theta} u(s)^{2}+2 u(s) d s\right| \leq \int_{0}^{\theta} 3|u(s)| d s \leq 3 \sqrt{\int_{0}^{\theta}|u(s)|^{2} d s} \sqrt{\theta}<3 \sqrt{\int_{0}^{\frac{2 \pi}{m}}|u(s)|^{2} d s} \sqrt{\frac{2 \pi}{m}} \\
& \leq 3 \frac{\sqrt{2 \pi}}{m} \sqrt{\int_{\mathbb{T}}|u(s)|^{2} d s}
\end{aligned}
$$

Thus, (Equation 6.2.22) yields that

$$
\begin{equation*}
\int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \frac{2 c_{2} \delta}{m} \int_{\mathbb{T}}|u|^{2} d \theta+\frac{18 \pi^{2} \delta}{c_{2} m} \int_{\mathbb{T}}|u|^{2} d \theta=\frac{1}{m}\left(2 c_{2} \delta+\frac{18 \pi^{2} \delta}{c_{2}}\right) \int_{\mathbb{T}}|u|^{2} d \theta \tag{6.2.23}
\end{equation*}
$$

For the left-hand side of (Equation 6.1.2), we use (Equation 6.1.36) to obtain

$$
\begin{equation*}
\frac{2}{c_{3}} \int_{\mathbb{T}}|u|^{2} d \theta \leq \int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi} \tag{6.2.24}
\end{equation*}
$$

Hence it follows from (Equation 6.2.23), (Equation 6.2.24) and (Equation 6.1.2) that

$$
\frac{2 \lambda}{c_{3}} \int_{\mathbb{T}}|u|^{2} d \theta \leq \lambda\left(\int_{D}|x|^{2} d x-\frac{|D|^{2}}{2 \pi}\right)=\frac{1}{2} \int_{D}\left|x-2 \nabla\left(1_{D} * \mathcal{N}\right)\right|^{2} d x \leq \frac{1}{m}\left(c_{2} \delta+\frac{9 \pi^{2} \delta}{c_{2}}\right) \int_{\mathbb{T}}|u|^{2} d \theta
$$

Therefore we have

$$
\lambda \leq \frac{c_{3}}{2}\left(c_{2} \delta+\frac{9 \pi^{2} \delta}{c_{2}}\right) \frac{1}{m} \leq \frac{c}{m}
$$

where the last inequality follows from our choice for $c$ in (Equation 6.2.18). This contradicts our assumption (Equation 6.2.20), thus completes the proof.

By simple maximum principle type argument, Theorem H gives a upper bound for $r_{\max }$.

Corollary 6.2.3. There exist constants $c>0$ and $m_{1} \geq 3$ such that if $(D, \Omega)$ is a solution to
(Equation 1.2.4) that is simply-connected, $m$-fold symmetric for some $m \geq m_{1}$ and $|D|=\pi$, then $r_{\max }-1 \leq \frac{c}{m}$.

Proof. Thanks to Theorem H, we can pick a constants $C_{1}$ and $m_{0}$ such that if $m \geq m_{0}$, then $\lambda<$ $\frac{C_{1}}{m}$. Moreover, it follows from Lemma 6.2 .1 that there exists $C_{2}>0$ such that $\left|\nabla \varphi_{m}(r, \theta)\right| \leq \frac{C_{2} r}{m}$. Now, let us choose

$$
m_{1}:=\max \left\{m_{0}, 2\left(C_{1}+C_{2}\right)\right\}+1
$$

Since $\Delta \Psi=2 \lambda>0$ in $D$, the maximum principle for subharmonic functions implies that $\partial_{r} \Psi\left(r_{\max }, 0\right)>0$. Thus it follows from Lemma 6.2.1 that

$$
\begin{aligned}
0 & <\partial_{r} \varphi^{r}\left(r_{\max }\right)+C_{2} \frac{r_{\max }}{m}-\Omega r_{\max } \\
& =\frac{\left\lvert\, D \cap B_{r_{\max } \mid}^{2 \pi r_{\max }}+C_{2} \frac{r_{\max }}{m}-\Omega r_{\max }\right.}{} \\
& =\frac{1}{2 r_{\max }}+\left(\frac{C_{2}}{m}-\frac{1}{2}+\lambda\right) r_{\max } \\
& \leq \frac{1}{2 r_{\max }}+\left(\frac{C_{1}+C_{2}}{m}-\frac{1}{2}\right) r_{\max }
\end{aligned}
$$

where we used $\left|D \cap B_{r_{\max }}\right|=|D|=\pi$ to get the second equality and the last inequality follows from $\lambda \leq \frac{C_{1}}{m}$. Since $\frac{C_{1}+C_{2}}{m}<\frac{1}{2}$, we obtain,

$$
r_{\max }-1 \leq \sqrt{\frac{\frac{1}{2}}{\frac{1}{2}-\frac{C_{1}+C_{2}}{m}}}-1 \lesssim \frac{1}{m}
$$

### 6.2.2 Patches along bifurcation curves

This subsection is devoted to the proof of Theorem I. Since we are interested in a curve $\mathscr{C}_{m}$ that satisfies (A1)-(A4), we will make the following assumptions for the patches throughout this
subsection.
(a) $D$ is star-shaped, that is $\partial D=\{(1+u(\theta))(\cos \theta, \sin \theta): \theta \in \mathbb{T}\}$ for some $u \in C^{2}(\mathbb{T})$.
(b) $u$ is even and $\frac{2 \pi}{m}$-periodic for some $m \geq 3$, that is, $u(-\theta)=u(\theta)$ and $u\left(\theta+\frac{2 \pi}{m}\right)=u(\theta)$.
(c) $\partial_{\theta} u(\theta)<0$ for all $\theta \in\left(0, \frac{\pi}{m}\right)$.

For such a patch, we denote $r_{\text {min }}:=\min _{x \in \partial D}|x|=u\left(\frac{\pi}{m}\right)$ and $r_{\text {max }}:=\max _{x \in \partial D}|x|=u(0)$. Furthermore, we denote $\eta:=(1+u)^{-1}:\left(r_{\min }, r_{\max }\right) \mapsto\left(0, \frac{\pi}{m}\right)$. By the symmetry, we only need to focus on the fundamental sector $S:=\left\{(r, \theta): r \geq 0, \theta \in\left(0, \frac{\pi}{m}\right)\right\}$. See Figure Figure 6.2 for an illustration of these definitions.


Figure 6.2: Illustration of the definitions of $r_{\min }, r_{\max }, \eta(\rho)$ and $S$ on a 6 -fold vortex patch

Note that we will establish several lemmas with assuming $|D|=|B|=\pi$. Certainly this is not satisfied by the solutions on the curve $\mathscr{C}_{m}$ but we will resolve this issue in the proof of the theorem

Our proof for Theorem I relies on Theorem H. Roughly speaking, we will show that if $\|u\|_{L^{\infty}(\mathbb{T})}$ is large compared to $\frac{1}{m}$, then $\lambda\left(=\frac{1}{2}-\Omega\right)$ must be large enough to contradict Theorem H. However, the main difficulty comes from the fact that lower bounds for $\lambda$ that we can derive from the identities (Equation 1.2.8) and (Equation 1.2.9) are not comparable with $\|u\|_{L^{\infty}}$ (Lemma 6.2.4).

Thus, the scenario that we want to rule out is that for large $m, \partial D$ is so spiky that $\int_{\mathbb{T}}|u|^{2} d \theta$ is small while $\|u\|_{L^{\infty}}$ is relatively large.

Since $r_{\max }$ can be estimated as in Corollary 6.2 .3 , we will mainly focus on estimating $r_{\text {min }}$. Using the identity (Equation 6.1.1), we derive a lower bound for $\lambda$ in the next lemma.

Lemma 6.2.4. If $(D, \Omega)$ is a solution to (Equation 1.2.4) with $|D|=\pi$ then

$$
\lambda \gtrsim \frac{\int_{\mathbb{T}}|u|^{2} d \theta}{r_{\max }\|u\|_{L^{\infty}(\mathbb{T})}}
$$

Proof. we use (Equation 6.1.1) in Lemma 6.1.1 to obtain

$$
\lambda\left(\int_{D}|x|^{2}-2 p(x) d x\right)=\left(\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}\right) .
$$

For a moment, let us assume that

$$
\begin{equation*}
\int_{D}|x|^{2}-2 p(x) d x \lesssim r_{\max }\|u\|_{L^{\infty}(\mathbb{T})} . \tag{6.2.25}
\end{equation*}
$$

Then it follows from (Equation 6.1.36) that

$$
\lambda=\frac{\int_{D} \frac{|x|^{2}}{2} d x-\frac{|D|^{2}}{4 \pi}}{\int_{D}|x|^{2}-2 p(x) d x} \gtrsim \frac{\int_{\mathbb{T}}|u|^{2} d \theta}{r_{\max }\|u\|_{L^{\infty}}(\mathbb{T})},
$$

which implies the desired result.
Now let us prove (Equation 6.2.25). Note that $p(x) \geq \frac{r_{\text {min }}^{2}-|x|^{2}}{2}$ in $B_{r_{m i n}}$. Indeed, $p-\frac{r_{\text {min }}^{2}-|x|^{2}}{2}$ is harmonic in $B_{r_{m i n}}$ and non-negative on $\partial B_{r_{m i n}}$ since $p$ is non-negative in $D$. Therefore the inequality follows from the maximum principle. From this, we obtain

$$
\int_{D} 2 p(x) d x \geq \int_{B_{r_{\min }}} r_{\min }^{2}-|x|^{2} d x=\frac{\left|B_{r_{\min }}\right|^{2}}{2 \pi}
$$

Since $\int_{D}|x|^{2} d x<\int_{B_{r_{\text {max }}}}|x|^{2} d x=\frac{\left|B_{r_{\text {max }}}\right|^{2}}{2 \pi}$, it follows that

$$
\int_{D}|x|^{2}-2 p(x) d x=\frac{1}{2 \pi}\left(\left|B_{r_{\max }}\right|^{2}-\left|B_{r_{\min }}\right|^{2}\right) \lesssim r_{\max }\|u\|_{L^{\infty}(\mathbb{T})},
$$

which proves (Equation 6.2.25).

Thanks to Lemma 6.2.4, we only need to rule out the case where $\int_{\mathbb{T}}|u|^{2} d \theta$ is too small, compared to $\|u\|_{L^{\infty}}$. To this end, we will pick $r_{1}$, and $r_{2}$ so that $r_{\text {min }}<r_{1}<r_{2}<1$ and find a lower bound for $\frac{\pi}{m}-\eta\left(r_{2}\right)$ by showing that $\left|u^{\prime}(\theta)\right|$ is bounded from above for $1+u(\theta) \in$ $\left(r_{\text {min }}, r_{1}\right)$. Since the relative stream function $\Psi$ is constant on $\partial D$, we have $\frac{d}{d \theta}(\Psi((1+u(\theta)), \theta))=$ 0 . Therefore (Equation 6.2.1) yields that

$$
\begin{equation*}
u^{\prime}(\theta)=-\frac{\partial_{\theta} \Psi(r, \theta)}{\partial_{r} \Psi(r, \theta)}=-\frac{\partial_{\theta} \varphi_{m}(r, \theta)}{\left(\partial_{r} \varphi^{r}(r)-\Omega r\right)+\partial_{r} \varphi_{m}(r, \theta)}, \quad \text { where } r=1+u(\theta) \tag{6.2.26}
\end{equation*}
$$

In the next two lemmas, we will estimate the denominator and numerator in (Equation 6.2.26) but the proofs will be postponed to the end of this subsection.

Lemma 6.2.5. Let $(D, \Omega)$ be a solution to (Equation 1.2.4) that satisfies the assumptions (a)-(c) for some $m \geq 3$ and $|D|=\pi$. Let $r_{1}, r_{2}>0$ be such that $r_{\text {min }}<r_{1}<r_{2}<1$ and let $\delta:=\frac{\pi}{m}-\eta\left(r_{2}\right)$. Then there exist constants $c, C>0$ such that if $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, it holds that

$$
\partial_{r} \varphi^{r}(r)-\Omega r \geq m \delta\left(c \frac{\left(1-r_{2}\right)^{2}}{\|u\|_{L^{\infty}(\mathbb{T})}}-C\left(r_{1}-r_{m i n}\right)\right)
$$

for all $r \in\left(r_{\text {min }}, r_{1}\right)$.

Lemma 6.2.6. Let $D$ be a patch that satisfies the assumptions (a)-(c) for some $m \geq 3$. Let us pick
$r_{1}, r_{2}>0$ so that $r_{\text {min }}<r_{1}<r_{2}<1$ and $r_{1}^{2} \leq r_{\text {min }} r_{2}$. If $\delta:=\frac{\pi}{m}-\eta\left(r_{2}\right)<\frac{\pi}{4 m}$, then it holds that

$$
\begin{align*}
\partial_{r} \varphi_{m}(r, \eta(r)) & \geq-\frac{c r}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \delta  \tag{6.2.27}\\
\partial_{\theta} \varphi_{m}(r, \eta(r)) & \leq \frac{c r^{2}}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \delta \tag{6.2.28}
\end{align*}
$$

for all $r \in\left(r_{\text {min }}, r_{1}\right)$, where $c$ is a universal constant that does not depend on any variables..

Note that the linear dependence on $\delta$ in (Equation 6.2.27) and (Equation 6.2.28) is crucial in the proof of the next lemma, since this allows us to bound $u^{\prime}$ independently of $\delta$ when we plug the above bounds into (Equation 6.2.26).

Now we can rule out the scenario that $\partial D$ is too spiky inwards.

Lemma 6.2.7. There exist $c, C>0$ and $m_{1} \geq 3$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) that satisfies the assumptions (a)-(c) for some $m \geq m_{1},|D|=\pi$ and $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, then

$$
1-r_{\min } \leq \frac{c}{m} \quad \text { or } \quad \int_{\mathbb{T}}|u|^{2} d \theta \geq C\left(1-r_{\min }\right)^{2}
$$

Proof. Thanks to Lemma 6.2.5 and 6.2.6, we can choose $c_{1}$ and $c_{2}>0$ such that if $r_{\text {min }}<r_{1}<$ $r_{2}<1$ and $r_{1}^{2} \leq r_{\text {min }} r_{2}$, then

$$
\begin{align*}
& \partial_{r} \varphi_{r}(r)-\Omega r \geq m \delta\left(c_{1} \frac{\left(1-r_{2}\right)^{2}}{\|u\|_{L^{\infty}(\mathbb{T})}}-c_{2}\left(r_{1}-r_{m i n}\right)\right),  \tag{6.2.29a}\\
& \partial_{r} \varphi_{m}(r, \eta(r)) \geq-\frac{c_{2}}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \delta,  \tag{6.2.29b}\\
& \partial_{\theta} \varphi_{m}(r, \eta(r)) \leq \frac{c_{2}}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \delta, \tag{6.2.29c}
\end{align*}
$$

for $r \in\left(r_{\text {min }}, r_{1}\right)$. We will pick

$$
\begin{equation*}
c:=\max \left\{12, \frac{24 c_{2}}{c_{1}}\right\}, \quad C:=\frac{\pi}{18} \quad \text { and } \quad m_{1}:=2 c \tag{6.2.30}
\end{equation*}
$$

Now let us assume that $D$ is a solution to (Equation 1.2.4) that satisfies the assumptions (a)-(c) for some $m \geq m_{1},|D|=\pi$ and $1-r_{\text {min }}=\|u\|_{L^{\infty}(\mathbb{T})}<\frac{1}{2}$. If $1-r_{\text {min }} \leq \frac{c}{m}$, then there is nothing to prove, thus let us assume that

$$
\begin{equation*}
1-r_{\min }>\frac{c}{m} \tag{6.2.31}
\end{equation*}
$$

Let $\tilde{c}:=m\left(1-r_{\text {min }}\right)$ so that $1-r_{\text {min }}=\frac{\tilde{c}}{m}$. From (Equation 6.2.31), we have

$$
\begin{equation*}
\tilde{c}>c \tag{6.2.32}
\end{equation*}
$$

We choose

$$
\begin{equation*}
r_{1}:=1-\frac{\tilde{c}-2}{m} \quad \text { and } \quad r_{2}:=1-\frac{\tilde{c}}{3 m} . \tag{6.2.33}
\end{equation*}
$$

And we consider two cases: $\frac{\pi}{m}-\eta\left(r_{2}\right) \geq \frac{\pi}{4 m}$ and $\frac{\pi}{m}-\eta\left(r_{2}\right)<\frac{\pi}{4 m}$.
Case1. Let us assume that $\frac{\pi}{m}-\eta\left(r_{2}\right) \geq \frac{\pi}{4 m}$.
Since $\frac{\pi}{m}-\eta\left(r_{2}\right) \geq \frac{\pi}{4 m}$, it follows from the monotonicity of $u$ that $u(\theta)<r_{2}-1=-\frac{\tilde{c}}{3 m}$ for $\theta \in\left(\frac{3 \pi}{4 m}, \frac{\pi}{m}\right)$. Using $m$-fold symmetry of $u$, we obtain

$$
\begin{equation*}
\int_{\mathbb{T}}|u|^{2} d \theta \geq 2 m \int_{0}^{\frac{\pi}{m}}|u|^{2} d \theta>2 m \int_{\frac{3 \pi}{4 m}}^{\frac{\pi}{m}}|u|^{2} d \theta \geq \frac{\pi \tilde{c}^{2}}{18 m^{2}}=C\left(1-r_{\text {min }}\right)^{2} \tag{6.2.34}
\end{equation*}
$$

where the last equality follows from (Equation 6.2.30), which says $C=\frac{\pi}{18}$.
Case2. Now we assume $\frac{\pi}{m}-\eta\left(r_{2}\right)<\frac{\pi}{4 m}$.
We first check whether $r_{1}$ and $r_{2}$ in (Equation 6.2.33) satisfy the hypotheses in Lemma 6.2.6.

Clearly $r_{\text {min }}<r_{1}<r_{2}<1$, since $\tilde{c}>c \geq 4$, which follows from (Equation 6.2.30) and (Equation 6.2.32). To show $r_{1}^{2} \leq r_{\text {min }} r_{2}$, we compute

$$
\begin{aligned}
r_{1}^{2}-r_{m i n} r_{2} & =\left(1-\frac{\tilde{c}-2}{m}\right)^{2}-\left(1-\frac{\tilde{c}}{m}\right)\left(1-\frac{\tilde{c}}{3 m}\right) \\
& =-\frac{2 \tilde{c}}{3 m}+\frac{4}{m}+\frac{2 \tilde{c}^{2}-12 \tilde{c}+12}{3 m^{2}} \\
& \leq-\frac{\tilde{c}}{3 m}+\frac{4}{m}+\frac{-12 \tilde{c}+12}{3 m^{2}} \\
& \leq 0
\end{aligned}
$$

where the first ineqiality follows from $\frac{\tilde{c}}{m}=1-r_{\text {min }}=\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, and the last inequality follows from $\tilde{c}>12$.

Since $\partial D$ is a level set of $\Psi, \Psi(1+u(\theta), \theta)=$ does not depend on $\theta$. Therefore,

$$
-u^{\prime}(\theta)=\frac{\partial_{\theta} \Psi(1+u(\theta), \theta)}{\partial_{r} \Psi(1+u(\theta), \theta)}=\frac{\partial_{\theta} \varphi_{m}(1+u(\theta), \theta)}{\partial_{r} \varphi^{r}(1+u(\theta))-\Omega r+\partial_{r} \varphi_{m}(1+u(\theta), \theta)} .
$$

Hence, it follows from (Equation 6.2.29) that

$$
\begin{equation*}
-u^{\prime}(\theta) \leq \frac{\frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}}{\frac{c_{1}}{c_{2}} m \frac{\left(1-r_{2}\right)^{2}}{\|u\|_{L^{\infty}(\mathrm{T})}}-m\left(r_{1}-r_{\min }\right)-\frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}} \quad \text { for } \quad(1+u(\theta))<r_{1} . \tag{6.2.35}
\end{equation*}
$$

Now we assume for a moment that

$$
\begin{equation*}
\frac{\frac{c_{1}}{c_{2}} m\left(1-r_{2}\right)^{2}}{3\|u\|_{L^{\infty}}} \geq \max \left\{m\left(r_{1}-r_{\text {min }}\right), \frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}\right\} . \tag{6.2.36}
\end{equation*}
$$

Then (Equation 6.2.35) yields that

$$
0 \leq-u^{\prime}(\theta) \leq 1 \quad \text { for } \quad(1+u(\theta))<r_{1}
$$

which implies that

$$
\begin{equation*}
\frac{2}{m}=r_{1}-r_{\min }=\int_{\eta\left(r_{1}\right)}^{\frac{\pi}{m}}-u^{\prime}(\theta) d \theta \leq \frac{\pi}{m}-\eta\left(r_{1}\right) \tag{6.2.37}
\end{equation*}
$$

Furthermore, monotonicity of $u$ implies that $|u(\theta)|>\left|r_{1}-1\right|>\left|r_{2}-1\right|$ for $\theta \in\left(\eta\left(r_{1}\right), \frac{\pi}{m}\right)$. Therefore we obtain that
$\int_{\mathbb{T}}|u|^{2} d \theta>2 m \int_{\eta\left(r_{1}\right)}^{\frac{\pi}{m}}|u|^{2} d \theta \geq 2 m\left(\frac{\pi}{m}-\eta\left(r_{1}\right)\right)\left(1-r_{2}\right)^{2} \geq \frac{4 \tilde{c}^{2}}{9 m^{2}}=\frac{4}{9}\left(1-r_{m i n}\right)^{2} \geq C\left(1-r_{m i n}\right)^{2}$,
where the third inequality follows from (Equation 6.2.33) and (Equation 6.2.37) and the last inequality follows from $C=\frac{\pi}{18}<\frac{4}{9}$. Thus the desired result follows from (Equation 6.2.34) and (Equation 6.2.38).

To complete the proof, we need to show (Equation 6.2.36). It follows from (Equation 6.2.33) that

$$
\begin{equation*}
\frac{\frac{c_{1}}{c_{2}} m\left(1-r_{2}\right)^{2}}{3\|u\|_{L^{\infty}(\mathbb{T})}}=\frac{\frac{c_{1}}{c_{2}} m\left(1-r_{2}\right)^{2}}{3\left(1-r_{\min }\right)}=\frac{c_{1} \tilde{c}}{12 c_{2}} \geq 2 \tag{6.2.39}
\end{equation*}
$$

where the last inequality follows from (Equation 6.2.30) and (Equation 6.2.32) which imply that $\tilde{c} \geq \frac{24 c_{2}}{c_{1}}$. We also have

$$
\begin{equation*}
m\left(r_{1}-r_{\min }\right)=2 \tag{6.2.40}
\end{equation*}
$$

which follows from (Equation 6.2.33). To estimate $\frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}$, let us use an elementary inequality that for any $0<b<a<m$, it holds that

$$
\begin{equation*}
\left(\frac{1-\frac{a}{m}}{1-\frac{b}{m}}\right)^{m} \leq e^{-(a-b)} \tag{6.2.41}
\end{equation*}
$$

Indeed, by taking logarithm in the left-hand side, we can compute,

$$
\log \left(\left(\frac{1-\frac{a}{m}}{1-\frac{b}{m}}\right)^{m}\right)=m \log \left(1-\frac{a-b}{m-b}\right) \leq m \log \left(1-\frac{a-b}{m}\right) \leq-(a-b),
$$

where the last inequality follows from $\log (1-x)<-x$ for all $x>0$. This proves (Equation 6.2.41). Then we use (Equation 6.2.33) and obtain

$$
\begin{equation*}
\frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}=\frac{1}{1-\left(\frac{1-\frac{\tilde{c}-2}{m}}{1-\frac{\varepsilon}{3 m}}\right)^{m}} \leq \frac{1}{1-e^{-\left(\frac{\tilde{\tilde{c}}}{3}-2\right)}} \leq 2 \tag{6.2.42}
\end{equation*}
$$

where the last inequality follows from (Equation 6.2.30) and (Equation 6.2.32), which imply $\tilde{c} \geq$ 12.

Thus (Equation 6.2.39), (Equation 6.2.40) and (Equation 6.2.42) yield

$$
\begin{equation*}
\frac{\frac{c_{1}}{c_{2}} m\left(1-r_{2}\right)^{2}}{3\|u\|_{L^{\infty}(\mathbb{T})}} \geq 2 \geq \max \left\{m\left(r_{1}-r_{\min }\right), \frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}\right\} . \tag{6.2.43}
\end{equation*}
$$

This proves (Equation 6.2.36) and finishes the proof.

Now we can estimate $\|u\|_{L^{\infty}(\mathbb{T})}$ whose corresponding patch has area $\pi$, that is, $|D|=\pi$.

Proposition 6.2.8. There exist constants $c>0$ and $m_{0} \geq 3$ such that if $(D, \Omega)$ is a solution to (Equation 1.2.4) that satisfies assumptions (a)-(c) for some $m \geq m_{0},|D|=\pi$ and $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, then

$$
\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{c}{m} .
$$

Proof. In order to use the previous lemmas, let us fix some constants. We fix constants $c_{i}^{\prime} s$ and $m_{1}$ so that if $(D, \Omega)$ is a solution to (Equation 1.2.4) that satisfies the assumptions (a)-(c) for some $m \geq m_{1}$ and $|D|=\pi$, then
(B1) (From Theorem H) $\lambda \leq \frac{c_{1}}{m}$.
(B2) (From Corollary 6.2.3) $r_{\max }-1 \leq \frac{c_{2}}{m}$.
(B3) (From Lemma 6.2.4) $\lambda \geq c_{3} \frac{\int_{\mathbb{T}}|u|^{2} d \theta}{r_{\max }\|u\|_{L^{\infty}(\mathbb{T})}}$.
(B4) (From Lemma 6.2.7) if $1-r_{\text {min }}=\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$ then

$$
1-r_{\min } \leq \frac{c_{4}}{m} \quad \text { or } \quad \int_{\mathbb{T}}|u|^{2} d \theta \geq c_{5}\left(1-r_{\min }\right)^{2}
$$

Let us set

$$
\begin{equation*}
c:=\max \left\{c_{2}, c_{4}, \frac{2 c_{1}}{c_{3} c_{5}}\right\}+1, \quad \text { and } \quad m_{0}:=\max \left\{m_{1}, 2 c\right\}+1 \tag{6.2.44}
\end{equation*}
$$

Then we will prove that if $(D, \Omega)$ is a solution to (Equation 1.2.4) and satisfies the assumptions (a)-(c) for some $m \geq m_{0}$, then

$$
\begin{equation*}
\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{c}{m} \tag{6.2.45}
\end{equation*}
$$

Let us assume for a contradiction that

$$
\begin{equation*}
\|u\|_{L^{\infty}(\mathbb{T})}>\frac{c}{m} . \tag{6.2.46}
\end{equation*}
$$

Then we have that

$$
\begin{equation*}
\int_{\mathbb{T}}|u|^{2} d \theta \geq c_{5}\left(1-r_{\min }\right)^{2} \tag{6.2.47}
\end{equation*}
$$

Indeed,

$$
\|u\|_{L^{\infty}(\mathbb{T})}>\frac{c}{m}>\frac{c_{2}}{m} \geq r_{\max }-1,
$$

where we used (Equation 6.2.44), (Equation 6.2.46) and (B2), therefore $\|u\|_{L^{\infty}(\mathbb{T})}=1-r_{\text {min }}$. Thus (B4) and (Equation 6.2.46) imply (Equation 6.2.47). Furthermore (B2) and (Equation 6.2.44) also imply that

$$
r_{\max } \leq 1+\frac{c_{2}}{m} \leq 2
$$

Thus we use (B3) and (Equation 6.2.47) to obtain

$$
\lambda \geq c_{3} \frac{c_{5}\left(1-r_{\text {min }}\right)^{2}}{2\|u\|_{L^{\infty}(\mathbb{T})}} \geq \frac{c_{3} c_{5}\|u\|_{L^{\infty}(\mathbb{T})}}{2} \geq \frac{c c_{3} c_{5}}{2 m}>\frac{c_{1}}{m}
$$

where the third inequality follows from (Equation 6.2.46) and the last inequality follows from (Equation 6.2.44). However this contradicts (B1).

Now we are ready to prove the main theorem of this subsection

Proof of Theorem I: Thanks to Proposition 6.2 .8 , we can pick $c_{1}$ and $m_{1}$ so that $2 c_{1}<m_{1}$ and if $(D, \Omega)$ is a solution to (Equation 1.2.4), that satisfies the assumptions (a)-(c) for some $m \geq m_{1}$ and $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, then

$$
\begin{equation*}
\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{c_{1}}{m} . \tag{6.2.48}
\end{equation*}
$$

Now let us consider a curve $\mathscr{C}_{m}$, that satisfies the properties (A1)-(A4) for some $m \geq m_{1}$. We will
show that

$$
\begin{equation*}
\sup _{s \in[0, \infty)}\left\|\tilde{u}_{m}(s)\right\|_{L^{\infty}(\mathbb{T})} \lesssim \frac{1}{m} . \tag{6.2.49}
\end{equation*}
$$

To do so, let us define $\hat{u}_{m}(s)$ so that,

$$
\begin{equation*}
1+\hat{u}_{m}(s)=\frac{\sqrt{\pi}}{\sqrt{\left|D^{\tilde{u}_{m}(s)}\right|}}\left(1+\tilde{u}_{m}(s)\right) \tag{6.2.50}
\end{equation*}
$$

where the definition of $D^{\tilde{u}_{m}(s)}$ is as in (A2). Clearly, $s \mapsto \hat{u}_{m}(s)$ is a continuous curve in $C^{2}(\mathbb{T})$ such that $\left|D^{\hat{u}_{m}(s)}\right|=\pi$. Since $\hat{u}_{m}(0)=0$, it follows from the continuity of the curve and (Equation 6.2.48) that

$$
\begin{equation*}
\sup _{s \in[0, \infty)}\left\|\hat{u}_{m}(s)\right\|_{L^{\infty}(\mathbb{T})} \leq \frac{c_{1}}{m} . \tag{6.2.51}
\end{equation*}
$$

Now let us pick an arbitrary $s \in[0, \infty)$ and denote $\tilde{u}:=\tilde{u}_{m}(s)$ and $\hat{u}:=\hat{u}_{m}(s)$. Then it follows from (A1) and (Equation 6.2.50) that

$$
\begin{aligned}
0=\int_{\mathbb{T}} \tilde{u}(\theta) d \theta & =2 \pi\left(\frac{\sqrt{\mid D^{\tilde{u} \mid}}}{\sqrt{\pi}}-1\right)+\frac{\sqrt{\left|D^{\tilde{u}}\right|}}{\sqrt{\pi}} \int_{\mathbb{T}} \hat{u}(\theta) d \theta \\
& =2 \pi\left(\frac{\sqrt{\left|D^{\tilde{u}}\right|}}{\sqrt{\pi}}-1\right)-\frac{1}{2} \frac{\sqrt{\left|D^{\tilde{u}}\right|}}{\sqrt{\pi}} \int_{\mathbb{T}} \hat{u}(\theta)^{2} d \theta
\end{aligned}
$$

where the last equality follows from (Equation 6.1.32). Hence (Equation 6.2.51) implies that $\frac{\sqrt{\left|D^{\tilde{u} \mid}\right|}}{\sqrt{\pi}}=1+O\left(\frac{1}{m^{2}}\right)$. Therefore (Equation 6.2.50) and (Equation 6.2.51) yield that

$$
\|\tilde{u}\|_{L^{\infty}(\mathbb{T})}=\frac{\sqrt{\left|D^{\tilde{u}}\right|}}{\sqrt{\pi}}\left(1+\|\hat{u}\|_{L^{\infty}(\mathbb{T})}\right)-1=\|\hat{u}\|_{L^{\infty}(\mathbb{T})}+O\left(\frac{1}{m^{2}}\right) \lesssim \frac{1}{m},
$$

where the last equality follows from (Equation 6.2.51). This proves (Equation 6.2.49) and the theorem.

Proof of Lemma 6.2.5: From Lemma 6.2.1, it follows that

$$
\begin{equation*}
\partial_{r} \varphi^{r}(r)-\Omega r=r\left(\frac{1}{2}-\Omega\right)-\frac{\left|B_{r} \backslash D\right|}{2 \pi r}=r \lambda-\frac{\left|B_{r} \backslash D\right|}{2 \pi r}=: J_{1}-J_{2} \tag{6.2.52}
\end{equation*}
$$

where we used $\lambda=\frac{1}{2}-\Omega$. Note that we have $\|u\|_{L^{\infty}(\mathbb{T})} \leq \frac{1}{2}$, therefore $J_{1} \sim \lambda$ and $J_{2} \sim\left|B_{r} \backslash D\right|$. Let us estimate $J_{2}$ first. Since $u$ is even and $m$-periodic, we have that for all $r \in\left(r_{\text {min }}, r_{1}\right)$,

$$
\left|B_{r} \backslash D\right|=2 m \int_{r_{m i n}}^{r}\left(\frac{\pi}{m}-\eta(\rho)\right) \rho d \rho \lesssim m r \delta\left(r-r_{\min }\right) \leq m r \delta\left(r_{1}-r_{m i n}\right)
$$

where we used $\frac{\pi}{m}-\eta(\rho)<\delta$ for $\rho<r<r_{2}$ to get the first inequality. Hence we obtain

$$
\begin{equation*}
J_{2} \lesssim m \delta\left(r_{1}-r_{\min }\right) \tag{6.2.53}
\end{equation*}
$$

To estimate $J_{1}$, we use Lemma 6.2.4 and obtain

$$
\begin{equation*}
\lambda \gtrsim \frac{\int_{\mathbb{T}}|u|^{2} d \theta}{r_{\max }\|u\|_{L^{\infty}}} \gtrsim \frac{\int_{\mathbb{T}}|u|^{2} d \theta}{\|u\|_{L^{\infty}}}, \tag{6.2.54}
\end{equation*}
$$

where we we used $r_{\max } \leq 1+\|u\|_{L^{\infty}(\mathbb{T})} \lesssim 1$ to get the last inequality. From periodicity of $u$, it follows that

$$
\begin{equation*}
\int_{\mathbb{T}}|u|^{2} d \theta=2 m \int_{0}^{\frac{\pi}{m}}|u|^{2} d \theta \geq 2 m \int_{\eta\left(r_{2}\right)}^{\frac{\pi}{m}}|u|^{2} d \theta \geq 2 m\left(1-r_{2}\right)^{2}\left(\frac{\pi}{m}-\eta\left(r_{2}\right)\right)=2 m \delta\left(1-r_{2}\right)^{2} \tag{6.2.55}
\end{equation*}
$$

where we used $1+u(\theta)<r_{2}$ for $\theta \in\left(\eta\left(r_{2}\right), \frac{\pi}{m}\right)$ by monotonicity of $u$ to get the second inequality.

Hence it follows from (Equation 6.2.54) and (Equation 6.2.55) that

$$
\begin{equation*}
J_{1} \gtrsim \lambda \gtrsim \frac{m \delta\left(1-r_{2}\right)^{2}}{\|u\|_{L^{\infty}(\mathbb{T})}} \tag{6.2.56}
\end{equation*}
$$

Thus the desired result follows from (Equation 6.2.52), (Equation 6.2.53) and (Equation 6.2.56).

Now we prove Lemma 6.2.6. The proof is based on the formulae given in (B.1.3).
Proof of Lemma 6.2.6: Let us assume that $\delta<\frac{\pi}{4 m}$. We will prove (Equation 6.2.27) first. By monotonicity of $\eta$ (assumption (c)), it follows from Lemma B.1.3 that for all $r \in\left(r_{\text {min }}, r_{1}\right)$,

$$
\begin{aligned}
\partial_{r} \varphi_{m}(r, \eta(r)) \geq & \frac{1}{2 \pi} \int_{r_{\min }}^{r} \frac{\rho}{r} \arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right) d \rho \\
+ & \frac{1}{2 \pi} \int_{r}^{r_{\max }} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)-\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)-\eta(\rho)))}\right)\right. \\
& \left.-\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right)\right) d \rho
\end{aligned}
$$

where we used that $\sin (m(\eta(r)-\eta(\rho))) \leq 0$ for $\rho<r$ so we can drop one of the integrands for free. Note that the integrand in the second integral is positive for $\rho \in\left(r, r_{2}\right)$, since $\sin (m(\eta(r)-\eta(\rho)))>$ 0 and $\sin (m(\eta(r)+\eta(\rho)))<0$ for $r<\rho<r_{2}$, which follows from $\frac{\pi}{m}-\eta\left(r_{2}\right)<\frac{\pi}{4 m}$. We will use the second integrand in the second integral to cancel the first integral, that is, we have that

$$
\begin{aligned}
\partial_{r} \varphi_{m}(r, \eta(r)) \geq & \frac{1}{2 \pi}\left(\int_{r_{\text {min }}}^{r} \frac{\rho}{r} \arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right) d \rho\right. \\
& \left.-\int_{r}^{r_{2}} \frac{\rho}{r} \arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right) d \rho\right) \\
+ & \frac{1}{2 \pi} \int_{r_{2}}^{r_{\text {max }}} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)-\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)-\eta(\rho)))}\right)\right.
\end{aligned}
$$

$$
\begin{align*}
& \left.-\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right)\right) d \rho \\
= & : \frac{1}{2 \pi} K_{1}+\frac{1}{2 \pi} K_{2} . \tag{6.2.57}
\end{align*}
$$

To estimate $K_{1}$, we use that (note that $\frac{\pi}{m}-\eta(r)<\frac{\pi}{4 m}$ for $r \leq r_{2}$ )
and obtain

$$
K_{1} \geq \int_{r_{m i n}}^{r} \frac{\rho}{r} \arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (2 m \eta(r))}{1-\left(\frac{\rho}{r}\right)^{m}}\right) d \rho-\int_{r}^{r_{2}} \frac{\rho}{r} \arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (2 m \eta(r))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (2 m \delta)}\right) d \rho
$$

From the change of variables $\left(\frac{\rho}{r}\right)^{m} \mapsto x$ for the first integral and $\left(\frac{r}{\rho}\right)^{m} \mapsto x$ for the second integral, it follows that

$$
\begin{aligned}
K_{1} & \geq \frac{r}{m}\left(\int_{\left(\frac{r_{\text {min }}}{r}\right)^{m}}^{1} x^{-1+\frac{2}{m}} \arctan \left(\frac{x \sin (2 m \eta(r))}{1-x}\right) d x-\int_{\left(\frac{r}{r_{2}}\right)^{m}}^{1} x^{-1-\frac{2}{m}} \arctan \left(\frac{x \sin (2 m \eta(r))}{1-x \cos (2 m \delta)}\right) d x\right) \\
& \geq \frac{r}{m} \int_{\left(\frac{r}{r_{2}}\right)^{m}}^{1} x^{-1-\frac{2}{m}}\left(\arctan \left(\frac{x \sin (2 m \eta(r))}{1-x}\right)-\arctan \left(\frac{x \sin (2 m \eta(r))}{1-x \cos (2 m \delta)}\right)\right) d x .
\end{aligned}
$$

where we used $\frac{r_{\text {min }}}{r}>\frac{r}{r_{2}}$ for $r \in\left(r_{\text {min }}, r_{1}\right)$, which follows from $r_{1}^{2}<r_{\text {min }} r_{2}$, and $x^{-1-\frac{2}{m}}>$ $x^{-1+\frac{2}{m}}$ for $0<x<1$ to get the second inequality (note that the first integrand is negative). Therefore it follows from Lemma B.2.2 that (note that $\sin (2 m \eta(r))<0$ for $r<r_{1}$ ),

$$
\begin{equation*}
K_{1} \gtrsim-\frac{r}{m}(1-\cos (2 m \delta)) \gtrsim-\frac{r}{m} \sin (2 m \delta) \gtrsim-r \delta \tag{6.2.58}
\end{equation*}
$$

where the second inequality is due to $1-\cos (2 m \delta)<\sin (2 m \delta)$ for $\delta<\frac{\pi}{4 m}$.

Now let us estimate $K_{2}$. Note that the integrand in $K_{2}$ is non-negative if $\eta(\rho)>\frac{\pi}{m}-\eta(r)$. Indeed, by monotonicity of $\eta$, we have $\eta(r)>\eta(\rho)$ for any $\rho \geq r_{2}>r$, which implies the first integrand in $K_{2}$ is positive for all $\rho \in\left(r_{2}, r_{\max }\right)$. Thus, if we choose $r_{3}:=\eta^{-1}\left(\frac{\pi}{m}-\eta(r)\right)>r_{2}$ then the integrand in $K_{2}$ is strictly positive for $\rho \in\left(r_{2}, r_{3}\right)$. Hence we have

$$
\begin{aligned}
& K_{2} \geq \int_{r_{3}}^{r_{\max }} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)-\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)-\eta(\rho))}\right)\right. \\
&\left.-\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right)\right) d \rho
\end{aligned}
$$

Note that $\left|\partial_{\theta}\left(\arctan \left(\frac{x \sin (m \theta)}{1-x \cos (m \theta)}\right)\right)\right| \lesssim \frac{m x}{1-x}$ for all $0 \leq x<1$ and $\theta \in\left[-\frac{\pi}{m}, \frac{\pi}{m}\right]$. Indeed,

$$
\partial_{\theta}\left(\arctan \left(\frac{x \sin (m \theta)}{1-x \cos (m \theta)}\right)\right)=\frac{m x(\cos (m \theta)-1+(1-x))}{(x-1)^{2}+2 x(1-\cos (m \theta))}
$$

Hence either $(1-x) \leq 1-\cos (m \theta)$ or $(1-x)>1-\cos (m \theta)$, one can easily see that $\left|\frac{m x(\cos (m \theta)-1+(1-x))}{(x-1)^{2}+2 x(1-\cos (m \theta))}\right| \lesssim \frac{m x}{1-x}$. Since we also have $\frac{r}{\rho}<\frac{r_{1}}{r_{3}}<\frac{r_{1}}{r_{2}}$ for $\rho>r_{3}$, and it follows from the mean-value theorem that the integrand can be bounded as

$$
\begin{aligned}
\left\lvert\, \arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)-\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)-\eta(\rho))}\right)\right. & \left.=\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(r)+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(r)+\eta(\rho)))}\right) \right\rvert\, \\
& \lesssim \frac{m \eta(\rho)\left(\frac{r}{\rho}\right)^{m}}{1-\left(\frac{r}{\rho}\right)^{m}} \\
& <\frac{m \eta(\rho)\left(\frac{r}{\rho}\right)^{m}}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}
\end{aligned}
$$

Therefore we obtain

$$
\begin{aligned}
K_{2} \gtrsim-\frac{1}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \int_{r_{3}}^{r_{\max }}\left(\frac{r}{\rho}\right)^{m-1} m \eta(\rho) d \rho & \gtrsim-m \frac{\left(\frac{\pi}{m}-\eta(r)\right)}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \int_{r_{3}}^{r_{\max }}\left(\frac{r}{\rho}\right)^{m-1} d \rho \\
& \gtrsim-\frac{r m}{(m-2)} \frac{\left(\frac{\pi}{m}-\eta(r)\right)}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}
\end{aligned}
$$

where we used $\eta(\rho)<\frac{\pi}{m}-\eta(r)$ for $\rho>r_{3}$ to get the second inequality. Since $m \geq 3$ and $\frac{\pi}{m}-\eta(r)<\frac{\pi}{m}-\eta\left(r_{2}\right)=\delta$, the above inequality implies

$$
\begin{equation*}
K_{2} \gtrsim-\frac{r}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}}\left(\frac{\pi}{m}-\eta(r)\right) \gtrsim-\frac{r}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \delta . \tag{6.2.59}
\end{equation*}
$$

Thus, (Equation 6.2.57), (Equation 6.2.58) and (Equation 6.2.59) yield (Equation 6.2.27).

Now, let us prove (Equation 6.2.28). Since $\sin (m \eta(\rho))<\sin (m \delta)$ for all $\rho<r_{2}$, it follows from Lemma B.1.3 that

$$
\begin{aligned}
\partial_{\theta} \varphi_{m}(r, \eta(r)) & <\frac{1}{4 \pi} \underbrace{\int_{r_{\min }}^{r} \rho \log \left(1+\frac{4\left(\frac{\rho}{r}\right)^{m} \sin ^{2}(m \delta)}{\left(1-\left(\frac{\rho}{r}\right)^{m}\right)^{2}}\right) d \rho}_{=: K_{3}}+\frac{1}{4 \pi} \underbrace{\int_{r}^{r_{2}} \rho \log \left(1+\frac{4\left(\frac{r}{\rho}\right)^{m} \sin ^{2}(m \delta)}{\left(1-\left(\frac{r}{\rho}\right)^{m}\right)^{2}}\right) d \rho}_{=: K_{4}} \\
& +\frac{1}{4 \pi} \underbrace{\int_{r_{2}}^{r_{\max }} \rho \log \left(1+\frac{4\left(\frac{r}{\rho}\right)^{m} \sin (m \eta(\rho)) \sin (m \delta)}{\left(1-\left(\frac{r}{\rho}\right)^{m}\right)^{2}}\right) d \rho}_{=: K_{5}} .
\end{aligned}
$$

Again, the change of variables, $\left(\frac{\rho}{r}\right)^{m} \mapsto x$ and $\left(\frac{r}{\rho}\right)^{m} \mapsto x$ yields that

$$
\begin{aligned}
& K_{3} \lesssim \frac{r^{2}}{m} \int_{0}^{1} x^{-1+\frac{2}{m}} \log \left(1+\frac{4 x \sin ^{2}(m \delta)}{(1-x)^{2}}\right) d x \\
& K_{4} \lesssim \frac{r^{2}}{m} \int_{0}^{1} x^{-1-\frac{2}{m}} \log \left(1+\frac{4 x \sin ^{2}(m \delta)}{(1-x)^{2}}\right) d x
\end{aligned}
$$

Thus it follows from Lemma B.2.3 that

$$
\begin{equation*}
K_{3}, K_{4} \lesssim \frac{r^{2}}{m}|\sin (m \delta)|<r^{2} \delta . \tag{6.2.60}
\end{equation*}
$$

To estimate $K_{5}$, recall that $r<r_{1}$ and $\log (1+x) \leq x$, which yields that

$$
\begin{align*}
K_{5} & \lesssim \int_{r_{2}}^{r_{\max }} \rho\left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m \eta(\rho)) \sin (m \delta)}{\left(1-\left(\frac{r}{\rho}\right)^{m}\right)^{2}}\right) d \rho \\
& =\frac{r^{2}}{m} \int_{\left(\frac{r}{r_{m a x}}\right)^{m}}^{\left(\frac{r}{r_{2}}\right)^{m}} x^{-1-\frac{2}{m}}\left(\frac{x \sin (m \eta(\rho)) \sin (m \delta))}{(1-x)^{2}}\right) d x \\
& \lesssim \frac{r^{2}}{m} \int_{0}^{\left(\frac{r_{1}}{r_{2}}\right)^{m}} x^{-\frac{2}{m}} \frac{\sin (m \delta)}{(1-x)^{2}} d x \\
& =\frac{r^{2}}{m}\left(\int_{0}^{\min \left\{\frac{1}{2},\left(\frac{r_{1}}{r_{2}}\right)^{m}\right\}} x^{-\frac{2}{m}} \frac{\sin (m \delta)}{(1-x)^{2}} d x+\int_{\min \left\{\frac{1}{2},\left(\frac{r_{1}}{r_{2}}\right)^{m}\right\}}^{\left(\frac{r_{1}}{r_{2}}\right)^{m}} x^{-\frac{2}{m}} \frac{\sin (m \delta)}{(1-x)^{2}} d x\right) \\
& \lesssim \frac{r^{2}}{m}\left(m \delta+\frac{\sin (m \delta)}{\left(1-\left(\frac{r_{1}}{r_{2}}\right)^{m}\right)}\right) \\
& \lesssim \frac{\delta}{1-\left(\frac{r_{1}}{r_{2}}\right)^{m}} \tag{6.2.61}
\end{align*}
$$

Therefore (Equation 6.2.60) and (Equation 6.2.61) yield (Equation 6.2.28). This finishes the proof.

Appendices

## APPENDIX A

## FUNTIONAL DERIVATIVES AND SOME BASIC INTEGRALS

## A. 1 Derivatives of the Functional

## A.1.1 Functional derivatives

Recall that $\mathcal{F}(b, g, r)=\left(\mathcal{F}_{1}, \mathcal{F}_{2}\right)$ is given in (Equation 5.1.4). For simplicity, we denote

$$
\begin{aligned}
& A_{1}:=(b+g(\eta)) \\
& A_{2}:=\left(r^{\prime}(\theta) \cos (\theta-\eta)-(1+r(\theta)) \sin (\theta-\eta)\right)(1+r(\eta))-(1+r(\theta)) r^{\prime}(\theta), \\
& A_{3}:=\frac{1}{(1+r(\theta))^{2}+(1+r(\eta))^{2}-2(1+r(\theta))(1+r(\eta)) \cos (\theta-\eta)} \\
& A_{4}:=r^{\prime}(\theta)(1+r(\theta)) \\
& A_{5}:=(b+g(\eta))(b+g(\theta)), \\
& A_{6}:=(1+r(\theta))^{2}-\left(r^{\prime}(\theta) \sin (\theta-\eta)+(1+r(\theta)) \cos (\theta-\eta)\right)(1+r(\eta)), \\
& A_{7}:=\frac{1}{r^{\prime}(\theta)^{2}+(1+r(\theta))^{2}}, \\
& A_{8}:=\frac{(1+r(\theta))^{2}(b+g(\theta))}{r^{\prime}(\theta)^{2}+(1+r(\theta))^{2}} .
\end{aligned}
$$

We also denote the average integral by $f f(\theta) d \theta:=\frac{1}{2 \pi} \int_{-\pi}^{\pi} f(\theta) d \theta$. Therefore the functional $\mathcal{F}$ can be written as

$$
\begin{align*}
& \mathcal{F}_{1}=f A_{1} A_{2} A_{3} d \eta+\Omega A_{4}  \tag{A.1.1}\\
& \tilde{\mathcal{F}}_{2}=f A_{5} A_{6} A_{7} A_{3} d \eta-\Omega A_{8} \tag{A.1.2}
\end{align*}
$$

We will expand $A_{i}(g, \theta, \eta)$ and $A_{i}(r, \theta, \eta)$ up to quadratic/cubic order in $g$ and $r$.

Lemma A.1.1. Let $A_{i}^{\prime} s$ be as above. We have

$$
\begin{aligned}
A_{1} & =b+g(\eta) \\
A_{2} & =-\sin (\theta-\eta)+\left[r^{\prime}(\theta)(\cos (\theta-\eta)-1)-(r(\theta)+r(\eta)) \sin (\theta-\eta)\right] \\
& +\left[r^{\prime}(\theta) r(\eta)(\cos (\theta-\eta)-1)-r^{\prime}(\theta)(r(\theta)-r(\eta))-r(\theta) r(\eta) \sin (\theta-\eta)\right] \\
A_{3} & =\frac{1}{2-2 \cos (\theta-\eta)}-\frac{r(\theta)+r(\eta)}{2-2 \cos (\theta-\eta)} \\
& +\frac{1}{2-2 \cos (\theta-\eta)}\left[r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}-\frac{(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)}\right] \\
& +\frac{1}{2-2 \cos (\theta-\eta)}\left[\frac{(r(\theta)+r(\eta))(r(\theta)-r(\eta))^{2}}{1-\cos (\theta-\eta)}-\left[r(\theta)^{3}+r(\theta)^{2} r(\eta)+r(\theta) r(\eta)^{2}+r(\eta)^{3}\right]\right]+O\left(r^{4}\right), \\
A_{4} & =r^{\prime}(\theta)+r(\theta) r^{\prime}(\theta), \\
A_{5} & =b^{2}+b(g(\theta)+g(\eta))+g(\theta) g(\eta), \\
A_{6} & =(1-\cos (\theta-\eta))+\left[(r(\theta)+r(\eta))(1-\cos (\theta-\eta))+(r(\theta)-r(\eta))-r^{\prime}(\theta) \sin (\theta-\eta)\right] \\
& +\left[r(\theta)(r(\theta)-r(\eta))+r(\theta) r(\eta)(1-\cos (\theta-\eta))-r^{\prime}(\theta) r(\eta) \sin (\theta-\eta)\right] \\
A_{7} & =1-2 r(\theta)+\left[3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right]+\left[4 r(\theta) r^{\prime}(\theta)^{2}-4 r(\theta)^{3}\right]+O\left(r^{4}\right), \\
A_{8} & =b+g(\theta)-b r^{\prime}(\theta)^{2}+\left[2 b r(\theta) r^{\prime}(\theta)^{2}-g(\theta) r^{\prime}(\theta)^{2}\right]+O\left(r^{4}+g^{4}\right) .
\end{aligned}
$$

Proof. Straightforward.

## Linear parts

We denote by $A_{i}^{j}$ the $j$ th order term in $A_{i}$. For example, $A_{3}^{1}=-\frac{r(\theta)+r(\eta)}{2-2 \cos (\theta-\eta)}$.
Lemma A.1.2. Let $\mathcal{F}_{i}$ 's and $A_{i}$ 's be defined as before. Then

$$
\begin{align*}
\left.\frac{d}{d t} \mathcal{F}_{1}(b, t g, t r)\right|_{t=0} & =-f \frac{g(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta),  \tag{A.1.3}\\
\left.\frac{d}{d t} \tilde{\mathcal{F}}_{2}(b, t g, t r)\right|_{t=0} & =\left(\frac{b}{2}-\Omega\right) g(\theta)+b^{2}\left(f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)} d \eta-r(\theta)\right) . \tag{A.1.4}
\end{align*}
$$

Proof. We compute $\mathcal{F}_{1}$ first. In view of (Equation A.1.1), we collect the linear terms in $A_{1} A_{2} A_{3}+$ $\Omega A_{4}$ from Lemma A.1.1. Hence we have

$$
\begin{aligned}
\mathcal{F}_{1}(b, t g, t r) & =f-\frac{t g(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}+b t \frac{r^{\prime}(\theta)(\cos (\theta-\eta)-1)-(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} \\
& +\frac{b t \sin (\theta-\eta)(r(\theta)+r(\eta))}{2-2 \cos (\theta-\eta)} d \eta+t \Omega r^{\prime}(\theta)+O\left(t^{2}\right) \\
& =-f \frac{t g(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+t\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta)+O\left(t^{2}\right) .
\end{aligned}
$$

Thus we obtain (Equation A.1.3) by differentiating with respect to $t$.
In order to compute the derivative of $\tilde{\mathcal{F}}_{2}$, we collect the linear terms in (Equation A.1.2) from Lemma A.1.1 and obtain

$$
\begin{aligned}
\tilde{\mathcal{F}}_{2}(b, t g, t r)= & f \frac{b t(g(\theta)+g(\eta))}{2} \\
& +t b^{2}\left[\frac{(r(\theta)+r(\eta))}{2}+\frac{(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}-\frac{r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
& -t b^{2} r(\theta)-\frac{t b^{2}}{2}(r(\theta)+r(\eta))-t \Omega g(\theta) d \eta+O\left(t^{2}\right) \\
= & t\left(\frac{b}{2}-\Omega\right) g(\theta)+t b^{2}\left(f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)} d \eta-r(\theta)\right)+O\left(t^{2}\right),
\end{aligned}
$$

where we used $f g(\eta) d \eta=f r(\eta) d \eta=0$ and $f \frac{\sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=0$. By differentiating in $t$, we obtain the desired result (Equation A.1.4).

## Quadratic parts

Now we compute the quadratic expansion of $\mathcal{F}_{1}$ and $\tilde{\mathcal{F}}_{2}$.

Lemma A.1.3. Let $\mathcal{F}_{i}$ 's and $A_{i}$ 's be defined as before. Then

$$
\begin{align*}
\mathcal{F}_{1}= & \left(\frac{b}{2}+\Omega\right) r(\theta) r^{\prime}(\theta)-b f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta+b f \frac{(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& + \text { linear terms }+O\left(r^{3}+g^{3}\right)  \tag{A.1.5}\\
\tilde{\mathcal{F}}_{2}= & \frac{3 b^{2}}{2} r(\theta)^{2}-b g(\theta) r(\theta)+b\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta)^{2} \\
& -\frac{b^{2}}{2} f \frac{(r(\theta)-r(\eta))(5 r(\theta)+r(\eta))}{2-2 \cos (\theta-\eta)} d \eta+b f \frac{(g(\theta)+g(\eta))(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& -b f \frac{g(\eta) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+\text { linear terms }+O\left(r^{3}+g^{3}\right) \tag{A.1.6}
\end{align*}
$$

Proof. We compute $\mathcal{F}_{1}$ first. By collecting quadratic terms in (Equation A.1.1) from Lemma A.1.1, we have (we will have, for example, $A_{1}^{2} A_{2}^{0} A_{3}^{0}+A_{1}^{0} A_{2}^{2} A_{3}^{0}+A_{1}^{0} A_{2}^{0} A_{3}^{2}+A_{1}^{1} A_{2}^{1} A_{3}^{0}+$ $\left.A_{1}^{1} A_{2}^{0} A_{3}^{1}+A_{1}^{0} A_{2}^{1} A_{3}^{1}+\Omega A_{4}^{2}\right)$.

$$
\begin{aligned}
\mathcal{F}_{1}= & f b\left[-\frac{1}{2} r^{\prime}(\theta) r(\eta)-\frac{r^{\prime}(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}-\frac{r(\theta) r(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
& -\frac{b \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\left[r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}-\frac{(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)}\right] \\
& +\left[-\frac{g(\eta) r^{\prime}(\theta)}{2}-\frac{g(\eta)(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right]+\frac{g(\eta)(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} \\
& +\left[\frac{b r^{\prime}(\theta)(r(\theta)+r(\eta))}{2}+\frac{b(r(\theta)+r(\eta))^{2} \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right]+\Omega r(\theta) r^{\prime}(\theta) d \eta+O\left(r^{3}+g^{3}\right) \\
= & f\left(\frac{b}{2}+\Omega\right) r(\theta) r^{\prime}(\theta)-\frac{b r^{\prime}(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}+\frac{b(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& + \text { linear terms }+O\left(r^{3}+g^{3}\right),
\end{aligned}
$$

which yields (Equation A.1.5). Now we will expand $\tilde{\mathcal{F}}_{2}$ up to the quadratic order. By collecting all quadratic terms in (Equation A.1.2) from Lemma A.1.1, we obtain (we will have $A_{5}^{2} A_{6}^{0} A_{7}^{0} A_{3}^{0}+$ $A_{5}^{0} A_{6}^{2} A_{7}^{0} A_{3}^{0}+A_{5}^{0} A_{6}^{0} A_{7}^{2} A_{3}^{0}+A_{5}^{0} A_{6}^{0} A_{7}^{0} A_{3}^{2}+A_{5}^{1} A_{6}^{1} A_{7}^{0} A_{3}^{0}+A_{5}^{1} A_{6}^{0} A_{7}^{1} A_{3}^{0}+A_{5}^{1} A_{6}^{0} A_{7}^{0} A_{3}^{1}+A_{5}^{0} A_{6}^{1} A_{7}^{1} A_{3}^{0}+$

$$
\begin{aligned}
&\left.A_{5}^{0} A_{6}^{1} A_{7}^{0} A_{3}^{1}+A_{5}^{0} A_{6}^{0} A_{7}^{1} A_{3}^{1}-\Omega A_{8}^{2}\right) \\
& \tilde{\mathcal{F}}_{2}= f \frac{g(\theta) g(\eta)}{2}+\left[b^{2} \frac{r(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}+b^{2} \frac{r(\theta) r(\eta)}{2}-b^{2} \frac{r^{\prime}(\theta) r(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
&+b^{2} \frac{\left(3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right)}{2}+\left[\frac{b^{2}\left(r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}\right)}{2}-\frac{b^{2}}{2} \frac{(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)}\right] \\
&+\left[b \frac{(g(\theta)+g(\eta))(r(\theta)+r(\eta))}{2}+b \frac{(g(\theta)+g(\eta))(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}-b \frac{(g(\theta)+g(\eta)) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
&-b((g(\theta)+g(\eta))) r(\theta)-b \frac{(g(\theta)+g(\eta))(r(\theta)+r(\eta))}{2} \\
&+\left[-b^{2} r(\theta)(r(\theta)+r(\eta))-2 b^{2} \frac{r(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}+2 b^{2} \frac{r(\theta) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
&+\left[-b^{2} \frac{(r(\theta)+r(\eta))^{2}}{2}-b^{2} \frac{\left(r(\theta)^{2}-r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)}+b^{2} \frac{r^{\prime}(\theta)(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)}\right] \\
&+b^{2} r(\theta)(r(\theta)+r(\eta))+b \Omega r^{\prime}(\theta)^{2} d \eta+O\left(r^{3}+g^{3}\right) \\
&= f b^{2} \frac{\left(3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right)}{2}-b g(\theta) r(\theta)+b \Omega r^{\prime}(\theta)^{2} \\
&+\frac{1}{2-2 \cos (\theta-\eta)}\left(-\frac{b^{2}}{2}(r(\theta)-r(\eta))(5 r(\theta)+r(\eta))+b(g(\theta)+g(\eta))(r(\theta)-r(\eta))\right) \\
&-\frac{b g(\eta) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+\operatorname{linear} \operatorname{terms}+O\left(r^{3}+g^{3}\right)
\end{aligned}
$$

where we used $f g(\eta) d \eta=f r(\eta) d \eta=f \frac{\sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=0$. This yields the desired result (Equation A.1.6).

Lemma A.1.4. Let $\mathcal{F}_{i}$ 's and $A_{i}$ 's be defined as before. Then

$$
\begin{align*}
\frac{d^{2}}{d t d s} \mathcal{F}_{1}(b, t g+s \tilde{g}, t r & +s \tilde{r})\left.\right|_{t=s=0}=\left(\frac{b}{2}+\Omega\right)\left(r(\theta) \tilde{r}^{\prime}(\theta)+\tilde{r}(\theta) r^{\prime}(\theta)\right) \\
& -b f \frac{r^{\prime}(\theta)(\tilde{r}(\theta)-\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta-b f \frac{\tilde{r}^{\prime}(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& +2 b f \frac{(r(\theta)-r(\eta))(\tilde{r}(\theta)-\tilde{r}(\eta)) \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} \tag{A.1.7}
\end{align*}
$$

and

$$
\begin{align*}
\frac{d^{2}}{d t d s} \tilde{\mathcal{F}}_{2}(b, t g+s \tilde{g}, t r & +s \tilde{r})\left.\right|_{t=s=0}=3 b^{2} r(\theta) \tilde{r}(\theta)-b(g(\theta) \tilde{r}(\theta)+\tilde{g}(\theta) r(\theta))+2 b\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta) \tilde{r}^{\prime}(\theta) \\
& -\frac{b^{2}}{2} f \frac{(r(\theta)-r(\eta))(5 \tilde{r}(\theta)+\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta-\frac{b^{2}}{2} f \frac{(\tilde{r}(\theta)-\tilde{r}(\eta))(5 r(\theta)+r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& +b f \frac{(g(\theta)+g(\eta))(\tilde{r}(\theta)-\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta+b f \frac{(\tilde{g}(\theta)+\tilde{g}(\eta))(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& -b f \frac{g(\eta) \tilde{r}^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta-b f \frac{\tilde{g}(\eta) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta \tag{A.1.8}
\end{align*}
$$

Proof. We compute $\mathcal{F}_{1}$ first. From (Equation A.1.5) in Lemma A.1.3, we collect all st terms and obtain

$$
\begin{aligned}
\mathcal{F}_{1}(b, t g+s \tilde{g}, t r+s \tilde{r})= & s t\left[\left(\frac{b}{2}+\Omega\right)\left(r(\theta) \tilde{r}^{\prime}(\theta)+\tilde{r}(\theta) r^{\prime}(\theta)\right)\right. \\
& -b f \frac{r^{\prime}(\theta)(\tilde{r}(\theta)-\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta-b f \frac{\tilde{r}^{\prime}(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& \left.+2 b f \frac{(r(\theta)-r(\eta))(\tilde{r}(\theta)-\tilde{r}(\eta))}{(2-2 \cos (\theta-\eta))^{2}}\right]+ \text { linear terms }+O\left(t^{2}+s^{2}\right)
\end{aligned}
$$

Once we differentiate the above equation with respect to $t$ and $s$, the desired result (Equation A.1.7) follows immediately. Similarly, we collect all st terms from (Equation A.1.6) and obtain

$$
\begin{aligned}
\tilde{\mathcal{F}}_{2}(b, t g+s \tilde{g}, t r+s \tilde{r})= & s t\left[3 b^{2} r(\theta) \tilde{r}(\theta)-b(g(\theta) \tilde{r}(\theta)+\tilde{g}(\theta) r(\theta))+2 b\left(\Omega-\frac{b}{2}\right) r^{\prime}(\theta) \tilde{r}^{\prime}(\theta)\right. \\
& -\frac{b^{2}}{2} f \frac{(r(\theta)-r(\eta))(5 \tilde{r}(\theta)+\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta-\frac{b^{2}}{2} f \frac{(\tilde{r}(\theta)-\tilde{r}(\eta))(5 r(\theta)+r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& +b f \frac{(g(\theta)+g(\eta))(\tilde{r}(\theta)-\tilde{r}(\eta))}{2-2 \cos (\theta-\eta)} d \eta+b f \frac{(\tilde{g}(\theta)+\tilde{g}(\eta))(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& \left.-b f \frac{g(\eta) \tilde{r}^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta-b f \frac{\tilde{g}(\eta) r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right] \\
& + \text { linear terms }+O\left(t^{2}+s^{2}\right) .
\end{aligned}
$$

Once we differentiate the above equation with respect to $t$ and $s$, the desired result (Equation A.1.8)
follows immediately.

## Cubic parts

We will expand $\tilde{\mathcal{F}}_{2}$ up to cubic order with respect to the $r$ variable (we fix $g=0$ ). We denote $B:=A_{3} A_{6}$ so that (Equation A.1.2) can be written as (with $g=0$ )

$$
\begin{equation*}
\tilde{\mathcal{F}}_{2}=b^{2} f B d \eta A_{7}-\Omega A_{8} \tag{A.1.9}
\end{equation*}
$$

We will first expand $B$ up to cubic order.

Lemma A.1.5. Let $\tilde{\mathcal{F}}_{2}, A_{i}$ and $B$ be as defined as before. Then

$$
\begin{aligned}
f B d \eta= & \frac{1}{2}+f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)} d \eta-\frac{1}{2} f \frac{r(\theta)^{2}-r(\eta)^{2}}{2-2 \cos (\theta-\eta)} d \eta \\
+ & {\left[\frac{1}{2} f \frac{(r(\theta)-r(\eta))\left(r(\theta)^{2}+r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta\right.} \\
& \left.+f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta\right]+O\left(r^{4}\right)
\end{aligned}
$$

Proof. Using (A.1.1), we will compute the constant $\left(=: B^{0}\right)$, linear $\left(=: B^{1}\right)$, quadratic $\left(=: B^{2}\right)$ and cubic ( $=: B^{3}$ ) terms of $B=A_{3} A_{6}$ separately. It is straightforward that

$$
\begin{equation*}
f B^{0} d \eta=\frac{1}{2} \tag{A.1.10}
\end{equation*}
$$

For $B^{1}$, we compute $B^{1}\left(=A_{3}^{1} A_{6}^{0}+A_{3}^{0} A_{6}^{1}\right)$,

$$
\begin{equation*}
f B^{1} d \eta=f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)}-\frac{r^{\prime}(\theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=f \frac{r(\theta)-r(\eta)}{2-2 \cos (\theta-\eta)} d \eta \tag{A.1.11}
\end{equation*}
$$

where we used $f \frac{\sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=0$. For $B^{2}$, we compute $A_{3}^{2} A_{6}^{0}+A_{3}^{0} A_{6}^{2}+A_{3}^{1} A_{6}^{1}$, hence

$$
\begin{align*}
f B^{2} d \eta= & {\left[\frac{r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}}{2}-\frac{1}{2} f \frac{(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)} d \eta\right] } \\
& +\left[f \frac{r(\theta)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)}+\frac{r(\theta) r(\eta)}{2}-\frac{r^{\prime}(\theta) r(\eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right] \\
& +\left[f-\frac{1}{2}(r(\theta)+r(\eta))^{2}-\frac{r(\theta)^{2}-r(\eta)^{2}}{2-2 \cos (\theta-\eta)}+\frac{r^{\prime}(\theta)(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right] \\
= & -\frac{1}{2} f \frac{\left(r(\theta)^{2}-r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta \tag{A.1.12}
\end{align*}
$$

where we used $f \frac{\sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=0$. For $B^{3}$, we compute $B^{3}\left(=A_{3}^{3} A_{6}^{0}+A_{3}^{2} A_{6}^{1}+A_{3}^{1} A_{6}^{2}+A_{3}^{0} A_{6}^{3}\right)$,

$$
\begin{align*}
f B^{3} d \eta= & {\left[f \frac{(r(\theta)+r(\eta))(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)} d \eta-\frac{1}{2} f r(\theta)^{3}+r(\theta)^{2} r(\eta)+r(\theta) r(\eta)^{2}+r(\eta)^{3} d \eta\right] } \\
& +\left[f \frac{(r(\theta)+r(\eta))\left(r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}\right)}{2} d \eta-\frac{1}{2} f \frac{(r(\theta)+r(\eta))(r(\theta)-r(\eta))^{2}}{2-2 \cos (\theta-\eta)} d \eta\right. \\
& +f \frac{r(\theta)^{3}-r(\eta)^{3}}{2-2 \cos (\theta-\eta)} d \eta-f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& \left.-f \frac{r^{\prime}(\theta)\left(r(\theta)^{2}+r(\theta) r(\eta)+r(\eta)^{2}\right) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta+f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta\right] \\
& +\left[-f \frac{r(\theta)\left(r(\theta)^{2}-r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-f \frac{r(\theta) r(\eta)(r(\theta)+r(\eta))}{2} d \eta\right. \\
& \left.+f \frac{r^{\prime}(\theta) r(\eta)(r(\theta)+r(\eta)) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right] \\
= & \frac{1}{2} f \frac{(r(\theta)-r(\eta))\left(r(\theta)^{2}+r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& +f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} . \tag{A.1.13}
\end{align*}
$$

Thus the desired result follows from (Equation A.1.10), (Equation A.1.11), (Equation A.1.12) and (Equation A.1.13).

Lemma A.1.6. Let $\tilde{\mathcal{F}}_{2}, A_{i}$ 's and $B$ be as defined as before. Then

$$
\begin{aligned}
\left.\frac{1}{6} \frac{d^{3}}{d t^{3}} \tilde{\mathcal{F}}_{2}(b, 0, t r)\right|_{t=0}= & \frac{b^{2}}{2} f \frac{(r(\theta)-r(\eta))\left(3 r(\theta)^{2}+2 r(\theta) r(\eta)+r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-b^{2} f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& +b^{2} f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} \eta+b^{2} f \frac{\left(3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& +2 b(b-\Omega) r(\theta) r^{\prime}(\theta)^{2}-2 b^{2} r(\theta)^{3}
\end{aligned}
$$

Proof. We first collect all cubic terms of $\tilde{\mathcal{F}}_{2}(b, 0, r)$ in $r$. From (Equation A.1.9), we have that the cubic terms consist of $b^{2}\left(B^{3} A_{7}^{0}+B^{2} A_{7}^{1}+B^{1} A_{7}^{2}+B^{0} A_{7}^{3}\right)-\Omega A_{8}^{3}$. Using Lemma A.1.1 and A.1.5 and the fact that $A_{7}$ does not depend on $\eta$, we obtain

$$
\begin{aligned}
\tilde{\mathcal{F}}_{2}(b, 0, r)= & b^{2}\left[\frac{1}{2} f \frac{(r(\theta)-r(\eta))\left(r(\theta)^{2}+r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta\right. \\
& \left.+f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta\right]+b^{2} f \frac{r(\theta)\left(r(\theta)^{2}-r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta \\
& +b^{2} f \frac{\left(3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta+b^{2}\left[2 r(\theta) r^{\prime}(\theta)^{2}-2 r(\theta)^{3}\right] \\
& -2 b \Omega r(\theta) r^{\prime}(\theta)^{2}+\text { lower order terms }+O\left(r^{4}\right) \\
= & \frac{b^{2}}{2} f \frac{(r(\theta)-r(\eta))\left(3 r(\theta)^{2}+2 r(\theta) r(\eta)+r(\eta)^{2}\right)}{2-2 \cos (\theta-\eta)} d \eta-b^{2} f \frac{(r(\theta)-r(\eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& +b^{2} f \frac{r^{\prime}(\theta)(r(\theta)-r(\eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} \eta+b^{2} f \frac{\left(3 r(\theta)^{2}-r^{\prime}(\theta)^{2}\right)(r(\theta)-r(\eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& +2 b(b-\Omega) r(\theta) r^{\prime}(\theta)^{2}-2 b^{2} r(\theta)^{3}+\text { lower order terms }+O\left(r^{4}\right)
\end{aligned}
$$

Therefore, the desired result follows immediately.

## A.1.2 Derivatives of the reduced functional

We denote $v:=(0, \cos (2 \theta))$. Given a pair of functions $(g, r)$, we denote $Q$ be the projection to the second mode of $r$, that is, $Q(g, r):=\left(\frac{1}{\pi} \int r(\theta) \cos (2 \theta) d \theta\right) \cos (2 \theta)$.

Lemma A.1.7. Let $\mathcal{F}, v, Q$ be defined as before. We fix $b=2$ and $\Omega=1$. Then,

$$
\begin{align*}
& \partial_{b} D \mathcal{F}(2,0) v=(\sin (2 \theta), 0),  \tag{A.1.14}\\
& \left.\frac{1}{2} \frac{d^{2}}{d t^{2}} \mathcal{F}(2, t v)\right|_{t=0}=(-2 \sin (4 \theta),-3 \cos (4 \theta)),  \tag{A.1.15}\\
& \left.\partial_{b} Q \frac{d^{2}}{d t^{2}} \mathcal{F}(b, t v)\right|_{b=2, t=0}=(0,0),  \tag{A.1.16}\\
& \tilde{v}:=-[D \mathcal{F}(2,0)]^{-1}(I-Q) \partial_{b} D \mathcal{F}(2,0) v=(2 \cos (2 \theta), 0),  \tag{A.1.17}\\
& \hat{v}:=-\left.[D \mathcal{F}(2,0)]^{-1} \frac{d^{2}}{d t^{2}}[(I-Q) \mathcal{F}(2, t v)]\right|_{t=0}=\left(-8 \cos (4 \theta), \frac{3}{2} \cos (4 \theta)\right),  \tag{A.1.18}\\
& \left.\frac{d^{2}}{d t d s} Q \mathcal{F}(2, t v+s \hat{v})\right|_{t=s=0}=(0,-12 \cos (2 \theta)),  \tag{A.1.19}\\
& \left.\frac{1}{3} \frac{d^{3}}{d t^{3}} Q \mathcal{F}(2, t v)\right|_{t=0}=(0,4 \cos (2 \theta)),  \tag{A.1.20}\\
& Q \frac{d^{2}}{d t d s} \mathcal{F}(2, t v+s \tilde{v})=(0,0),  \tag{A.1.21}\\
& \frac{1}{2} Q \partial_{b} D \mathcal{F}(2,0) \hat{v}=(0,0),  \tag{A.1.22}\\
& 2 Q \partial_{b} D \mathcal{F}(2,0) \tilde{v}=(0,2 \cos (2 \theta)) . \tag{A.1.23}
\end{align*}
$$

Proof. To prove (Equation A.1.14), it follows from Lemma A.1.2 that

$$
\partial_{b} D \mathcal{F}(b, 0) v=\left(\sin (2 \theta), 2 b\left(f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta-\cos (2 \theta)\right)\right)=(\sin (2 \theta), 0)
$$

where the last equality follows from (Equation A.1.30).
To prove (Equation A.1.15), note that $\left.\frac{1}{2} \frac{d^{2}}{d t^{2}} \mathcal{F}(2, t v)\right|_{t=0}=$

$$
\begin{aligned}
&\left(\left.\frac{1}{2} \frac{d^{2}}{d t^{2}} \mathcal{F}_{1}(2, t v)\right|_{t=0},\left.\frac{1}{2}\left(I-P_{0}\right) \frac{d^{2}}{d t^{2}} \tilde{\mathcal{F}}_{2}(2, t v)\right|_{t=0}\right) . \text { Hence it follows from Lemma A.1.3 that } \\
&\left.\frac{1}{2} \frac{d^{2}}{d t^{2}} \mathcal{F}_{1}(2, t v)\right|_{t=0}=-4 \cos (2 \theta) \sin (2 \theta)+4 \sin (2 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta \\
&+2 f \frac{(\cos (2 \theta)-\cos (2 \eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
&=-4 \cos (2 \theta) \sin (2 \theta)+4 \sin (2 \theta) \cos (2 \theta)-2 \sin (4 \theta) \\
&=-2 \sin (4 \theta)
\end{aligned}
$$

where the second equality follows from (Equation A.1.30) and Lemma A.1.11. Also, Lemma A.1.3 gives

$$
\begin{aligned}
\frac{1}{2}\left(I-P_{0}\right) \frac{d^{2}}{d t^{2}} & \left.\tilde{\mathcal{F}}_{2}(2, t v)\right|_{t=0} \\
& =\left(I-P_{0}\right)\left[6 \cos ^{2}(2 \theta)\right]-2\left(I-P_{0}\right) f \frac{(\cos (2 \theta)-\cos (2 \eta))(5 \cos (2 \theta)+\cos (2 \eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& =-4\left(I-P_{0}\right) \cos ^{2}(2 \theta)-2\left(I-P_{0}\right) f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta \\
& =-3 \cos (4 \theta)
\end{aligned}
$$

where the second equality follows from (Equation A.1.30) and the last equality follows from Lemma A.1.9. Therefore we obtain (Equation A.1.15).

To prove (Equation A.1.16), we can repeat the above computation and find that $\frac{d^{2}}{d t^{2}} \tilde{\mathcal{F}}_{2}(b, t v) \in$ span $\{\cos (4 \theta)\}$, independently of $b$. By projecting it to the space of the second mode, we obtain (Equation A.1.16).

To prove (Equation A.1.17), note that $(I-Q) \partial_{b} D \mathcal{F}(2,0) v=(\sin (2 \theta), 0)$, which follows from (Equation A.1.14). Also, it follows from Lemma A.1.2 and (Equation A.1.29) that

$$
D \mathcal{F}(2,0)(-2 \cos (2 \theta), 0)=(\sin (2 \theta), 0)=(I-Q) \partial_{b} D \mathcal{F}(2,0) v
$$

This immediately implies (Equation A.1.17).
To prove (Equation A.1.18), we use Lemma A.1.2 and (Equation A.1.30) and (Equation A.1.29) to obtain

$$
D \mathcal{F}(2,0)\left(8 \cos (4 \theta),-\frac{3}{2} \cos (4 \theta)\right)=(-4 \sin (4 \theta),-6 \cos (4 \theta))=\left.\frac{d^{2}}{d t^{2}} \mathcal{F}(2, t v)\right|_{t=0},
$$

where the last equality follows from (Equation A.1.15). Therefore we obtain

$$
-\left.[D \mathcal{F}(2,0)]^{-1} \frac{d^{2}}{d t^{2}}[(I-Q) \mathcal{F}(2, t v)]\right|_{t=0}=\left(-8 \cos (4 \theta), \frac{3}{2} \cos (4 \theta)\right)
$$

which proves (Equation A.1.18).
To prove (Equation A.1.19), note that $\frac{d^{2}}{d t d s} Q \mathcal{F}=\left(0, P_{2} \frac{d^{2}}{d t d s} \tilde{\mathcal{F}}_{2}\right)$. Therefore, it follows from (Equation A.1.18) and (Equation A.1.8) in Lemma A.1.4 that (plugging $g=0, \tilde{g}=$ $-8 \cos (4 \theta), r=\cos (2 \theta)$, and $\left.\tilde{r}=\frac{3}{2} \cos (4 \theta)\right)$

$$
\begin{aligned}
\frac{d^{2}}{d t d s} P_{2} \mathcal{F}_{2}(2, t v+s \tilde{v})= & P_{2}[(18 \cos (2 \theta) \cos (4 \theta)+16 \cos (2 \theta) \cos (4 \theta)) \\
& -2 f \frac{(\cos (2 \theta)-\cos (2 \eta))\left(\frac{15}{2} \cos (4 \theta)+\frac{3}{2} \cos (4 \eta)\right)}{2-2 \cos (\theta-\eta)} d \eta \\
& -3 f \frac{(\cos (4 \theta)-\cos (4 \eta))(5 \cos (2 \theta)+\cos (2 \eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& -16 f \frac{(\cos (4 \theta)+\cos (4 \eta))(\cos (2 \theta)-\cos (2 \eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& \left.-32 f \frac{\cos (4 \eta) \sin (2 \theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right] \\
= & P_{2} K_{1}+P_{2} K_{2}+P_{2} K_{3}+P_{2} K_{4}+P_{2} K_{5} .
\end{aligned}
$$

For $K_{1}$, we compute

$$
\begin{equation*}
P_{2} K_{1}=P_{2}(34 \cos (2 \theta) \cos (4 \theta))=17 P_{2}(\cos (2 \theta)+\cos (6 \theta))=17 \cos (2 \theta) \tag{A.1.24}
\end{equation*}
$$

For $K_{2}$ we compute

$$
\begin{align*}
P_{2} K_{2} & =P_{2}\left(-15 \cos (4 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta-3 f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (4 \eta)}{2-2 \cos (\theta-\eta)} d \eta\right) \\
& =P_{2}\left(-15 \cos (2 \theta) \cos (4 \theta)+\frac{3}{2}(\cos (2 \theta)-\cos (6 \theta))\right) \\
& =P_{2}(-6 \cos (2 \theta)-9 \cos (6 \theta)) \\
& =-6 \cos (2 \theta), \tag{A.1.25}
\end{align*}
$$

where the second equality follows from (Equation A.1.30) and Lemma A.1.9. For $K_{3}$, we compute

$$
\begin{align*}
P_{2} K_{3} & =P_{2}\left(-15 \cos (2 \theta) f \frac{\cos (4 \theta)-\cos (4 \eta)}{2-2 \cos (\theta-\eta)} d \eta-3 f \frac{(\cos (4 \theta)-\cos (4 \eta)) \cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta\right) \\
& =P_{2}(-30 \cos (2 \theta) \cos (4 \theta)-3 \cos (6 \theta)) \\
& =P_{2}(-15 \cos (2 \theta)-18 \cos (6 \theta)) \\
& =-15 \cos (2 \theta), \tag{A.1.26}
\end{align*}
$$

where the second equality follows from (Equation A.1.30) and Lemma A.1.9. For $K_{4}$, we compute

$$
\begin{align*}
P_{2} K_{4} & =P_{2}\left(-16 \cos (4 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta-16 f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (4 \eta)}{2-2-\cos (\theta-\eta)} d \eta\right) \\
& =P_{2}(-16 \cos (2 \theta) \cos (4 \theta)+8 \cos (2 \theta)-8 \cos (6 \theta)) \\
& =P_{2}(-16 \cos (6 \theta)) \\
& =0 \tag{A.1.27}
\end{align*}
$$

where the second equality follows from (Equation A.1.30) and Lemma A.1.9. For $K_{5}$, we compute

$$
\begin{align*}
P_{2} K_{5} & =-16 P_{2}(\sin (2 \theta) \sin (4 \theta)) \\
& =-8 P_{2}(\cos (2 \theta)-\cos (6 \theta)) \\
& =-8 \cos (2 \theta), \tag{A.1.28}
\end{align*}
$$

where the first equality follows from (Equation A.1.29). Hence it follows from (Equation A.1.24), (Equation A.1.25), (Equation A.1.26), (Equation A.1.27) and (Equation A.1.28) that

$$
\frac{d^{2}}{d t d s} Q \mathcal{F}=\left(0, P_{2} \frac{d^{2}}{d t d s} \tilde{\mathcal{F}}_{2}\right)=(0,-12 \cos (2 \theta))
$$

which proves (Equation A.1.19).
To prove (Equation A.1.20), note that $\frac{1}{3} Q \frac{d^{3}}{d t^{3}} \mathcal{F}=\left(0, \frac{1}{3} P_{2} \frac{d^{3}}{d t^{3}} \tilde{\mathcal{F}}_{2}\right)$. We use Lemma A.1.6 with $r=\cos (2 \theta)$ and obtain

$$
\begin{aligned}
\frac{1}{3} P_{2} \frac{d^{3}}{d t^{3}} \tilde{\mathcal{F}}_{2}(2, t v)= & P_{2}\left[4 f \frac{(\cos (2 \theta)-\cos (2 \eta))\left(3 \cos ^{2}(2 \theta)+2 \cos (2 \theta) \cos (2 \eta)+\cos ^{2}(2 \eta)\right)}{2-2 \cos (\theta-\eta)} d \eta\right. \\
& -8 f \frac{(\cos (2 \theta)-\cos (2 \eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& -16 \sin (2 \theta) f \frac{(\cos (2 \theta)-\cos (2 \eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta \\
& +8 f \frac{\left(3 \cos ^{2}(2 \theta)-4 \sin ^{2}(2 \theta)\right)(\cos (2 \theta)-\cos (2 \eta))}{2-2 \cos (\theta-\eta)} d \eta \\
& \left.+\left(32 \cos (2 \theta) \sin ^{2}(2 \theta)-16 \cos ^{3}(2 \theta)\right)\right] \\
& =P_{2} L_{1}+P_{2} L_{2}+P_{2} L_{3}+P_{2} L_{4}+P_{2} L_{5}
\end{aligned}
$$

For $L_{1}$, we compute

$$
\begin{aligned}
P_{2} L_{1}= & P_{2}\left(12 \cos ^{2}(2 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta+8 \cos (2 \theta) f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta\right. \\
& \left.+2 f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta+2 f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (4 \eta)}{2-2 \cos (\theta-\eta)} d \eta\right) \\
= & P_{2}\left(12 \cos ^{3}(2 \theta)+8 \cos (2 \theta)\left(-\frac{1}{2}+\frac{1}{2} \cos (4 \theta)\right)+2 \cos (2 \theta)+(-\cos (2 \theta)+\cos (6 \theta))\right) \\
= & P_{2}(8 \cos (2 \theta)+6 \cos (6 \theta)) \\
= & 8 \cos (2 \theta)
\end{aligned}
$$

where the second equality follows from (Equation A.1.30) and A.1.9. For $L_{2}$, we use Lemma A.1.10 and obtain

$$
\begin{aligned}
P_{2} L_{2} & =-8 P_{2}\left(\frac{9}{4} \cos (2 \theta)-\cos (6 \theta)\right) \\
& =-18 \cos (2 \theta)
\end{aligned}
$$

For $L_{3}$, we use Lemma A.1.11 and obtain

$$
P_{2} L_{3}=16 P_{2} \sin (2 \theta) \sin (4 \theta)=8 P_{2}(\cos (2 \theta)-\cos (6 \theta))=8 \cos (2 \theta)
$$

For $L_{4}$, we compute

$$
\begin{aligned}
P_{2} L_{4} & =P_{2}\left(24 \cos ^{2}(2 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta-32 \sin ^{2}(2 \theta) f \frac{\cos (2 \theta)-\cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta\right) \\
& =P_{2}\left(24 \cos ^{3}(2 \theta)-32 \sin ^{2}(2 \theta) \cos (2 \theta)\right) \\
& =P_{2}(10 \cos (2 \theta)+14 \cos (6 \theta)) \\
& =10 \cos (2 \theta)
\end{aligned}
$$

where the second equality follows from (Equation A.1.30). For $L_{5}$, it follows immediately that

$$
P_{2} L_{5}=P_{2}(-4 \cos (2 \theta)-12 \cos (6 \theta))=-4 \cos (2 \theta)
$$

Collecting the above results, we obtain

$$
\frac{1}{3} P_{2} \frac{d^{3}}{d t^{3}} \tilde{\mathcal{F}}_{2}(2, t v)=4 \cos (2 \theta)
$$

which implies (Equation A.1.20).
To prove (Equation A.1.21), we use (Equation A.1.17) and (Equation A.1.8) in Lemma A.1.4 with $g=0, \tilde{g}=2 \cos (2 \theta), r=\cos (2 \theta)$ and $\tilde{r}=0$ and obtain

$$
\begin{aligned}
P_{2} \frac{d^{2}}{d t d s} \tilde{\mathcal{F}}_{2}(2, t v+s \tilde{v})= & P_{2}\left(-4 \cos ^{2}(2 \theta)+4 f \frac{(\cos (2 \theta)+\cos (2 \eta))(\cos (2 \theta)-\cos (2 \eta))}{2-2 \cos (\theta-\eta)} d \eta\right. \\
& \left.+8 f \frac{\cos (2 \eta) \sin (2 \theta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta\right) \\
= & P_{2}\left(-4 \cos ^{2}(2 \theta)+4 \cos ^{2}(2 \theta)+4\left(-\frac{1}{2}+\frac{1}{2} \cos (4 \theta)\right)+4 \sin ^{2}(2 \theta)\right) \\
= & 0
\end{aligned}
$$

where the second equality follows from (Equation A.1.30), (Equation A.1.29) and (A.1.9). This implies (Equation A.1.21).

To prove (Equation A.1.22), it follows from (Equation A.1.18) and Lemma A.1.2 that

$$
\partial_{b} D \mathcal{F}(2,0)\left(-8 \cos (4 \theta), \frac{3}{2} \cos (4 \theta)\right)=(3 \sin (4 \theta), 2 \cos (4 \theta)) .
$$

By projecting it to the image of $Q$, we obtain (Equation A.1.22).
To prove (Equation A.1.23), we note that $Q \partial_{b} D \mathcal{F}(2,0) \tilde{v}=\left.\left(0, P_{2} \partial_{b} \frac{d}{d t} \tilde{\mathcal{F}}_{2}(2, t \tilde{v})\right)\right|_{t=0}$. Hence it follows from (Equation A.1.4) in Lemma A.1.2 and (Equation A.1.17) that $\left.P_{2} \partial_{b} \frac{d}{d t} \tilde{\mathcal{F}}_{2}(2, t \tilde{v})\right|_{t=0}=$
$\cos (2 \theta)$. This implies (Equation A.1.23).

## A.1.3 Basic Integrals

Lemma A.1.8. For $\mathbb{N} \ni m \geq 1$, it holds that

$$
\begin{align*}
& f \frac{\cos (m \eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=\frac{1}{2} \sin (m \theta),  \tag{A.1.29}\\
& f \frac{\cos (m \theta)-\cos (m \eta)}{2-2 \cos (\theta-\eta)} d \eta=\frac{m}{2} \cos (m \theta) \tag{A.1.30}
\end{align*}
$$

Proof. For (Equation A.1.29), it is clear that $f \frac{\cos (m \eta) \sin (\theta-\eta)}{2-2 \cos (\theta-\eta)} d \eta=\frac{1}{2} f \cos (m \eta) \cot \left(\frac{\theta-\eta}{2}\right) d \eta=$ $\frac{1}{2} H(\cos (m \theta))(\theta)$, where $H$ denotes the Hilbert transform in the periodic domain. Therefore the result follows immediately since $H(\cos (m \theta))(\theta)=\sin (m \theta)$.

For (Equation A.1.30), we recall that $f \frac{f(\theta)-f(\eta)}{1-\cos (\theta-\eta)} d \eta=\Lambda f(\theta)=:(-\Delta)^{\frac{1}{2}} f(\theta)$. Thus (Equation A.1.30) follows immediately.

## Lemma A.1.9.

$$
\begin{align*}
& f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (4 \eta)}{2-2 \cos (\theta-\eta)} d \eta=-\frac{1}{2} \cos (2 \theta)+\frac{1}{2} \cos (6 \theta),  \tag{A.1.31}\\
& f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta=-\frac{1}{2}+\frac{1}{2} \cos (4 \theta),  \tag{A.1.32}\\
& f \frac{(\cos (4 \theta)-\cos (4 \eta)) \cos (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta=\cos (6 \theta),  \tag{A.1.33}\\
& f \frac{(\cos (2 \theta)-\cos (2 \eta)) \sin (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta=\frac{1}{2} \sin (4 \theta) . \tag{A.1.34}
\end{align*}
$$

Proof. We will show (Equation A.1.31) only. (Equation A.1.32), (Equation A.1.33) and
(Equation A.1.34) can be proved in the same way. For (Equation A.1.31), one can write

$$
\begin{aligned}
f \frac{(\cos (2 \theta)-\cos (2 \eta)) \cos (4 \eta)}{2-2 \cos (\theta-\eta)} d \eta & =\frac{1}{2} \Lambda(\cos (2 \theta) \cos (4 \theta))(\theta)-\frac{1}{2} \cos (2 \theta) \Lambda(\cos (4 \theta))(\theta) \\
& =-\frac{1}{2} \cos (2 \theta)+\frac{1}{2} \cos (6 \theta)
\end{aligned}
$$

where the last equality follows from (Equation A.1.30).

Lemma A.1.10.

$$
f \frac{(\cos (2 \theta)-\cos (2 \eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta=\frac{9}{4} \cos (2 \theta)-\cos (6 \theta)
$$

Proof. We compute

$$
\begin{aligned}
f & \frac{(\cos (2 \theta)-\cos (2 \eta))^{3}}{(2-2 \cos (\theta-\eta))^{2}} d \eta=f \frac{(-2 \sin (\theta-\eta) \sin (\theta+\eta))^{3}}{\left(4 \sin ^{2}\left(\frac{\theta-\eta}{2}\right)\right)^{2}} d \eta=4 f-\frac{\cos ^{3}\left(\frac{\theta-\eta}{2}\right) \sin ^{3}(\theta+\eta)}{\sin \left(\frac{\theta-\eta}{2}\right)} d \eta \\
& =4 f \frac{\cos ^{3} \frac{\eta}{2} \sin ^{3}(2 \theta+\eta)}{\sin \frac{\eta}{2}} d \eta=4 f \frac{\cos ^{3} \frac{\eta}{2}}{\sin \frac{\eta}{2}}\left(3 \sin ^{2}(2 \theta) \cos (2 \theta) \cos ^{2} \eta \sin \eta+\cos ^{3}(2 \theta) \sin ^{3} \eta\right) d \eta \\
& =12 \sin ^{2}(2 \theta) \cos (2 \theta) f \frac{\cos ^{3} \frac{\eta}{2} \cos ^{2} \eta \sin \eta}{\sin \frac{\eta}{2}} d \eta+4 \cos ^{3}(2 \theta) f \frac{\cos ^{3} \frac{\eta}{2} \sin ^{3} \eta}{\sin \frac{\eta}{2}} d \eta \\
& =\frac{21}{4} \sin ^{2}(2 \theta) \cos (2 \theta)+\frac{5}{4} \cos ^{3}(2 \theta) \\
& =\frac{9}{4} \cos (2 \theta)-\cos (6 \theta)
\end{aligned}
$$

which proves the lemma.

Lemma A.1.11.

$$
f \frac{(\cos (2 \theta)-\cos (2 \eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta=-\sin (4 \theta) .
$$

Proof. Using the integration by parts, we compute

$$
f \frac{(\cos (2 \theta)-\cos (2 \eta))^{2} \sin (\theta-\eta)}{(2-2 \cos (\theta-\eta))^{2}} d \eta=-2 f \frac{(\cos (2 \theta)-\cos (2 \eta)) \sin (2 \eta)}{2-2 \cos (\theta-\eta)} d \eta
$$

Therefore the result follows from (Equation A.1.34).

## APPENDIX B

## DERIVATIVES OF STREAM FUNCTION

## B. 1 Derivatives of the stream function

In this appendix, we will derive some formulae for zero-mean stream function by using Fourier series.

Lemma B.1.1. For $\rho>0$, let $h \in L^{2}\left(\partial B_{\rho}\right)$ such that $\int_{|y|=\rho} h(y) d \mathcal{H}^{1}(y)=0$. Then it holds that for $x=(r \cos \theta, r \sin \theta)$,

$$
\frac{1}{2 \pi} \int_{|y|=\rho} h(y) \log |x-y| d \mathcal{H}^{1}(y)= \begin{cases}-\sum_{n=-\infty n \neq 0}^{\infty} \frac{\rho}{2|n|} \hat{h}(\rho, n)\left(\frac{r}{\rho}\right)^{|n|} e^{i n \theta} & \text { if } \rho \geq r, \\ -\sum_{n=-\infty n \neq 0}^{\infty} \frac{\rho}{2|n|} \hat{h}(\rho, n)\left(\frac{\rho}{r}\right)^{|n|} e^{i n \theta} & \text { if } \rho<r\end{cases}
$$

where $\hat{h}(\rho, n)=\frac{1}{2 \pi} \int_{\partial B} h(\rho y) e^{-i n y} d \mathcal{H}^{1}(y)$.
Proof. By adapting the abuse of notation $h(y)=h(\rho, \eta)$, for $y=(\rho \cos \eta, \rho \sin \eta)$, we have that

$$
\begin{aligned}
\frac{1}{2 \pi} \int_{|y|=\rho} h(y) \log |x-y| d \mathcal{H}^{1}(y) & =\frac{\rho}{4 \pi} \int_{\mathbb{T}} h(\rho, \eta) \log |(r \cos \theta, r \sin \theta)-(\rho \cos \eta, \rho \sin \eta)|^{2} d \eta \\
& =\frac{\rho}{4 \pi} \int_{\mathbb{T}} h(\rho, \eta) \log \left(r^{2}+\rho^{2}-2 r \rho \cos (\theta-\eta)\right) d \eta .
\end{aligned}
$$

Using the Fourier expansion $h(\rho, \eta):=\sum_{n=-\infty, n \neq 0}^{\infty} \hat{h}(\rho, n) e^{i n \eta}$ where $\hat{h}(\rho, n) \quad:=$ $\frac{1}{2 \pi} \int_{\mathbb{T}} h(\rho, \eta) e^{-i n \eta} d \eta$, we have

$$
\begin{aligned}
\frac{\rho}{4 \pi} \int_{\mathbb{T}} h(\rho, \eta) & \log \left(r^{2}+\rho^{2}-2 r \rho \cos (\theta-\eta)\right) d \eta \\
& =\sum_{n=-\infty, n \neq 0}^{\infty} \frac{\rho}{4 \pi} \hat{h}(\rho, n) \underbrace{\int_{\mathbb{T}} e^{i n \eta} \log \left(r^{2}+\rho^{2}-2 r \rho \cos (\theta-\eta)\right) d \eta}_{=: A_{n}(r, \rho, \theta)}
\end{aligned}
$$

where we used $\hat{h}(\rho, 0)=0$ since $h$ has zero mean on $\partial B_{\rho}$. To compute $A_{n}$, we recall from [24, Lemma A.1] that for $0 \leq x \leq 1$ and $\mathbb{Z} \ni n \neq 0$, it holds that

$$
\begin{equation*}
\int_{\mathbb{T}} e^{i n \eta} \log \left(1+x^{2}-2 x \cos (\theta-\eta)\right) d \eta=-\frac{2 \pi}{|n|} e^{i n \theta} x^{|n|} \tag{B.1.1}
\end{equation*}
$$

Then it directly follows from (Equation B.1.1) that

$$
A_{n}(r, \rho, \theta)= \begin{cases}-\frac{2 \pi}{|n|} e^{i n \theta}\left(\frac{r}{\rho}\right)^{|n|} & \text { if } \rho \geq r \\ -\frac{2 \pi}{|n|} e^{i n \theta}\left(\frac{\rho}{r}\right)^{|n|} & \text { if } \rho<r\end{cases}
$$

Plugging this into the above equation, the desired result follows immediately.

Lemma B.1.2. For a bounded $m$-fold symmetric domain $D$ in $\mathbb{R}^{2}$, let us consider a decomposition of $1_{D} * \mathcal{N}$,

$$
1_{D} * \mathcal{N}(r, \theta)=g * \mathcal{N}(r)+\left(1_{D}-g\right) * \mathcal{N}(r, \theta)=: \varphi^{r}(r)+\varphi_{m}(r, \theta)
$$

where $g(r):=\frac{1}{2 \pi r} \mathcal{H}^{1}\left(\partial B_{r} \cap D\right)$. Then,

$$
\begin{align*}
\partial_{r} \varphi_{m}(r, \theta) & =\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{\rho}{r}\right)^{n m+1} \cos (n m \eta)\right) d \rho d \eta \\
& -\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1} \cos (n m \eta)\right) d \rho d \eta  \tag{B.1.2}\\
\partial_{\theta} \varphi_{m}(r, \theta) & =-\frac{r}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{\rho}{r}\right)^{n m+1} \sin (n m \eta)\right) d \rho d \eta \\
& -\frac{r}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1} \sin (n m \eta)\right) d \rho d \eta \tag{B.1.3}
\end{align*}
$$

where $h(\rho, \theta):=1_{D}(\rho(\cos \theta, \sin \theta))-g(\rho)$.

Proof. We compute that for $x:=(r \cos \theta, r \sin \theta)$,

$$
\begin{aligned}
\varphi_{m}(r, \theta) & =\frac{1}{2 \pi} \int_{\mathbb{R}^{2}}\left(1_{D}(y)-g(|y|)\right) \log |x-y| d y \\
& =\int_{0}^{\infty}(\frac{1}{2 \pi} \int_{|y|=\rho} \underbrace{\left(1_{D}(y)-g(|y|)\right)}_{=: h(y)} \log |x-y| d \mathcal{H}^{1}(y)) d \rho
\end{aligned}
$$

By adapting the abuse of notation $h(y)=h(\rho, \eta)$ for $y=(\rho \cos \eta, \rho \sin \eta)$, we have $\int_{\mathbb{T}} h(\rho, \eta) d \eta=$ 0 for all $\rho>0$. Since $D$ is $m$-fold symmetric, we also have that $\eta \mapsto h(\rho, \eta)$ is $\frac{2 \pi}{m}$-periodic function for each fixed $\rho$. Therefore, it follows from Lemma B.1.1 that

$$
\varphi_{m}(r, \theta):=-\sum_{n=-\infty, n \neq 0}^{\infty} \frac{1}{2|n m|}\left(\int_{0}^{r} \rho \hat{h}(\rho, n m)\left(\frac{\rho}{r}\right)^{|n m|} e^{i n m \theta} d \rho+\int_{r}^{\infty} \rho \hat{h}(\rho, n m)\left(\frac{r}{\rho}\right)^{|n m|} e^{i n m \theta} d \rho\right)
$$

where $\hat{h}(\rho, n m):=\frac{1}{2 \pi} \int_{\mathbb{T}} h(\rho, \eta) e^{-i n m \eta} d \eta$. Therefore we have

$$
\begin{align*}
& \partial_{r} \varphi_{m}(r, \theta)=-\sum_{n=-\infty, n \neq 0}^{\infty} \frac{1}{2}\left(-\int_{0}^{r} \hat{h}(\rho, n m)\left(\frac{\rho}{r}\right)^{|n m|+1} e^{i n m \theta} d \rho+\int_{r}^{\infty} \hat{h}(\rho, n m)\left(\frac{r}{\rho}\right)^{|n m|-1} e^{i n m \theta} d \rho\right), \\
& \partial_{\theta} \varphi_{m}(r, \theta)=-\sum_{n=-\infty, n \neq 0}^{\infty} \frac{i n}{2|n|} r\left(\int_{0}^{r} \hat{h}(\rho, n m)\left(\frac{\rho}{r}\right)^{|n m|+1} e^{i n m \theta} d \rho+\int_{r}^{\infty} \hat{h}(\rho, n m)\left(\frac{r}{\rho}\right)^{|n m|-1} e^{i n m \theta} d \rho\right) . \tag{B.1.4}
\end{align*}
$$

To simplify the radial derivative, we use the definition of $\hat{h}$ and (Equation B.1.4) to obtain

$$
\begin{aligned}
\partial_{r} \varphi_{m}(r, \theta) & =\frac{1}{4 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta)\left(\sum_{n \neq 0}\left(\frac{\rho}{r}\right)^{|n m|+1} e^{i n m(\theta-\eta)}\right) d \rho d \eta \\
& -\frac{1}{4 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta)\left(\sum_{n \neq 0}\left(\frac{r}{\rho}\right)^{|n m|-1} e^{i n m(\theta-\eta)}\right) d \rho d \eta
\end{aligned}
$$

$$
\begin{aligned}
& =\frac{1}{4 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n \neq 0}\left(\frac{\rho}{r}\right)^{|n m|+1} e^{-i n m \eta}\right) d \rho d \eta \\
& -\frac{1}{4 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n \neq 0}\left(\frac{r}{\rho}\right)^{|n m|-1} e^{-i n m \eta}\right) d \rho d \eta \\
& =\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{\rho}{r}\right)^{n m+1} \cos (n m \eta)\right) d \rho d \eta \\
& -\frac{1}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1} \cos (n m \eta)\right) d \rho d \eta,
\end{aligned}
$$

where we used the change of variables, $\eta \mapsto \eta+\theta$ to get the second inequality. This proves (Equation B.1.2). In the same way, we use (Equation B.1.5) and the change of variables to obtain

$$
\begin{aligned}
\partial_{\theta} \varphi_{m}(r, \theta) & =-\frac{r}{4 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta)\left(\sum_{n \neq 0} \frac{i n}{|n|}\left(\frac{\rho}{r}\right)^{|n m|+1} e^{i n m(\theta-\eta)}\right) d \rho d \eta \\
& -\frac{r}{4 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta)\left(\sum_{n \neq 0} \frac{i n}{|n|}\left(\frac{r}{\rho}\right)^{|n m|-1} e^{i n m(\theta-\eta)}\right) d \rho d \eta \\
& =-\frac{r}{4 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n \neq 0} \frac{i n}{|n|}\left(\frac{\rho}{r}\right)^{|n m|+1} e^{-i n m \eta}\right) d \rho d \eta \\
& -\frac{r}{4 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n \neq 0} \frac{i n}{|n|}\left(\frac{r}{\rho}\right)^{|n m|-1} e^{-i n m \eta}\right) d \rho d \eta \\
& =-\frac{r}{2 \pi} \int_{\mathbb{T}} \int_{0}^{r} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{\rho}{r}\right)^{n m+1} \sin (n m \eta)\right) d \rho d \eta \\
& -\frac{r}{2 \pi} \int_{\mathbb{T}} \int_{r}^{\infty} h(\rho, \eta+\theta)\left(\sum_{n=1}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1} \sin (n m \eta)\right) d \rho d \eta
\end{aligned}
$$

which proves (Equation B.1.3).

Lemma B.1.3. For a patch $D$ that satisfies the assumptions (a)-(c) in subsection subsection 6.2.2,
it holds that for $r \in\left(r_{\min }, r_{\max }\right)$ and $\eta:=u^{-1}$,

$$
\begin{aligned}
\partial_{r} \varphi_{m}(r, \theta) & =\int_{r_{\min }}^{r_{\max }} f_{1}(\rho, r, \theta) d \rho \\
\partial_{\theta} \varphi_{m}(r, \theta) & =\int_{r_{\min }}^{r_{\max }} f_{2}(\rho, r, \theta) d \rho
\end{aligned}
$$

where
$f_{1}(\rho, r, \theta)= \begin{cases}\frac{1}{2 \pi} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\theta-\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\theta-\eta(\rho)))}\right)-\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\theta+\eta(\rho)))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\theta+\eta(\rho)))}\right)\right) & \text { if } \rho \geq r \\ \frac{1}{2 \pi} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\theta+\eta(\rho)))}{1-\left(\frac{\rho}{r}\right)^{m} \cos (m(\theta+\eta(\rho)))}\right)-\arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\theta-\eta(\rho)))}{1-\left(\frac{\rho}{r}\right)^{n} \cos (m(\theta-\eta(\rho)))}\right)\right) & \text { if } \rho<r,\end{cases}$
$f_{2}(\rho, r, \theta)= \begin{cases}\frac{\rho}{4 \pi} \log \left(1+\frac{4\left(\frac{r}{\rho}\right)^{m} \sin (m \eta(\rho)) \sin (m \theta)}{1+\left(\frac{\Gamma}{\rho}\right)^{m}-2\left(\frac{r}{\rho}\right)^{m} \cos (m(\theta-\eta(\rho)))}\right) & \text { if } \rho \geq r \\ \frac{\rho}{4 \pi} \log \left(1+\frac{4\left(\frac{\rho}{r}\right)^{m} \sin (m \eta(\rho)) \sin (m \theta)}{1+\left(\frac{\rho}{r}\right)^{m}-2\left(\frac{\rho}{r}\right)^{m} \cos (m(\theta-\eta(\rho)))}\right) & \text { if } \rho<r .\end{cases}$
Proof. The proof is based on Lemma B.1.2. Using $m$-fold symmetry and evenness of the patch, we will compute the series.

From Lemma B.1.2 and Fubini theorem, it follows that

$$
\begin{align*}
\partial_{r} \varphi_{m}(r, \theta) & =\frac{1}{2 \pi} \sum_{n=1}^{\infty} \int_{0}^{r}\left(\frac{\rho}{r}\right)^{n m+1}\left(\int_{\mathbb{T}} h(\rho, s+\theta) \cos (n m s) d s\right) d \rho \\
& -\frac{1}{2 \pi} \sum_{n=1}^{\infty} \int_{r}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1}\left(\int_{\mathbb{T}} h(\rho, s+\theta) \cos (n m s) d s\right) d \rho \tag{B.1.8}
\end{align*}
$$

and

$$
\begin{align*}
\partial_{\theta} \varphi_{m}(r, \theta) & =-\frac{r}{2 \pi} \sum_{n=1}^{\infty} \int_{0}^{r}\left(\frac{\rho}{r}\right)^{n m+1}\left(\int_{\mathbb{T}} h(\rho, s+\theta) \sin (n m s) d s\right) d \rho \\
& -\frac{r}{2 \pi} \sum_{n=1}^{\infty} \int_{r}^{\infty}\left(\frac{r}{\rho}\right)^{n m-1}\left(\int_{\mathbb{T}} h(\rho, s+\theta) \sin (n m s) d s\right) d \rho \tag{B.1.9}
\end{align*}
$$

where $h(\rho, s):=1_{D}(\rho \cos s, \rho \sin s)-g(\rho)$. Using the definition of $\eta=u^{-1}$, the following holds for $s \in\left[-\frac{\pi}{m}, \frac{\pi}{m}\right]$ :

$$
h(\rho, s)= \begin{cases}1-g(\rho) & \text { if } s \in(-\eta(\rho), \eta(\rho)) \\ -g(\rho) & \text { if } s \in\left[-\frac{\pi}{m}, \frac{\pi}{m}\right] \backslash(-\eta(\rho), \eta(\rho)) .\end{cases}
$$

Therefore $m$-fold symmetry of $D$ yields that

$$
\begin{align*}
\int_{\mathbb{T}} h(\rho, s+\theta) \cos (n m s) d s & =m \int_{-\frac{\pi}{m}}^{\frac{\pi}{m}} h(\rho, s) \cos (n m(s-\theta)) d s \\
& =m \int_{-\eta(\rho)}^{\eta(\rho)} \cos (n m(s-\theta)) d s \\
& =\frac{1}{n}(\sin (n m(\eta(\rho)-\theta))+\sin (n m(\eta(\rho)+\theta))) \tag{B.1.10}
\end{align*}
$$

Similarly, we have

$$
\begin{equation*}
\int_{\mathbb{T}} h(\rho, s+\theta) \sin (n m s) d s=-\frac{1}{n}(\cos (n m(\eta(\rho)-\theta))-\cos (n m(\eta(\rho)+\theta))) \tag{B.1.11}
\end{equation*}
$$

Hence (Equation B.1.8) and (Equation B.1.10) yield that

$$
\begin{aligned}
& \partial_{r} \varphi_{m}(r, \theta) \\
& \quad=\frac{1}{2 \pi} \int_{0}^{r} \sum_{n=1}^{\infty}\left(\frac{1}{n}\left(\frac{\rho}{r}\right)^{n m+1}(\sin (n m(\eta(\rho)-\theta))+\sin (n m(\eta(\rho)+\theta)))\right) d \rho \\
& \\
& -\frac{1}{2 \pi} \int_{r}^{\infty} \sum_{n=1}^{\infty}\left(\frac{1}{n}\left(\frac{r}{\rho}\right)^{n m-1}(\sin (n m(\eta(\rho)-\theta))+\sin (n m(\eta(\rho)+\theta)))\right) d \rho \\
& \quad=\frac{1}{2 \pi} \int_{0}^{r} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\eta(\rho)-\theta))}{1-\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right)+\arctan \left(\frac{\left(\frac{\rho}{r}\right)^{m} \sin (m(\eta(\rho)+\theta))}{1-\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(\rho)+\theta))}\right)\right) d \rho \\
& \\
& \quad-\frac{1}{2 \pi} \int_{r}^{\infty} \frac{\rho}{r}\left(\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(\rho)-\theta))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right)\right.
\end{aligned}
$$

$$
+\arctan \left(\frac{\left(\frac{r}{\rho}\right)^{m} \sin (m(\eta(\rho)+\theta))}{1-\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(\rho)+\theta))}\right) d \rho
$$

where the last equality follows from (Equation B.2.2) in Lemma B.2.1. Since the integrands in the above integrals are zero if $\rho<r_{\min }$ or $\rho>r_{\max }$, we can replace 0 and $\infty$ in integration limits by $r_{\text {min }}$ and $r_{\text {max }}$, respectively. This proves (Equation B.1.6). To prove (Equation B.1.7), we use (Equation B.1.9),(Equation B.1.11) and (Equation B.2.1) to obtain

$$
\begin{aligned}
& \partial_{\theta} \varphi_{m}(r, \theta) \\
&=\frac{r}{2 \pi} \int_{0}^{r} \sum_{n=1}^{\infty}\left(\frac{1}{n}\left(\frac{\rho}{r}\right)^{n m+1}(\cos (n m(\eta(\rho)-\theta))-\cos (n m(\eta(\rho)+\theta)))\right) d \rho \\
&+\frac{r}{2 \pi} \int_{r}^{\infty} \sum_{n=1}^{\infty}\left(\frac{1}{n}\left(\frac{r}{\rho}\right)^{n m-1}(\cos (n m(\eta(\rho)-\theta))-\cos (n m(\eta(\rho)+\theta)))\right) d \rho \\
&=\frac{r}{4 \pi} \int_{0}^{r} \frac{\rho}{r} \log \left(\frac{1+\left(\frac{\rho}{r}\right)^{2 m}-2\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(\rho)+\theta))}{1+\left(\frac{\rho}{r}\right)^{2 m}-2\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right) d \rho \\
&+\frac{r}{4 \pi} \int_{r}^{\infty} \frac{\rho}{r} \log \left(\frac{1+\left(\frac{r}{\rho}\right)^{2 m}-2\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(\rho)+\theta))}{1+\left(\frac{r}{\rho}\right)^{2 m}-2\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right) d \rho \\
&=\frac{1}{4 \pi} \int_{0}^{r} \rho \log \left(1+\frac{2\left(\frac{\rho}{r}\right)^{m}(\cos (m(\eta(\rho)-\theta))-\cos (m(\eta(\rho)+\theta)))}{1+\left(\frac{\rho}{r}\right)^{2 m}-2\left(\frac{\rho}{r}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right) d \rho \\
&+\frac{1}{4 \pi} \int_{r}^{\infty} \rho \log \left(1+\frac{2\left(\frac{r}{\rho}\right)^{m}(\cos (m(\eta(\rho)-\theta))-\cos (m(\eta(\rho)+\theta)))}{1+\left(\frac{r}{\rho}\right)^{2 m}-2\left(\frac{r}{\rho}\right)^{m} \cos (m(\eta(\rho)-\theta))}\right) d \rho \\
&=\frac{1}{4 \pi} \int_{0}^{r} \rho \log \left(1+\frac{4\left(\frac{\rho}{r}\right)^{m} \sin (m \eta(\rho)) \sin (m \theta)}{1+\left(\frac{\rho}{r}\right)^{2 m}-2\left(\frac{\rho}{r}\right)^{m} \cos (m(\theta-\eta(\rho)))}\right) d \rho \\
&+\frac{1}{4 \pi} \int_{r}^{\infty} \rho \log \left(1+\frac{4\left(\frac{r}{\rho}\right)^{m} \sin (m \eta(\rho)) \sin (m \theta)}{1+\left(\frac{r}{\rho}\right)^{2 m}-2\left(\frac{r}{\rho}\right)^{m} \cos (m(\theta-\eta(\rho)))}\right) d \rho,
\end{aligned}
$$

where the last equality follows from $\cos (x-y)-\cos (x+y)=2 \sin x \sin y$. This proves (Equation B.1.7).

## B. 2 Helpful lemmas

Lemma B.2.1. For $|x|<1$ and $y \in(-\pi, \pi)$, it holds that

$$
\begin{gather*}
\sum_{n=1}^{\infty} \frac{1}{n} x^{n} \cos (n y)=-\frac{1}{2} \log \left(1+x^{2}-2 x \cos y\right)  \tag{B.2.1}\\
\sum_{n=1}^{\infty} \frac{1}{n} x^{n} \sin (n y)=\arctan \left(\frac{x \sin y}{1-x \cos y}\right) \tag{B.2.2}
\end{gather*}
$$

Consequently, we have

$$
\begin{align*}
& \sum_{n=1}^{\infty} x^{n} \cos (n y)=\frac{x(\cos y-x)}{(1-x)^{2}+2 x(1-\cos y)}  \tag{B.2.3}\\
& \sum_{n=1}^{\infty} x^{n} \sin (n y)=\frac{x \sin y}{(1-x)^{2}+2 x(1-\cos y)} \tag{B.2.4}
\end{align*}
$$

Proof. Let $f(x, y):=\sum_{n=1}^{\infty} \frac{1}{n} x^{n} e^{i n y}$. Then we compute

$$
\begin{aligned}
\partial_{x} f(x, y) & =\frac{1}{x} \sum_{n=1}^{\infty}\left(x e^{i y}\right)^{n}=\frac{e^{i y}}{1-x e^{i y}}=\frac{(\cos y-x)+i \sin y}{(1-x \cos y)^{2}+x^{2} \sin ^{2} y} \\
& =\partial_{x}\left(-\frac{1}{2} \log \left(1+x^{2}-2 x \cos y\right)+i \arctan \left(\frac{x \sin y}{1-x \cos y}\right)\right)
\end{aligned}
$$

Since $f(0, y)=0$, we have $f(x, y)=-\frac{1}{2} \log \left(1+x^{2}-2 x \cos y\right)+i \arctan \left(\frac{x \sin y}{1-x \cos y}\right)$. Equating the real and imaginary parts separately, we can obtain (Equation B.2.1) and (Equation B.2.2). By differentiating (Equation B.2.1) and (Equation B.2.2) and multiplying by $x$, one can easily obtain (Equation B.2.3) and (Equation B.2.4).

Lemma B.2.2. For $m \geq 3$ and $a, b \in(0,1)$, it holds that

$$
\int_{0}^{1} x^{-1-\frac{2}{m}}\left(\arctan \left(\frac{a x}{1-x}\right)-\arctan \left(\frac{a x}{1-b x}\right)\right) d x \lesssim 1-b
$$

Proof. By the change of variables, $b x \mapsto x$, we have

$$
\begin{aligned}
\int_{0}^{1} x^{-1-\frac{2}{m}} \arctan \left(\frac{a x}{1-b x}\right) d x & =\int_{0}^{b} b^{\frac{2}{m}} x^{-1-\frac{2}{m}} \arctan \left(\frac{\frac{a x}{b}}{1-x}\right) d x \\
& \geq \int_{0}^{b} b x^{-1-\frac{2}{m}} \arctan \left(\frac{a x}{1-x}\right) d x
\end{aligned}
$$

where we used $b^{\frac{2}{m}} \geq b$ for $0<b<1$ and $m \geq 3$. Therefore it follows that

$$
\begin{aligned}
\int_{0}^{1} x^{-1-\frac{2}{m}}\left(\arctan \left(\frac{a x}{1-x}\right)-\arctan \left(\frac{a x}{1-b x}\right)\right) d x & \leq \int_{b}^{1} x^{-1-\frac{2}{m}} \arctan \left(\frac{a x}{1-x}\right) d x \\
& +(1-b) \int_{0}^{b} x^{-1-\frac{2}{m}} \arctan \left(\frac{a x}{1-x}\right) d x \\
& \lesssim 1-b
\end{aligned}
$$

which proves the desired inequality.

Lemma B.2.3. For $m \geq 3$ and $a \in(0,1)$, it holds that

$$
\begin{equation*}
\int_{0}^{1} x^{-1-\frac{2}{m}} \log \left(1+\frac{a x}{(1-x)^{2}}\right) d x \lesssim \sqrt{a} \tag{B.2.5}
\end{equation*}
$$

Proof. If $x<\frac{1}{2}$, then $\log \left(1+\frac{a x}{(1-x)^{2}}\right) \lesssim a x$. Therefore,

$$
\begin{align*}
\int_{0}^{1} x^{-1-\frac{2}{m}} \log \left(1+\frac{a x}{(1-x)^{2}}\right) d x & \lesssim \int_{0}^{\frac{1}{2}} a x^{-\frac{2}{m}} d x+\int_{\frac{1}{2}}^{1} \log \left(1+\frac{a x}{(1-x)^{2}}\right) d x \\
& \lesssim a+\int_{\frac{1}{2}}^{1} \log \left(1+\frac{a}{(1-x)^{2}}\right) d x \tag{B.2.6}
\end{align*}
$$

where we used $m \geq 3$ to estimate the first integral and $x \in\left(\frac{1}{2}, 1\right)$ for the second integral. To
estimate the second integral, we compute

$$
\begin{align*}
\int_{\frac{1}{2}}^{1} \log \left(1+\frac{a}{(1-x)^{2}}\right) d x & =\int_{\frac{1}{2}}^{1} \frac{d}{d x}(x-1) \log \left(1+\frac{a}{(1-x)^{2}}\right) d x \\
& =\frac{1}{2} \log (1+4 a)+\int_{\frac{1}{2}}^{1} \frac{2 a}{(1-x)^{2}+a} d x \\
& \lesssim a+\int_{\frac{1}{2}}^{1-\sqrt{a}} \frac{a}{(1-x)^{2}} d x+\int_{1-\sqrt{a}}^{1} 1 d x \\
& \lesssim \sqrt{a} . \tag{B.2.7}
\end{align*}
$$

Thus the desired result follows from (Equation B.2.6) and (Equation B.2.7).

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[^0]:    ${ }^{1}$ For an open curve $i=n+1, \ldots, n+m$, note that (H2) does not require $\gamma_{i}$ to be $C^{1}$ up to the boundary of $S_{i}$, and its derivative is allowed to blow up at the endpoints. This is motivated by the fact that in the explicit uniformly-rotating solution (Equation 1.2.3), its strength $\gamma$ is Hölder continuous in $[-a, a]$ and smooth in the interior, but its derivative blows up at the endpoints.

[^1]:    ${ }^{2}$ In fact, (Equation 4.2.2) also holds (with a slightly smaller $\epsilon_{0}$ and $c_{0}$ ) for $\eta, \eta^{\prime} \in(-2,2)$, even though such $R_{i}^{\epsilon}$ may not belong to $D_{i}^{\epsilon}$. We will use this fact later in the proof of Lemma 4.2.5.

